

Quantification of Volcano Deformation caused by Volatile Accumulation and Release

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Key Points:

- Exsolving volatiles accumulate at the roof of a magma storage zone and contribute to surface deformation through buoyancy forces
- 3D numerical models show that surface deformation is a function of the volatiles' volume, density and depth as well as crustal rigidity
- Volatile release during eruption can cause syn-eruptive subsidence of a few cm, which is 20% of the observed signal at Calbuco in 2015

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Abstract

Crustal-stored magma reservoirs contain exsolved volatiles which accumulate in the reservoir roof, exerting a buoyancy force on the crust. This produces surface uplift and sudden loss of volatiles through eruption results in syn-eruptive subsidence. Here, we present three-dimensional, visco-elasto-plastic, numerical modeling results which quantify the ground deformation arising from the growth and release of a volatile reservoir. Deformation is independent of crustal thermal distribution and volatile reservoir shape, but is a function of volatile volume, density and depth and crustal rigidity. We present a scaling law for the volatiles' contribution to syn-eruptive subsidence and show this contributes $\sim 20\%$ of the observed subsidence associated with the 2015 Calbuco eruption. Our results highlight the key role that volatile-driven buoyancy can have in volcano deformation, shows a new link between syn-eruptive degassing and deflation, and highlights that shallow gas accumulation and release may have a major impact on ground deformation of volcanoes.

Plain Language Summary

Magma contains a lot of gases which separate from it when it approaches the surface. These gases can collect right above the magma storage region a few kilometers below the surface. They have a much lower density than the rocks surrounding them and push upwards like a balloon filled with air that is pressed under water. In this study, we use computer models to understand how much a volcano would grow from the push of the gases below and how much it would shrink when the gases escape because of an eruption. We find that the gases can cause the volcano to grow and shrink up to a few centimeters during accumulation and release, respectively. The amount of surface movement depends on the volume, density and depth of the gas reservoir as well as on the toughness of the rocks above it. We derive a simple equation which allows us to compute the surface movement using the aforementioned parameters. With this equation and estimates about the amount of accumulated gas at the 2015 Calbuco eruption, we can assume that about 20% of the observed surface movement was caused by the release of the magmatic gases.

Index Terms

8145 Physics of magma and magma bodies

48 1211 Non-tectonic deformation
49 0545 Modeling
50 8430 Volcanic gases
51 1036 Magma chamber processes

52 **Keywords**

53 3D modeling, Buoyancy, Volcanic uplift, Scaling analysis, Magmatic gases, Calbuco

54 **1 Introduction**

55 Volcano deformation is most frequently interpreted in terms of models of surface
56 deformation due to processes in magma bodies of various geometries. The most widely
57 applied model is that of a point source of pressure embedded within a uniform elastic
58 halfspace (Mogi, 1958), but a range of more advanced models and approaches exist (e.g.
59 Fialko et al., 2001a; Hickey et al., 2016). As liquid magma flows in/out of the these “de-
60 formation sources“, they expand/contract. Most often, such magma flow is considered
61 to cause uniform pressure change on the boundary of the magma body, and the density
62 difference between magma and host rock is not considered specifically. It has, however,
63 been demonstrated in a number of studies that magma buoyancy can cause significant
64 stresses in volcano roots and contribute to failure of magma bodies (e.g. Sigmundsson
65 et al., 2020). A particular phenomena not considered by traditional volcano deforma-
66 tion models is the effect of accumulated exsolved volatiles in volcano roots and their re-
67 lease during eruptions.

68 During major explosive eruptions an excess of gas may be observed, beyond that
69 which can be explained by a petrological calculation of the original dissolved volatile amounts
70 and the volume of erupted lavas. Excess gas was observed in the 1991 eruption of Pinatubo,
71 Philippines and an analysis from Wallace and Gerlach (1994) showed that this could be
72 explained by a pre-existing gas/volatile phase representing 0.7 to 1.3 wt% of the erupted
73 magma. Volatile accumulation was proposed to occur in the roof zone of the system. On
74 22 April 2015, the Chilean volcano Calbuco produced a sub-Plinian eruption (Castruccio
75 et al., 2016; Romero et al., 2016; Arzilli et al., 2019) with two explosive phases. The first
76 was found to be powered by an excess gas phase with three times the amount of SO₂ es-
77 timated to be produced by the erupted mass (Pardini et al., 2018). In highly silicic sys-
78 tems, the volume of erupted products may be only a fraction of the magma reservoir vol-

79 ume, as eruptible magma is extracted from a large crystal mush (e.g. Bachmann & Bergantz,
80 2004). This creates the possibility that a voluminous volatile body is created within mag-
81 matic systems prior to eruption, ponding in the roof zone, producing both observed ex-
82 cess gas and a buoyancy force on the crust, arising from the volatiles' lower density (\sim
83 500 kg m^{-3}) compared with melt and crust. At a depth of 8 km and pressure of 200 MPa
84 the solubility of CO_2 in a basalt is ~ 700 ppm (Newman & Lowenstern, 2002), while the
85 initial CO_2 contents may be 1 wt% (10,000 ppm) or greater (Blundy et al., 2010). So
86 a significant free gas phase can be expected in magma reservoirs if the volatiles exsolve
87 but cannot escape to the surface. The purpose of this study is to examine the impact
88 of the sudden release of a large volume of exsolved volatiles and the associated loss of
89 buoyancy on the deformation field of a volcano.

90 To do that, we utilize the three-dimensional (3D) thermomechanical finite differ-
91 ences code LaMEM (Kaus et al., 2016) to model the stresses and deformation that a sud-
92 den change in the density field induces in the overlying crust and at the surface. LaMEM
93 solves the density dependent Stokes equations for (nearly) incompressible visco-elasto-
94 plastic fluid flow and runs on massively parallel clusters, allowing us to use high reso-
95 lutions, even in 3D. The code has already been applied to magmatic systems before (e.g.
96 Reuber et al., 2018; Piccolo et al., 2020; Spang et al., 2021).

97 **2 Methods**

98 Section 2.1 introduces the software used for modeling as well as the physics and
99 governing equations. Section 2.2 presents the model setup and the parameters used. Sec-
100 tion 2.3 describes the key parameters that we identified and our approach to deriving
101 a scaling law for the surface deformation due to volatile release. In section 2.4, we in-
102 troduce our area of application, the Chilean volcano Calbuco.

103 **2.1 Thermomechanical Code**

104 The 3D thermomechanical finite differences code LaMEM (Kaus et al., 2016) was
105 used to calculate deformation due to magmatic sources hosted in a finite-size model do-
106 main. The code solves for the conservation of momentum, mass and energy (eq. 1-3),
107 using a staggered grid in combination with a marker-in-cell approach (Harlow & Welch,
108 1965).

$$\frac{\partial \tau_{ij}}{\partial x_j} - \frac{\partial p}{\partial x_i} + \rho g_i = 0 \quad (1)$$

$$\frac{\partial v_i}{\partial x_i} = 0 \quad (2)$$

$$\rho C_p \frac{DT}{Dt} = \frac{\partial}{\partial x_i} \left(k \frac{\partial T}{\partial x_i} \right) + H \quad (3)$$

109 τ_{ij} is the Cauchy stress deviator, x_i ($i = 1, 2, 3$) denotes the Cartesian coordinates, p
 110 is pressure (positive in compression), ρ density, g_i gravitational acceleration, v_i the ve-
 111 locity vector, C_p the specific heat capacity, T the temperature, k the thermal conduc-
 112 tivity, H the volumetric heat source and D/Dt is the material time derivative.

113 Free slip conditions are applied to the boundaries of the model domain, allowing
 114 movement parallel to the domain edges while setting perpendicular velocities to 0. At
 115 the top of the setup, we include 1 km of sticky air above the stabilized free surface (Duretz
 116 et al., 2011; Kaus et al., 2010). The rocks are characterized by a temperature- and strain
 117 rate-dependent visco-elasto-plastic rheology where the strain rate is the sum of the elas-
 118 tic, viscous and plastic components:

$$\dot{\epsilon}_{ij} = \dot{\epsilon}_{ij}^{el} + \dot{\epsilon}_{ij}^{vi} + \dot{\epsilon}_{ij}^{pl} \quad (4)$$

119 $\dot{\epsilon}_{ij}$ denotes the total deviatoric strain rate tensor, while $\dot{\epsilon}_{ij}^{el}$, $\dot{\epsilon}_{ij}^{vi}$ and $\dot{\epsilon}_{ij}^{pl}$ represent the elas-
 120 tic, viscous and plastic strain rate components. A detailed discussion of this equation
 121 and all of its components is given by Kaus et al. (2016), but here we will focus on the
 122 material parameters which impact the three components.

123 The elastic component $\dot{\epsilon}_{ij}^{el}$ is inverse proportional to the shear modulus G :

$$\dot{\epsilon}_{ij}^{el} = \frac{1}{2G} \frac{D\tau_{ij}}{Dt}, \quad (5)$$

124 where $D\tau_{ij}/Dt$ corresponds to the objective derivative of the stress tensor.

125 The viscous strain rate component $\dot{\epsilon}_{ij}^{vi}$ is governed by the viscosity η , which follows
 126 the temperature- and strain rate-dependent powerlaw relationship of dislocation creep:

$$\eta = \frac{1}{2} (B_n)^{-\frac{1}{n}} (\dot{\epsilon}_{II})^{\frac{1}{n}-1} \exp\left(\frac{E_n}{nRT}\right), \quad (6)$$

127 where B_n is the creep constant, $\dot{\epsilon}_{II}$ the square root of the second invariant of the strain
 128 rate ($\dot{\epsilon}_{II} = (\frac{1}{2}\dot{\epsilon}_{ij}\dot{\epsilon}_{ij})^{1/2}$), E_n the activation energy, n the powerlaw exponent, R the uni-
 129 versal gas constant and T the temperature.

130 The plastic component is characterized by the Drucker-Prager failure criterion (Drucker
 131 & Prager, 1952) which is a good approximation of Byerlee's law (Byerlee, 1978):

$$\tau_{II} \leq \sin(\phi)p + \cos(\phi)c_0 \quad (7)$$

132 τ_{II} is the square root of the second invariant of the stress tensor ($\tau_{II} = (\frac{1}{2}\tau_{ij}\tau_{ij})^{1/2}$),
 133 ϕ is the friction angle, p the pressure and c_0 the cohesion. Equation 7 describes how much
 134 stress can be accommodated with visco-elastic deformation.

135 As buoyancy is the driving force in our model, we need densities to be independent
 136 of temperature (i.e. no thermal expansion) and pressure (i.e. incompressible). For the
 137 volatile reservoir, we use the ideal gas law to estimate density (see supplementary text
 138 S1).

139 2.2 Model Setup and Parameter Selection

140 Obtaining a quantitative understanding of ground deformation requires the use of
 141 3D models, but as they are computationally expensive, we do initial testing in 2D which
 142 allows an efficient evaluation of the respective importance of various model parameters.

143 Our reference model uses a homogeneous crust, hosting a spherical, low-viscosity,
 144 non-buoyant magma reservoir with a radius of 1 km. On top of the magma body, we place
 145 a reservoir of exsolved and accumulated volatiles which we approximate as a sphere (r
 146 $= 250$ m) of low density, viscosity and rigidity. We use a non-buoyant magma body to
 147 focus on the volatiles' contribution to surface deformation. To maximize the resolution,
 148 we use only one quarter (half for 2D) of the perfectly symmetric domain. Figures S1a,
 149 S1b and S2a show that we do not introduce any effects through this simplification. Us-
 150 ing 384 cells in each direction yields roughly 56.6 million cells and a vertical resolution
 151 of about 40 meters (Figure 1a). As the model extends 15 km in the vertical and 50 km
 152 in horizontal directions, the horizontal cell size is only 40 meters in the central 2.5 km
 153 and then increases towards the horizontal edges of the domain. Resolution tests confirm
 154 that this is sufficient (Figures S1a, S1b and S2b). We run the model for 40 years with

155 a constant time step of 1 year. This domain width and time stepping allows the eval-
 156 uation of reliable models (see section 3.1).

157 To approximate the release of the exsolved volatiles from the system during erup-
 158 tion, they are instantaneously replaced by non-buoyant magma after 20 years (the time
 159 of eruption in the model). In reality, this (and the eruption of magma) would lead to a
 160 volume change in the magma reservoir. However, in this study we solely focus on the ef-
 161 fect of a loss of buoyancy associated with the release of exsolved volatiles and therefore
 162 do not consider other contributions to surface deformation.

163 Supplementary table S1 shows the parameters we use for the different model ma-
 164 terials. The rheology of the crust follows the powerlaw relationship of dislocation creep
 165 of wet quartzite (Ranalli, 1995) while magma and volatile reservoir are linear visco-elasto-
 166 plastic. We use a shear modulus of 2 GPa, in line with upscaled values from laboratory
 167 experiments on volcanic rocks (Heap et al., 2020). Cohesion and friction angle of intact
 168 rocks are typically estimated in the range of a few MPa and 30° respectively (Hoek &
 169 Brown, 1997), so we use 5 MPa and 20° for the presumably pre-damaged crust of a mag-
 170 matic system. The thermal conductivity is 3 W (mK)^{-1} and the heat capacity $1000 \text{ J (molK)}^{-1}$
 171 for all materials. We employ a background thermal gradient of 30 K km^{-1} and set the
 172 initial temperature of volatiles and magma to 800°C . Before we start the mechanical
 173 model ($t = 0$), we allow for 50 ka of thermal diffusion to account for the heated crust
 174 around a magma body which we keep at a constant temperature during this heating phase.

175 **2.3 Scaling Law for Deflation**

176 Through initial testing, we identified four key parameters that influence the ground
 177 deformation (Figure 1b). The radius of the volatile reservoir (r_{vol}), the depth of the volatile
 178 reservoir (d_{vol}), the density contrast between volatiles and crust ($\Delta\rho$) and the shear mod-
 179 ulus of the crust (G_{crust}). In both 2D and 3D, we run a set of systematic parameter com-
 180 binations to derive a scaling law for the deflation at the surface due to the volatile re-
 181 moval. For each parameter, we test 5 (9 for d_{vol}) different values while keeping the oth-
 182 ers constant.

2.4 Calbuco

The Chilean stratovolcano Calbuco erupted on April 22, 2015 and interferometric analysis of synthetic aperture radar images (InSAR analysis) from the Sentinel-1 satellites revealed a co-eruptive subsidence of about 12 cm (Nikkhoo et al., 2016; Delgado et al., 2017). Using different analytical solutions, Nikkhoo et al. (2016) and Delgado et al. (2017) reproduce the surface deformation with deflating sources at a depth of around 8 km. Petrological estimates for the location of the magma storage zone range from 5.5 to 12 km depth (Morgado et al., 2019; Arzilli et al., 2019; Namur et al., 2020). Namur et al. (2020) also suggest that magma moved to shallower levels weeks or month prior to eruption.

Pardini et al. (2018) found that a pre-existing volatile phase must have been present to explain $1.5 \cdot 10^8$ kg of excess SO_2 produced by the Calbuco eruption. Assuming a typical arc-magma H_2O abundance of 3 wt% (100 times the SO_2 content (Pardini et al., 2018)), we expect that this pre-exsolved volatile phase would contain $1.5 \cdot 10^{10}$ kg of H_2O (Spilliaert et al., 2006). CO_2 is much less soluble than H_2O or SO_2 and therefore to calculate the pre-eruptive CO_2 content we conservatively estimate that the total mass of the magma reservoir was ten times (Annen et al., 2008) the erupted mass of $4.9 \cdot 10^{11}$ kg (Pardini et al., 2018). Assuming that 0.5–1.5 wt% of CO_2 (Blundy et al., 2010) was exsolved prior to eruption leads to pre-exsolved CO_2 masses of $2.5 \cdot 10^{10}$ – $7.4 \cdot 10^{10}$ kg. We therefore estimate that the total pre-exsolved gas mass was $3.9 \cdot 10^{10}$ – $8.8 \cdot 10^{10}$ kg. We consider two scenarios, with volatiles stored at 5.5 km and 8 km depth and use the ideal gas law and lithostatic pressure to estimate the density of the volatile reservoir (see supplementary text S1).

3 Results

Section 3.1 describes the general behavior of the model and discusses dependencies on time stepping and size of the model domain. In section 3.2, we discuss the effects of changing the geometry of the magma body and volatile reservoir as well as the thermal structure of the crust. Section 3.3 describes the derivation of the scaling law in 3D. The 2D scaling law is discussed in supplementary text S2 and the differences between 2D and 3D in supplementary text S3. In section 3.4 we apply the scaling law derived in section 3.3 to the case of the 2015 Calbuco eruption.

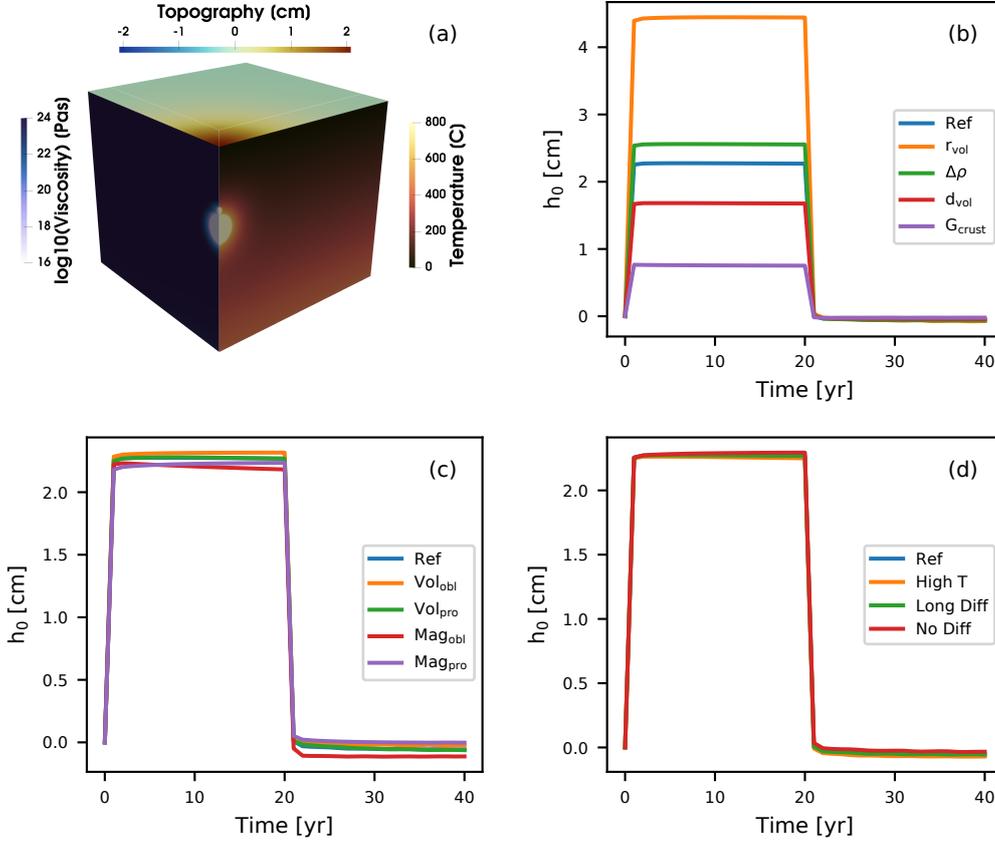


Figure 1. (a) 3D model setup showing the temperature on the right panel and viscosity on the left panel. Top panel shows surface elevation before the release of the volatiles. The 2D setup is a slice along the boundary of the 3D model. 2D and 3D model both extend 50 km in lateral direction(s) but that part of the figure was cut to enlarge the relevant features. (b-d) Surface level directly above the sources calculated in 3D models. (b) Effect of perturbing one crucial material parameter compared to the reference model. (c) Effect of changing the shape of the magma and volatile reservoir in comparison to reference model. Subscript *obl* means oblate, subscript *pro* means prolate. The volume was preserved in all cases. (d) Effect of changing the temperature structure in the crust through higher reservoir temperature or longer/no initial diffusion time. See section 3.2 for details.

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3.1 General Behavior

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At the start of each simulation, the surface above the buoyant volatile reservoir undergoes immediate uplift, and quickly (within 2 time steps) reaches a steady state. Upon replacing the volatiles with non-buoyant magma (i.e. an eruption), the surface quickly (within 2 time steps) returns to the original level. Independently of the time step we employ (0.1 - 10 years), the surface reaches the same level after 2 steps with the first step being very close to it already (Figures S1c and S2d). We observe the same behavior after removing the volatile reservoir. We therefore conclude that the surface response is immediate and has no time dependence. The small adjustment, necessary in the second time step, is inferred to be of numerical origin. To minimize computational cost and enable us to observe any potential time dependencies, we use a time step of 1 year for all our models. In reality, the uplift or inflation phase may take place over a long time as the volatile reservoir grows gradually, but will reach the same magnitude as in our models. Volcano deflation, however happens often on timescales of eruption as all volatiles are expected to reach the surface, once a pathway has been established.

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Supplementary figures S1d and S2c show that the surface displacement depends on the width of the model domain. The displacement increases with increasing model width but at 50 km width, the effect levels off. We therefore ran all models with a width of 50 km.

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We do not observe plastic failure in any of our models. Even after reducing cohesion (c_0) by an order of magnitude to 0.5 MPa and friction angle (ϕ) to 10° while increasing r_{gas} to 500 meters and G_{crust} to 10 GPa to maximize crustal stresses, stresses due to changes in buoyancy never exceed a few MPa which is insufficient to exceed the Drucker-Prager failure criterion.

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3.2 Influence of Source Geometry and Thermal Structure

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In Figure 1c, we show the results of testing different shapes for the magma and volatile reservoirs. Both the oblate and prolate shapes have an aspect ratio of 2 while preserving the volume of the spherical version. None of the geometrical variations lead to a significant difference in vertical displacement.

243 Figure 1d shows the effect of changing the thermal structure of the crust. In the
 244 'No Diff' example, we omit the 50 ka of thermal diffusion and start with a crust that only
 245 has the background temperature gradient while in the 'Long Diff' example, we double
 246 the temperature diffusion time from 50 to 100 ka. For the 'High T' example, we set the
 247 magma and volatile temperature to 1000 °C instead of 800 °C. The surface response is
 248 almost identical with the reference model for all cases.

249 3.3 3D Scaling law

250 Figure 1b shows the effect of varying four material parameters that have a consid-
 251 erable effect on the surface response. The radius and depth of the volatile reservoir (d),
 252 the density contrast between volatiles and crust ($\Delta\rho$) and the shear modulus of the crust
 253 (G). We performed a systematic parameter variation, testing 5 different values for each
 254 parameter (9 for d_{fl}) while keeping the other parameters constant. Supplementary fig-
 255 ure S5 shows the results for individual parameters. From this, we are able to derive the
 256 following scaling relationship:

$$\Delta h_0 = A \frac{r^3 \Delta\rho g}{d^{3/2} G} \quad (8)$$

257 where Δh_0 is the vertical displacement at the surface above the source, g is the grav-
 258 itational acceleration and A is a pre-factor of 12π with units of $\text{m}^{0.5}$ to satisfy the equa-
 259 tion. The scaling law is similar to the one derived for the 2D case with the only excep-
 260 tion being the exponents of r and d (see equation S4). Fits for the individual paramet-
 261 ers are shown in Figure S5.

262 Figure 2a shows how equation 8 predicts the results of the 3D models and Figure
 263 2b shows the error which is generally smaller than 10%. Green crosses in Figures 2a and
 264 2b show models where we changed multiple parameters to validate equation 8. Subsidi-
 265 dence in models with a shallow ($d \leq 2 \text{ km}$) volatile reservoir are overestimated. An-
 266 alytical solutions for the gravity anomaly of buried cylinders or spheres have the same
 267 issue of only being applicable when the depth of the body is much larger than its radius
 268 (Turcotte & Schubert, 2002). The same is true for simple models relating the inflation
 269 of magma bodies to surface deformation (e.g. Mogi, 1958; Yang et al., 1988).

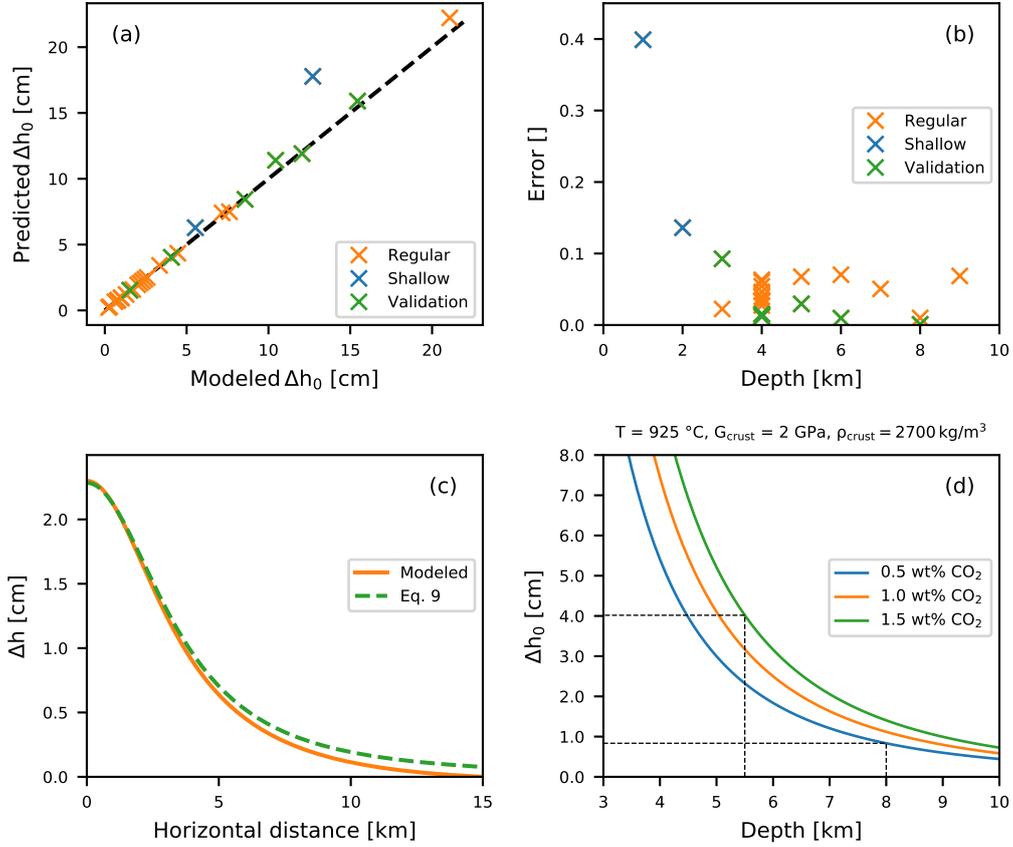


Figure 2. (a) Comparison between modeled subsidence and subsidence predicted by equation 8. Black dashed line shows 1:1 correlation. Orange crosses show models that were used to derive equation 8 and green crosses show models used to validate the scaling law. Blue crosses show models with a shallow (≤ 2 km) volatile reservoir that were excluded. (b) Error between the modeled subsidence and the predicted subsidence as a function of depth. (c) Subsidence along a radial profile for the reference case in orange in comparison to the prediction made by equation 9. (d) Coupling between the ideal gas law and equation 8 to show the full dependence of maximum subsidence Δh_0 on reservoir depth d . Solid lines denote different volatile mass estimates as discussed in section 2.4. Dashed black lines show how we determine the minimum and maximum of subsidence in section 3.4.

270 Equation 8 only describes the vertical displacement directly above the center of the
 271 volatile reservoir. Figure 2c shows a profile of the vertical change along the surface. Our
 272 numerical models show that we can modify equation 8 to:

$$\Delta h(x) = A \frac{r^3 \Delta \rho g}{G} \frac{d}{(d^2 + x^2)^{5/4}} \quad (9)$$

273 where x is the horizontal distance from the center of the volatile reservoir projected to
 274 the surface. Figure 2c shows that the modeled surface displacement is fit well by equa-
 275 tion 9. In this form, our scaling law is very similar to the analytical solution of ground
 276 deformation due to a point source of pressure within an elastic halfspace, the "Mogi model"
 277 (Mogi, 1958). The most notable difference being the exponent of 5/4 instead of 3/2, which
 278 stems from the depth dependence of $\frac{1}{d^{3/2}}$ (see equation 8 and Supplementary figure S5c)
 279 while the "Mogi model" has a depth dependence of $\frac{1}{d^2}$.

280 3.4 Calbuco

281 For Calbuco, we use equation 8 with $\rho_{\text{crust}} = 2700 \text{ kg m}^{-3}$ and $G_{\text{crust}} = 2 \text{ GPa}$ to
 282 predict a maximum surface subsidence of 4 cm due to the loss of buoyancy from $8.8 \cdot 10^{10}$
 283 kg of exsolved volatiles for the case of storage at 5.5 km depth. For the 8 km depth sce-
 284 nario, and a lower limit estimate of the erupted gas mass ($3.9 \cdot 10^{10}$ kg), we predict 1 cm.
 285 Equations 8 and 9 imply that the surface displacement depends on the reservoir depth
 286 to the power of 1.5. In reality, r and $\Delta \rho$ are also functions of the pressure in the volatile
 287 reservoir and thereby of the depth. Figure 2d illustrates this nonlinear dependence and
 288 shows how we arrive at our minimum and maximum estimates.

289 4 Discussion

290 4.1 Rheology

291 Given that, even for rocks with considerably lowered plastic strength, the stresses
 292 caused by the changes in buoyancy are not sufficient to exceed the failure criterion, plas-
 293 ticity is not a factor for the process we are investigating. Figure 1d also suggests that
 294 on the timescales of an eruption, viscous components have no impact on the deforma-
 295 tion, even with the weakening caused by heating of the crust. The process of surface sub-
 296 sidence caused by the loss of a buoyant volatile reservoir due to eruption can therefore

297 be considered as quasi-elastic, and as a result it is possible to derive a scaling law for the
 298 problem.

299 **4.2 Surface Subsidence due to Buoyancy Loss**

300 Instantly (on the timescale of an eruption), removing a buoyant volatile reservoir
 301 from the top of an upper crustal magma body leads to an instantaneous subsidence. The
 302 magnitude of subsidence decays with radial distance from the reservoir center, but is sig-
 303 nificant in a radius of several kilometers (Figure 2c). The surface response is insensitive
 304 to the temperature structure (Figure 2d) of the crust which allows us to derive a scal-
 305 ing law for the expected subsidence (equations 8 and 9). As the surface deformation is
 306 also independent of the shape of the volatile reservoir (Figure 1c), we suggest this alter-
 307 native form of equation 9:

$$\Delta h(x) = \frac{9V\Delta\rho g}{G} \frac{d}{(d^2 + x^2)^{5/4}} \quad (10)$$

308 where V is the volume of the volatile reservoir. As other analytical solutions for the sur-
 309 face effects of buried bodies, the scaling law's accuracy decreases when the ratio between
 310 radius and depth of the body exceeds 0.1 (Figure 2b).

311 The inferred scaling law (equations 8, 9 and 10) has a similar structure to the Mogi
 312 model including a pre-factor, a cubic dependence on radius, an elastic property of the
 313 crust and a term describing the decay of the signal with distance. One difference is the
 314 term of the driving force of deformation. In the Mogi model, it is either a pressure or
 315 a volume change, while in our scaling law, it is buoyancy. The other notable difference
 316 is the exponent of the depth dependence (2 for Mogi and 1.5 in our model). This could
 317 be caused by the different mechanisms that are at work. The pressure point source of
 318 the Mogi model applies a pressure to the surrounding crust in all directions, while in our
 319 case, buoyancy is expected to exert a cumulative upwards force in line with Archimedes'
 320 principle (e.g. Sigmundsson et al., 2020).

321 Another difference to common scaling laws for volcano deformation (e.g. Mogi, 1958;
 322 McTigue, 1987) is the lack of compressibility in our models because of its complex in-
 323 terplay with densities. As vertical displacement is usually multiplied by the term $(1 -$

324 ν), our scaling law might provide a minimum estimate as a commonly used Poisson's ra-
325 tio of $\nu = 0.25$ results in a larger factor than incompressibility ($\nu = 0.5$).

326 **4.3 Calbuco**

327 Applying our scaling law to the case of the 2015 Calbuco eruption, yields a sub-
328 sidence of 1–4 cm (Figure 2d). With an incidence angle of 33° (Delgado et al., 2017),
329 these vertical velocities can be projected into line-of-sight displacement (Fialko et al.,
330 2001b) and represent 7% to 28% of the observed surface deformation. This is an indi-
331 cation that the majority of co-eruptive subsidence was caused by the volumetric loss of
332 material (volatiles and magma) but a significant part of the signal may originate from
333 the loss of buoyancy provided by a body of exsolved volatiles.

334 In fact, the best-fit sphere and spheroid models of Delgado et al. (2017) have a resid-
335 ual of about 3 cm in the center of subsidence. The mechanism described in our work pro-
336 vides an additional source of uplift, large enough to cover this misfit entirely.

337 **4.4 Implications for Modeling Volcanic Deformation**

338 The release of a buoyant body of exsolved volatiles from the top of an upper crustal
339 magma reservoir can lead to significant (on the order of a few cm) syn-eruptive defla-
340 tion at the surface. This effect is likely smaller than the effect of volume change in vol-
341 canic roots during eruptions as magma moves to the surface. In the case of Calbuco, the
342 contributions may have a ratio between 3:1 and 10:1 in favor of the volume loss. This
343 ratio depends, however, on the quantity of pre-exsolved volatiles.

344 Adding equation 10 to existing models could be a simple way of achieving a bet-
345 ter fit to the observed deformation while also providing an explanation for the excess gas
346 that is detected for a number of eruptions.

347 As Figure 1 shows, the presence of a buoyant body of exsolved volatiles also causes
348 surface uplift of the same magnitude as its removal causes subsidence. That means that
349 inflation of a few centimeters over time, which is traditionally interpreted to be a sign
350 of magma intrusion at depth, could also be caused by the formation of a body of exsolved
351 volatiles at the top of the magma reservoir.

352 Furthermore, magma is usually buoyant at the depth where it intrudes. So even
353 if the intruded magma does not form a significant volatile reservoir, it still exerts a buoy-
354 ancy force on the crust that adds to the surface deformation caused by displacing host
355 rock. Although the effect of magma buoyancy on surface deformation was not explic-
356 itly investigated here, it is likely that equation 10 also gives a good estimate of its ef-
357 fect and could be added to existing solutions for surface uplift.

358 **5 Conclusions**

359 We conducted a series of 3D visco-elasto-plastic models to investigate the surface
360 deformation caused by the instant removal of a buoyant reservoir of exsolved volatiles
361 from the top of a magma body, as would be the case during an eruption. Our results show
362 that the removal of the reservoir causes subsidence at the surface which is independent
363 of the shape of the volatile and magma reservoirs as well as from the thermal state of
364 the crust. Instead, the process is quasi-elastic, allowing us to derive an analytical solu-
365 tion for the surface subsidence including the volume and depth of the reservoir, the den-
366 sity contrast between volatiles and crust, as well as the shear modulus of the crust. This
367 analytical solution predicts surface deformations on the order of up to a few centime-
368 ters.

369 We applied our scaling law to the case of the 2015 Calbuco eruption and, depend-
370 ing on the depth of the reservoir and volatile mass, predict subsidence of 1–4 cm, which
371 is about 20% of the observed signal. We expect that most of the observed surface de-
372 formation is caused by the volume loss of volatiles and magma.

373 Adding our scaling law to existing models for volcano deformation would present
374 a step forward, towards models that include all the relevant mechanisms that occur in
375 volcanic roots.

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385 **Open Research Section**

386 Software for this research is available on zenodo at:

387 LaMEM (Kaus et al., 2016):
 388 <http://doi.org/10.5281/zenodo.5734975>

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