

1 **AMOC and water-mass transformation in high- and**
2 **low-resolution models: Climatology and variability**

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6 **Key Points:**

- 7 • A high-resolution coupled simulation reproduces subpolar North Atlantic water-
- 8 mass transformation from a reanalysis-forced ocean simulation
- 9 • Low-resolution simulations have larger biases in sea-surface heat fluxes, tem-
- 10 perature and salinity than the high-resolution simulations
- 11 • Despite climatological differences between the low- and high-resolution mod-
- 12 els, mechanisms of low-frequency AMOC variability are similar

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Abstract

Water-mass transformation in the North Atlantic plays an important role in the Atlantic Meridional Overturning Circulation (AMOC) and its variability. Here we analyze subpolar North Atlantic water-mass transformation in high- and low-resolution versions of the Community Earth System Model (CESM1) and investigate whether differences in resolution and climatological water-mass transformation impact low-frequency AMOC variability. We find that high-resolution simulations reproduce the water-mass transformation found in a reanalysis-forced high-resolution ocean simulation more accurately than low-resolution simulations. We also find that the low-resolution CESM1 simulations, including one forced with the same atmospheric reanalysis data, have larger biases in surface heat fluxes, sea-surface temperatures and salinities compared to the high-resolution simulations. Despite these major climatological differences, the mechanisms of low-frequency AMOC variability are similar in the high- and low-resolution versions of CESM1. The Labrador Sea WMT plays a major role in driving AMOC variability, and a similar NAO-like sea-level pressure pattern leads AMOC changes. However, the high-resolution simulation shows a more pronounced atmospheric response to the AMOC variability. The consistent role of Labrador Sea WMT in low-frequency AMOC variability across high- and low-resolution coupled simulations, including a simulation which accurately reproduces the WMT found in an atmospheric reanalysis-forced high-resolution ocean simulation, suggests that the mechanisms are similar in the real world.

Plain Language Summary

Water-mass transformation, which refers to the process of converting a water parcel from one density to another, plays an important role in the Atlantic Meridional Overturning Circulation (AMOC). Here we use high- and low-resolution climate models to investigate whether differences in the model resolution and time-mean water-mass transformation patterns impact AMOC fluctuations. We find that high-resolution coupled simulations reproduce the water-mass transformation found in a high-resolution ocean simulation driven by atmospheric reanalysis data, which we take as our closest analogue to observations. We also find that the low-resolution simulations, including one forced with the same atmospheric reanalyses, have larger discrepancies in surface properties compared to the high-resolution coupled simulation. Despite these ma-

(~1) global climate models (e.g. (Langehaug, Rhines, et al., 2012)). However, low-resolution global climate models vary in their representation of which deep water formation regions dominate the AMOC structure and variability (e.g. Langehaug, Rhines, et al. (2012); Menary et al. (2015); Heuze (2017); Oldenburg et al. (2021)). The biases in the deep water formation regions coincide with biases in subpolar temperature and salinity relative to observations (Langehaug, Rhines, et al., 2012). In addition, Nordic Seas overflow processes, which are responsible for producing the dense water masses that make up the southward flowing branch of AMOC and take place at comparatively small spatial scales (Treguier et al., 2005; Langehaug, Medhaug, et al., 2012), tend to be weak in low-resolution ocean models compared to observations (Bailey et al., 2005). This results in a deficit in the volume transport of these water masses. Moreover, low-resolution models do not resolve ocean mesoscale eddies, which are known to contribute to water-mass transformation via convection and lateral buoyancy fluxes, particularly in the Labrador Sea (Garcia-Quintana et al., 2019).

In low-resolution simulations, low-frequency AMOC variability appears to be driven primarily by Labrador Sea WMT changes, regardless of where the climatological WMT is concentrated (Oldenburg et al., 2021). The mechanism of the low-frequency AMOC variability involves upper ocean cooling and densification in the Labrador Sea, driven by northwesterly winds off eastern North America. This increases deep convection there, which later strengthens AMOC. The strengthened AMOC carries anomalous warm water northward into the subpolar regions, weakening deep water formation and hence AMOC. This mechanism, dominated by Labrador Sea WMT variability, holds true across three low-resolution models with distinct representations of deep water formation in subpolar regions (Oldenburg et al., 2021). However, one concern with these results is that low-resolution simulations likely overestimate deep water formation and subduction in the Labrador Sea region compared to high-resolution ocean simulations (Garcia-Quintana et al., 2019). This is because of the large role that convective eddies play during the restratification phase in the spring and summer months. Mixed-layer depths are also likely too deep in low-resolution models owing to the absence of eddies (Garcia-Quintana et al., 2019). This raises several interesting questions: (1) Do the mechanisms of low-frequency AMOC and OHT variability found in low-resolution models, where the Labrador Sea appears to be the most important region for initiating AMOC variability (Oldenburg et al., 2021), still hold in a high-resolution

109 model? (2) How does the ocean resolution of a model affect the partitioning of WMT
110 between the different deep water formation regions?

111 In this paper, we aim to evaluate how well a high-resolution coupled model re-
112 produces the surface-forced WMT found in a high-resolution atmospheric reanalysis-
113 forced ocean simulation, which we consider as an approximation to observations, and
114 compare that to what is found in a low-resolution version of the same model. We then
115 analyze the factors that set the magnitude of WMT in these simulations. Finally, we
116 examine the mechanisms of low-frequency AMOC variability in the high- and low-
117 resolution versions of the coupled model. We focus in particular on the link between
118 the AMOC variability and the WMT variability in the different deep-water formation
119 regions and on how the variability is affected by the differences in resolution and mean
120 state.

121 In Section 2, we describe the model simulations used in this analysis. In Section
122 3, we compute the WMT and AMOC in the different simulations and analyze the fac-
123 tors that explain the differences between them. In Section 4, following the methods
124 of Oldenburg et al. (2021), we use a low-frequency component analysis (LFCA) to elu-
125 cidate the mechanisms of low-frequency AMOC variability in the high- and low-resolution
126 versions of the coupled model. In Section 5, we summarize and discuss the overall
127 results and conclusions.

128 **2 Description of models**

129 We use output from a 1800-year pre-industrial control simulation of the Com-
130 munity Earth System Model Version 1.1 (CESM1.1, Hurrell (2013)), with a nominal hor-
131 izontal resolution of 1° in the atmosphere and ocean. We henceforth refer to this low-
132 resolution CESM1 simulation as CESM1-LR. We also use output from a 500-year pre-
133 industrial control simulation of CESM1.3 by the International Laboratory for High-
134 Resolution Earth System Prediction (iHESP) (Chang et al., 2020), which uses an eddy-
135 resolving 0.1° version of the Parallel Ocean Component version 2 (POP2) and a 0.25°
136 version of the Community Atmosphere Model version 5 (CAM5). We henceforth re-
137 fer to this high-resolution CESM1 simulation as CESM1-HR. Unlike its low-resolution
138 counterpart, this model does not include a parameterization for overflows of deep wa-
139 ter from the Nordic Seas into the North Atlantic while still not fully resolving the over-

140 flow processes. Here we analyze the last 350 years of the 500-year simulation, because
141 the first 150 years are considered spin-up.

142 For our analysis of reanalysis-forced ocean-sea-ice simulations, we use output
143 from 1° and 0.1° POP2 ocean simulations, respectively, both forced with atmospheric
144 reanalysis data from the Japanese 55-year Reanalysis (JRA-55, Kobayashi et al. (2015);
145 Harada et al. (2016); Kim et al. (2021)), spanning years 1958-2018. Henceforth, we re-
146 fer to these low- and high-resolution simulations as JRA55-LR and JRA55-HR, respec-
147 tively. Here we are seeking an analogue to observations which still provides full ocean
148 output data. Given that historical ocean observations are limited to particular regions
149 or require reconstruction from proxies, an atmospheric reanalysis-forced ocean sim-
150 ulation, which includes an ocean constrained at the surface to best estimates of his-
151 torical atmospheric states, is a useful alternative. It would be possible to instead use
152 ocean assimilation data. However, they typically do not have closed heat and salt bud-
153 gets, which are important when linking WMT to the interior ocean state. Also, his-
154 torical ocean observations are fairly limited compared to atmospheric observational
155 data, which reduces the reliability of assimilation products. Hence, we take JRA55-
156 HR as our closest analogue to observations.

157 Here we compare the rest of the simulations to JRA55-HR to determine whether
158 increasing the ocean and atmospheric resolution of a coupled model leads to a more
159 accurate representation of WMT and AMOC. Comparing JRA55-LR with CESM1-LR
160 illustrates the role of atmospheric forcing (reanalysis data versus a coupled atmosphere)
161 at the same ocean model resolution, while comparing JRA55-LR with JRA55-HR il-
162 lustrates the role of ocean model resolution (parameterized versus resolved mesoscale
163 eddies) under the same atmospheric forcing.

164 **3 Comparison of WMT and AMOC climatologies**

165 Before analyzing WMT and AMOC, it is helpful to consider the time-mean win-
166 ter (January-February-March) mixed-layer depth to determine where the deep con-
167 vection and deep water formation occur in the different models. In JRA55-HR, deep
168 mixed layers are concentrated mostly in the Labrador Sea and Irminger and Iceland
169 Basins (IIB), with some deep mixed layers in the Greenland-Iceland-Norwegian (GIN)
170 Seas as well (Fig. 1a). In JRA55-LR, the mixed layers overall are deeper, and the deep-

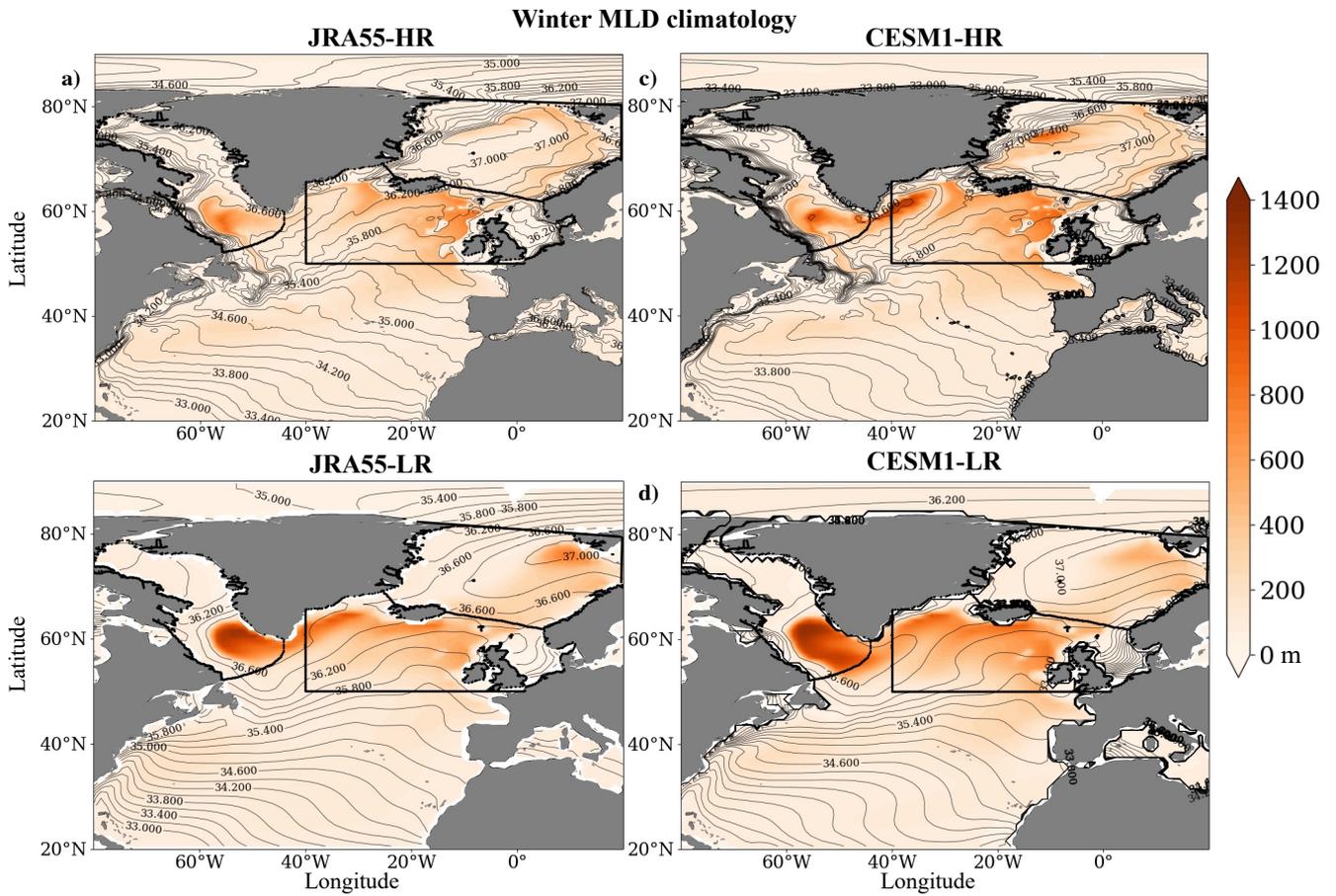


Figure 1: Climatological mixed-layer depth (colors) and sea-surface potential density referenced to 2000 m (contours) both averaged over January, February and March in **a)** JRA55-HR, **b)** JRA55-LR, **c)** CESM1-HR and **d)** CESM1-LR. The thick black lines represent the region masks for the Labrador Sea (left), Irminger-Iceland Basins (lower right) and GIN Seas (upper right).

171 est mixed layers are concentrated in the Labrador Sea, though there are still deep mixed
 172 layers in the IIB and GIN Seas (Fig. 1b). In CESM1-HR, the mixed-layer depth pat-
 173 terns look similar to JRA55-HR, but the mixed-layer depths are deeper in all of the
 174 deep water formation regions (Fig. 1c). In CESM1-LR, the deepest mixed layers are
 175 mostly concentrated in the Labrador Sea, even more so than in JRA55-LR, which shows
 176 similar overall patterns (Fig. 1b, d). It is noteworthy that CESM1-HR captures the mixed-
 177 layer depth patterns found in JRA55-HR much better than either of the low-resolution
 178 models, despite JRA55-LR being forced with the same atmospheric reanalysis data as
 179 JRA55-HR.

180 Throughout our analysis, we use AMOC calculated in density coordinates, rather
 181 than AMOC calculated in depth coordinates, because it is more appropriate for an-
 182 alyzing subpolar AMOC variability and is strongly connected to the the analysis of
 183 WMT as a function of density class (Straneo, 2006; Pickart & Spall, 2007). We first look
 184 at the AMOC climatology to determine how well the coupled simulations (and JRA55-
 185 LR) reproduce the AMOC from the reanalysis-forced high-resolution dataset, JRA55-
 186 HR. To compute AMOC, we use Eq. (1) from Newsom et al. (2016).

187 In JRA55-HR, the maximum AMOC is located at $\sigma_2 = 36.48 \text{ kg m}^{-3}$, where it
 188 reaches 21.8 Sv (Fig. 2a). In JRA55-LR, the maximum is located at $\sigma_2 = 36.58 \text{ kg m}^{-3}$
 189 and is 20.7 Sv (Fig. 2b). AMOC in CESM1-HR reaches a maximum of 25.4 Sv at $\sigma_2 =$
 190 36.53 kg m^{-3} (Fig. 2c). In CESM1-LR, AMOC reaches a maximum of 28.6 Sv at $\sigma_2 =$
 191 36.64 kg m^{-3} (Fig. 2d). Hence, in terms of maximum magnitude, JRA55-LR reproduces
 192 the AMOC found in JRA55-HR the best of all the other model simulations, though
 193 CESM1-HR reproduces the density where the maximum occurs most accurately. Sur-
 194 prisingly, the maximum AMOC is actually smaller in JRA55-LR than in JRA55-HR;
 195 we would expect a higher resolution simulation to yield a weaker AMOC, as in CESM1-
 196 HR and CESM1-LR, and also what was found in other studies of coupled GCMs (Winton,
 197 2014; Sein et al., 2018). All of the simulations have AMOC maxima located at higher
 198 densities than JRA55-HR. CESM1-HR has a maximum AMOC at a density closest to
 199 the JRA55-HR maximum, while CESM1-LR has a maximum AMOC at a density fur-
 200 thest from the JRA55-HR maximum. These results indicate that although increasing
 201 the resolution of the atmosphere and ocean yields an AMOC substantially closer to
 202 reanalysis-forced ocean data, there are likely biases in the atmospheric component of
 203 the coupled simulations even at high resolution.

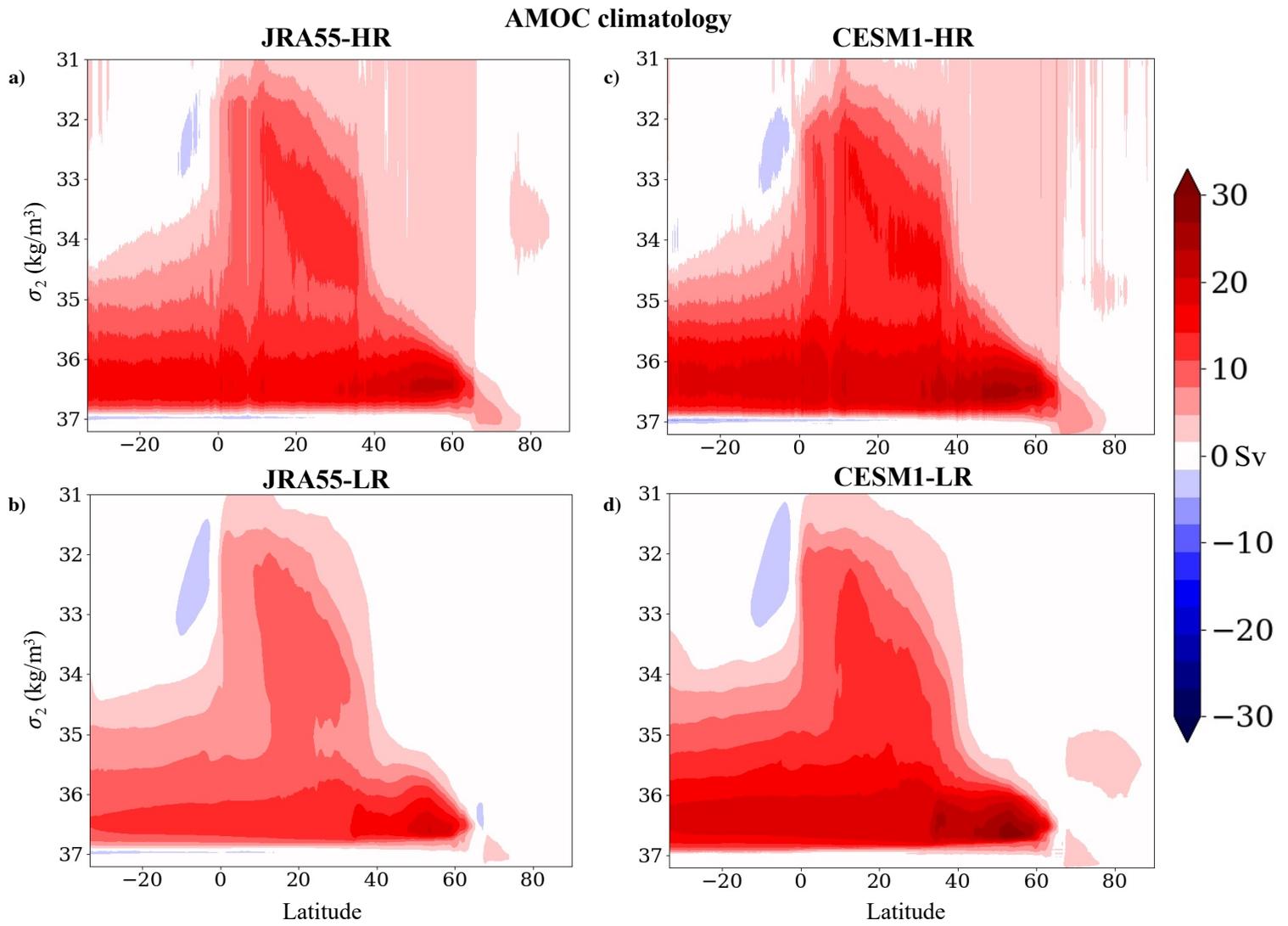


Figure 2: Climatological AMOC in a) JRA55-HR, b) JRA55-LR, c) CESM1-HR and d) CESM1-LR.

204 To compute the surface-forced WMT, we follow the methodology used in Oldenburg
 205 et al. (2021) among others. Here, as in Oldenburg et al. (2021), we neglect the mixing
 206 contributions because the time resolution of the model output data is not high enough
 207 to calculate them. We compute the WMT separately in the Labrador Sea, Irminger and
 208 Iceland Basins (IIB) and GIN Seas using the region masks shown in the boxes in Fig.
 209 1 to determine each region's contribution to the total WMT.

210 In all four simulations, the thermal WMT component dominates over the haline
 211 contribution. However, the partitioning of WMT in the different regions varies sub-
 212 stantially among the simulations. In JRA55-HR, none of the peaks in WMT in the dif-
 213 ferent regions align with the density of maximum AMOC. The IIB contributes the most
 214 to the WMT at densities lower than the density of maximum AMOC (Fig. 3a), reach-
 215 ing a maximum value of 14.2 Sv at $\sigma_2 = 36 \text{ kg/m}^3$. At densities higher than the max-
 216 imum AMOC, the WMT is dominated by contributions from the Labrador Sea and
 217 GIN Seas, with a much narrower peak in the Labrador Sea. The Labrador Sea has a
 218 peak of 7.7 Sv at $\sigma_2 = 36.7 \text{ kg/m}^3$, and the GIN Seas WMT peaks at 4.6 Sv at $\sigma_2 =$
 219 36.56 kg/m^3 . Though these densities are further away from the maximum AMOC,
 220 they are likely still important for AMOC given that internal mixing acts to reduce the
 221 density of the densest water masses. In JRA55-LR, the peaks in the IIB and GIN Seas
 222 WMT occur closer to the maximum AMOC, reaching maxima equal to 14.5 and 6.2
 223 Sv at $\sigma_2 = 36.32$ and $\sigma_2 = 36.62$, respectively, and the IIB dominates the WMT near
 224 the AMOC maximum (Fig. 3b). The Labrador Sea peak in WMT is located at about
 225 the same density as in JRA55-HR, with a peak value of 11.4 Sv at $\sigma_2 = 36.7 \text{ kg/m}^3$.
 226 Furthermore, the peaks in the Labrador Sea and GIN Seas WMT are narrower in JRA55-
 227 LR than they are in JRA55-HR.

228 The WMT in CESM1-HR looks the most similar to JRA55-HR of all the other sim-
 229 ulations, with the most notable difference being that the WMT peaks in the IIB and
 230 Labrador Sea WMT are larger than in JRA55-HR (Fig. 3c), with the IIB WMT reach-
 231 ing a maximum value of 17.4 Sv at $\sigma_2 = 36 \text{ kg/m}^3$, the Labrador Sea WMT reach-
 232 ing a maximum of 8.3 Sv at $\sigma_2 = 36.74 \text{ kg/m}^3$, and the GIN Seas WMT peaking at
 233 5.0 Sv at $\sigma_2 = 36.74 \text{ kg/m}^3$. However, the partitioning of the WMT between the dif-
 234 ferent regions remains similar to JRA55-HR. In CESM1-LR, on the other hand, the WMT
 235 looks quite different, with much larger WMT peaks in the IIB and the Labrador Sea
 236 WMT than in any of the other simulations (Fig. 3d), reaching maxima equal to 19.6

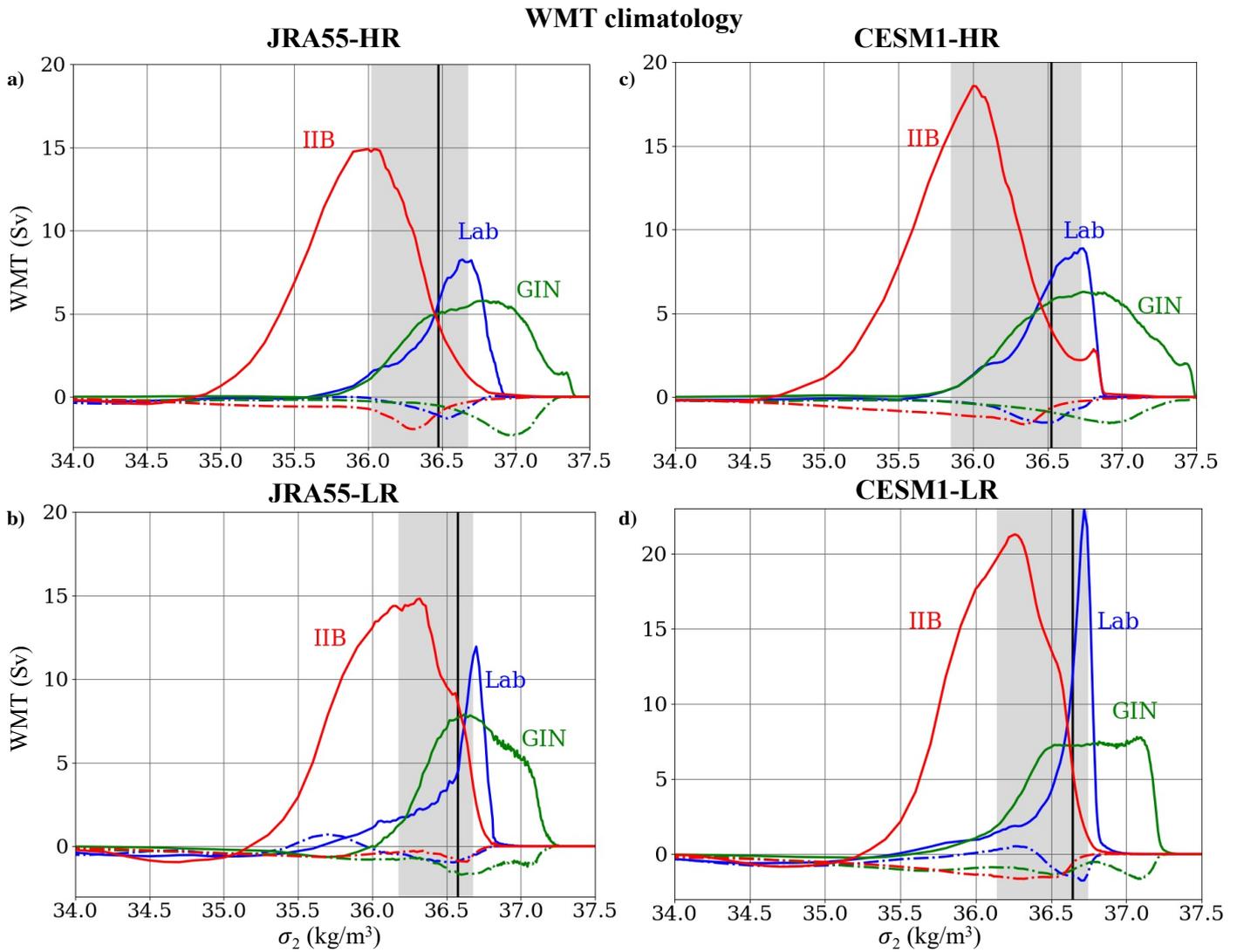


Figure 3: Climatological water-mass transformation thermal (solid lines) and freshwater (dashed lines) components in the Labrador Sea (Lab), GIN Seas and Irminger and Iceland Basins (IIB) for **a) JRA55-HR**, **b) JRA55-LR**, **c) CESM1-HR** and **d) CESM1-LR**. The black vertical lines indicate the density where the climatological AMOC reaches its maximum in each model. The grey shaded areas represent the density range where AMOC is within 25% of its maximum value. A more detailed illustration of what particular areas of the deep water formation regions contribute to the surface density flux over different density classes is shown in Figures 4-5, as well as Fig. S1.

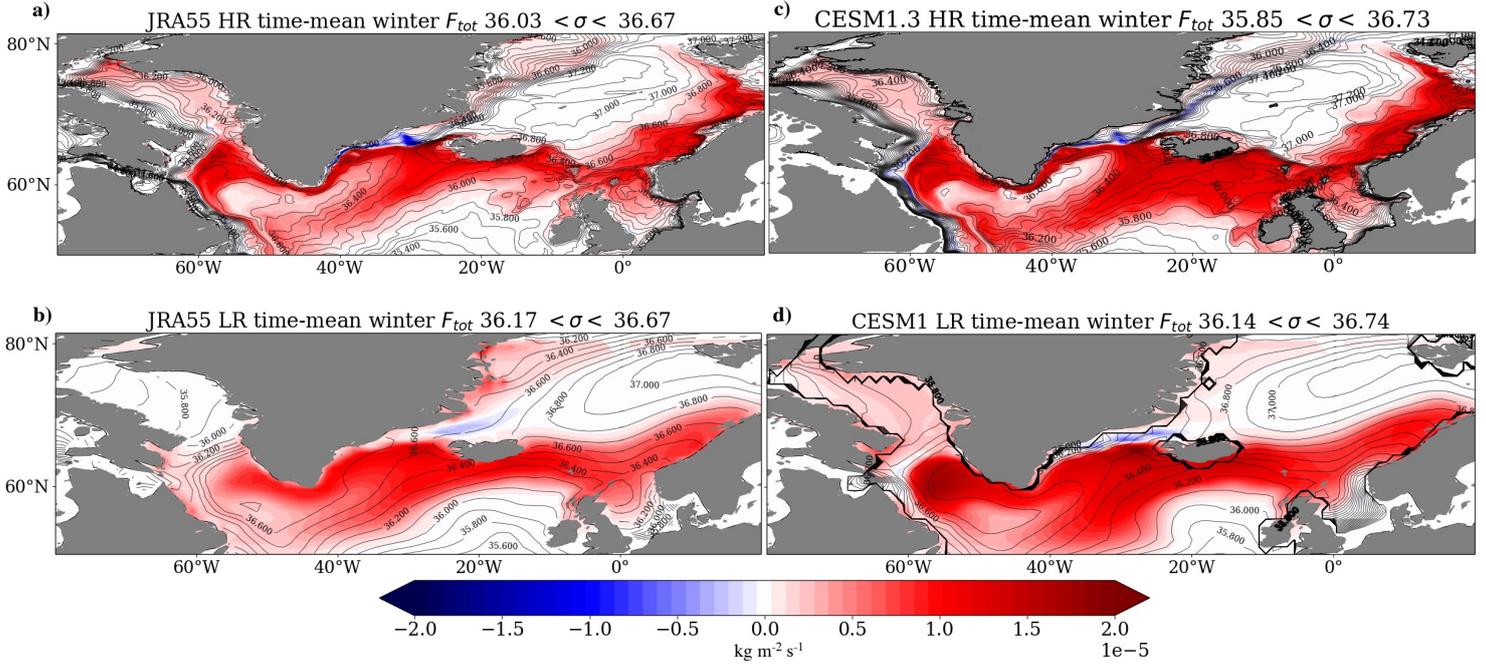


Figure 4: Colors: Total climatological winter surface density flux calculated using the methodology from Oldenburg et al. (2021) over densities where AMOC is at least 75% of its maximum. Contours: Time-mean winter sea-surface potential density referenced to 2000 m for **a** JRA55-HR, **b** JRA55-LR, **c** CESM1-HR and **d** CESM1-LR.

237 and 21.2 Sv at $\sigma_2 = 36.26$ and $\sigma_2 = 36.72$, respectively. The peak in Labrador Sea
 238 WMT is also much narrower than in JRA55-HR and CESM1-HR, and looks more sim-
 239 ilar to JRA55-LR. The GIN Seas WMT peaks at $\sigma_2 = 36.82$ kg/m³, where it reaches
 240 a maximum of 6.9 Sv. This seems to indicate that increasing the atmospheric and ocean
 241 resolution in a coupled model yields a fairly realistic representation of WMT in the
 242 different deep water formation regions, certainly much more realistic than an equiv-
 243 alent low-resolution coupled model. The major discrepancies between JRA55-LR and
 244 JRA55-HR indicate that a higher ocean model resolution is essential in order to pro-
 245 vide an accurate representation of WMT; having correct atmospheric surface forcing
 246 alone is insufficient.

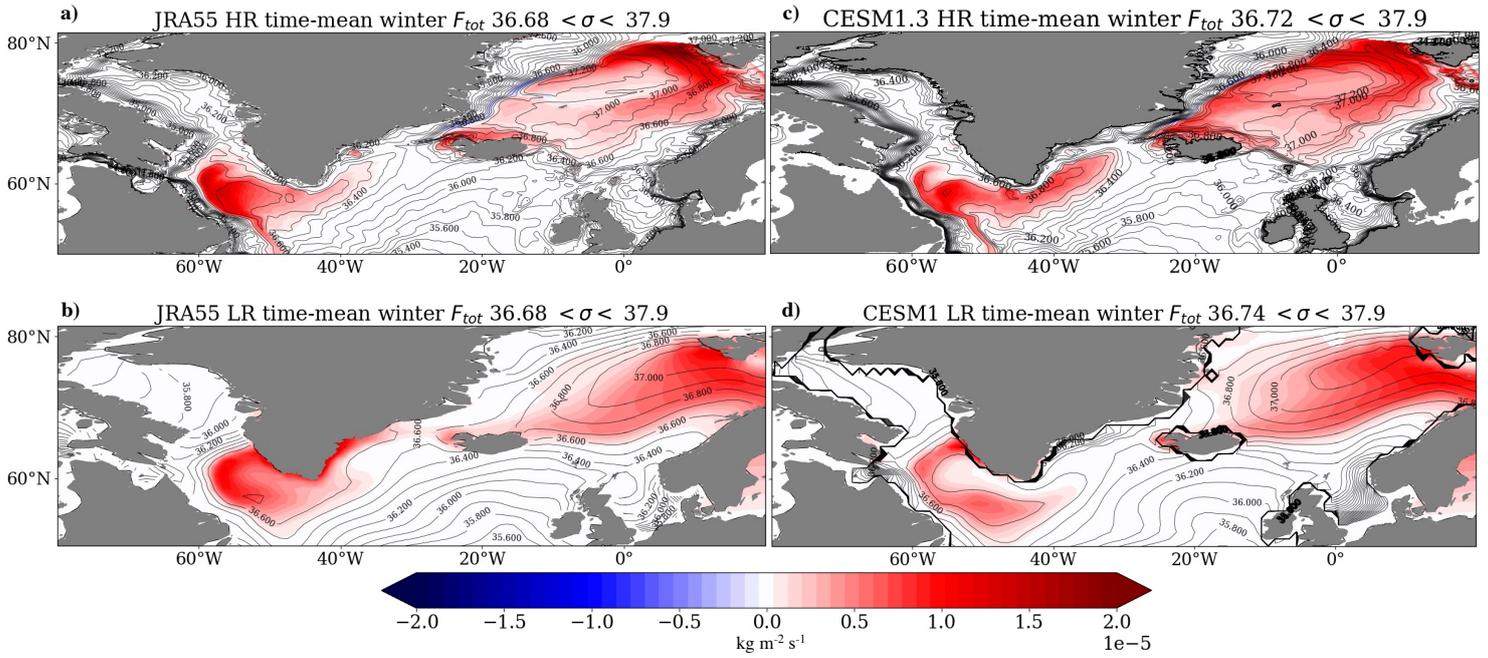


Figure 5: Colors: Total climatological winter surface density flux calculated using the methodology from Oldenburg et al. (2021) over densities above the maximum density where AMOC reaches 75% of its maximum. Contours: Time-mean winter sea-surface potential density referenced to 2000 m for **a** JRA55-HR, **b** JRA55-LR, **c** CESM1-HR and **d** CESM1-LR.

247 To illustrate which parts of each region contribute to the WMT in different den-
248 sity classes, it is useful to look at the full surface-density flux calculated using the method-
249 ology of Oldenburg et al. (2021) and others. Since we are interested in the density classes
250 relevant for AMOC, we isolate the density flux for densities lower than the minimum
251 density where AMOC reaches 75% of its maximum (Fig. S1), densities within the den-
252 sity range where AMOC is at least 75% of its maximum (Fig. 4), and densities above
253 that density range (Fig. 5). In the lowest density range, the surface-density flux is con-
254 centrated in the Irminger and Iceland Basins, with small contributions from the other
255 regions, mainly near coastlines where the water is fresher and lighter than the inte-
256 rior areas (Fig. S1). Because interior mixing tends to reduce the density of water parcels,
257 the surface-density fluxes in this density range are unlikely to contribute to AMOC.

258 In the density range near the AMOC maximum, CESM1-HR reproduces the den-
259 sity flux patterns found in JRA55-HR fairly well. In both of these simulations, most
260 of the Labrador Sea surface density flux is concentrated in the northern section of the
261 Labrador Sea rather than in the southern section, where density fluxes are weaker (Fig.
262 4a, c). The patterns found in the GIN Seas are also similar; however, the surface den-
263 sity fluxes in the southern part of the IIB are much higher in CESM1-HR than in JRA55-
264 HR (Fig. 4a, c). The low-resolution simulations show similar overall patterns to JRA55-
265 HR, but lack several key features (Fig. 4b, d). For example, Labrador Sea fluxes are
266 more concentrated in the central and southern sections compared to JRA55-HR and
267 CESM1-HR, particularly in CESM1-LR (Fig. 4d). JRA55-LR reproduces the flux pat-
268 terns in the IIB fairly well (Fig. 4b). However, neither low-resolution simulation has
269 an accurate representation of the more complex smaller scale density structures found
270 in JRA55-HR and CESM1-HR, where the densities are less uniform, particularly near
271 coastlines. For the highest density range, the interior and southern parts of the Labrador
272 Sea contribute more to WMT in JRA55-HR and CESM1-HR compared to the lower den-
273 sity classes (Fig. 5a, c). There are also larger contributions from the interior and north-
274 ern parts of the GIN Seas. The same overall patterns are found in the low-resolution
275 simulations (Fig. 5b, d). However, in JRA55-LR the surface density fluxes in the Labrador
276 Sea are shifted to the east relative to JRA55-HR and CESM1-HR, and the northern part
277 of the GIN Seas is not emphasized as much as in the high-resolution simulations, with
278 a much more uniform pattern in the eastern GIN Seas (Fig. 5b). In CESM1-LR, the con-

279 tributions to WMT from the Labrador Sea are smaller, and the eastern area of the GIN
280 Seas is more emphasized compared to in JRA55-LR (Fig. 5d).

281 To allow for a more direct comparison between AMOC and the WMT in the dif-
282 ferent regions, we also calculate the surface-forced overturning streamfunction follow-
283 ing the methodology of Marsh (2000); Oldenburg et al. (2021) and others.

284 Here we calculate the surface-forced overturning streamfunction for each of the
285 three regions separately, which allows us to quantify how much the surface-forced WMT
286 in each region contributes to AMOC (neglecting mixing). CESM1-HR reproduces the
287 surface-forced overturning found in JRA55-HR far better than either low-resolution
288 simulation in all regions (Fig. 6a-d, i-l). In JRA55-LR and CESM1-LR, the overturn-
289 ing is too strong in all the regions, especially in the Labrador Sea and IIB (Fig. 6e-h,
290 m-p). Also, the Labrador Sea surface-forced overturning is concentrated over a smaller
291 density range in the LR models compared to the HR versions (Fig. 6b, f, j, n). For the
292 IIB, overturning in the HR simulations is shifted towards lower densities compared
293 to the LR versions (Fig. 6c, g, k, o). Overturning in the GIN Seas is also concentrated
294 over a smaller density range in the LR models than in the HR models (Fig. 6d, h, l,
295 p).

296 To determine what is responsible for the discrepancies in the WMT between JRA55-
297 HR and the other simulations, we discuss the climatologies of several surface prop-
298 erties used in the WMT calculation, including the sea-surface heat fluxes as well as
299 the sea-surface potential temperatures, salinities and densities. Although the fresh-
300 water fluxes also contribute to the WMT, the freshwater components of WMT are very
301 small in all four simulations (Fig. 3). Hence we do not show them here, but rather in
302 the supplementary section (Fig. S2). For these quantities, we present the climatology
303 in JRA55-HR (Fig. 7e) and the anomalies for the other simulations relative to JRA55-
304 HR. CESM1-HR shows a much more accurate representation of the time-mean den-
305 sity structure compared to both low-resolution simulations, particularly in the Labrador
306 Sea and near all coastlines (Fig. 7f). CESM1-HR anomalies in sea-surface temperatures
307 and salinities relative to JRA55-HR are more substantial than its density anomalies (Fig.
308 8b, f), but they are mostly density compensating, yielding smaller density anomalies.
309 These anomalies lead to small positive density anomalies in the GIN Seas, IIB and Labrador
310 Sea, except near the coastlines (Fig. 7f), likely due to increased freshwater runoff com-

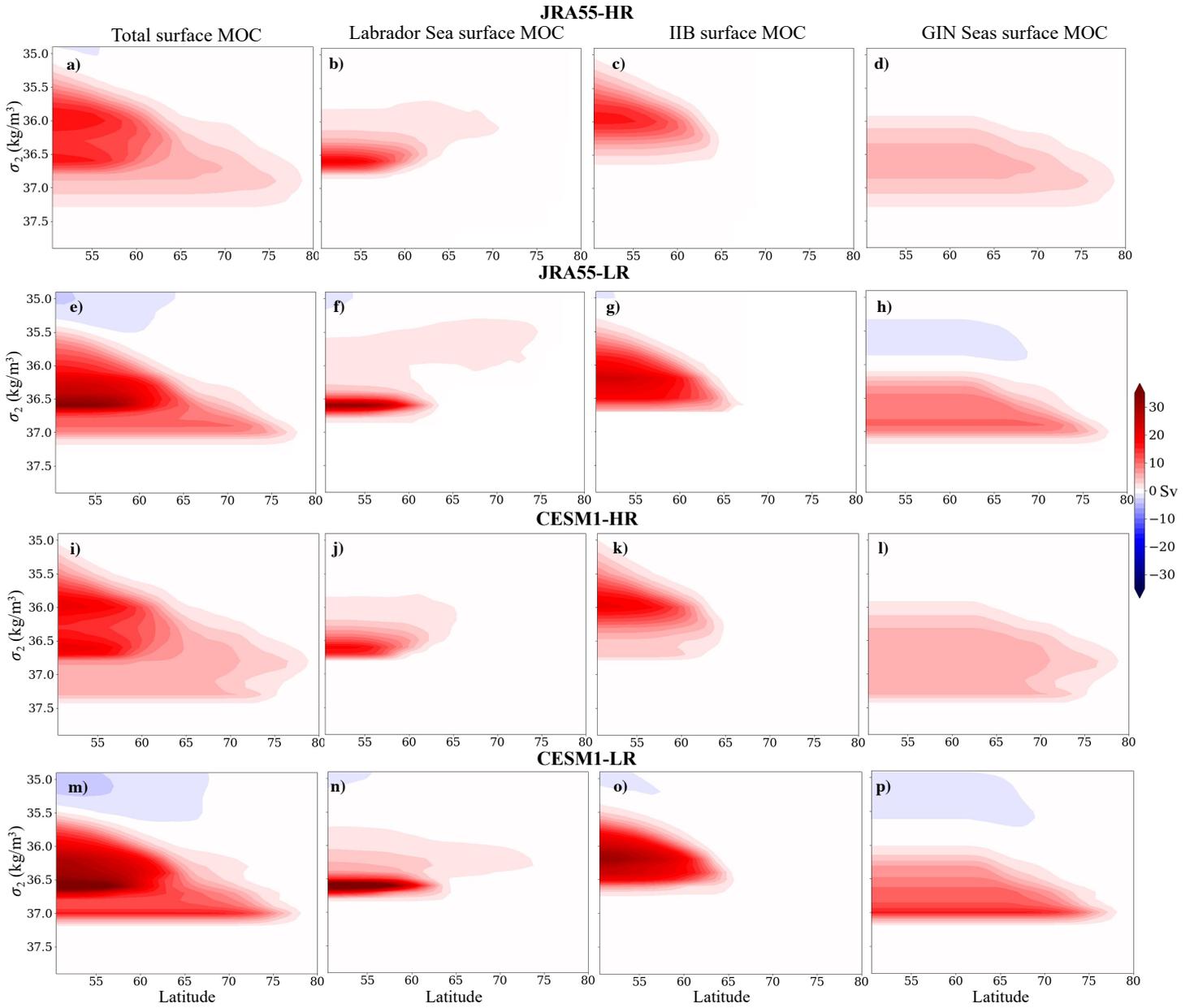


Figure 6: Climatological surface-forced overturning streamfunction in **a-d)** JRA55-HR, **e-h)** JRA55-LR, **i-l)** CESM1-HR and **m-p)** CESM1-LR computed over all regions (first column), the Labrador Sea (second column), the Irminger-Iceland Basins (IIB, third column) and GIN Seas (fourth column).

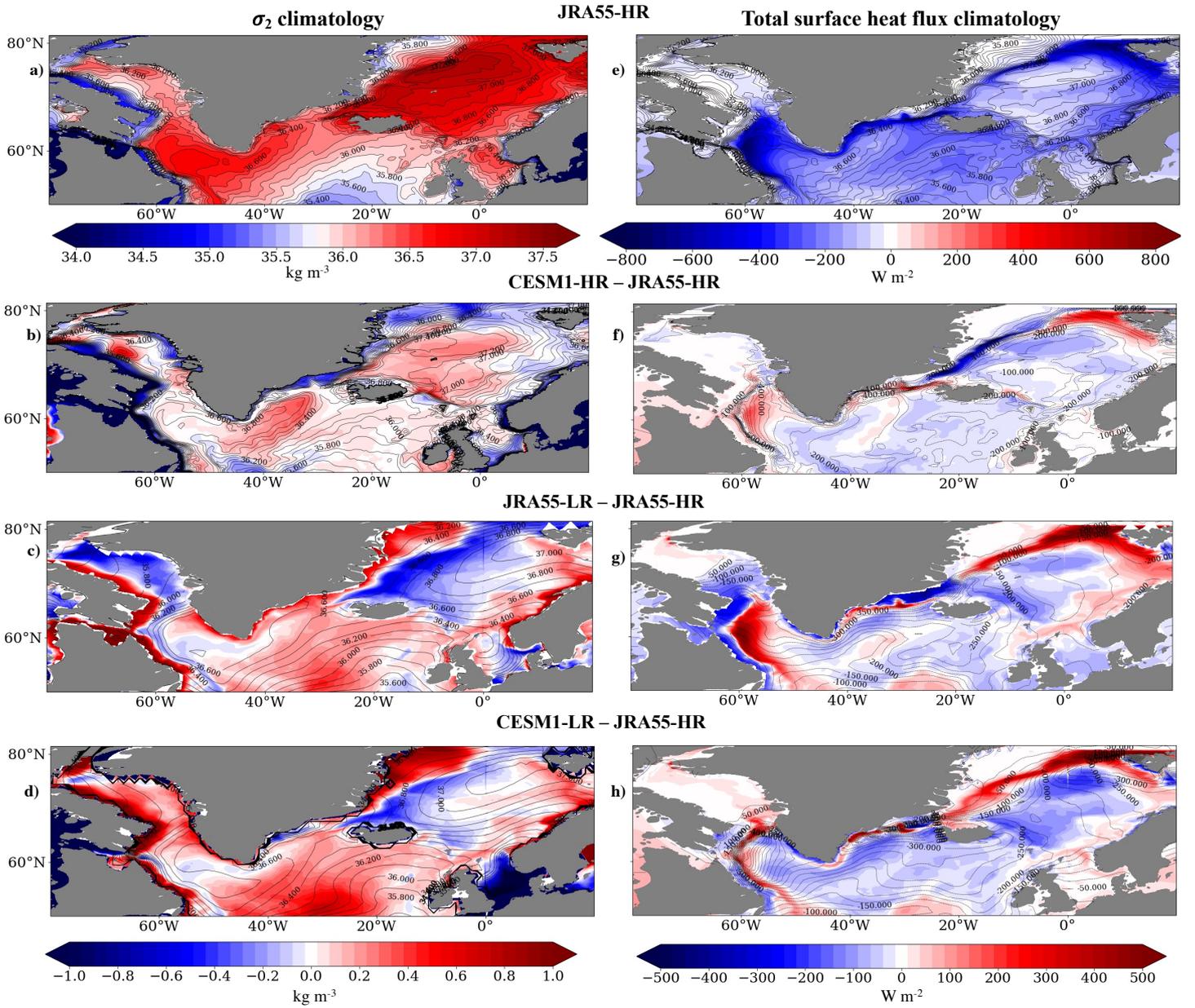


Figure 7: **a)** JRA55-HR climatology of sea-surface potential density, referenced to 2000m. **b-d)** Sea-surface potential density climatologies (contours) and anomalies relative to JRA55-HR (colors) for **b)** CESM1-HR, **c)** JRA55-LR and **d)** CESM1-LR. **e)** JRA55-HR total sea-surface heat flux climatology. **f-h)** Sea-surface heat flux climatologies (contours) and anomalies relative to JRA55-HR (colors) for **f)** CESM1-HR, **g)** JRA55-LR and **h)** CESM1-LR.

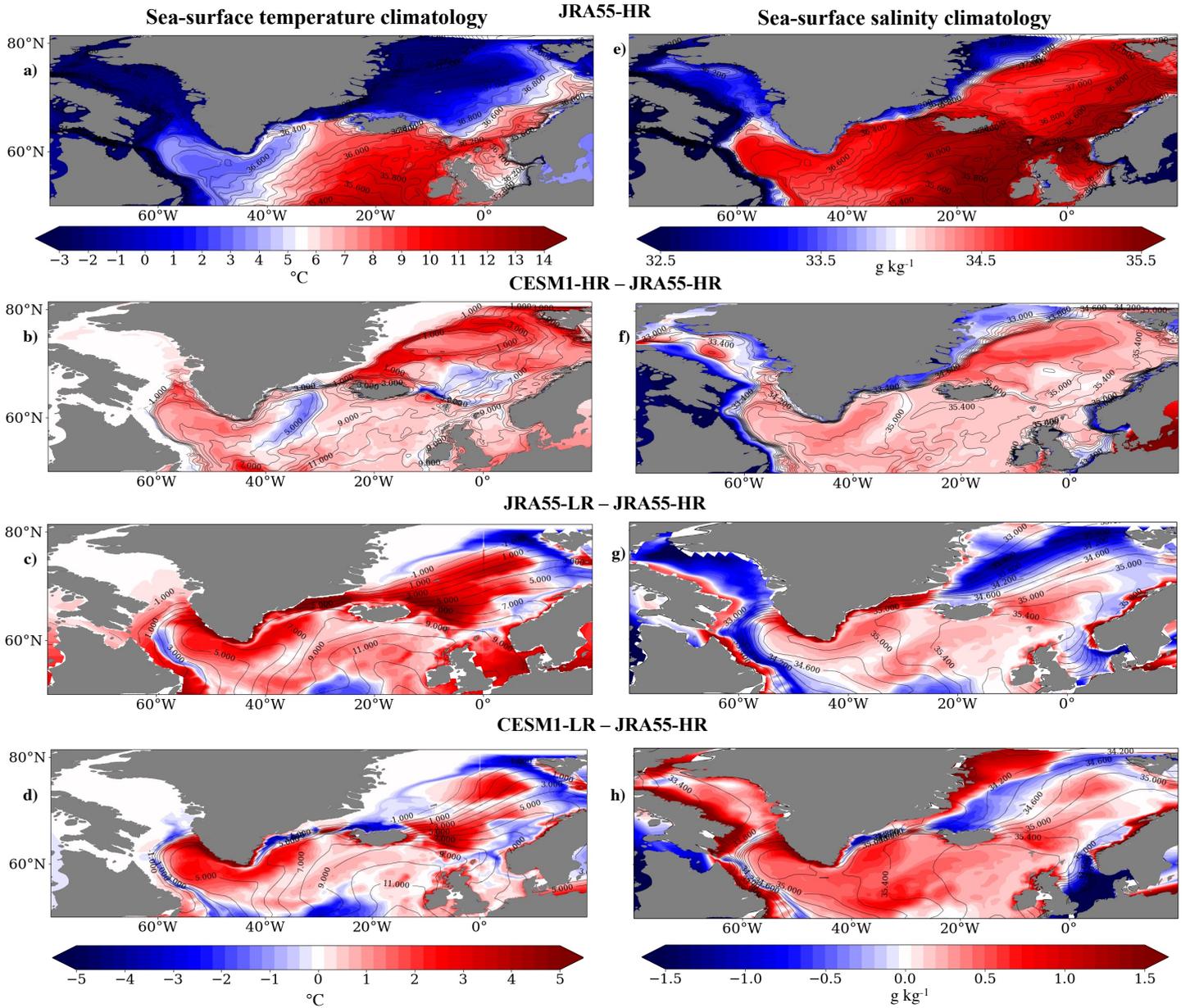


Figure 8: **a)** JRA55-HR sea-surface potential temperature climatology. **b-d)** Sea-surface potential temperature climatologies (contours) and anomalies relative to JRA55-HR (colors) for **b)** CESM1-HR, **c)** JRA55-LR and **d)** CESM1-LR. **e)** JRA55-HR sea-surface salinity climatology. **f-h)** Sea-surface salinity climatologies (contours) and anomalies relative to JRA55-HR (colors) for **f)** CESM1-HR, **g)** JRA55-LR and **h)** CESM1-LR.

311 compared to JRA55-HR (Fig. S2). JRA55-LR, on the other hand, shows large negative den-
312 sity anomalies in the central GIN Seas, but positive anomalies near the coastlines (Fig.
313 7g). There are also positive anomalies in the eastern subpolar gyre and in the north-
314 ern Labrador Sea. The density structure looks similar in CESM1-LR, with similar anoma-
315 lies relative to JRA55-HR in most regions, except for in the northern Labrador Sea where
316 there are actually positive anomalies (Fig. 7h), due to a fairly salty Labrador Sea com-
317 pared to the other simulations (Fig. 8h). The higher densities in the low-resolution sim-
318 ulations explain why the WMT and AMOC peaks occur at higher densities than in
319 JRA55-HR and CESM1-HR (Fig. 3), and the generally more uniform density fields in
320 the Labrador Sea explain the narrower WMT peaks in the LR simulations compared
321 to JRA55-HR and CESM1-HR. Also, the high densities in the GIN Seas in CESM1-HR
322 explain why there is positive WMT in that region at higher densities than what is seen
323 in the other models (Fig. 3c).

324 CESM1-HR best reproduces the surface heat fluxes found in JRA55-HR (Fig. 7a,
325 b), with some positive anomalies in the central and northern Labrador Sea and broad
326 negative anomalies throughout the IIB and GIN Seas, aside from the far north, which
327 exhibits positive anomalies (Fig. 7b). The larger (more negative) heat fluxes in the IIB
328 and GIN Seas explain the larger IIB and GIN WMT in CESM1-HR compared to JRA55-
329 HR, given that stronger heat fluxes drive higher WMT. JRA55-LR exhibits larger pos-
330 itive anomalies in the Labrador Sea and northern GIN Seas compared to CESM1-HR
331 (Fig. 7c). In CESM1-LR, there is a mix of positive and negative anomalies in the Labrador
332 Sea, and larger negative anomalies in the central GIN Seas (Fig. 7d).

333 Surprisingly, CESM1-HR reproduces the WMT, sea-surface heat fluxes, sea-surface
334 temperatures and salinities of JRA55-HR far better than JRA55-LR does, which high-
335 lights the importance of ocean resolution in accurately representing these variables.
336 It also indicates that simply forcing an ocean model with atmospheric reanalyses is
337 insufficient if the ocean is low-resolution.

338 **4 Mechanisms of low-frequency AMOC variability in high- and low-resolution** 339 **versions of CESM**

340 We next turn our attention to the mechanisms driving low-frequency AMOC vari-
341 ability. Following the methods of Oldenburg et al. (2021), we apply a low-frequency
342 component analysis (LFCA; R. C. Wills et al. (2018); R. C. J. Wills et al. (2019)) to AMOC

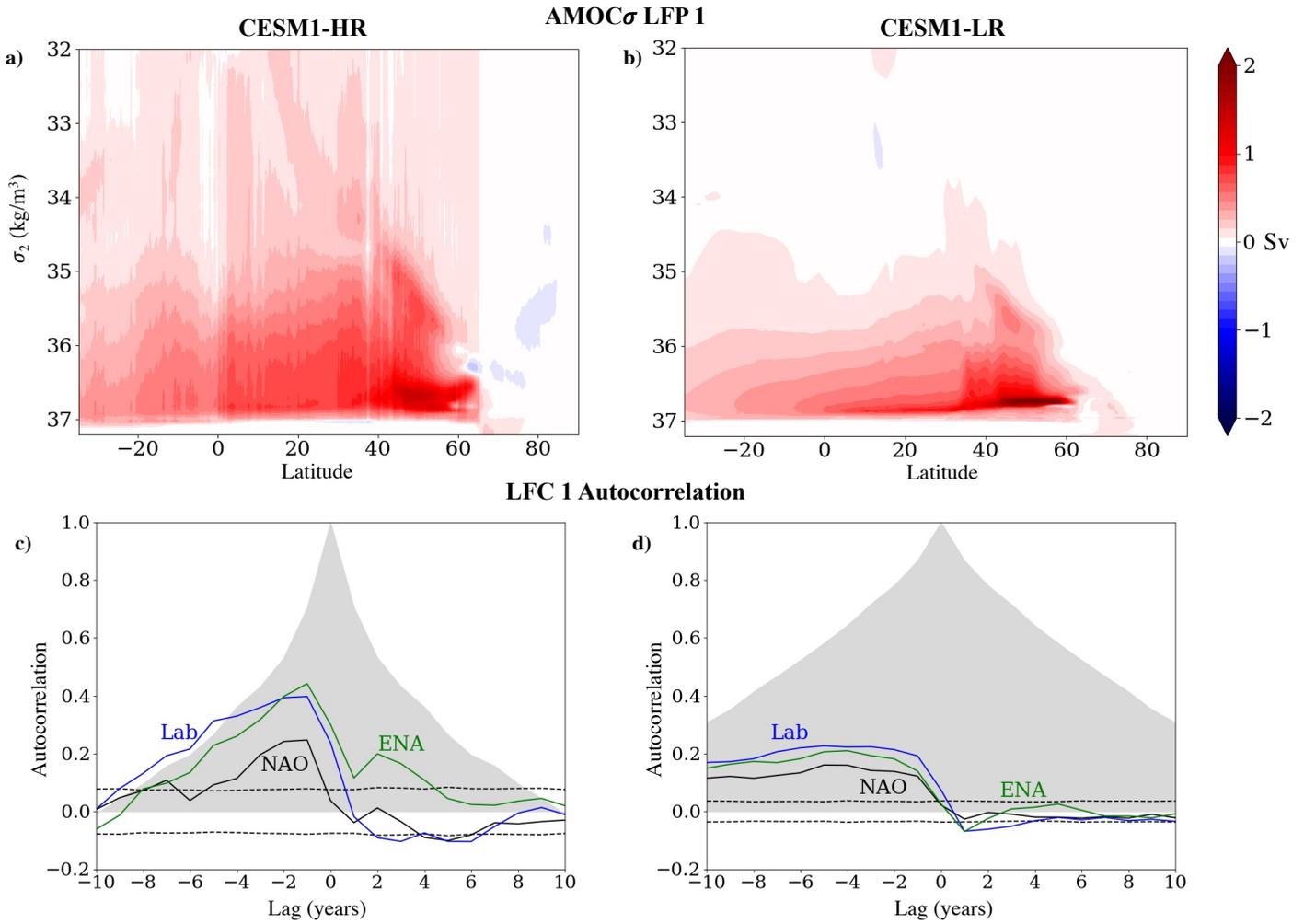


Figure 9: Top row: LFP 1 of AMOC for **a) CESM1-HR** and **b) CESM1-LR**. Bottom row: Autocorrelations of LFC 1 (shaded), correlation of the NAO with LFC 1 (solid black lines) and significance levels (dashed black lines), and correlations of both the Labrador Sea (blue lines) and Eastern North Atlantic (ENA; green lines) winter mixed-layer depths with LFC 1 for **c) CESM1-HR** and **d) CESM1-LR**. NAO here is defined as the difference between the sea-level pressure between the Azores (25.5°W, 37.5°N) and Iceland (21.5°W, 64.5°N). The ENA here includes both the Irminger and Iceland Basins and the GIN Seas.

343 in density coordinates in CESM1-HR and CESM1-LR's pre-industrial control simula-
 344 tions. We focus on the first low-frequency pattern (LFP) of AMOC, which is the lin-
 345 ear combination of the leading empirical orthogonal functions (EOFs) with the high-
 346 est ratio of low-frequency variance to total variance in its corresponding timeseries
 347 (i.e., its low-frequency component; LFC). The LFP (Fig. 9) represents the AMOC anomaly
 348 associated with a one standard deviation (1σ) anomaly in the corresponding LFC. Here,
 349 low-frequency variance is defined as the variance at 10-year and longer timescales,
 350 computed via the point-wise application of a Lanczos filter with a low-pass cutoff of
 351 10 years. Low-pass filtering is only used to identify the LFP, and all information about
 352 high-frequency variations is preserved. For both models, we include the six leading
 353 EOFs in the LFCA. The choice of the number of EOFs does not substantially influence
 354 the results for either model.

355 In our previous analysis of low-resolution coupled model simulations (Oldenburg
 356 et al., 2021), we found that WMT in the Labrador Sea plays a more substantial role
 357 in driving AMOC and OHT variability than would be expected based on its role in
 358 driving the climatology of AMOC and OHT. Here, we examine whether the model
 359 resolution affects this result, given that higher resolution models represent Labrador
 360 Sea processes much better than low-resolution ones (see section 3). Hence, here we
 361 carry out an analysis similar to Oldenburg et al. (2021) with a focus entirely on AMOC
 362 instead of Atlantic OHT. Our goal is to determine whether the mechanisms of low-
 363 frequency AMOC variability in low-resolution simulations still hold in high-resolution
 364 models. We first compute the LFPs and LFCs of annual-mean AMOC in CESM1-HR
 365 and CESM1-LR, then calculate lead-lag regressions between the first LFC and other
 366 fields, including winter mixed-layer depth, surface-forced WMT, winter sea-level pres-
 367 sure (SLP) and AMOC. Although the LFPs already give the AMOC anomaly at lag-
 368 0, the pattern of AMOC anomalies evolves over time and therefore can look differ-
 369 ent at lead and lag times.

370 The first LFPs of AMOC in CESM1-HR and CESM1-LR share some common fea-
 371 tures, with maxima in the mid to subpolar latitudes. In CESM1-HR, the maximum value
 372 is equal to 1.48 Sv and is located at 47° N and $\sigma_2 = 36.675 \text{ kg/m}^3$. In CESM1-LR,
 373 the maximum value is equal to 2.51 Sv and is located at 53.5° N and $\sigma_2 = 36.74 \text{ kg/m}^3$.
 374 This is substantially stronger and at a higher latitude and density than in CESM1-HR.
 375 The peak is also broader in CESM1-HR. The other major difference is that the pos-

376 itive values extend to lower densities in CESM1-HR compared to CESM1-LR. The ra-
377 tios of low-frequency to total variance for the LFPs are equal to 0.70 and 0.87 for CESM1-
378 HR and CESM1-LR, respectively. The LFC autocorrelations remain high for much longer
379 lag times in CESM1-LR compared to CESM1-HR (Fig. 9c, d). In CESM1-HR, the au-
380 tcorrelation drops off more quickly, reaching zero by lag 10 years (Fig. 9c). The lower
381 ratio of low-frequency to total variance in CESM1-HR indicates that that model's LFC
382 includes more high-frequency variability, and the lower autocorrelation is consistent
383 with an AMOC that changes more rapidly over lead and lag times (Fig. S3).

384 In CESM1-HR, there is a persistent SLP pattern linked to anomalous northwest-
385 erly winds off eastern North America starting at about lead 4 years (Fig. 10b). This
386 SLP pattern remains until lag zero, which is the time of maximum AMOC (Fig. 10a-
387 d). This pattern's persistence must result from processes in the ocean, because the per-
388 sistence time scale of atmospheric anomalies is less than a month (Ambaum & Hoskins,
389 2002). At lag zero, the SLP pattern becomes more zonal and the eastern SLP inten-
390 sifies (Fig. 10d). After lag zero, the pattern reverses (Fig. 10e, f) with a pattern that looks
391 similar to the negative phase of the NAO. In CESM1-LR, there is a similar SLP pat-
392 tern at lead times and at lag-0 (Fig. 10g-j). In both HR and LR models, the effect of
393 the SLP pattern at lead times on the subpolar winter mixed-layer depths can be seen
394 in Fig. S4, which shows deepening mixed-layer depths, particularly in the Labrador
395 Sea. The time evolution of Labrador Sea mixed-layer depth mirrors that of the NAO
396 (Fig. 9c, d). The ENA mixed-layer depth does follow the NAO to some degree, espe-
397 cially in CESM1-LR, but it doesn't mirror it to the same degree as the Labrador Sea
398 in CESM1-HR (Fig. 9c, d). After lag zero, the SLP pattern dissipates completely in CESM1-
399 LR (Fig. 10l). However, unlike many low-resolution models, including CESM1-LR and
400 the LR models discussed in Oldenburg et al. (2021), CESM1-HR shows a coherent SLP
401 pattern after the time of maximum AMOC. This indicates an atmospheric response
402 to the low-frequency AMOC variability not seen in the equivalent low-resolution model.
403 This response can also be seen in the negative lagged correlation of the NAO with LFC
404 1 (Fig. 9c), which peaks at a lag of 5 years.

405 The AMOC changes before and after the time of maximum AMOC can be seen
406 in Fig. S3, which shows a strengthening of AMOC at lead times and a weakening at
407 lag times in both models. In CESM1-HR, WMT in the Labrador Sea strengthens in the
408 years leading up to maximum AMOC, reaching its maximum at lead 2, concurrent

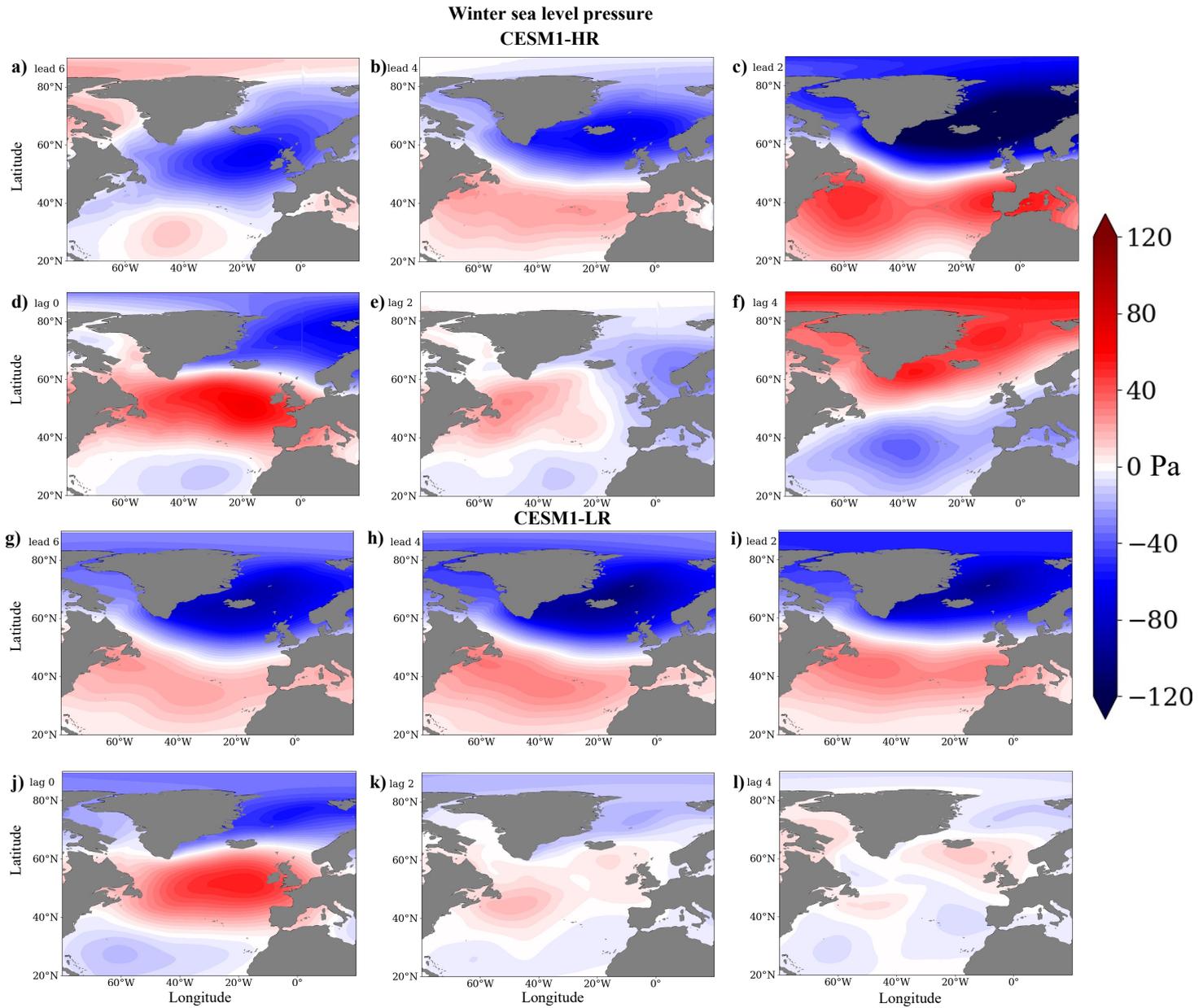


Figure 10: Lead-lag regressions of sea-level pressure averaged over January, February and March onto the first LFC of AMOC for (a-f) CESM1-HR and (g-l) CESM1-LR. Lead times represent anomalies that lead the LFC, i.e., before the time of maximum AMOC. Because the LFCs are unitless, the units of the regressions are Pa (N/m^2).

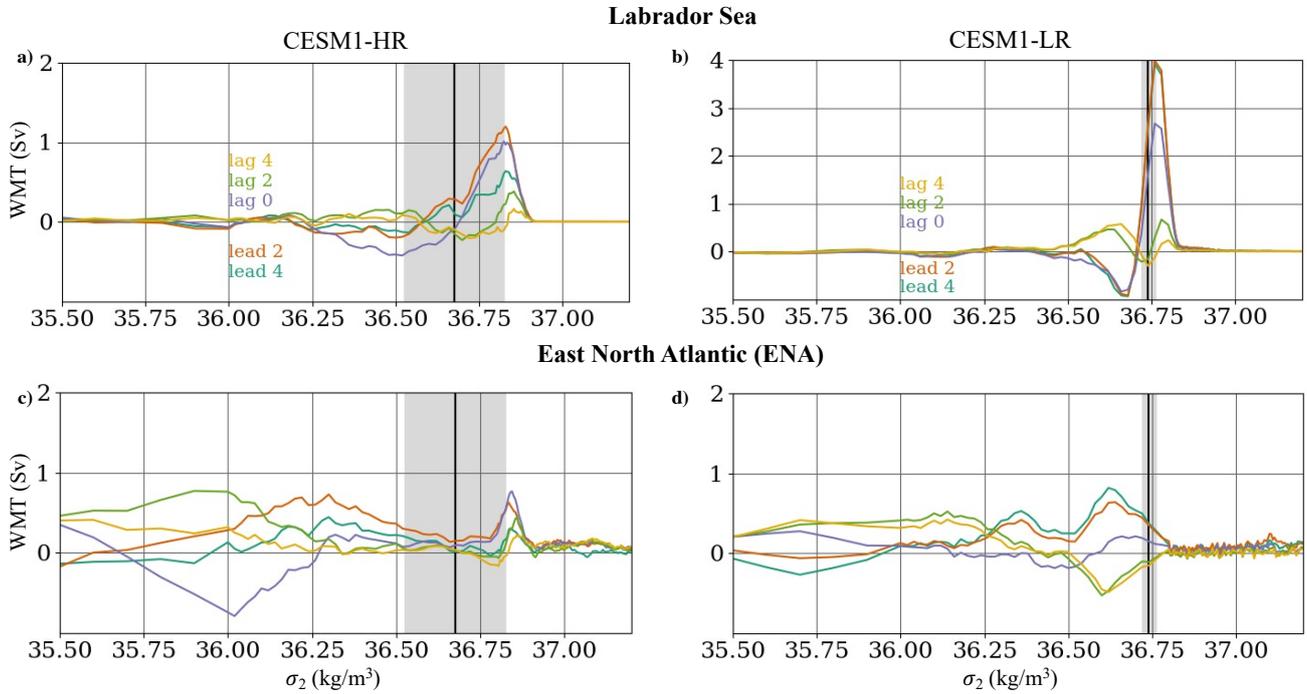


Figure 11: Lead-lag regressions of water mass transformation (WMT) onto the first LFC of AMOC for CESM1-HR (left column) and CESM1-LR (right column). **a, b**) WMT integrated throughout the Labrador Sea section. **c, d**) WMT integrated throughout the Eastern North Atlantic (ENA) region. The black vertical lines represent the density where the AMOC regression at lag zero is at its maximum in each model. The grey shaded areas show the density range where the AMOC regression at lag zero is within 25% of its maximum value. The black lines in Fig. 1 show what we consider to be the Labrador Sea, the Irminger and Iceland Basins, and the GIN Seas in this calculation. The ENA here includes both the GIN Seas and the Irminger and Iceland Basins. Lead times represent anomalies that lead the LFC, i.e., before the time of maximum AMOC. Because the LFCs are unitless, the units of the regressions are Sv.

409 with the strengthening of AMOC and the mixed layer deepening in the Labrador Sea,
410 IIB and GIN Seas (Fig. 11a). This peak is equal to 1.29 Sv and is located at $\sigma_2 = 36.83$
411 kg/m^3 , which is at a substantially higher density than the location of the maximum
412 AMOC anomaly at lag zero, but is still within the density range of the broad positive
413 AMOC anomaly. After lead 2, the WMT rapidly decreases. The Eastern North Atlantic
414 (ENA) WMT, which includes both the GIN Seas and IIB, also increases at lead times,
415 peaking at lead one years (Fig. 11c). This peak is equal to 0.83 Sv and is located at
416 $\sigma_2 = 36.84$, which is further from the peak in AMOC than the Labrador Sea WMT
417 peak. The peak in ENA WMT is mostly due to changes in IIB WMT rather than the
418 GIN Seas (not shown).

419 In CESM1-LR, the Labrador Sea WMT also increases at lead times, reaching its
420 maximum at lead 2 years (Fig. 11b). This maximum is equal to 3.99 Sv and is located
421 at $\sigma_2 = 36.76 \text{ kg/m}^3$, which is at a slightly higher density than the maximum AMOC
422 anomaly. The ENA WMT also strengthens at lead times, but already peaks by lead
423 4 years (Fig. 11d). This peak is equal to 0.82 Sv and is located at $\sigma_2 = 36.62 \text{ kg/m}^3$,
424 which is at a substantially lower density than the maximum AMOC anomaly. This
425 WMT increase is mostly due to changes in the IIB rather than the GIN Seas (not shown).

426 Based on these results, the mechanisms of AMOC variability between CESM1-
427 HR and CESM1-LR are qualitatively similar but still have quantitative differences. In
428 both models, the Labrador Sea plays a dominant role in driving low-frequency AMOC
429 variability, and the leading sea-level pressure patterns are similar. The primary dif-
430 ferences are that CESM1-HR, unlike CESM1-LR, shows a substantial atmospheric re-
431 sponse after the time of maximum AMOC, and that the Labrador Sea does not dom-
432 inate the WMT variability as much as it does in CESM1-LR.

433 5 Discussion and Conclusions

434 Based on the results from Section 3, a coupled model with increased atmospheric
435 and ocean resolutions accurately reproduces the WMT, sea-surface temperatures and
436 sea-surface salinities found in a reanalysis-forced high-resolution ocean simulation.
437 The ocean resolution appears to be particularly important, as even a low-resolution
438 ocean simulation forced with atmospheric reanalysis data doesn't represent the WMT
439 as accurately as the high-resolution coupled model simulation. This illustrates the im-

440 portance of resolving, rather than parameterizing, mesoscale eddies for the ability to
441 accurately represent mixed-layer depth and deep water formation, particularly in the
442 Labrador Sea.

443 The better representation of WMT is explained by a more accurate representa-
444 tion of the density structure in the high-resolution simulation compared to the low-
445 resolution simulations, which have relatively uniform density fields in comparison,
446 particularly in the Labrador Sea. Smaller discrepancies in surface heat fluxes in the
447 deep water formation regions in the high-resolution simulation also help explain why
448 it captures the climatological WMT better than the low-resolution simulations.

449 In section 4, we used LFCA to assess the mechanisms of low-frequency AMOC
450 variability in high- and low-resolution versions of the same model, finding that the
451 mechanisms are qualitatively similar but quantitatively different. The Labrador Sea
452 WMT still plays a major role in the WMT and AMOC variability in the high-resolution
453 model despite the fact that it shows a smaller role for the Labrador Sea in climato-
454 logical WMT and AMOC than the low-resolution version. The analysis here neglects
455 interior ocean mixing. However, despite the fact that most of the Labrador Sea WMT
456 changes occur at higher densities than the AMOC changes, the Labrador Sea's dom-
457 inance in AMOC variability likely still holds because mixing tends to make the dens-
458 est water lighter.

459 One noteworthy difference between the simulations is that the high-resolution
460 model shows a substantial atmospheric response to the AMOC variability not seen
461 in the low-resolution version. This type of atmospheric response has been seen in a
462 study of a medium-resolution coupled model, but with a longer lag time between the
463 AMOC change and the negative NAO response (Wen et al., 2016). NAO-like responses
464 of differing signs to AMOC variability have also been found in other studies (Dong
465 & Sutton, 2003; Gastineau & Frankignoul, 2012; Gastineau et al., 2013; Frankignoul et
466 al., 2013, 2015). The model simulations we analyzed here do not give insight into whether
467 the atmospheric or oceanic resolution is responsible for the increased atmospheric re-
468 sponse to AMOC variability in CESM1-HR, but recent work suggests that the atmo-
469 spheric response to near-surface ocean anomalies is larger at higher atmospheric res-
470 olution (e.g., Czaja et al. (2019)). Overall, it appears that the mode of AMOC variabil-
471 ity in the high-resolution model is associated with stronger anomalies in atmospheric

472 fields (i.e., sea-level pressure), while the low-resolution version is associated with stronger
 473 anomalies in ocean fields, namely in the water-mass transformation, particularly in
 474 the Labrador Sea.

475 Our results suggest that increasing the ocean and atmospheric resolution of a cou-
 476 pled model substantially improves the representation of climatological AMOC and
 477 WMT. However, the mechanisms driving low-frequency AMOC variability remain qual-
 478 itatively similar even though the climatologies differ. This is consistent with what was
 479 found in three low-resolution coupled models with distinct representations of WMT
 480 in the different subpolar North Atlantic deep water formation regions, which all showed
 481 similar mechanisms of AMOC and OHT variability, with the Labrador Sea playing
 482 a dominant role (Oldenburg et al., 2021).

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 488 able. CESM1-LR data can be downloaded at <https://earthsystemgrid.org>. CESM1-HR
 489 data can be found at <https://ihesp.tamu.edu/>. JRA55-LR and JRA55-HR data are avail-
 490 able from the NCAR server. Post-processed data used to create some of the figures
 491 in this paper can be found at <https://faculty.washington.edu/karmour/code/Oldenburg/>.
 492 MATLAB and Python code for LFCA is available at <https://github.com/rcjwills/lfca>.

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