

1 Diagnosing the thickness-weighted averaged eddy-mean 2 flow interaction in an eddying North Atlantic ensemble

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8 **Key Points:**

- 9 • Mesoscale-resolving ensemble runs of the North Atlantic Ocean are used to diag-
10 nose the thickness-weighted averaged eddy-mean flow interaction.
- 11 • The Eliassen-Palm flux divergence, which is directly related to the eddy Ertel po-
12 tential vorticity (PV) flux, tends to meridionally decelerate the Gulf Stream.
- 13 • The eddy Ertel PV flux can be parametrized as a local-gradient flux of the residual-
14 mean Ertel PV via an anisotropic eddy diffusivity tensor.

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15 **Abstract**

16 The thickness-weighted average (TWA) framework, which treats the residual-mean flow
 17 as the prognostic variable, has provided us with a clear theoretical understanding of the
 18 eddy feedback onto the residual-mean flow. The averaging operator involved in the TWA
 19 framework, although in theory being an ensemble mean, in practice has often been ap-
 20 proximated by a temporal mean, which conflates the temporal variability with the ed-
 21 dies. Here, we analyze an ensemble of North Atlantic simulations at mesoscale resolv-
 22 ing resolution ($1/12^\circ$). We therefore recognize means and eddies in terms of ensemble
 23 means and fluctuations about those means, in keeping with the TWA formalism proposed
 24 by Young (2012). Eddy-mean flow feedbacks are encapsulated in the Eliassen-Palm (E-
 25 P) flux tensor and its divergence indicates that the eddies contribute to the zonal me-
 26 andering of the Gulf Stream and its deceleration in the meridional direction. We also
 27 show that the eddy Ertel potential vorticity (PV) flux can be parametrized as an isopy-
 28 cnic local-gradient flux of the residual-mean Ertel PV via an anisotropic eddy diffusiv-
 29 ity tensor. As the E-P flux divergence and eddy Ertel PV flux are directly related to one
 30 another, this provides a new pathway forward for a unified mesoscale eddy closure scheme.

31 **Plain Language Summary**

32 We have greatly benefited from global climate simulations in gaining insight into
 33 what the climate would look like in an ever warming future. Due to computational con-
 34 straints, however, the oceanic component of such simulations have been poorly constrained;
 35 the storm systems in the ocean, often referred to as eddies, have the spatial scales of roughly
 36 50 km and simulating this accurately on a global scale, which is on the order of 1000 km,
 37 has remained challenging. Although relatively small in scale compared to the global Earth,
 38 eddies have been known to modulate the climate by transporting heat from the equa-
 39 tor to the poles. By running a regional simulation of the North Atlantic Ocean and tak-
 40 ing advantage of recent theoretical development, we provide a new pathway in improv-
 41 ing the representation of these eddies and as such, improving global ocean and climate
 42 simulations.

43 **1 Introduction**

44 Eddy-mean flow interaction has been a key framework in understanding jet forma-
 45 tion in geophysical flows such as in the atmosphere and ocean (e.g. Vallis, 2017, Chap-

ters 12 and 15). A prominent example of such a jet in the North Atlantic ocean is the Gulf Stream. Previous studies have shown how eddies fluxing buoyancy and momentum back into the mean flow energize the Gulf Stream (Lévy et al., 2010; Waterman & Lilly, 2015; Chassignet & Xu, 2017; Aluie et al., 2018). Basin-scale simulations, however, often lack sufficient spatial resolution to accurately resolve the eddies and hence, result in underestimating the eddy fluxes of momentum and tracers (Capet et al., 2008; Arbic et al., 2013; Kjellsson & Zanna, 2017; Balwada et al., 2018; Uchida et al., 2019; Schubert et al., 2020). Due to computational constraints, we will continue to rely on models which only partially resolve the mesoscale, a scale roughly on the order of $O(20\text{-}200\text{ km})$ at which the ocean currents are most energetic (Stammer, 1997; Xu & Fu, 2011, 2012; Ajayi et al., 2020), for global ocean and climate simulations. As a result, there has been an ongoing effort to develop energy-backscattering eddy parametrizations which incorporate the dynamical effects of eddy momentum fluxes due to otherwise unresolved mesoscale turbulence (e.g. Kitsios et al., 2013; Anstey & Zanna, 2017; Zanna & Bolton, 2020; Bachman et al., 2018; Bachman, 2019; Jansen et al., 2019; Perezhogin, 2019; Juricke et al., 2020).

There has been less emphasis, however, on quantifying the spatial and temporal characteristics of the eddy buoyancy and momentum fluxes themselves, which the parametrizations are deemed to represent. The focus of this study is, therefore, to examine the dynamical effects of mesoscale turbulence on the mean flow in realistic partially air-sea coupled eddying ensemble runs of the North Atlantic. To achieve this goal, we employ the thickness-weighted average (TWA) framework developed by De Szoeke and Bennett (1993), Young (2012), J. R. Maddison and Marshall (2013) and Aoki (2014), which treats the residual-mean velocity as a prognostic variable and allows for a straightforward theoretical understanding of the eddy feedback onto the (residual) mean flow; the framework has been fruitful in examining eddy-mean flow interaction in idealized modelling studies (e.g. D. P. Marshall et al., 2012; Cessi & Wolfe, 2013; Ringler et al., 2017; Bire & Wolfe, 2018).

To our knowledge, Aiki and Richards (2008), Aoki et al. (2016) and Zhao and Marshall (2020) are the only studies that diagnose the TWA framework in realistic ocean simulations. Aiki and Richards (2008), however, recompute the hydrostatic pressure using potential density for their off-line diagnosis in defining their buoyancy coordinate, which can result in significant discrepancies from the pressure field used in their on-line cal-

79 culation and consequently errors in the diagnosed geostrophic shear. Aoki et al. (2016)
80 negate this complication between the buoyancy coordinate and pressure field by analyz-
81 ing their outputs in geopotential coordinates but with the limitation of second-order ac-
82 curacy in eddy fluxes, and all three studies assume ergodicity. The ergodic assumption
83 of treating a temporal mean equivalent to an ensemble mean, although a pragmatic one,
84 prevents examining the temporal evolution of the residual mean fields and conflates tem-
85 poral variability with the eddies, which can have leading-order consequences in quan-
86 tifying the energy cycle (c.f. Aiki & Richards, 2008, Table 2 in their paper); by adjust-
87 ing the temporal mean from monthly to annual, they show that the amount of kinetic
88 and potential energy stored in the mean and eddy reservoirs can change up to a factor
89 of four. Eddy-mean flow interaction in the TWA framework, hence, warrants further in-
90 vestigation, and we believe our study is the first to strictly implement an ensemble mean
91 in this context.

92 When discussing *eddy* versus *mean flow*, one of the ambiguities lies in how the two
93 are decomposed (Bachman et al., 2015). As noted above, often, the eddies are defined
94 from a practical standpoint as the deviation from a temporally and/or spatially coarse-
95 grained field regardless of the coordinate system (e.g. Aiki & Richards, 2008; Lévy et
96 al., 2012; Sasaki et al., 2014; Griffies et al., 2015; Aoki et al., 2016; Uchida et al., 2017;
97 Zhao & Marshall, 2020), which leaves open the question of how the filter affects the de-
98 composition. Due to the ensemble averaging nature of the TWA framework, we are uniquely
99 able to define the two; the *mean flow* (ensemble mean) is the predictable field determined
100 by the surface boundary forcings and the *eddy* (residual from the ensemble mean) the
101 field due to intrinsic variability of mesoscale turbulence (Sérazin et al., 2017; Leroux et
102 al., 2018).

103 The paper is organized as follows: We describe the model configuration in section 2
104 and briefly provide an overview of the TWA framework in section 3. The results are given
105 in section 4. In particular, we highlight in section 4.2 how the Eliassen-Palm (E-P) flux
106 divergence is related to the Ertel potential vorticity (PV) and that it can be parametrized
107 via a local-gradient flux closure. Discussion and conclusions are given in section 5.

108 2 Model description

109 The model configuration is similar to the realistic air-sea coupled runs in Jamet
 110 et al. (2019a); Jamet et al. (2019b) and Jamet et al. (2020) where we run 12 air-sea par-
 111 tially coupled ensemble members of the North Atlantic ocean at mesoscale resolving res-
 112 olution ($1/12^\circ$) using the hydrostatic configuration of the Massachusetts Institute of Tech-
 113 nology general circulation model (MITgcm; J. Marshall et al., 1997). We have 46 ver-
 114 tical levels increasing from 6 m near the surface to 250 m at depth. Harmonic, bihar-
 115 monic horizontal and vertical viscosity values of $A_{h2} = 20 \text{ m}^2 \text{ s}^{-1}$, $A_{h4} = 10^{10} \text{ m}^4 \text{ s}^{-1}$
 116 and $A_v = 10^{-5} \text{ m}^2 \text{ s}^{-1}$ were used respectively. For completeness, we provide a brief sum-
 117 mary of the configuration below.

118 Figure 1 shows the bathymetry of the modelled domain extending from 20°S to 55°N .
 119 In order to save computational time and memory allocation, the North Atlantic basin
 120 was configured to zonally wrap around periodically. Open boundary conditions are ap-
 121 plied at the north and south boundaries of our domain and Strait of Gibraltar, such that
 122 oceanic velocities (\mathbf{u}) and tracers (θ, s) are restored with a 36 minutes relaxation time
 123 scale toward a state derived by an ocean-only global Nucleus for European Modelling
 124 of the Ocean (NEMO) simulation (Molines et al., 2014, ORCA12.L46-MJM88 run in their
 125 paper). A sponge layer is applied to the two adjacent grid points at the open boundaries
 126 where model variables are restored toward boundary conditions with a one-day relax-
 127 ation time scale. Although relatively short, no adverse effects were apparent upon in-
 128 spection in response to these relaxation time scales (not shown). The open boundary con-
 129 ditions are applied every five days and linearly interpolated in between.

130 The 12-member ensemble was constructed as follows: From a simulation run un-
 131 der yearly repeating atmospheric forcing and boundary conditions taken from the (Jamet
 132 et al., 2020, Open boundary conditions Climatologic and Atmosphere Climatologic (OCAC)
 133 ensemble runs in their paper), 12 snapshots of the OCAC model state with a four-year
 134 interval between each are selected and used to initialise each ensemble member used in
 135 this study. Such initial conditions are meant to reflect the growth of dynamically con-
 136 sistent oceanic perturbations de-correlated at both seasonal and interannual time scales;
 137 they are usually referred to as macro initial conditions (Stainforth et al., 2007; Hawkins
 138 et al., 2016). At the surface, the ocean is partially coupled to an atmospheric boundary
 139 layer model (CheapAML; Deremble et al., 2013). In CheapAML, atmospheric surface

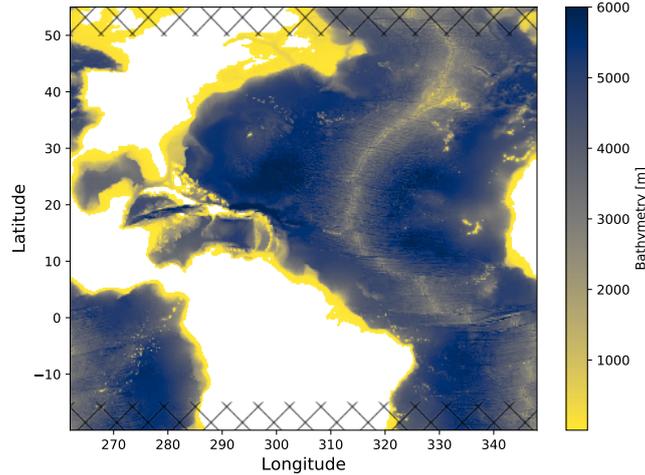


Figure 1. Bathymetry of the modelled domain. The domain is configured to wrap around zonally. The hatches indicate the northern and southern regions excluded from our analysis.

140 temperature and relative humidity respond to ocean surface structures by exchanges of
 141 heat and humidity computed according to the Coupled Ocean–Atmosphere Response Ex-
 142 periment (COARE3; Fairall et al., 2003) flux formula, but are strongly restored toward
 143 prescribed values over land; there are no zonally propagating signals of climate telecon-
 144 nection. The 12 ensemble members are then integrated forward in time for five years (1963-
 145 1967), and exposed to the same realistic forcing across all ensemble members; the sur-
 146 face forcing is taken from the Drakkar forcing set and boundary forcing from the ORCA12.L46-
 147 MJM88 run (details are given in Jamet et al., 2019a). During this interval, the oceanic
 148 state and the atmospheric boundary layer temperature and humidity evolve in time. In
 149 the following, we interpret the ensemble mean as the ocean response to the prescribed
 150 atmospheric forcing, while the ensemble spread is attributed to intrinsic ocean dynam-
 151 ics that develop at mesoscale-resolving resolution (Sérazin et al., 2017; Leroux et al., 2018;
 152 Jamet et al., 2019b).

153 The model outputs were saved as instantaneous snapshots every five days and five-
 154 day averages. We examined whether the five-day temporal smoothing would affect the
 155 terms in the E-P flux described below, upon which we found the difference between the
 156 eddy-eddy correlation terms diagnosed from instantaneous and temporally averaged fields
 157 to be up to same order of the total variance of the instantaneous field (not shown). The
 158 additional dimension of ensembles also negates the necessity for any temporal averag-
 159 ing to define the mean. As such, in the following analysis, we will only use the instan-

160 taneous snapshot outputs. We also exclude the northern and southern extent of 5° and
 161 first year of integration from our analysis to avoid effects from the open boundary con-
 162 ditions, sponge layer and initialization shock, and to maximize the intrinsic variability
 163 amongst the ensemble members respectively.

164 **3 Theory and implementation of thickness-weighted averaging**

The ocean is a stratified fluid, and the circulation and advection of tracers tend to align themselves along the stratified density surfaces. Hence, the most natural way to understand the circulation is to consider the variables in a thickness-weighted form and the residual-mean flow rather than the Eulerian mean flow. We leave the detailed derivation of the TWA framework to Young (2012) and here, only provide a brief summary; the primitive equations in geopotential coordinates are first transformed to buoyancy coordinates upon which a thickness weighting and ensemble averaging along constant buoyancy surfaces is applied to obtain the TWA governing equations. Following Young’s notation, the TWA horizontal momentum equations in the buoyancy coordinate system $(\tilde{t}, \tilde{x}, \tilde{y}, \tilde{b})$ are:

$$\hat{u}_{\tilde{t}} + \hat{u}\hat{u}_{\tilde{x}} + \hat{v}\hat{u}_{\tilde{y}} + \hat{\omega}\hat{u}_{\tilde{b}} - f\hat{v} + \overline{m}_{\tilde{x}} = -\bar{\mathbf{e}}_1 \cdot (\nabla \cdot \mathbf{E}) + \hat{\mathcal{X}} \quad (1)$$

$$\hat{v}_{\tilde{t}} + \hat{u}\hat{v}_{\tilde{x}} + \hat{v}\hat{v}_{\tilde{y}} + \hat{\omega}\hat{v}_{\tilde{b}} + f\hat{u} + \overline{m}_{\tilde{y}} = -\bar{\mathbf{e}}_2 \cdot (\nabla \cdot \mathbf{E}) + \hat{\mathcal{Y}} \quad (2)$$

165 where $\widehat{(\cdot)}$ and $\overline{(\cdot)}$ are the TWA and ensemble-mean variables respectively and the sub-
 166 scripts denote partial derivatives. The Montgomery potential is $m = \phi - \tilde{b}\zeta$ where ϕ
 167 is the dynamically active part of hydrostatic pressure and ζ is the isopycnal depth. The
 168 vectors $\bar{\mathbf{e}}_1 = \mathbf{i} + \bar{\zeta}_{\tilde{x}}\mathbf{k}$ and $\bar{\mathbf{e}}_2 = \mathbf{j} + \bar{\zeta}_{\tilde{y}}\mathbf{k}$ form the basis vectors spanning the buoyancy
 169 horizontal space where \mathbf{i} , \mathbf{j} and \mathbf{k} are the Cartesian geopotential unit vectors (Young, 2012;
 170 Ringler et al., 2017), and \mathbf{E} is the E-P flux tensor described in detail in Section 4.1. \mathcal{X}
 171 and \mathcal{Y} are the viscous and forcing terms.

One subtle yet important point involves the buoyancy coordinate (\tilde{b}) for a realistic, non-linear equation of state (EOS) for density (Jackett & McDougall, 1995). The analysis in Young (2012) implicitly assumes a linear EOS. With a realistic EOS the vertical coordinate can no longer “naively” be defined by neutrally-surfaced isopycnals which allow for adiabatic adjustment, such as potential density, and is the subject of some debate. We argue for the use of in-situ density (detailed arguments are given in Appendix A); other choices can be made (e.g. Stanley, 2019). The consequence of using in-situ buoy-

ancy as the coordinate ($\tilde{b} = -g(\rho - \rho_0)/\rho_0$ where ρ is the in-situ density and $\rho_0 = 999.8 \text{ kg m}^{-3}$ is the Boussinesq reference density) is that buoyancy is no longer conserved due to compressibility effects even under adiabatic conditions, i.e.

$$\frac{D\tilde{b}}{Dt} = \frac{wg^2}{c_s^2} \quad (3)$$

172 where c_s is the speed of sound and the vertical velocity is $w = \frac{D\zeta}{Dt}$. The right-hand side
 173 of eqn. (3) can be included with the diapycnal velocity $\varpi = \tilde{b}_\theta \dot{\theta} + \tilde{b}_s \dot{s} - \tilde{b}_\Phi \rho_0 \omega g$ where
 174 $\dot{\theta}$ and \dot{s} are the net diabatic contributions on potential temperature and practical salin-
 175 ity respectively calculated online and outputted as diagnostics, and $\tilde{b}_\Phi \stackrel{\text{def}}{=} -\frac{g}{\rho_0 c_s^2}$ is the
 176 compressibility effect due to the dynamically non-active part of the hydrostatic pressure.
 177 The use of in-situ buoyancy maintains the desirable properties of a unique, statically sta-
 178 ble vertical coordinate, a simple hydrostatic balance ($\sigma = \zeta_{\tilde{b}} = -m_{\tilde{b}\tilde{b}}$), the imperme-
 179 ability of Ertel PV ($\Pi = \sigma^{-1}(f + v_{\tilde{x}} - u_{\tilde{y}})$) and nonacceleration conditions (Young,
 180 2012) with a correction due to compressibility (the pressure term in ϖ). Discussion re-
 181 garding the energetics are given in Appendix B.

The raw simulation outputs were in geopotential coordinates so we first remapped all of the variables in eqns. (1) and (2) onto 60 buoyancy levels spread linearly across the range of $\tilde{b} \in [-0.196, -0.491] \text{ [m s}^{-2}\text{]}$:

$$(\mathbf{u}, z, \nabla_{\mathbf{h}}\phi, \theta, s, \rho_\theta \dot{\theta}, \rho_s \dot{s}, \frac{wg^2}{c_s^2})(t, x, y, z) \mapsto (\mathbf{u}, \zeta, \tilde{\nabla}_{\mathbf{h}}m, \theta, s, \rho_\theta \dot{\theta}, \rho_s \dot{s}, \frac{wg^2}{c_s^2})(\tilde{t}, \tilde{x}, \tilde{y}, \tilde{b}) \quad (4)$$

using the `fastjmd95` Python package to compute the in-situ density and its partial derivatives (Abernathey, 2020), and the `xlayers` Python package (Jones, 2019; Jones et al., 2020) which implements the MITgcm layers package off-line and allows for coordinate remapping consistent with the finite-volume discretization. For the horizontal pressure gradient, we have invoked the identity:

$$\nabla_{\mathbf{h}}\phi(z) \mapsto \nabla_{\mathbf{h}}\phi(\tilde{b}) = \tilde{\nabla}_{\mathbf{h}}m \quad (5)$$

182 where the subscript $(\cdot)_{\mathbf{h}}$ represents the horizontal gradient and $\tilde{\nabla}_{\mathbf{h}} = (\partial_{\tilde{x}}, \partial_{\tilde{y}})$. Note iden-
 183 tity (5) only holds when in-situ buoyancy is used as the coordinate system, see eqns. (A1-
 184 A3).

185 4 Results

We start by showing the time series of domain-averaged horizontal kinetic energy (KE) and potential temperature (Fig. 2a). The OCAC run, from which the initial con-

ditions were taken, had a long-term warming trend so we corrected for this in our off-line analysis by removing the domain-integrated residual heat content from the ensemble mean at the initial time step defined as:

$$\Delta\theta_i \stackrel{\text{def}}{=} \frac{\int C_p(\theta_i - \bar{\theta})dV}{\int C_p dV} \quad (6)$$

186 where the subscript i ($= [1, 2, \dots, 12]$) is the ensemble index and $C_p = 3929.245 \text{ J kg}^{-1} \text{ K}^{-1}$
 187 is the specific heat capacity (McDougall, 2003); the residual per ensemble member ($\Delta\theta_i$)
 188 was removed from the potential temperature throughout the five years of output prior
 189 to the following analysis. We also corrected for salinity in the same manner. Figure 2a
 190 shows the simulation has a prominent seasonal cycle with a slight cooling trend and in-
 191 crease in KE.

192 In Fig. 2, we also show the (residual) mean fields on December 26, 1963, the last
 193 day of the spin-up period. The depth of the buoyancy level shown in Fig. 2c is below the
 194 ensemble-mean mixed-layer depth (MLD; Fig. 2b) basin wide where diabatic effects are
 195 small. We focus on this buoyancy level for the remainder of this study as it is below the
 196 MLD and the isopycnal does not outcrop but is shallow enough to capture the imprint
 197 of the Gulf Stream and eddies. The ensemble-mean MLD was computed as the depth
 198 at which the potential density computed from ensemble-mean temperature and salin-
 199 ity fields increased by 0.03 kg m^{-3} from the density at 10 m depth ($\overline{\text{MLD}} \stackrel{\text{def}}{=} \text{MLD}(\bar{\theta}, \bar{s})$;
 200 de Boyer Montégut et al., 2004). The mean KE fields ($K^\# \stackrel{\text{def}}{=} |\hat{\mathbf{u}}|^2/2$; Fig. 2d) show the
 201 characteristic features of the Gulf Stream, North Brazil Current and equatorial under-
 202 current. The mean Rossby number ($\text{Ro}^\# \stackrel{\text{def}}{=} f^{-1}(\hat{v}_x - \hat{u}_y)$) shown in Fig. 2e is smaller
 203 than unity except for near the equator where the Coriolis parameter becomes small, in-
 204 dicated that over most of the North Atlantic basin, the mean flow in the interior is bal-
 205 anced. The kinematics of discretizing the gradients in buoyancy coordinates are given
 206 in Appendix C. We now move on to examine the eddy feedback onto the mean flow.

207 4.1 The Eliassen-Palm flux tensor

The E-P flux tensor (\mathbf{E}) in the TWA framework (eqns. (1) and (2)) is:

$$\mathbf{E} = \begin{pmatrix} \widehat{u''u''} + \frac{1}{2\bar{\sigma}}\overline{\zeta'^2} & \widehat{u''v''} & 0 \\ \widehat{v''u''} & \widehat{v''v''} + \frac{1}{2\bar{\sigma}}\overline{\zeta'^2} & 0 \\ \overline{\varpi''u''} + \frac{1}{\bar{\sigma}}\overline{\zeta'm'_x} & \overline{\varpi''v''} + \frac{1}{\bar{\sigma}}\overline{\zeta'm'_y} & 0 \end{pmatrix} \quad (7)$$

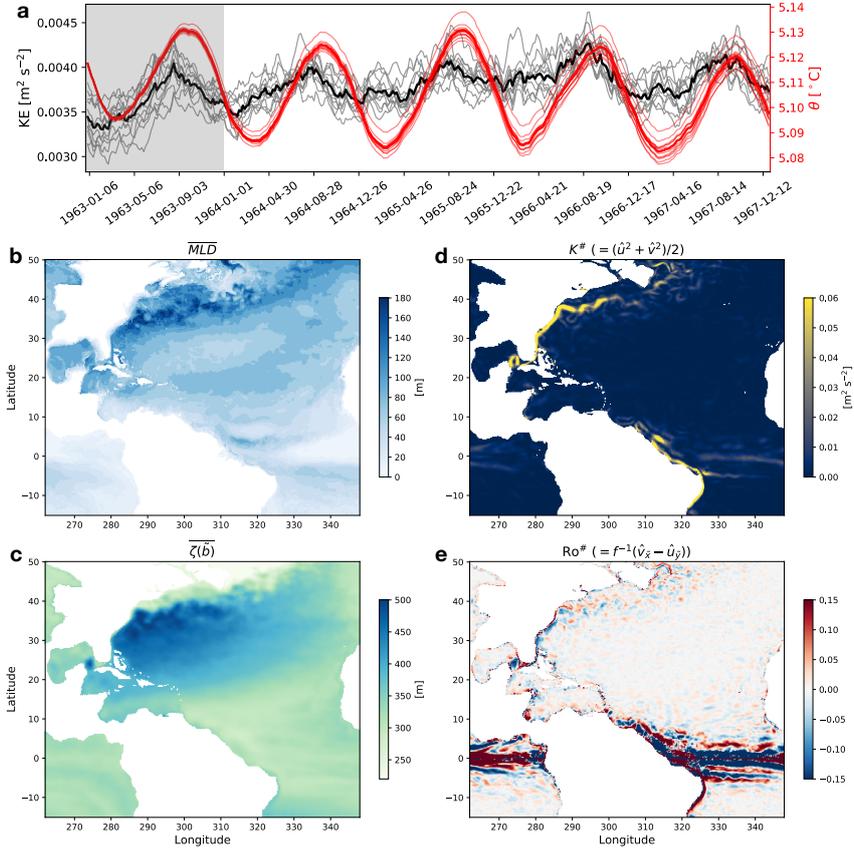


Figure 2. Time series of the domain-averaged KE (black) and temperature corrected for its residual heat content (red) for the 12 ensemble members between 15S-50N. The thick lines represent the ensemble mean and the grey shading indicates the one-year spin up period **a**. **b,c** The ensemble-mean MLD on December 26, 1963 and isopycnal depth with the buoyancy $\tilde{b} = -0.276$ [m s⁻²]. **d,e** The TWA kinetic energy ($K^\#$) and Rossby number ($Ro^\#$) on the isopycnal.

where $(\cdot)'' = (\cdot) - \widehat{(\cdot)}$ and $(\cdot)' = (\cdot) - \overline{(\cdot)}$ are the residual of instantaneous snapshot outputs from the thickness-weighted and ensemble averages respectively (J. R. Maddison & Marshall, 2013; Aoki, 2014; Ringler et al., 2017). The two are related via the quasi-Stokes velocity (Greatbatch, 1998; McDougall & McIntosh, 2001):

$$\begin{aligned}\mathbf{u}'' &= \mathbf{u} - \frac{\overline{\mathbf{u}\sigma}}{\overline{\sigma}} = \overline{\mathbf{u}} + \mathbf{u}' - \frac{(\overline{\mathbf{u}} + \mathbf{u}')(\overline{\sigma} + \sigma')}{\overline{\sigma}} \\ &= \mathbf{u}' + \frac{\overline{\mathbf{u}'\sigma'}}{\overline{\sigma}}.\end{aligned}$$

208 We show each term in eqn. (7) in Fig. 3. The Reynolds stress term $\widehat{u''v''}$ is associated
 209 with barotropic processes (e.g. Vallis, 2017, Chapter 15). The eddy momentum flux terms
 210 $|\widehat{\mathbf{u}''}|^2$ in Fig. 3b,e are seen to exchange momentum between eddies and the mean flow,
 211 i.e. to accelerate or decelerate the Gulf Stream. The interfacial form stress $(\zeta'\tilde{\nabla}_h m')$; Fig. 3c,f)
 212 associated with baroclinic instability is “deceivingly” orders of magnitude smaller than
 213 the other terms. The contribution from the adiabatic and compressibility effects (i.e. the
 214 terms with ϖ) were smaller than the interfacial form stress by another order of magni-
 215 tude or more in the subtropics (not shown). It is important to keep in mind, however,
 216 that it is the divergence of the E-P flux and not the flux itself that goes into the momen-
 217 tum equation, and the horizontal ($\tilde{\nabla}_h$) and vertical gradient ($\partial_{\tilde{b}}$) differ by roughly $O(10^6)$.

218

Writing out the E-P flux divergence in eqns. (1) and (2) gives:

$$-\mathbf{e}_1 \cdot (\tilde{\nabla} \cdot \mathbf{E}) = -\overline{\sigma}^{-1} \left([\overline{\sigma}(\widehat{u''u''} + \frac{1}{2\overline{\sigma}}\zeta'^2)]_{\tilde{x}} + [\overline{\sigma}(\widehat{v''u''})]_{\tilde{y}} + [\overline{\sigma}(\widehat{\varpi''u''} + \frac{1}{\overline{\sigma}}\zeta'm'_x)]_{\tilde{b}} \right) \quad (8)$$

$$= -\overline{\sigma}^{-1} \left([\overline{\sigma}u''u'']_{\tilde{x}} + \zeta'^2/2]_{\tilde{x}} + [\overline{\sigma}v''u'']_{\tilde{y}} + [\overline{\sigma}\varpi''u'' + \zeta'm'_x]_{\tilde{b}} \right), \quad (9)$$

$$-\mathbf{e}_2 \cdot (\tilde{\nabla} \cdot \mathbf{E}) = -\overline{\sigma}^{-1} \left([\overline{\sigma}(\widehat{u''v''})]_{\tilde{x}} + [\overline{\sigma}(\widehat{v''v''} + \frac{1}{2\overline{\sigma}}\zeta'^2)]_{\tilde{y}} + [\overline{\sigma}(\widehat{\varpi''v''} + \frac{1}{\overline{\sigma}}\zeta'm'_y)]_{\tilde{b}} \right) \quad (10)$$

$$= -\overline{\sigma}^{-1} \left([\overline{\sigma}u''v'']_{\tilde{x}} + [\overline{\sigma}v''v'' + \zeta'^2/2]_{\tilde{y}} + [\overline{\sigma}\varpi''v'' + \zeta'm'_y]_{\tilde{b}} \right). \quad (11)$$

219 Figure 4 shows that the divergence of interfacial form stress becomes a leading-order term
 220 along with the eddy momentum fluxes in the Gulf Stream region. The magnitude of the
 221 divergence of Reynolds stress term attributable to barotropic instability is the smallest
 222 (Fig. 4b,e). In the North Brazil Current region, the signal of baroclinic instability is in-
 223 significant (Fig. 4c,g) relative to barotropic instability and eddy momentum fluxes (Fig. 4a,b,e,f).
 224 It is quite surprising that the signal of the equatorial undercurrents, although having high
 225 KE (Fig. 2d), is significantly smaller than in the Gulf Stream and North Brazil Current
 226 regions, virtually not visible in Figs. 3 and 4. This implies that the residual-mean flow
 227 dominates over the eddies in the equatorial region.

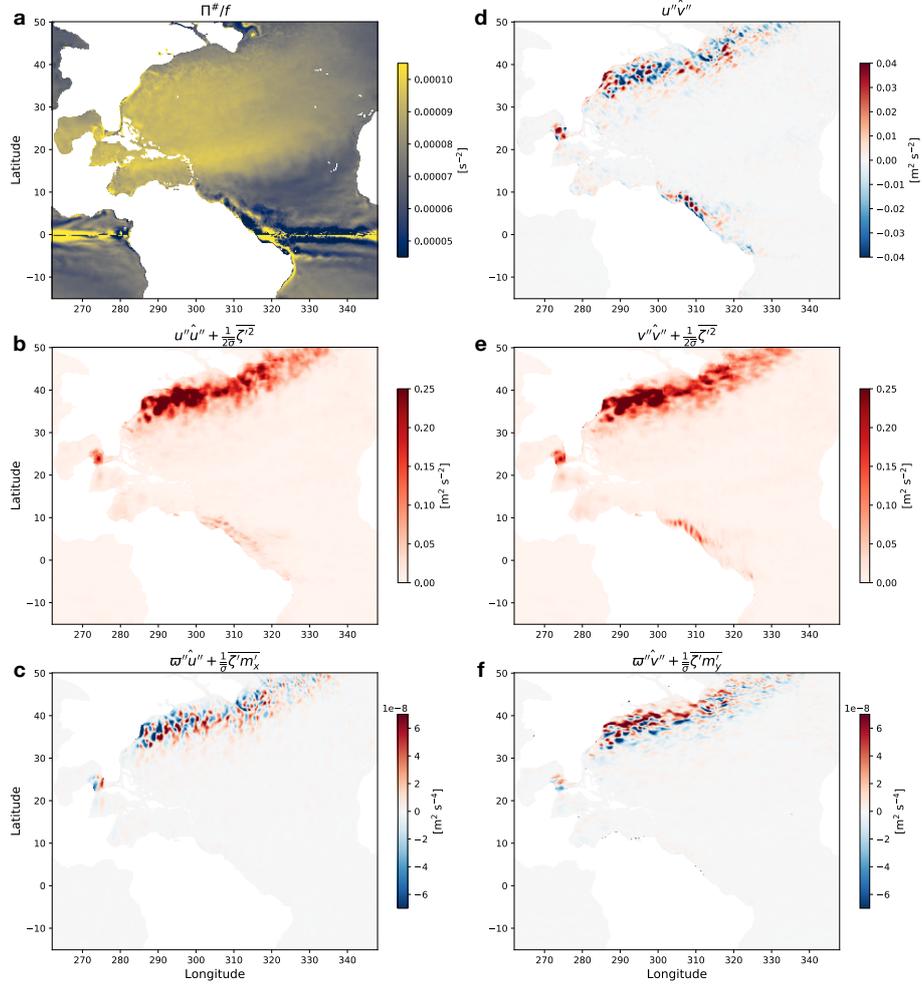


Figure 3. The residual-mean Ertel potential vorticity normalized by the local Coriolis parameter ($\Pi^\# / f \stackrel{\text{def}}{=} \bar{\sigma}^{-1}(1 + \text{Ro}^\#)$) **a** and terms in the E-P flux tensor **b-f** on December 26, 1963 on the isopycnal layer as in Fig. 2.

228 We now examine further details in the Gulf Stream region. The dipole features in
 229 the zonal direction of Reynolds stress and interfacial form stress likely contribute to the
 230 jet meandering (Fig. 4a,c). In the meridional direction, the eddy momentum flux diver-
 231 gence tends to smooth out the Gulf Stream (accelerate the Gulf Stream on the north-
 232 ern flank and decelerate it on the southern flank; Fig. 4f), while the divergence of inter-
 233 facial form stress (i.e. baroclinic instability) counteracts to sharpen it (Fig. 4g). The di-
 234 vergence of the eddy momentum fluxes and interfacial form stress largely cancel each other
 235 out (Fig. 4a,c,f,g), however, with the residual generally having dipole features in the zonal
 236 direction (Fig. 4d), and fluxing momentum out of the mean flow in the meridional di-
 237 rection (Fig. 4h). This net zonal dipole feature and meridional deceleration of the Gulf
 238 Stream were the case for any randomly chosen day in our five years of ensembles (not
 239 shown).

240 4.2 The Ertel potential vorticity flux

As was noted by Young (2012), the E-P flux divergence is directly related to the eddy Ertel PV flux and can be written as:

$$\bar{\mathbf{e}}_1 \cdot (\tilde{\nabla} \cdot \mathbf{E}) = -\overline{\sigma \mathbf{u}'' \Pi^\star} \cdot \mathbf{j}, \quad \bar{\mathbf{e}}_2 \cdot (\tilde{\nabla} \cdot \mathbf{E}) = \overline{\sigma \mathbf{u}'' \Pi^\star} \cdot \mathbf{i}, \quad (12)$$

241 where $\widehat{\mathbf{u}'' \Pi^\star} \stackrel{\text{def}}{=} \overline{\sigma^{-1} [\{\bar{\mathbf{e}}_2 \cdot (\tilde{\nabla} \cdot \mathbf{E})\} \bar{\mathbf{e}}_1 - \{\bar{\mathbf{e}}_1 \cdot (\tilde{\nabla} \cdot \mathbf{E})\} \bar{\mathbf{e}}_2]}$, $\Pi^\star \stackrel{\text{def}}{=} \Pi - \Pi^\#$ and $\Pi^\# \stackrel{\text{def}}{=} \overline{\sigma^{-1} (f +$
 242 $\hat{v}_x - \hat{u}_y)}$ are the eddy Ertel PV flux, eddy and residual-mean Ertel PV respectively. Note
 243 $\Pi^\#$, computed from the residual-mean velocities, is different from the thickness-weighted
 244 Ertel PV, viz. $\hat{\Pi} = \frac{\overline{\Pi \sigma}}{\overline{\sigma}} = \overline{\sigma^{-1} (f + \bar{v}_x - \bar{u}_y)}$. Equation (12) implies that if we are able
 245 to parametrize the eddy Ertel PV flux, equivalently we have parametrized the eddy feed-
 246 back onto the mean flow encapsulated in the E-P flux divergence.

It is well known that the governing equation for Ertel PV is similar to that of pas-
 sive tracers (S. K. Smith & Marshall, 2009; Vallis, 2017, Chapter 5), and that mesoscale
 eddies stir passive tracers along isopycnals (Redi, 1982; Gnanadesikan et al., 2015; Naveira Gara-
 bato et al., 2017; Griffies, 2018; Jones & Abernathey, 2019; Uchida et al., 2020). One sig-
 nificant difference between Ertel PV and passive tracers, however, is in its dynamical sig-
 nificance; the Ertel PV feeds back onto the dynamics in the form of eddy fluxes perhaps
 most well known in the transformed-Eulerian mean framework (e.g. Vallis, 2017, Chap-
 ter 10). This has led to the idea that the dynamical effect of mesoscale turbulence may
 be parametrized as a local gradient flux of the mean Ertel PV (Killworth, 1997; Great-

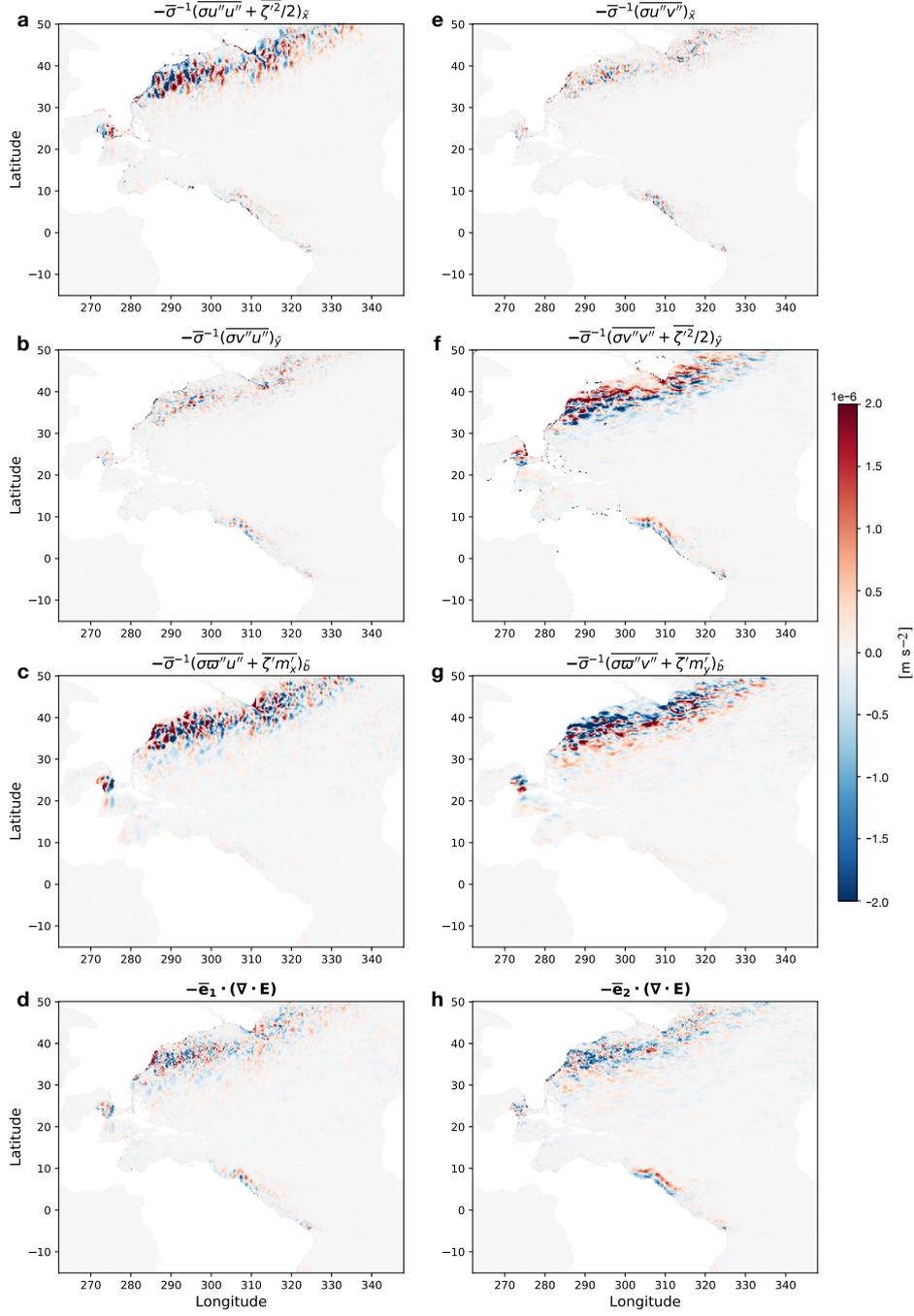


Figure 4. The terms in the divergence of E-P flux tensor on December 26, 1963 on the isopycnal layer as in Fig. 2. The red shadings indicate momentum being fluxed from the eddies to the mean flow and visa versa **a-c,e-g**. The panels are laid out so that summing up the top three rows per column yields the total zonal ($\bar{\mathbf{e}}_1 \cdot (\nabla \cdot \mathbf{E})$) **d** and meridional E-P flux divergence ($\bar{\mathbf{e}}_2 \cdot (\nabla \cdot \mathbf{E})$) **h** respectively.

batch, 1998; D. P. Marshall et al., 1999, 2012), i.e.

$$\widehat{\mathbf{u}''\Pi^\star} = -\kappa\tilde{\nabla}_h\Pi^\# . \quad (13)$$

247 where κ is the eddy diffusivity. Equations (1), (2), (12) and (13) provide a pathway for
 248 a unique solution for the eddy closure problem as the E-P flux divergence is gauge in-
 249 variant (J. R. Maddison & Marshall, 2013).

While it is tempting to directly infer a scalar eddy diffusivity from eqn. (13), as-
 suming an isotropic diffusivity for an anisotropic flow as in our realistic simulation is a
 poor approximation (R. D. Smith & Gent, 2004; Ferrari & Nikurashin, 2010; Fox-Kemper
 et al., 2013). We, therefore, take the approach of estimating the eddy diffusivity tensor
 (\mathbf{K}) from a least-squares best fit to (Plumb & Mahlman, 1987; Abernathy et al., 2013;
 Bachman & Fox-Kemper, 2013):

$$\underbrace{\begin{pmatrix} \widehat{u''\theta''} & \widehat{v''\theta''} \\ \widehat{u''s''} & \widehat{v''s''} \\ \widehat{u''\Pi^\star} & \widehat{v''\Pi^\star} \end{pmatrix}}_{\mathbf{F}} = - \underbrace{\begin{pmatrix} \hat{\theta}_{\bar{x}} & \hat{\theta}_{\bar{y}} \\ \hat{s}_{\bar{x}} & \hat{s}_{\bar{y}} \\ \Pi_{\bar{x}}^\# & \Pi_{\bar{y}}^\# \end{pmatrix}}_{\mathbf{G}} \cdot \underbrace{\begin{pmatrix} \kappa^{uu} & \kappa^{vu} \\ \kappa^{uv} & \kappa^{vv} \end{pmatrix}}_{\mathbf{K}} . \quad (14)$$

250 In studies trying to parametrize the eddy-induced fluxes of isopycnal thickness in the buoy-
 251 ancy equation, they have always had the freedom to parametrize the flux itself or its di-
 252 vergence. This has caused some ambiguity regarding the whether the rotational com-
 253 ponent of eddy fluxes, often referred to as the gauge freedom, should be parametrized
 254 (discussed in depth by Griffies, 2018). However, since the TWA equations are forced di-
 255 rectly by the eddy Ertel PV flux itself and not its divergence, we do not need to consider
 256 the discussion centred around rotational fluxes. In other words, eqn. (12) makes the case
 257 for parametrizing the *total* eddy flux, as opposed to solely its divergent component, when
 258 formulating a closure scheme for Ertel PV. The assumption that goes into eqn. (14) is
 259 that the eddy flux of temperature, salinity and Ertel PV behave statistically in a simi-
 260 lar manner (Bachman et al., 2015). Since they are all active tracers, we would expect
 261 this assumption to hold to a good degree.

262 The least-squares fit can be estimated as $\mathbf{K} = \mathbf{G}^+\mathbf{F}$ where \mathbf{G}^+ is the Moore-Penrose
 263 pseudo inverse of \mathbf{G} for each data point (Bachman et al., 2015). The gradients of the mean
 264 field, however, tended to be noisy due to errors accumulating from the remapping pro-
 265 cess (eqn. (4)). Therefore, we applied a convolutional spatial smoothing to the mean fields
 266 ($\hat{\theta}, \hat{s}, \Pi^\#$) prior to taking their gradient and eddy terms (viz. each element in \mathbf{F}) with

267 a 5×5 Hann filter in the horizontal grid points using the `xscale` Python package (Sérazin,
 268 2019). The spatial smoothing can be considered similar to a numerical convergence of
 269 the fields with an increase in the number of ensemble members. Each row in \mathbf{F} and \mathbf{G}
 270 was then normalized by horizontal median of the magnitude of each eddy fluxes (i.e. $\frac{\langle \mathbf{u}'' C'', \mathbf{F}_C^{\text{param}} \rangle}{\text{median}[\|\mathbf{u}'' C''\|]}$)
 271 where $\mathbf{F}_C^{\text{param}}$ ($= \mathbf{G}_C \cdot \mathbf{K}_C$) is the parametrized flux of an arbitrary tracer C) prior to
 272 the inversion so that each tracer had roughly equal weighting in inverting eqn. (14).

273 From Fig. 4, it is evident that the equatorial region contributes little to the Gulf
 274 Stream, so we will focus on north of 20N in this section. Figure 5a,d shows the diagnosed
 275 non-smoothed eddy Ertel PV flux, which we refer to as the “true” flux ($\mathbf{F}_{\Pi}^{\text{true}}$), and its
 276 parametrized equivalent via eqn. (14) as a local-gradient flux of the mean Ertel PV (Fig. 5b,e).
 277 We see that the local-gradient flux closure successfully captures the spatial features of
 278 the true flux with the residual between the two being small (Fig. 5c,f). The residual comes
 279 from the smoothing we have applied prior to inverting eqn. (14) and/or errors in the remap-
 280 ping and discretization, but it is likely that this residual would decrease with an increase
 281 in the number of ensemble members. One may argue that since we are fitting the eddy
 282 diffusivities, the agreement is to be expected. It is nevertheless encouraging to see how
 283 well the eddy Ertel PV fluxes can be represented via an anisotropic eddy diffusivity ten-
 284 sor (Fig. 6) compared to previous studies reconstructing the eddy tracer fluxes with a
 285 scalar diffusivity (e.g. Wilson & Williams, 2006; J. Maddison et al., 2015). This also pro-
 286 vides confidence to the assumption behind eqn. (14) that the Ertel PV behaves similarly
 287 to active tracers along isopycnals. In other words, along with the TWA framework, we
 288 have chosen the appropriate regression model to relate the total eddy transport of ac-
 289 tive tracers to their mean fields. Although it is possible to invert eqn. (14) with just two
 290 tracers, the inversion becomes ill defined unless their distributions are orthogonal to each
 291 other (Bachman et al., 2015). We have, therefore, kept it over-determined using three
 292 tracers.

The diffusivities presented in Fig. 6 are roughly on the same order as previous es-
 timates based on satellite products (J. Marshall et al., 2006; Abernathey & Marshall,
 2013; Klocker & Abernathey, 2014; Busecke et al., 2017; Bolton et al., 2019) and mod-
 elling studies (Wilson & Williams, 2006; Abernathey et al., 2013; Bachman & Fox-Kemper,
 2013), which range spatially between $O(10^2-10^4)$ [$\text{m}^2 \text{s}^{-1}$]. The negative values, how-
 ever, may come as a surprise. One of the key differences from the satellite-based esti-
 mates is that we do not assume an isotropic down-gradient flux closure with a scalar dif-

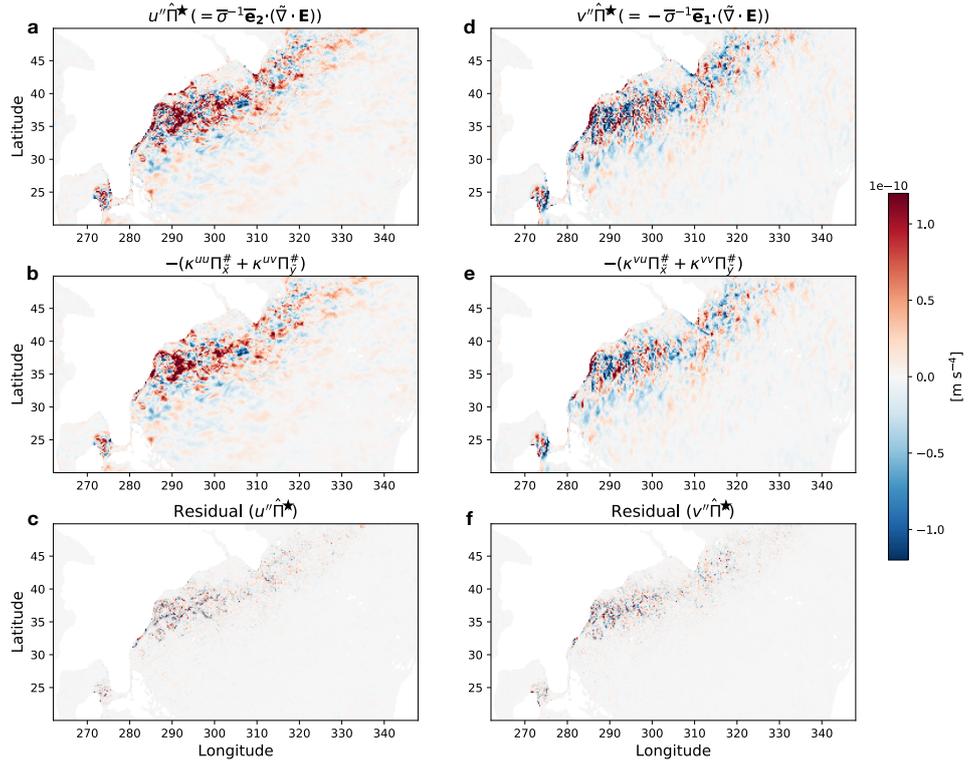


Figure 5. The diagnosed zonal and meridional eddy PV flux on December 26, 1963 on the isopycnal layer as in Fig. 2. We see a strong signal in the Gulf Stream region **a,d**. **b,e** The parametrized eddy PV flux via eqn. (14). **c,f** The residual between the true and parametrized eddy PV flux.

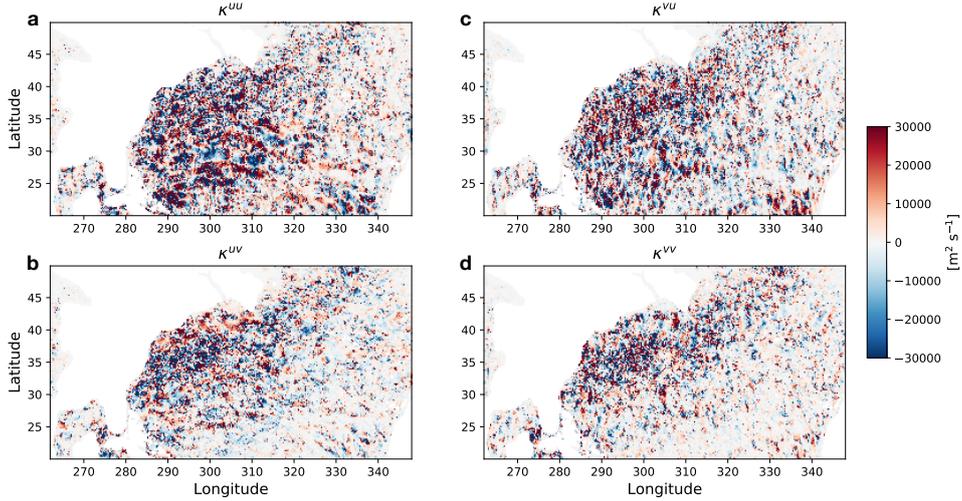


Figure 6. The diagnosed eddy diffusivity parameters via eqn. (14) in the diffusivity tensor \mathbf{K} on December 26, 1963 on the isopycnal layer as in Fig. 2.

fusivity. In other words, the negative "κ"s do not necessarily translate to up-gradient tracer fluxes as, based on eqn. (14), the closure is a linear combination of the zonal and meridional gradients; the fluxes could be down gradient in the two-dimensional sense. On the other hand, in cases where the eddy fluxes are locally oriented up gradient of the mean tracer field, negative "κ"s would be a faithful representation of this. We show the inner angle between the eddy flux and gradient of the mean field:

$$\varphi_C = \arccos \left[\frac{\mathbf{F}_C^{\text{true}} \cdot \mathbf{G}_C}{|\mathbf{F}_C^{\text{true}}| |\mathbf{G}_C|} \right] \quad (15)$$

293 in Fig. 7 for each tracer; a down-gradient eddy flux would result in $\varphi \sim 0$. There are
 294 regions of both down-gradient and up-gradient eddy fluxes for all three tracers (Fig. 7).
 295 Although the eddy fluxes should be down gradient of the mean field in the global sense
 296 in order to allow for the homogenization of tracers (D. P. Marshall et al., 2012; J. R. Mad-
 297 dison & Marshall, 2013), a locally up-gradient eddy flux is associated with an up-gradient
 298 transfer of tracer variance. It should not be surprising that in a realistic simulation, in-
 299 stantaneous fields of tracer variance can be spatially inhomogenous with sources, sinks
 300 and transport of variance (Wilson & Williams, 2006). In the context of energy-backscattering
 301 eddy parametrizations, when the tracer is Ertel PV, an up-gradient eddy flux is equiv-
 302 alent to the eddies fluxing momentum back into the mean flow, which is precisely the
 303 effect we would want to represent.

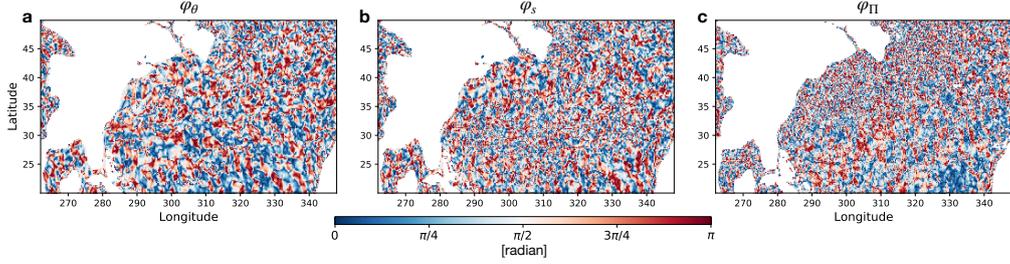


Figure 7. The inner angle between the eddy flux and horizontal gradient of the mean on December 26, 1963 for potential temperature (φ_θ) **a**, practical salinity (φ_s) **b** and Ertel PV (φ_Π) **c** on the isopycnal layer as in Fig. 2. The angles are close to zero when the eddy flux is oriented down gradient of the mean Ertel PV and close to π when oriented up gradient.

304 It is also informative to examine the diffusive component of the diffusivity tensor
 305 in regards to isopycnic tracer mixing, i.e. the eigenvalues of the symmetric part of the
 306 tensor ($\mathbf{S} \stackrel{\text{def}}{=} \frac{1}{2}(\mathbf{K} + \mathbf{K}^T)$ where \mathbf{K}^T is the transpose). The spatial median of the eigen-
 307 values along the major-axis (λ^M) and minor-axis (λ^m) of eigenvectors on December 26,
 308 1963 (Fig. 6) are 1286 (12119) $\text{m}^2 \text{s}^{-1}$ and 56 (2247) $\text{m}^2 \text{s}^{-1}$ respectively with a long tail
 309 in both positive and negative values. The values in curly brackets show the median of
 310 the normed diffusivities $|\lambda^M|$ and $|\lambda^m|$ respectively. The negative values likely come from
 311 the mean flow being strongly inhomogeneous. The spatial median of the anisotropy pa-
 312 rameter ($|\lambda^M|/|\lambda^m|$) is around 4.6. Although the order of magnitude of the eigenvalues
 313 is similar to previous modelling studies (e.g. Abernathey et al., 2013), it is difficult to
 314 make a direct comparison due to the differences in the averaging operator and model con-
 315 figuration.

We end this section by showing the spatial correlation and error between the true and parametrized eddy flux along the temporal and buoyancy dimensions:

$$r_C = \frac{\sum[(F_C^{\text{true}} - \langle F_C^{\text{true}} \rangle)(F_C^{\text{param}} - \langle F_C^{\text{param}} \rangle)]}{\sqrt{\sum(F_C^{\text{true}} - \langle F_C^{\text{true}} \rangle)^2} \sqrt{\sum(F_C^{\text{param}} - \langle F_C^{\text{param}} \rangle)^2}}, \quad (16)$$

$$\mathcal{E}_C \stackrel{\text{def}}{=} \frac{|F_C^{\text{true}} - F_C^{\text{param}}|}{|F_C^{\text{true}}|} \quad (17)$$

316 where $\langle \cdot \rangle$ is the horizontal domain average. Equations (16) and (17) were calculated us-
 317 ing every three grid points in the zonal and meridional dimension between 20N-50N, and
 318 every two grid points in the buoyancy dimension across the range roughly correspond-
 319 ing to depths between 300–2000 m. The correlation is generally higher than 0.9 for po-
 320 tential temperature and 0.5 for practical salinity across all seasons in the quasi-adiabatic

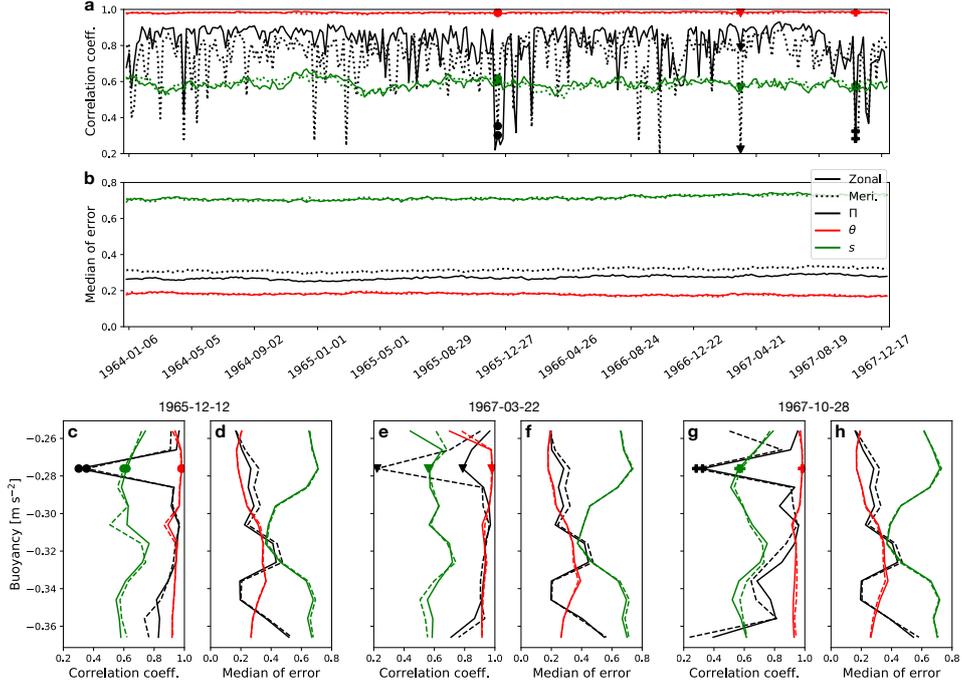


Figure 8. The correlation coefficient and spatial median of the error for potential temperature (red; $r_\theta, \mathcal{E}_\theta$), practical salinity (green; r_s, \mathcal{E}_s) and Ertel PV (black; r_Π, \mathcal{E}_Π). The zonal component is shown in solid lines and the meridional in dotted lines. **a,b** The correlation coefficients and error on the isopycnal as in Fig. 2. The circle, triangle and plus markers indicate the dates we show the vertical profiles. **c-h** The vertical profiles on December 12, 1965, March 22, 1967 and October 28, 1967.

interior for the latter four years of output we have (Fig. 8). The performance of Ertel PV is between the two ($r_\Pi \sim 0.8$) with large temporal fluctuations. The spatial correlation is sensitive to extrema due to its dependence on the spatial mean ($\langle F_\Pi \rangle$; eqn. (16)), likely responsible for the fluctuations as the spatial median of the error is stable over the entire time series ($\mathcal{E}_\Pi \sim 0.3$; Fig. 8b). The robustness of our parametrization can also be seen from the vertical structure of the error (Fig. 8d,f,h); it shows very little temporal variation regardless of the date.

5 Discussion and summary

By running a 12-member ensemble run of the North Atlantic Ocean at mesoscale resolving resolution ($1/12^\circ$), we have shown that the thickness-weighted average (TWA) framework can be employed successfully in diagnosing eddy-mean flow interactions in

332 a realistic ocean simulation. The ensemble approach negates the necessity for any tem-
 333 poral averaging in defining the residual-mean flow; we are able to exclude any tempo-
 334 ral variability, such as seasonal and interannual fluctuations, from the eddy term and ex-
 335 tract the intrinsic mesoscale variability of the ocean. We show that the Eliassen-Palm
 336 (E-P) flux divergence, which encapsulates the eddy feedback onto the mean flow (J. R. Mad-
 337 dison & Marshall, 2013), tends to meridionally decelerate the Gulf Stream (Fig. 4h). Mod-
 338 elling studies with varying spatial resolution have shown that the Gulf Stream tends to
 339 overshoot northwards in coarse resolution models (e.g. Lévy et al., 2010; Chassignet &
 340 Xu, 2017). The meridional deceleration implies that this overshooting may partially be
 341 attributable to mesoscale eddy feedback, in particular baroclinic instability (Fig. 4g), be-
 342 ing insufficiently resolved at such resolutions, in addition to submesoscale boundary layer
 343 processes (e.g. Renault et al., 2016; Schoonover et al., 2017). The overshooting is also
 344 apparent in our simulation (Fig. 2d) implying that even models with the resolution of
 345 $1/12^\circ$ could benefit from parametrizing the eddy momentum fluxes.

346 In the TWA framework, the eddy Ertel potential vorticity (Ertel PV) flux is di-
 347 rectly related to the E-P flux divergence (Young, 2012). In the context of eddy parametriza-
 348 tion, this implies that if we can parametrize the eddy Ertel PV flux, we have a solution
 349 for the mesoscale eddy closure problem, which we provide in Figs. 5 and 6. We would
 350 like to emphasize that the eddy diffusivities presented in this paper are diagnostic rather
 351 than prognostic variables. Future work would need to examine how each parameter in
 352 the eddy diffusivity tensor (\mathbf{K} ; eqn. (14)) is determined by the residual-mean field for
 353 a prognostic eddy closure scheme. Data-driven methods may be a viable way to discover
 354 such equations to constrain the “ κ ”s (e.g. Zhang & Lin, 2018; Zanna & Bolton, 2020).
 355 While it is beyond the scope of this study, it would also be interesting to examine the
 356 relation between the “ κ ”s and eddy shape, orientation and/or energy (e.g. D. P. Mar-
 357 shall et al., 2012; Waterman & Lilly, 2015; Bachman et al., 2017; Anstey & Zanna, 2017;
 358 Mak et al., 2018; Poulsen et al., 2019).

359 Nevertheless, we have shown that the eddy Ertel PV flux can be parametrized as
 360 an active tracer by a local-gradient flux closure across all seasons (Fig. 8). The appar-
 361 ent success of our diffusivity tensor lies on the fact that it relates the eddy fluxes to the
 362 residual-mean as opposed to the Eulerian-mean fields. As was noted by McDougall and
 363 McIntosh (2001) and Young (2012), the TWA framework allows one to shift the focus
 364 of eddy parametrization from the buoyancy equation to the momentum equations (1)

365 and (2). What follows is that the tensor \mathbf{K} includes information of not only the (eddy-
 366 induced) skew-diffusive flux of isopycnal thickness (Gent & McWilliams, 1990; Griffies,
 367 1998) and isopycnic tracer diffusivity (Redi, 1982), but also the eddy momentum fluxes,
 368 which energy backscattering eddy parametrizations are being developed to represent (e.g.
 369 Kitsios et al., 2013; Bachman et al., 2018; Bachman, 2019; Zanna & Bolton, 2020; Jansen
 370 et al., 2019; Perezhogin, 2019; Juricke et al., 2020). Although there are four parameters
 371 in the tensor \mathbf{K} , this is no more than assuming, for example, spatial variability and anisotropy
 372 in the isopycnal skew diffusivity (Gent & McWilliams, 1990; Griffies, 1998) and isopy-
 373 cnic tracer diffusivity (Redi, 1982). We believe our results provide a robust framework
 374 to evaluate such newly developed parametrizations in primitive equation models, i.e. they
 375 should be representing the E-P flux divergence, and a first step towards a unified mesoscale
 376 eddy closure scheme.

377 **Appendix A Neutral surfaces as the coordinate system**

Suppose we use neutrally-surfaced density (Stanley, 2019) to define the coordinate system, i.e. $\tilde{b} = -g\delta\rho/\rho_0$, which we will refer to as neutral buoyancy, where $\delta\rho$ is the neutrally-surfaced density anomaly. The Montgomery potential then becomes $m(\tilde{b}) = \phi(\tilde{b}) - b(\tilde{b})\zeta(\tilde{b})$ where b is the in-situ buoyancy defined by in-situ density and satisfies $\phi_{\tilde{\zeta}} = b$. Hence, the hydrostatic balance (eqn. (109) in Young, 2012) becomes:

$$\begin{aligned} m_{\tilde{b}} &= (\phi - b\zeta)_{\tilde{b}} \\ &= \phi_{\tilde{b}} - b_{\tilde{b}}\zeta - b\zeta_{\tilde{b}} \\ &= -b_{\tilde{b}}\zeta. \end{aligned}$$

Although it is possible to continue on by carrying around the Jacobian term keeping in mind that $b_{\tilde{b}} \neq 1$, the simplicity of the TWA framework is lost down the line due to the chain rule. An example being:

$$\tilde{\nabla}_{\mathbf{h}}m = \nabla_{\mathbf{h}}m - m_{\tilde{b}}\nabla_{\mathbf{h}}b \tag{A1}$$

$$= \nabla_{\mathbf{h}}m + b_{\tilde{b}}\zeta\nabla_{\mathbf{h}}b \tag{A2}$$

$$= \nabla_{\mathbf{h}}(\phi - b\zeta) + b_{\tilde{b}}\zeta\nabla_{\mathbf{h}}b \tag{A3}$$

378 where the right-hand side of (A3) is not equivalent to $\nabla_{\mathbf{h}}\phi$.

Appendix B Energetics under a non-linear equation of state

The TWA residual-mean horizontal momentum equation in geopotential coordinates neglecting dissipation is (Young, 2012):

$$\hat{\mathbf{u}}_t + \mathbf{v}^\# \cdot \nabla \hat{\mathbf{u}} + \mathbf{f} \times \hat{\mathbf{u}} = -\nabla_h \phi^\# - \nabla_h \cdot \mathbf{E}, \quad (\text{B1})$$

where $\mathbf{v}^\# \stackrel{\text{def}}{=} (\hat{\mathbf{u}}, w^\#)$ so the residual-mean kinetic energy ($K^\# = |\hat{\mathbf{u}}|^2/2$) budget becomes:

$$K_t^\# + \mathbf{v}^\# \cdot \nabla K^\# = -\hat{\mathbf{u}} \cdot \nabla_h \phi^\# - \hat{\mathbf{u}} \cdot (\nabla_h \cdot \mathbf{E}) \quad (\text{B2})$$

$$= -\hat{\mathbf{u}} \cdot \nabla_h \phi^\# - w^\# \phi_z^\# + w^\# b^\# - \hat{\mathbf{u}} \cdot (\nabla_h \cdot \mathbf{E}) \quad (\text{B3})$$

$$= -\mathbf{v}^\# \cdot \nabla \phi^\# + w^\# b^\# - \hat{\mathbf{u}} \cdot (\nabla_h \cdot \mathbf{E}). \quad (\text{B4})$$

We can now define the mean potential enthalpy as (McDougall, 2003):

$$h^\# \stackrel{\text{def}}{=} \int_{\Phi_0}^{\Phi} \frac{b^\#}{\rho_0 g} d\Phi^\# \quad (\text{B5})$$

where $\Phi^\# = \Phi_0 - \rho_0 g z$ is the dynamically non-active part of the hydrostatic pressure to be consistent with the Boussinesq approximation. It is important to keep in mind that the “ z ” here is the average depth of an isopycnal surface (i.e. $z = \bar{\zeta}(\tilde{t}, \tilde{x}, \tilde{y}, b^\#)$). The mean advective derivative of $h^\#$ is:

$$\mathbf{v}^\# \cdot \nabla h^\# = h_{\Phi^\#}^\# \mathbf{v}^\# \cdot \nabla \Phi^\# + h_{\hat{\theta}}^\# \mathbf{v}^\# \cdot \nabla \hat{\theta} + h_{\hat{s}}^\# \mathbf{v}^\# \cdot \nabla \hat{s} \quad (\text{B6})$$

$$= -w^\# b^\# + h_{\hat{\theta}}^\# \mathbf{v}^\# \cdot \nabla \hat{\theta} + h_{\hat{s}}^\# \mathbf{v}^\# \cdot \nabla \hat{s} \quad (\text{B7})$$

Therefore,

$$K_t^\# + \mathbf{v}^\# \cdot \nabla (K^\# + h^\#) = -\mathbf{v}^\# \cdot \nabla \phi^\# + J_h^\# - \hat{\mathbf{u}} \cdot (\nabla_h \cdot \mathbf{E}) \quad (\text{B8})$$

where $J_h^\# \stackrel{\text{def}}{=} h_{\hat{\theta}}^\# \mathbf{v}^\# \cdot \nabla \hat{\theta} + h_{\hat{s}}^\# \mathbf{v}^\# \cdot \nabla \hat{s}$.

On the other hand, the TWA budget of total kinetic energy is:

$$\frac{1}{2} (\widehat{|\mathbf{u}|^2}_t + \tilde{\nabla} \cdot \widehat{\mathbf{v}|\mathbf{u}|^2}) = -\tilde{\nabla} \cdot \widehat{\mathbf{v}\phi} + \tilde{w}\tilde{b} \quad (\text{B9})$$

$$= -\tilde{\nabla} \cdot \widehat{\mathbf{v}\phi} - \tilde{\nabla} \cdot \widehat{\mathbf{v}h} + h_\theta \widehat{\nabla \cdot (\mathbf{v}\theta)} + h_s \widehat{\nabla \cdot (\mathbf{v}s)}, \quad (\text{B10})$$

where $\tilde{\nabla} \cdot$ is the three-dimensional divergence in buoyancy coordinates, and using the relation $\overline{\sigma\phi\theta} = \overline{\sigma(\hat{\phi}\hat{\theta} + \phi''\theta'')}$ (eqn. (72) in Young, 2012) yields:

$$\frac{1}{2} [\widehat{|\mathbf{u}|^2}_t + \tilde{\nabla} \cdot (\widehat{\mathbf{v}|\mathbf{u}|^2} + \widehat{\mathbf{v}''|\mathbf{u}|^2})] = -\tilde{\nabla} \cdot \widehat{\mathbf{v}\phi} - \tilde{\nabla} \cdot (\widehat{\mathbf{v}h} + \widehat{\mathbf{v}''h''}) + \hat{J}_h, \quad (\text{B11})$$

where $\hat{\mathbf{v}} \stackrel{\text{def}}{=} \hat{u}\bar{\mathbf{e}}_1 + \hat{v}\bar{\mathbf{e}}_2 + \bar{\sigma}^{-1}(\bar{\zeta}_{\hat{t}} + \hat{\omega}\bar{\zeta}_{\hat{b}})\bar{\mathbf{e}}_3$ and $\hat{J}_h \stackrel{\text{def}}{=} h_\theta\widehat{\nabla \cdot (\mathbf{v}\theta)} + h_s\widehat{\nabla \cdot (\mathbf{v}s)}$. Now, realizing that $\widehat{|\mathbf{u}|^2} = |\hat{\mathbf{u}}|^2 + \widehat{|\mathbf{u}''|^2} \stackrel{\text{def}}{=} 2(K^\# + K'')$, we get:

$$K_t^\# + K_t'' + \tilde{\nabla} \cdot [\hat{\mathbf{v}}(K^\# + K'' + \hat{h}) + \widehat{\mathbf{v}''|\mathbf{u}|^2} + \widehat{\mathbf{v}''h''}] = -\tilde{\nabla} \cdot \widehat{\mathbf{v}\phi} + \hat{J}_h, \quad (\text{B12})$$

which can be rewritten using the coordinate-invariant differential operator as:

$$K_t^\# + K_t'' + \nabla \cdot [\mathbf{v}^\#(K^\# + K'' + \hat{h})] + \nabla \cdot (\widehat{\mathbf{v}''|\mathbf{u}|^2} + \widehat{\mathbf{v}''h''}) = -\nabla \cdot \widehat{\mathbf{v}\phi} + \hat{J}_h. \quad (\text{B13})$$

Hence, subtracting eqn. (B8) from (B13) yields:

$$K_t'' + \nabla \cdot [\mathbf{v}^\#(K'' + \hat{h} - h^\#)] = -\nabla \cdot [\widehat{\mathbf{v}''|\mathbf{u}|^2} + \widehat{\mathbf{v}''h''} + \widehat{\mathbf{v}\phi} - \mathbf{v}^\#\phi^\#] + \hat{J}_h - J_h^\# + \hat{\mathbf{u}} \cdot (\nabla_h \cdot \mathbf{E}). \quad (\text{B14})$$

381 Equations (B8) and (B14) are the relations derived by Aoki (2014) but for a non-linear
 382 equation of state (EOS) where the residual-mean flow and eddies exchange energy via
 383 the E-P flux divergence.

For a linear EOS, the eddy potential energy (EPE; $\hat{h} - h^\#$) in eqn. (B14) can be rewritten as:

$$\hat{h} - h^\# = -b^\#(\hat{\zeta} - \bar{\zeta}) \quad (\text{B15})$$

$$= -b^\# \frac{\overline{\sigma' \zeta'}}{\bar{\sigma}} \quad (\text{B16})$$

384 by taking advantage of $\hat{h} = -\tilde{b}\hat{\zeta}$, $h^\# = -b^\#\bar{\zeta}$ and $\tilde{b} = b^\#(t, x, y, \bar{\zeta})$. Equation (B15)
 385 provides the physical intuition of EPE being defined as the difference between potential
 386 energy defined at the thickness-weighted and ensemble-averaged isopycnal depths.

387 Appendix C Kinematics of discretization

As in Fig. C1, imagine u_1 and u_2 are on the same buoyancy contour. The relation between the two is:

$$u_2 = u_1 + u_x \Delta x + u_\zeta \Delta \zeta. \quad (\text{C1})$$

$$\therefore u_{\bar{x}} \stackrel{\text{def}}{=} u_x + u_\zeta \frac{\Delta \zeta}{\Delta x} = \frac{u_2 - u_1}{\Delta x} \quad (\text{C2})$$

so once all of the variables are remapped onto the buoyancy coordinate from geopotential, the discretized horizontal gradients can be taken along the original Cartesian grid. The gradients on the model outputs were taken using the `xgcm` Python package (Abernathey & Busecke, 2019; Busecke & Abernathey, 2020). In order to minimize the discretization error, we take the ensemble mean first whenever possible, e.g. $\bar{\sigma} = \overline{\partial_b \bar{\zeta}} = \partial_b \bar{\zeta}$, $\tilde{\nabla}_h \bar{\sigma} =$

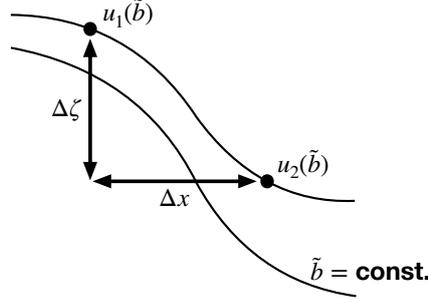


Figure C1. Schematic of discretized gradients.

$\partial_{\tilde{b}} \tilde{\nabla}_h \bar{\zeta}$ etc. The gradient operators commuting with the ensemble mean is also the case for the perturbations, i.e.

$$\tilde{\nabla}_h(\bar{m} + m') = \tilde{\nabla}_h m = \overline{\tilde{\nabla}_h m} + (\tilde{\nabla}_h m)'. \quad (\text{C3})$$

388 Hence, $\tilde{\nabla}_h m' = (\tilde{\nabla}_h m)'$ (c.f. J. R. Maddison & Marshall, 2013, Section 2.3 in their pa-
 389 per).

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