

Quantification of Volcano Deformation caused by Volatile Accumulation and Release

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Key Points:

- Exsolving volatiles accumulate at the roof of a magma storage zone and contribute to surface deformation through buoyancy forces
- 3D numerical models show that surface deformation is a function of the volatiles' volume, density and depth as well as crustal rigidity
- Volatile release during eruption can cause syn-eruptive subsidence of a few cm, which is 20% of the observed signal at Calbuco in 2015

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Abstract

Crustal-stored magma reservoirs contain exsolved volatiles which accumulate in the reservoir roof, exerting a buoyancy force on the crust. This produces surface uplift and sudden loss of volatiles through eruption results in syn-eruptive subsidence. Here, we present three-dimensional, visco-elasto-plastic, numerical modeling results which quantify the ground deformation arising from the growth and release of a volatile reservoir. Deformation is mostly independent of crustal thermal distribution and volatile reservoir shape, but is a function of volatile volume, density and depth and crustal rigidity. We present a scaling law for the volatiles' contribution to syn-eruptive subsidence and show this contributes $\sim 20\%$ of the observed subsidence associated with the 2015 Calbuco eruption. Our results highlight the key role that volatile-driven buoyancy can have in volcano deformation, show a new link between syn-eruptive degassing and deflation, and highlight that shallow volatile accumulation and release may have a significant impact on ground deformation of volcanoes.

Plain Language Summary

Magma contains a lot of gases which separate from it when it approaches the surface. These gases can collect right above the magma storage region a few kilometers below the surface. They have a much lower density than the rocks surrounding them and push upwards like a balloon filled with air that is pressed under water. In this study, we use computer models to understand how much a volcano would grow from the push of the gases below and how much it would shrink when the gases escape because of an eruption. We find that the gases can cause the volcano to grow and shrink up to a few centimeters during accumulation and release, respectively. The amount of surface movement depends on the volume, density and depth of the gas reservoir as well as on the toughness of the rocks above it. We derive a simple equation which allows us to compute the surface movement using the aforementioned parameters. With this equation and estimates about the amount of accumulated gas at the 2015 Calbuco eruption, we can assume that about 20% of the observed surface movement was caused by the release of the magmatic gases.

Index Terms

8145 Physics of magma and magma bodies

48 1211 Non-tectonic deformation
49 0545 Modeling
50 8430 Volcanic gases
51 1036 Magma chamber processes

52 **Keywords**

53 3D modeling, Buoyancy, Volcanic uplift, Scaling analysis, Magmatic gases, Calbuco

54 **1 Introduction**

55 Volcano deformation is most frequently interpreted in terms of models of surface
56 deformation due to processes in magma bodies of various geometries. The most widely
57 applied model is that of a point source of pressure embedded within a uniform elastic
58 halfspace (Mogi, 1958), but a range of more advanced models and approaches exist (e.g.
59 Fialko et al., 2001a; Hickey et al., 2016). As liquid magma flows in/out of the these de-
60 formation sources, they expand/contract. Most often, such magma flow is considered to
61 cause uniform pressure change on the boundary of the magma body, and the density dif-
62 ference between magma and host rock is not considered specifically. It has, however, been
63 demonstrated in a number of studies that magma buoyancy can cause significant stresses
64 in volcano roots and contribute to failure of magma bodies (e.g. Sigmundsson et al., 2020).
65 A particular phenomena not considered by traditional volcano deformation models is the
66 effect of accumulated exsolved volatiles in volcano roots and their release during erup-
67 tions.

68 During major explosive eruptions an excess of gas may be observed, beyond that
69 which can be explained by a petrological calculation of the original dissolved volatile amounts
70 and the volume of erupted lavas. Excess gas was observed in the 1991 eruption of Pinatubo,
71 Philippines and an analysis from Wallace and Gerlach (1994) showed that this could be
72 explained by a pre-existing gas/volatile phase representing 0.7 to 1.3 wt% of the erupted
73 magma. Volatile accumulation was proposed to occur in the roof zone of the system. On
74 22 April 2015, the Chilean volcano Calbuco produced a sub-Plinian eruption (Castruccio
75 et al., 2016; Romero et al., 2016; Arzilli et al., 2019) with two explosive phases. The first
76 was found to be powered by an excess gas phase with three times the amount of SO₂ es-
77 timated to be produced by the erupted mass (Pardini et al., 2018). In highly silicic sys-
78 tems, the volume of erupted products may be only a fraction of the magma reservoir vol-

79 ume, as eruptible magma is extracted from a large crystal mush (e.g. Bachmann & Bergantz,
 80 2004). This creates the possibility that a voluminous volatile body is created within mag-
 81 matic systems prior to eruption, ponding in the roof zone, producing both observed ex-
 82 cess gas and a buoyancy force on the crust, arising from the volatiles lower density (\sim
 83 500 kg m^{-3}) compared with melt and crust. At a depth of 8 km and pressure of 200 MPa,
 84 the solubility of CO_2 in a basalt is ~ 700 ppm (Newman & Lowenstern, 2002), while the
 85 initial CO_2 contents may be 1 wt% (10,000 ppm) or greater (Blundy et al., 2010). So
 86 a significant free volatile phase can be expected in magma reservoirs if the volatiles ex-
 87 solve but cannot escape to the surface. The purpose of this study is to examine the im-
 88 pact of the sudden release of a large volume of exsolved volatiles and the associated loss
 89 of buoyancy to estimate the significance of this process for volcano deformation mod-
 90 eling.

91 To do that, we utilize the three-dimensional (3D) thermomechanical finite differ-
 92 ences code LaMEM (Kaus et al., 2016) to model the stresses and deformation that a sud-
 93 den change in the density field induces in the overlying crust and at the surface. LaMEM
 94 solves the density dependent Stokes equations for (nearly) incompressible visco-elasto-
 95 plastic fluid flow and runs on massively parallel clusters, allowing us to use high reso-
 96 lutions, even in 3D. The code has already been applied to magmatic systems before (e.g.
 97 Reuber et al., 2018; Piccolo et al., 2020; Spang et al., 2021).

98 **2 Methods**

99 Section 2.1 introduces the software used for modeling as well as the physics and
 100 governing equations. Section 2.2 presents the model setup and the parameters used. De-
 101 tails on model resolution, time stepping and resolution tests are presented in supplemen-
 102 tary text S1. Section 2.3 describes the key parameters that we identified and our approach
 103 to deriving a scaling law for the surface deformation due to volatile release. In section
 104 2.4, we introduce our area of application, the Chilean volcano Calbuco.

105 **2.1 Thermomechanical Code**

106 The 3D thermomechanical finite differences code LaMEM (Kaus et al., 2016) was
 107 used to calculate deformation due to magmatic sources hosted in a finite-size model do-
 108 main. The code solves for the conservation of momentum, mass and energy (eq. 1-3),

109 using a staggered grid in combination with a marker-in-cell approach (Harlow & Welch,
110 1965).

$$\frac{\partial \tau_{ij}}{\partial x_j} - \frac{\partial p}{\partial x_i} + \rho g_i = 0 \quad (1)$$

$$\frac{\partial v_i}{\partial x_i} = 0 \quad (2)$$

$$\rho C_p \frac{DT}{Dt} = \frac{\partial}{\partial x_i} \left(k \frac{\partial T}{\partial x_i} \right) + H \quad (3)$$

111 τ_{ij} is the Cauchy stress deviator, x_i ($i = 1, 2, 3$) denotes the Cartesian coordinates, p
112 is pressure (positive in compression), ρ density, g_i gravitational acceleration, v_i the
113 velocity vector, C_p the specific heat capacity, T the temperature, k the thermal conduc-
114 tivity, H the volumetric heat source and D/Dt is the material time derivative.

115 Free slip conditions are applied to the boundaries of the model domain, allowing
116 movement parallel to the domain edges while setting perpendicular velocities to 0. At
117 the top of the setup, we include 1 km of sticky air above the stabilized free surface (Duretz
118 et al., 2011; Kaus et al., 2010). The rocks are characterized by a temperature- and strain
119 rate-dependent visco-elasto-plastic rheology where the strain rate is the sum of the elas-
120 tic, viscous and plastic components:

$$\dot{\epsilon}_{ij} = \dot{\epsilon}_{ij}^{el} + \dot{\epsilon}_{ij}^{vi} + \dot{\epsilon}_{ij}^{pl} \quad (4)$$

121 $\dot{\epsilon}_{ij}$ denotes the total deviatoric strain rate tensor, while $\dot{\epsilon}_{ij}^{el}$, $\dot{\epsilon}_{ij}^{vi}$ and $\dot{\epsilon}_{ij}^{pl}$ represent the elas-
122 tic, viscous and plastic strain rate components. A detailed discussion of this equation
123 and all of its components is given by Kaus et al. (2016), but here we will focus on the
124 material parameters which impact the three components.

125 The elastic component $\dot{\epsilon}_{ij}^{el}$ is inverse proportional to the shear modulus G :

$$\dot{\epsilon}_{ij}^{el} = \frac{1}{2G} \frac{D\tau_{ij}}{Dt}, \quad (5)$$

126 where $D\tau_{ij}/Dt$ corresponds to the objective derivative of the stress tensor.

127 The viscous strain rate component $\dot{\epsilon}_{ij}^{vi}$ is governed by the viscosity η , which follows
128 the temperature- and strain rate-dependent powerlaw relationship of dislocation creep:

$$\eta = \frac{1}{2}(B_n)^{-\frac{1}{n}}(\dot{\epsilon}_{II})^{\frac{1}{n}-1}\exp\left(\frac{E_n}{nRT}\right), \quad (6)$$

129 where B_n is the creep constant, $\dot{\epsilon}_{II}$ the square root of the second invariant of the strain
 130 rate ($\dot{\epsilon}_{II} = (\frac{1}{2}\dot{\epsilon}_{ij}\dot{\epsilon}_{ij})^{1/2}$), E_n the activation energy, n the powerlaw exponent, R the uni-
 131 versal gas constant and T the temperature.

132 The plastic component is characterized by the Drucker-Prager failure criterion (Drucker
 133 & Prager, 1952) which is a good approximation of Byerlee's law (Byerlee, 1978):

$$\tau_{II} \leq \sin(\phi)p + \cos(\phi)c_0 \quad (7)$$

134 τ_{II} is the square root of the second invariant of the stress tensor ($\tau_{II} = (\frac{1}{2}\tau_{ij}\tau_{ij})^{1/2}$),
 135 ϕ is the friction angle, p the pressure and c_0 the cohesion. Equation 7 describes how much
 136 stress can be accommodated with visco-elastic deformation.

137 As buoyancy is the driving force in our model, we need densities to be independent
 138 of temperature (i.e. no thermal expansion) and pressure (i.e. incompressible). For the
 139 volatile reservoir, we use the ideal gas law to estimate density (see supplementary text
 140 S2).

141 2.2 Model Setup and Parameter Selection

142 Obtaining a quantitative understanding of ground deformation requires the use of
 143 3D models, but as they are computationally expensive, we do initial testing in 2D which
 144 allows an efficient evaluation of the respective importance of various model parameters.

145 Our reference model (Figure 1a) uses a homogeneous crust, hosting a spherical, low-
 146 viscosity, non-buoyant magma reservoir with a radius of 1 km. As exsolved volatiles are
 147 expected to accumulate in the roof of the magmatic system, we approximate the volatile
 148 reservoir as a sphere ($r = 250$ m) of low density, viscosity and rigidity on top of the magma
 149 body. We use a non-buoyant magma body to focus on the volatiles' contribution to sur-
 150 face deformation. It still provides heat to the surrounding host rock and mechanically
 151 decouples the volatile reservoir from the underlying crust.

152 To approximate the release of the exsolved volatiles from the system during erup-
 153 tion, they are instantaneously replaced by non-buoyant magma after 20 years (the time

154 of eruption in the model). This is accomplished by copying the material properties (den-
 155 sity, viscosity and shear modulus) of the magma onto the volatile particles. The change
 156 in density induces a change in crustal stresses (see equation 1, supplementary Figure S1).

157 In reality, an eruption does not only remove the buoyancy forces of the volatile reser-
 158 voir but also the volatiles themselves as well as part of the magma. As magma injection
 159 is a commonly suggested trigger for eruptions (Canon-Tapia, 2014), the erupted volume
 160 may be replaced by intruding magma from a deep source. If this is not or only partly
 161 the case, the loss of volume leads to a depressurization of the remaining magma reser-
 162 voir. This likely triggers a combination of three processes. (i) The remaining magma ex-
 163 solves more volatiles due to the drop in confining pressure which expand upon exsolu-
 164 tion. (ii) The magma itself expands due to depressurization. (iii) The overburden sub-
 165 sides and contracts to close the space left behind. All of these processes influence sur-
 166 face deformation alongside the loss of buoyancy, but as the contribution of each of the
 167 three aforementioned processes and magma injection is not understood, we focus on the
 168 change in buoyancy forces. This way, we can constrain the magnitude of this individ-
 169 ual contributor to syn-eruptive subsidence and estimate whether it is significant.

170 Supplementary table S1 shows the parameters we use for the different model ma-
 171 terials. The rheology of the crust follows the powerlaw relationship of dislocation creep
 172 of wet quartzite (Ranalli, 1995) while magma and volatile reservoir are linear visco-elasto-
 173 plastic. We use a shear modulus of 2 GPa, in line with upscaled values from laboratory
 174 experiments on volcanic rocks (Heap et al., 2020). Cohesion and friction angle of intact
 175 rocks are typically estimated in the range of a few MPa and 30° respectively (Hoek &
 176 Brown, 1997), so we use 5 MPa and 20° for the presumably pre-damaged crust of a mag-
 177 matic system. The thermal conductivity is 3 W (m K)^{-1} and the heat capacity $1000 \text{ J (mol K)}^{-1}$
 178 for all materials. We employ a background thermal gradient of 30 K km^{-1} and set the
 179 initial temperature of volatiles and magma to 800°C . Before we start the mechanical
 180 model ($t = 0$), we allow for 50 kyr of thermal diffusion to account for the heated crust
 181 around a magma body which we keep at a constant temperature during this heating phase.

182 **2.3 Scaling Law for Deflation**

183 Through initial testing, we identified four key parameters that influence the ground
 184 deformation (Figure 1b). The radius of the volatile reservoir (r_{vol}), the depth of the volatile

185 reservoir (d_{vol}), the density contrast between volatiles and crust ($\Delta\rho$) and the shear mod-
 186 ulus of the crust (G_{crust}). In both 2D and 3D, we run a set of systematic parameter com-
 187 binations to derive a scaling law for the deflation at the surface due to the volatile re-
 188 moval. For each parameter, we test 5 (9 for d_{vol}) different values while keeping the oth-
 189 ers constant.

190 2.4 Calbuco

191 The Chilean stratovolcano Calbuco erupted on April 22, 2015 and interferomet-
 192 ric analysis of synthetic aperture radar images (InSAR analysis) from the Sentinel-1 satel-
 193 lites revealed a co-eruptive subsidence of about 12 cm (Nikkhoo et al., 2016; Delgado et
 194 al., 2017). Using different analytical solutions, Nikkhoo et al. (2016) and Delgado et al.
 195 (2017) reproduce the surface deformation with deflating sources at a depth of around
 196 8 km. Petrological estimates for the location of the magma storage zone range from 5.5
 197 to 12 km depth (Morgado et al., 2019; Arzilli et al., 2019; Namur et al., 2020). Namur
 198 et al. (2020) also suggest that magma moved to shallower levels weeks or month prior
 199 to eruption.

200 Pardini et al. (2018) found that a pre-existing volatile phase must have been present
 201 to explain $1.5 \cdot 10^8$ kg of excess SO_2 produced by the Calbuco eruption. Assuming a typ-
 202 ical arc-magma H_2O abundance of 3 wt% (100 times the SO_2 content (Pardini et al., 2018)),
 203 we expect that this pre-exsolved volatile phase would contain $1.5 \cdot 10^{10}$ kg of H_2O (Spilliaert
 204 et al., 2006). CO_2 is much less soluble than H_2O or SO_2 and therefore to calculate the
 205 pre-eruptive CO_2 content we conservatively estimate that the total mass of the magma
 206 reservoir was ten times (Annen et al., 2008) the erupted mass of $4.9 \cdot 10^{11}$ kg (Pardini
 207 et al., 2018). Assuming that 0.5–1.5 wt% of CO_2 (Blundy et al., 2010) was exsolved prior
 208 to eruption leads to pre-exsolved CO_2 masses of $2.5 \cdot 10^{10}$ – $7.4 \cdot 10^{10}$ kg. We therefore es-
 209 timate that the total pre-exsolved gas mass was $3.9 \cdot 10^{10}$ – $8.8 \cdot 10^{10}$ kg. We consider two
 210 scenarios, with volatiles stored at 5.5 km and 8 km depth and use the ideal gas law and
 211 lithostatic pressure to estimate the density of the volatile reservoir (see supplementary
 212 text S2).

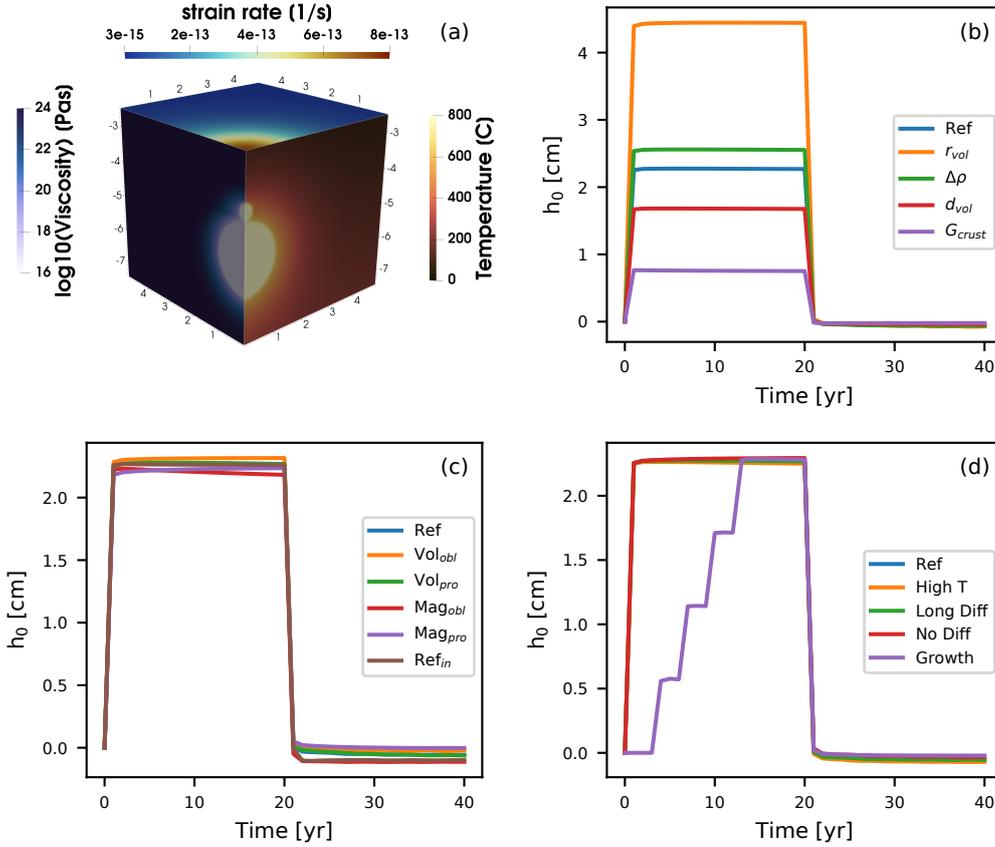


Figure 1. (a) Zoomed in part of 3D model setup showing the viscosity on the left panel, temperature on the right panel, and strain rate on the top panel before the release of the volatiles. The 2D setup is a slice along the boundary of the 3D model. 2D and 3D model both extend 50 km in lateral and 15 km in vertical direction but those parts of the figure were cut to enlarge the relevant features. Axes are in km. (b-d) Surface level directly above the sources calculated in 3D models. (b) Effect of perturbing one crucial material parameter compared to the reference model. (c) Effect of changing the shape of the magma and volatile reservoir in comparison to reference model. Subscript *obl* means oblate, subscript *pro* means prolate. The volume was preserved in all cases. Ref_{in} corresponds to a model where the volatile reservoir is placed inside the magma body. (d) Effect of changing the temperature structure in the crust through higher reservoir temperature or longer/no initial diffusion time. See section 3.2 for details. 'Growth' corresponds to a model where the volatile reservoir incrementally increases in size.

213 **3 Results**

214 Section 3.1 describes the general behavior of the model and discusses dependen-
 215 cies on time stepping and size of the model domain. In section 3.2, we discuss the effects
 216 of changing the geometry of the magma body and volatile reservoir as well as the ther-
 217 mal structure of the crust. Section 3.3 describes the derivation of the scaling law in 3D.
 218 The 2D is scaling law is discussed in supplementary text S3 and the differences between
 219 2D and 3D in supplementary text S4. In section 3.4 we apply the scaling law derived in
 220 section 3.3 to the case of the 2015 Calbuco eruption.

221 **3.1 General Behavior**

222 At the start of each simulation, the surface above the buoyant volatile reservoir un-
 223 dergoes immediate uplift, and quickly (within 2 time steps) reaches a steady state. Upon,
 224 replacing the volatiles with non-buoyant magma (i.e. an eruption), the surface quickly
 225 (within 2 time steps) returns to the original level. Independently of the time step we em-
 226 ploy (0.1 - 10 years), the surface reaches the same level after 2 steps with the first step
 227 being very close to it already (Figures S2c and S3d). We observe the same behavior af-
 228 ter removing the volatile reservoir. We therefore conclude that the surface response is
 229 immediate and has no time dependence. The small adjustment, necessary in the second
 230 time step, is inferred to be of numerical origin. To minimize computational cost and en-
 231 able us to observe any potential time dependencies, we use a time step of 1 year for all
 232 our models. In reality, the uplift or inflation phase may take place over a long time as
 233 the volatile reservoir grows gradually, but will reach the same magnitude as in our mod-
 234 els. Volcano deflation, however happens on timescales of eruptions as all volatiles are ex-
 235 pected to reach the surface, once a pathway has been established.

236 Supplementary Figures S2d and S3c show that the surface displacement depends
 237 on the width of the model domain. The displacement increases with increasing model
 238 width but at 50 km width, the effect levels off. We therefore ran all models with a width
 239 of 50 km.

240 We do not observe plastic failure in any of our models. Even after reducing cohe-
 241 sion (c_0) by an order of magnitude to 0.5 MPa and friction angle (ϕ) to 10° while increas-
 242 ing r_{vol} to 500 meters and G_{crust} to 10 GPa to maximize crustal stresses, stresses due

243 to changes in buoyancy never exceed a few MPa which is insufficient to exceed the Drucker-
 244 Prager failure criterion.

245 **3.2 Influence of Source Geometry and Thermal Structure**

246 In Figure 1c, we show the results of testing different shapes for the magma and volatile
 247 reservoirs. Both the oblate and prolate shapes have an aspect ratio of 2 while preserv-
 248 ing the volume of the spherical version. None of the geometrical variations lead to a sig-
 249 nificant difference in vertical displacement. Immersing the lower half or the entire volatile
 250 reservoir in the top of the magma body does not have significant effects either (Figure
 251 1c).

252 Figure 1d shows the effect of changing the thermal structure of the crust. In the
 253 'No Diff' example, we omit the 50 ka of thermal diffusion and start with a crust that only
 254 has the background temperature gradient while in the 'Long Diff' example, we double
 255 the temperature diffusion time from 50 to 100 ka. For the 'High T' example, we set the
 256 magma and volatile temperature to 1000 °C instead of 800 °C. The surface response is
 257 almost identical with the reference model for all cases.

258 **3.3 3D Scaling law**

259 Figure 1b shows the effect of varying four material parameters that have a consid-
 260 erable effect on the surface response. The radius and depth of the volatile reservoir (d),
 261 the density contrast between volatiles and crust ($\Delta\rho$) and the shear modulus of the crust
 262 (G). We performed a systematic parameter variation, testing 5 different values for each
 263 parameter (9 for d_{vol}) while keeping the other parameters constant. Supplementary Fig-
 264 ure S6 shows the results for individual parameters. From this, we are able to derive the
 265 following scaling relationship:

$$\Delta h_0 = A \frac{r^3 \Delta\rho g}{d^{3/2} G} \quad (8)$$

266 where Δh_0 is the vertical displacement at the surface above the source, g is the grav-
 267 itational acceleration and A is a pre-factor of 12π with units of $\text{m}^{0.5}$ to satisfy the equa-
 268 tion. The scaling law is similar to the one derived for the 2D case with the only excep-

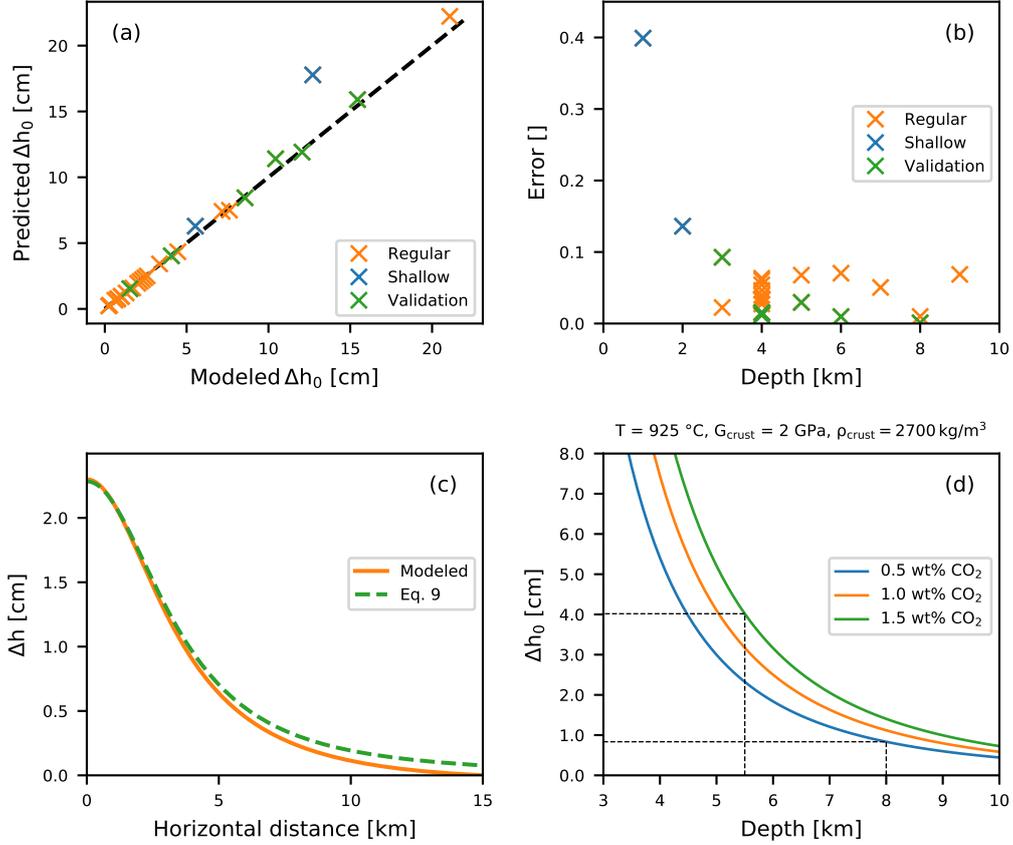


Figure 2. (a) Comparison between modeled subsidence and subsidence predicted by equation 8. Black dashed line shows 1:1 correlation. Orange crosses show models that were used to derive equation 8 and green crosses show models used to validate the scaling law. Blue crosses show models with a shallow (≤ 2 km) volatile reservoir that were excluded. (b) Error between the modeled subsidence and the predicted subsidence as a function of depth. (c) Subsidence along a radial profile for the reference case in orange in comparison to the prediction made by equation 9. Offset at the edge is due to the boundary conditions of the model as the uplift in the center needs to be balanced by subsidence along the edges. (d) Coupling between the ideal gas law and equation 8 to show the full dependence of maximum subsidence Δh_0 on reservoir depth d . Solid lines denote different volatile mass estimates as discussed in section 2.4. Dashed black lines show how we determine the minimum and maximum of subsidence in section 3.4.

269 tion being the exponents of r and d (see equation S4). Fits for the individual parame-
 270 ters are shown in Figure S6.

271 Figure 2a shows how equation 8 predicts the results of the 3D models and Figure
 272 2b shows the error which is generally smaller than 10%. Green crosses in Figures 2a and
 273 2b show models where we changed multiple parameters to validate equation 8. Subsidi-
 274 sidence in models with a shallow ($d \leq 2 \text{ km}$) volatile reservoir are overestimated. An-
 275 alytical solutions for the gravity anomaly of buried cylinders or spheres have the same
 276 issue of only being applicable when the depth of the body is much larger than its radius
 277 (Turcotte & Schubert, 2002). The same is true for simple models relating the inflation
 278 of magma bodies to surface deformation (e.g. Mogi, 1958; Yang et al., 1988).

279 Equation 8 only describes the vertical displacement directly above the center of the
 280 volatile reservoir. Figure 2c shows a profile of the vertical change along the surface. Our
 281 numerical models show that we can modify equation 8 to:

$$\Delta h(x) = A \frac{r^3 \Delta \rho g}{G} \frac{d}{(d^2 + x^2)^{5/4}} \quad (9)$$

282 where x is the horizontal distance from the center of the volatile reservoir projected to
 283 the surface. Figure 2c shows that the modeled surface displacement is fit well by equa-
 284 tion 9. In this form, our scaling law is very similar to the analytical solution of ground
 285 deformation due to a point source of pressure within an elastic halfspace, the "Mogi model"
 286 (Mogi, 1958). The most notable difference being the exponent of 5/4 instead of 3/2, which
 287 stems from the depth dependence of $\frac{1}{d^{3/2}}$ (see equation 8 and supplementary Figure S6c)
 288 while the "Mogi model" has a depth dependence of $\frac{1}{d^2}$.

289 3.4 Calbuco

290 For Calbuco, we use equation 8 with $\rho_{\text{crust}} = 2700 \text{ kg m}^{-3}$ and $G_{\text{crust}} = 2 \text{ GPa}$ to
 291 predict a maximum surface subsidence of 4 cm due to the loss of buoyancy from $8.8 \cdot 10^{10}$
 292 kg of exsolved volatiles for the case of storage at 5.5 km depth. For the 8 km depth sce-
 293 nario, and a lower limit estimate of the erupted gas mass ($3.9 \cdot 10^{10} \text{ kg}$), we predict 1 cm.
 294 Equations 8 and 9 imply that the surface displacement depends on the reservoir depth
 295 to the power of 1.5. In reality, r and $\Delta \rho$ are also functions of the pressure in the volatile

296 reservoir and thereby of the depth. Figure 2d illustrates this nonlinear dependence and
 297 shows how we arrive at our minimum and maximum estimates.

298 4 Discussion

299 4.1 Rheology

300 Given that, even for rocks with considerably lowered plastic strength, the stresses
 301 caused by the changes in buoyancy are not sufficient to exceed the failure criterion, plas-
 302 ticity is not a relevant factor in our models. Figure 1d also suggests that on the timescales
 303 of an eruption, viscous components have no impact on the deformation, even with the
 304 weakening caused by heating of the crust. The process of surface subsidence caused by
 305 the loss of a buoyant volatile reservoir due to eruption can therefore be considered as quasi-
 306 elastic, and as a result it is possible to derive a scaling law for the problem.

307 4.2 Surface Subsidence due to Buoyancy Loss

308 Instantly (on the timescale of an eruption) removing the buoyancy forces, exerted
 309 by a volatile reservoir, from the top of an upper crustal magma body leads to an instan-
 310 taneous subsidence. The magnitude of subsidence decays with radial distance from the
 311 reservoir center, but is significant in a radius of several kilometers (Figure 2c). The sur-
 312 face response is insensitive to the temperature structure (Figure 2d) of the crust which
 313 allows us to derive a scaling law for the expected subsidence (equations 8 and 9). As the
 314 shape of the volatile reservoir appears to play a minor role (Figure 1c), we suggest this
 315 alternative form of equation 9:

$$\Delta h(x) = \frac{9V \Delta \rho g}{G} \frac{d}{(d^2 + x^2)^{5/4}} \quad (10)$$

316 where V is the volume of the volatile reservoir. As other analytical solutions for the sur-
 317 face effects of buried bodies, the scaling law's accuracy decreases when the ratio between
 318 radius and depth of the body exceeds 0.1 (Figure 2b). The reduction to volume is in line
 319 with Archimedes' principle.

320 The inferred scaling law (equations 8, 9 and 10) has a similar structure to the Mogi
 321 model including a pre-factor, a cubic dependence on radius, an elastic property of the
 322 crust and a term describing the decay of the signal with distance. One difference is the

323 term of the driving force of deformation. In the Mogi model, it is either a pressure or
 324 a volume change, while in our scaling law, it is buoyancy. The other notable difference
 325 is the exponent of the depth dependence (2 for Mogi and 1.5 in our model). This could
 326 be caused by the different mechanisms that are at work. The pressure point source of
 327 the Mogi model applies a pressure to the surrounding crust in all directions, while in our
 328 case, buoyancy is expected to exert a cumulative upwards force in line with Archimedes'
 329 principle (e.g. Sigmundsson et al., 2020).

330 Another difference to common scaling laws for volcano deformation (e.g. Mogi, 1958;
 331 McTigue, 1987) is the lack of compressibility in our models because of its complex in-
 332 terplay with densities. As vertical displacement is usually multiplied by the term $(1 -$
 333 $\nu)$, our scaling law might provide a minimum estimate as a commonly used Poisson's ra-
 334 tio of $\nu = 0.25$ results in a larger factor than incompressibility ($\nu = 0.5$).

335 **4.3 Calbuco**

336 Applying our scaling law to the case of the 2015 Calbuco eruption, yields a sub-
 337 sidence of 1–4 cm (Figure 2d). With an incidence angle of 33° (Delgado et al., 2017),
 338 these vertical velocities can be projected into line-of-sight displacement (Fialko et al.,
 339 2001b) and represent 7% to 28% of the observed surface deformation. This is an indi-
 340 cation that the majority of co-eruptive subsidence was caused by the volumetric loss of
 341 material (volatiles and magma) but a significant part of the signal may originate from
 342 the loss of buoyancy provided by a body of exsolved volatiles.

343 In fact, the best-fit sphere and spheroid models of Delgado et al. (2017) have a resid-
 344 ual of about 3 cm in the center of subsidence. The mechanism described in our work pro-
 345 vides an additional source of uplift, large enough to cover this misfit entirely.

346 **4.4 Implications for Modeling Volcanic Deformation**

347 The release of a buoyant body of exsolved volatiles from the top of an upper crustal
 348 magma reservoir can lead to significant (on the order of a few cm) syn-eruptive subsi-
 349 dence at the surface. This effect is likely smaller than the effect of volume change in vol-
 350 canic roots during eruptions as magma moves to the surface. In the case of Calbuco, the
 351 contributions may have a ratio between 3:1 and 10:1 in favor of the volume loss. This
 352 ratio depends, however, on the quantity of pre-exsolved volatiles.

353 Adding equation 10 to existing models could be a simple way of achieving a bet-
354 ter fit to the observed deformation while also providing an explanation for the excess gas
355 that is detected for a number of eruptions.

356 As Figure 1 shows, the presence of a buoyant body of exsolved volatiles also causes
357 surface uplift of the same magnitude as its removal causes subsidence. That means that
358 inflation of a few centimeters over time, which is traditionally interpreted to be a sign
359 of magma intrusion at depth, could also be caused by the formation of a body of exsolved
360 volatiles at the top of the magma reservoir.

361 Furthermore, magma is usually buoyant at the depth where it intrudes. So even
362 if the intruded magma does not form a significant volatile reservoir, it still exerts a buoy-
363 ancy force on the crust that adds to the surface deformation caused by displacing host
364 rock. Although the effect of magma buoyancy on surface deformation was not explic-
365 itly investigated here, it is likely that equation 10 also gives a good estimate of its ef-
366 fect and could be added to existing solutions for surface uplift.

367 5 Conclusions

368 We conducted a series of 3D visco-elasto-plastic models to investigate the surface
369 deformation caused by the instantaneous removal of buoyancy forces, exerted by a reser-
370 voir of exsolved volatiles, from the top of a magma body, as would be the case during
371 an eruption. Our results show that the removal causes subsidence at the surface which
372 is mostly independent of the shape of the volatile and magma reservoirs as well as from
373 the thermal state of the crust. Instead, the process is quasi-elastic, allowing us to de-
374 rive an analytical solution for the surface subsidence including the volume and depth of
375 the reservoir, the density contrast between volatiles and crust, as well as the shear mod-
376 ulus of the crust. This analytical solution predicts surface deformations on the order of
377 up to a few centimeters.

378 We applied our scaling law to the case of the 2015 Calbuco eruption and, depend-
379 ing on the depth of the reservoir and volatile mass, predict subsidence of 1–4 cm, which
380 is about 20% of the observed signal. We expect that most of the observed surface de-
381 formation is caused by the volume loss of volatiles and magma.

382 Adding our scaling law to existing models for volcano deformation would present
383 a step forward, towards models that include all the relevant mechanisms that occur in
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395 **Open Research Section**

396 Software for this research is available on zenodo at:

397 LaMEM (Kaus et al., 2016):
398 <http://doi.org/10.5281/zenodo.5734975>

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