

1 **The delayed response of the troposphere-stratosphere-mesosphere**
2 **coupling to the 2019 southern SSW**

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16

17

Abstract

18 A strong Southern Hemisphere (SH) sudden stratospheric warming (SSW) event
19 occurred in September 2019 and significantly weakened the stratospheric polar vortex.
20 Due to the positive zonal wind anomalies in the troposphere, the barotropic/baroclinic
21 instability, primarily controlled by the horizontal/vertical wind shear, weakened in the
22 upper troposphere at midlatitudes in late September and early October. As a result,
23 planetary waves (PWs) were deflected equatorward near the tropopause rather than
24 upward into the stratosphere, resulting in less perturbation to the stratospheric polar
25 vortex. After October 15, the westward zonal wind anomalies propagate downward and
26 reach the troposphere, increasing the tropospheric barotropic/baroclinic instability. This
27 benefits the propagation of PWs into the stratosphere, leading to the early breaking of
28 the stratospheric polar vortex. In turn, the SH mesosphere becomes anomalously cold
29 due to the stratospheric wind filtering on the gravity waves (GWs), leading to the much
30 earlier onset of SH polar mesospheric clouds (PMCs).

31

Plain Language Summary

33 A rare sudden stratospheric warming event, characterized by the dramatic increase
34 in temperature and the weakening of the stratospheric circumpolar flow, occurred in
35 September 2019. The anomalous wind induced by the SSW event tends to propagate
36 downward in the following months. The induced anomalous wind shear can modulate
37 the atmospheric barotropic/baroclinic instability, guiding the propagation of the waves.

38 Along with the downward propagation of the SSW-induced perturbation, the
39 atmospheric instability increases and benefits the atmospheric waves propagating into
40 the stratosphere from late October to November. The waves propagate into the
41 stratosphere, interact with the mean flow, and contribute to the reversal of the
42 stratospheric zonal wind. The break of the stratospheric polar vortex can also affect the
43 mesosphere by filtering the small-scale gravity waves, resulting in the perturbation of
44 the temperature, water vapor distribution and the formation of clouds in the mesosphere.
45

46 **Key Points**

47 1 A rare Southern Hemisphere SSW event occurred in September 2019 and
48 contributed to the early onset of PMCs in November.

49 2 The downward propagation of the zonal wind anomaly affects the propagation
50 of PWs by modulating barotropic/baroclinic instability.

51 3 The secondary enhanced upward propagation of the PWs causes a delayed
52 response in both the polar stratosphere and mesosphere.

53

54 **1 Introduction**

55 Sudden stratospheric warming (SSW), one of the most dramatic stratospheric
56 events, is identified as minor warming when the stratospheric meridional temperature
57 gradient reverses or major warming when the stratospheric circumpolar westerly jet
58 completely reverses (Andrews et al., 1987; Butler et al., 2015). While major SSWs
59 occurred approximately six times per decade in the Northern Hemisphere (NH), there
60 was only one major SSW in 2002, and one minor but intense SSW in 2019 was
61 recorded thus far in the SH (Baldwin et al., 2003) due to relatively weak planetary
62 wave activity in the Southern Hemisphere (SH). Although classified as minor,
63 according to the standard World Meteorological Organization (WMO) definition
64 (Butler et al., 2015), the SSW that occurred in September 2019 in the SH was
65 associated with the strongest polar-cap warming and the second strongest circumpolar
66 westerly jet deceleration from 1979 to the present (Yamazaki et al., 2020; Shen et
67 al., 2020a and 2020b).

68 SSWs in the SH have been shown to significantly impact both the troposphere
69 and stratosphere despite their rarity (Thompson & Solomon, 2002; Thompson et al.,
70 2005). Via the downward control principle and wave-flow interaction, the influence
71 of SSW in the polar troposphere and stratosphere can persist for months (Baldwin &
72 Dunkerton, 2001; Plumb and Semeniuk, 2003; Jucker & Goyal, 2022). In particular,
73 stratospheric polar vortex variations and their downward coupling to the troposphere
74 are regarded as critical drivers of the SH surface temperature, southern annular mode

75 (SAM), and southern stratospheric polar vortex (SSPV) in austral spring and summer
76 (Thompson and Wallace, 2000; Thompson et al., 2005).

77 Previous studies have focused on the role of SSPV in driving climate variability
78 at the Antarctic surface (Thompson and Wallace, 2000, Kwok and Comiso, 2002,
79 Thompson and Solomon, 2002, Thomson et al., 2005). The variation in the lower
80 atmosphere could potentially affect the propagation and excitation of planetary waves
81 (PWs), which play an essential role in modulating vertical coupling from the
82 stratosphere to the mesosphere (Garcia-Herrera et al., 2006; Black & McDaniel, 2007;
83 Li et al., 2013; Yang et al., 2017). Stratospheric PW activity is influenced by, for
84 example, atmospheric temperature and wind patterns, which modulate the propagation
85 and refraction of PWs (Matsuno, 1970; Baldwin et al., 2021). The atmospheric
86 condition for the upward PW propagation is theoretically related to the variability of
87 the potential vorticity perturbations, which is controlled by the zonal wind and the
88 barotropic/baroclinic instability (Charney and Drazin, 1961; Matsuno, 1970; Hartman,
89 1983; Meyer & Forbes, 1997). Hence, persistent atmospheric perturbation induced by
90 SSW could potentially influence the propagation and refraction of upward PWs from
91 the lower troposphere to the stratosphere. The convergence or divergence of planetary
92 waves then affects the temperatures and wind patterns. This two-way feedback
93 between the waves and wind patterns is called the wave-mean-flow interaction
94 (Andrews et al., 1987).

95 Due to the filtering of gravity waves (GWs) by the stratospheric zonal wind

96 (Lindzen, 1981; McLandress, 1998), the variation in the stratospheric temperature
97 gradient and the winds could effectively modulate the mesospheric circulation and
98 thus temperature (e.g., Shepherd, 2000; Karlsson et al., 2011; Li et al., 2016). Polar
99 mesospheric clouds (PMCs), also known as noctilucent clouds (NLCs), are the
100 highest clouds on Earth that form in the polar summer mesopause region and are
101 considered to be important indicators of variations in temperature and circulation in
102 the mesosphere (Thomas et al., 1996; Hervig et al., 2009 and 2015). The earlier onset
103 of SH PMC (Solodovnik et al., 2021) also suggested irregular variation in the SH
104 middle atmosphere in the austral spring of 2019. The potential persistent influence of
105 SSW on vertical middle atmospheric coupling, however, has not been well
106 established.

107 The 2019 SH SSW provides an excellent opportunity to understand the coupling
108 process of different layers of the atmosphere in the seasonal evolution process. This
109 study explores the possible dynamic mechanism of the delayed impacts of the 2019
110 SH SSW on the vertical SH troposphere-stratosphere-mesosphere coupling from
111 September through November.

112

113 **2 Data and Method**

114 The Microwave Limb Sounder (MLS) onboard the Aura satellite, launched in July
115 2004, measures the middle atmosphere temperature and water vapor profiles between

116 261 and 0.001 hPa (~92 km) from 118- and 240-GHz radiances of O₂ spectra (Schwartz
 117 et al. 2008; Waters et al., 2006; Livesey et al., 2017). The latitudinal coverage of the
 118 Aura/MLS measurements is ~82°S-82°N. In this study, we calculate the daily zonal
 119 mean temperature and water vapor mixing ratio from the MLS version 4.2 dataset
 120 between August 2004 and December 2021 (available at
 121 https://disc.gsfc.nasa.gov/datasets/ML2T_005/summary).

122 Modern-Era Retrospective analysis for Research and Applications version 2
 123 (MERRA-2) (Gelaro et al., 2017) temperature and water vapor (obtained from the
 124 specific humidity) data are utilized to perform diagnostic analysis and illustrate the
 125 variations in the background atmosphere. The vertical coverage of the MERRA-2
 126 reanalysis data is from the surface to 0.01 hPa (~80 km). The Eliassen-Palm (EP) flux
 127 and its divergence were calculated according to the transformed Eulerian-mean (TEM)
 128 equations (Andrews et al., 1987; Eliassen & Palm, 1960):

$$129 \quad f_{\phi} = \rho_0 a \cos \phi (\bar{u}_z \overline{v'\theta'} / \bar{\theta}_z - \overline{v'u'}); \quad (1)$$

$$130 \quad f_p = \rho_0 a \cos \phi \{ [f - (a \cos \phi)^{-1} (\bar{u} \cos \phi)_{\phi}] \overline{v'\theta'} / \theta_z - \overline{w'u'} \}; \quad (2)$$

$$131 \quad Div \equiv (a \cos \phi)^{-1} \frac{\partial}{\partial \phi} (f_{\phi} \cos \phi) + \frac{\partial f_p}{\partial z}; \quad (3)$$

132 where u , v , w , and θ are the zonal, meridional and vertical wind, potential
 133 temperature, ρ_0 , a , ϕ , f represents the air density, Earth's radius, latitude, and Coriolis
 134 parameter, respectively; the subscripts ϕ and z denote the latitudinal gradient and the
 135 vertical gradient, respectively; the overbar indicates the zonal mean value, while prime
 136 indicates the zonal anomalies.

137 The residual mean meridional circulation was employed to characterize the
 138 mesospheric variation response to wave activities:

$$139 \quad \bar{v}^* \equiv \bar{v} - \rho^{-1} (\rho \overline{v'\theta'} / \bar{\theta}_z)_z; \quad (4)$$

$$140 \quad \bar{w}^* \equiv \bar{w} + (a \cos \phi)^{-1} (\cos \phi \overline{v'\theta'} / \bar{\theta}_z)_\phi; \quad (5)$$

141 The meridional gradient of the quasi-geostrophic potential vorticity (\bar{q}_ϕ) is used
 142 to indicate the atmospheric baroclinic/barotropic instability (Meyer & Forbes, 1997)
 143 and is expressed as:

$$144 \quad \bar{q}_\phi = 2 \Omega \cos \phi - \left(\frac{\bar{u} \cos \phi}{a \cos \phi} \right)_\phi - \frac{a}{\rho} \left(\frac{f^2}{N^2} \rho \bar{u}_z \right)_z; \quad (6)$$

145 where Ω is the angular velocity of the Earth's rotation and N^2 is the buoyant frequency
 146 ($N^2 = g^* \text{dln}\theta/\text{dz}$), which represents the static stability.

147 To offer guidance on the direction of wave propagation within the troposphere and
 148 stratosphere (Charney and Drazin, 1961), the index of refraction was calculated in the
 149 form given by Matsuno (1970):

$$150 \quad \text{RI} = \frac{\bar{q}_\phi}{a\bar{u}} - \frac{s^2}{a^2 \cos^2 \phi} - \frac{f^2}{4N^2 H^2}; \quad (7)$$

151 where s is the zonal wavenumber and $H = 7000$ m is the height scale.

152 According to the downward control principle, the latitudinal and vertical
 153 circulation patterns are approximately proportional to the gradients of the vertically
 154 integrated wave forces above that level (Haynes et al., 1991). Circulation is thus utilized
 155 to distinguish the contributions of GWs and PWs to the residual circulation anomaly.
 156 The meridional and vertical residual circulation patterns induced by PW and GW forces
 157 are proportional to the vertical and horizontal gradients of the corresponding stream

158 functions (Ψ_{pw}), (Ψ_{gw}) and can be calculated as follows (Haynes et al., 1991):

$$159 \quad v^*_{(pw,gw)} = -\frac{1}{\rho \cdot \cos\varphi} \frac{\partial \Psi_{(pw,gw)}}{\partial z}, \quad (8)$$

$$160 \quad w^*_{(pw,gw)} = -\frac{1}{a \cdot \rho \cdot \cos\varphi} \frac{\partial \Psi_{(pw,gw)}}{\partial \varphi}, \quad (9)$$

161 where g is the acceleration caused by gravity. Considering that the GW parameters
162 are difficult to present in the MERRA-2 reanalysis dataset, the GW-induced stream
163 function (Ψ_{gw}) can be calculated by the difference between the total (Ψ_{total}) and PW-
164 induced (Ψ_{pw}) stream functions (Karpechko and Manzini, 2012; Lubis et al., 2016),
165 which can be calculated by

$$166 \quad \Psi_{pw} = \int_z^\infty \left\{ \frac{a^{-1} \nabla \cdot \mathbf{F}}{(a \cdot \cos\varphi)^{-1} (\bar{u} \cdot \cos\varphi)_\varphi - f} \right\} dz', \quad (10)$$

$$167 \quad \Psi_{total} = \int_z^\infty \rho \cos\varphi \cdot v^* dz', \quad (11)$$

168 The meridional component of the total residual circulation v^* was calculated by
169 equation (4), and \mathbf{F} is the Eliassen–Palm flux (Equations 1 and 2). The anomalous
170 temperature, zonal wind, occurrence percentage of the PMCs, and the parameters
171 utilized to diagnose the wave activities are calculated by comparison to the
172 climatological mean from 2004 to 2021. The Cloud Imaging and Particle Size
173 instrument (CIPS), onboard the Aeronomy of Ice in the Mesosphere (AIM) satellite,
174 has been measuring the sunlight scattered by mesospheric clouds at a wavelength of
175 265 nm since 2007 (Russell et al., 2009; Bailey et al., 2009; Rusch et al., 2009; Benze
176 et al., 2009). The instrument consists of four nadir-viewing cameras covering
177 approximately 2000×1000 km in the polar region, with a horizontal resolution of ~2
178 km (McClintock et al., 2009). CIPS data were used to obtain the PMC frequency of

179 occurrence.

180

181 **3 Results**

182 As one of the strongest stratospheric warming events in the SH, the SSW in
183 September 2019 led to dramatic warming (with a maximum of ~ 40 K) in the
184 Antarctic polar stratosphere, associated with significant cooling in the polar
185 mesosphere (with a minimum of ~ -30 K), as suggested by both the MLS observation
186 and MERRS-2 reanalysis dataset (Figures 1a and 1b). Although temperature
187 anomalies are strong, the eastward zonal mean winds significantly weakened from 80
188 m/s to 20 m/s but did not reverse direction in the September 2019 SSW event (Figure
189 2b). In the mesosphere, the temperature variation is primarily controlled by adiabatic
190 heating due to upwelling and downwelling. The upper mesospheric temperature
191 decreased significantly between late August and mid-September, corresponding to
192 anomalous upwelling (Figure 2a), which is related to SSW-induced stratospheric
193 perturbations (Figure 2b). During the following months (from mid-September to
194 December), anomalous stratospheric warming and mesospheric cooling propagate
195 downward, resulting in ~ 15 K warming in the lower stratosphere and ~ 5 -10 K
196 cooling in the middle and upper stratosphere. The temperature in the upper
197 mesosphere of SH returned to normal during October and became anomalously
198 negative again in November (with a minimum of ~ 8 K).

199 Figure 1c presents the climatological mean PMC occurrence percentage
200 observed by the AIM satellite and the 2019 occurrence percentage. The PMC
201 occurrence usually becomes obvious (occurrence percentage > 20%) at the beginning
202 of December (approximately 20 days before the solstice). The Southern Hemisphere
203 PMC occurrence was significantly earlier in 2019, and the probability of occurrence
204 exceeded 20% at the end of November, seven days earlier than the climatological
205 mean.

206 As the SH polar temperature variation in MERRA2 agrees well with the MLS
207 observations, in the remainder of this study, the possible mechanism by which the
208 2019 SH SSW could affect the variation in the stratosphere and mesosphere in two
209 months will be investigated based on the MERRA2 reanalysis data.

210 The PWs, which affect the temperature and wind variation in the stratosphere by
211 providing energy and momentum via their convergence, play an essential role in
212 modulating the vertical coupling from the stratosphere to the mesosphere (Garcia-
213 Herrera et al., [2006](#); Black & McDaniel, [2007](#); Li et al., [2013](#); Yang et al., [2017](#)). In
214 August, the zonal mean eddy heat flux averaged from 45°S to 75°S at 100 hPa
215 (proportional to the vertical component of EP flux) decreased dramatically and
216 persisted until the peak of SSW 2019, indicating that upward PW propagation was
217 strengthened (Figure 2c). The upward propagation of PWs at 100 hPa was weaker
218 than the climatological average after the SSW (from mid-September to mid-October).
219 Meanwhile, the 2019 stratospheric eastward circumpolar flow remained unchanged at

220 20 m/s, which is different from the weakening of westerly zonal winds in the other
221 years due to seasonal variation (Figure 2b). Anomalous upwelling in the mesosphere
222 becomes much weaker, while anomalous temperature returns to normal by mid-
223 October (Figure 2a). The upward propagation of PWs in the stratosphere was again
224 strengthened from mid-October to November compared to the 2004-2021 mean
225 (Figure 2c). This led to the rapid weakening of the stratospheric zonal wind and a
226 reversal from eastward to westward in the middle of November 2019. The reverse of
227 the stratospheric zonal wind in 2019 occurred approximately half a month earlier than
228 the climatology, indicating an earlier break of the 2019 SH stratospheric polar vortex.
229 Simultaneously, the temperature increase and upwelling were suppressed in the SH
230 polar mesosphere. In summary, the variations in mesospheric temperature, circulation,
231 stratospheric zonal wind, and PW activity are in good agreement with each other two
232 months after SSW 2019 in the SH.

233 An abnormal upward propagating stratospheric planetary wave, which is
234 suppressed in the first month after SSW but enhanced in the second month after SSW,
235 is closely related to perturbations in the stratosphere and mesosphere from October to
236 November. Nonetheless, as shown in Supporting Information Figure S1, the planetary
237 wave activity in the lower southern troposphere (500 hPa) is stronger than usual
238 during the austral spring of 2019 (September-November) without significant
239 perturbations as in the stratosphere after the SSW event. It is implied that the upward
240 propagation of the PW to the stratosphere is not attributed to variations in the wave

241 source in the lower atmosphere.

242 Figure 2d shows the anomalous meridional gradient of the potential vorticity
243 (\bar{q}_ϕ) averaged over 50-70°S and 500-200 hPa, which characterizes the tropospheric
244 baroclinic/barotropic instability (when $\bar{q}_\phi < 0$) in the SH middle latitudes. The
245 instability ($\bar{q}_\phi < 0$) of the background atmosphere could interact strongly with PWs
246 by producing an in situ source of energy for the waves, benefiting the upward
247 propagation and amplification of the PWs (Matsuno, 1970; Hartman, 1983; Meyer &
248 Forbes, 1997).

249 From August to early September 2019, the tropospheric instability in the SH mid-
250 latitudes was stronger than usual. The atmospheric instability became weaker than usual
251 from mid-September to early October, consistent with the PW variability before and
252 after the SSW event. Since late September, the SH tropospheric instability became
253 enhanced (negative \bar{q}_ϕ anomalies) compared to the climatology mean and remained
254 stronger than average if a short-lived weakening was neglected in early November.
255 After mid-November 2019, although the tropospheric still has higher instability, the
256 early break of the polar vortex and the reversal of the circumpolar circulation (Figure
257 2a) prevent the upward propagation of PWs, and the upward-propagating planetary
258 waves in the tropopause region become weaker than usual (Figure 2b).

259 Due to the wave-mean flow interaction (Baldwin et al., 2003) and “downward
260 control” principle (Haynes et al., 1991; Garcia & Boville, 1994), the significant
261 variations during SSW events tend to progress downward from the upper stratosphere

262 to the lowermost stratosphere in 1-2 months (Baldwin and Dunkerton, 2001;
263 Christiansen, 2005; Sigmond et al., 2013). After the occurrence of SSW in September
264 2019, the eastward zonal mean zonal winds were suppressed in the midlatitude upper
265 stratosphere (approximately 30-50 km). The negative zonal wind anomalies
266 associated with the warmer-than-normal zonal mean temperature (Figures 1a and 1b)
267 propagated downward. They gradually decreased in October and November,
268 accompanied by increased atmospheric stability in the same region (Figure 3a). Since
269 mid-October 2019, the negative zonal wind anomalies in the stratosphere have
270 descended through the tropopause region and penetrated the lower troposphere,
271 resulting in negative zonal wind anomalies.

272 The downward propagation of zonal wind anomalies can lead to perturbations in
273 both strong meridional and vertical wind shear, which effectively modulate the
274 variability of atmospheric instability (Figure 3b). According to equation 6, either
275 meridional wind shear or vertical wind shear could contribute to the variability of
276 atmospheric instability. As shown in Figure 3b, the increasing atmospheric instability
277 benefits from perturbation of the vertical and meridional wind shears (term two and
278 term 3 in equation 6) when the anomalous zonal wind penetrates the troposphere
279 around October 15. At the beginning of November 2019, the \bar{q}_ϕ anomalies become
280 positive, primarily due to the vertical zonal wind shear variation. This suppressed
281 instability corresponds well to the 100 hPa eddy heat flux variations (Figure 2b).

282 In the SH spring of 2019, the enhanced activity of PWs persists in the lower

283 troposphere in the latitude range of 30-70°S (Figure S1, Figure 3c and Figure 3d).
284 However, during the first month following the SSW (September 17 to October 15),
285 anomalous Eliassen-Palm (EP) flux from the lower troposphere tends to propagate
286 equatorward to the area with higher atmospheric instability (with negative \bar{q}_ϕ) rather
287 than traveling upward into the stratosphere across the midlatitude upper stratosphere
288 (Figure 3c). At the lower latitudes (40°S and equatorward), the atmospheric instability
289 increases in the upper troposphere, which is related to the positive phase of the
290 tropospheric SAM in the SH mentioned by Jucker et al. (2022). The anomalous \bar{q}_ϕ in
291 the upper troposphere increased near 60°S, indicating higher barotropic/baroclinic
292 stability of the atmosphere and inhibiting the upward and poleward propagation of the
293 PWs. Due to the suppressed upward propagation of PWs into the stratosphere,
294 westward momentum transport into the SH stratosphere decreased, and the anomalous
295 westward zonal wind became less evident.

296 As discussed above, from October 15 to November 15, as the negative zonal
297 wind anomalies penetrate the troposphere, the atmospheric instability increases in the
298 midlatitude troposphere, benefiting the amplification of the PWs. Thus, the enhanced
299 EP flux is transported from the lower troposphere toward midlatitude and vertically
300 into the troposphere (green vector in Figure 3d).

301 To diagnose the effect of atmospheric variation on PW propagation and
302 refraction, the index of refraction (RI), which is a good indicator of the PW
303 propagation direction in the stratosphere, is also investigated. PWs are preferentially

304 ducted toward regions with a more positive index of refraction and refracted away
305 from regions with a more negative RI (Andrews et al. 1987). The variation in RI is
306 affected by barotropic/baroclinic instability, zonal wind (2nd term in equation 7), and
307 static stability (3rd term in equation 7).

308 Since the occurrence of SSW, the refractive index in the stratosphere has
309 decreased at mid-latitudes and increased at high latitudes due to variations in
310 barotropic/baroclinic instability caused by anomalies in wind shear, zonal wind, and
311 static stability (Jucker et al., 2022). From October 15 to November 15, the enhanced
312 PWs propagating upward into the stratosphere from mid-latitudes tend to deflect
313 poleward and modulate the circumpolar flow in the high latitudes. This accelerates the
314 seasonal reversal of the eastward wind to the westward wind in SH and leads to the
315 complete break of the SSPV in mid-November (Figure S2).

316 The climatological zonal mean zonal wind at the SH high latitudes in November
317 is characterized by a weak eastward wind in the lower stratosphere and an increased
318 westward wind in the upper stratosphere and lower mesosphere. Due to the early
319 break of the SH polar vortex in November 2019, the filtering of eastward and upward-
320 propagating GWs by eastward zonal wind is replaced by the filtering of the westward
321 GWs by the westward zonal wind in the lower stratosphere. In the upper stratosphere,
322 more westward-propagating GWs are filtered by the strengthened westward zonal
323 wind. As a result of the net effect of zonal wind filtering, the eastward GW forcing is
324 thus enhanced in the SH mesosphere, strengthening the SH mesospheric residual

325 meridional circulation with anomalous SH polar mesosphere upwelling (Figure 4a).
326 This suggests that the SH polar mesopause temperature is controlled by the
327 stratospheric zonal wind in the SH high latitudes via the gravity wave filtering process
328 (Karlsson et al., 2011; Li et al., 2016; Yang et al., 2017).

329 According to the downward control principle, the meridional circulation patterns
330 are approximately proportional to the gradients of the vertically integrated wave force
331 above that level (Haynes, 1991). Thus, the contributions of GWs and PWs to the
332 residual circulation anomaly could be distinguished by the vertical and horizontal
333 gradients of the corresponding stream functions, as shown in equations 8-11. Due to the
334 early break of the stratospheric polar vortex, the upwelling of the meridional circulation
335 was enhanced in the SH mesopause primarily due to the eastward GWs in the second
336 half of November (see Figure 4b). The enhanced upwelling in the SH polar region led
337 to as low as -10 K temperature anomalies from 60 to 80 km, 60°-90°S (~ -10 K) through
338 adiabatic cooling. It increased the water vapor mixing ratio from 70 to 80 km (with an
339 increase of ~0.2 ppmv, as high as 10% of background H₂O) through dynamic transport.
340 The early onset of PMCs in the SH mesosphere in November 2019 thus benefited from
341 both the temperature and water vapor variation in the upper mesosphere.

342

343 **4. Summary and Discussion**

344 The emerging picture of the mechanisms can be summarized as follows: as the

345 SSW event occurred in September 2019, the stratospheric polar vortex significantly
346 weakened with the much weaker circumpolar eastward zonal wind, while the zonal
347 wind in the troposphere, however, was mainly eastward. The barotropic/baroclinic
348 instability, primarily controlled by the vertical and meridional wind shear, is weaker
349 (positive anomalous meridional gradient of the potential vorticity) in the mid-latitudes
350 of the upper troposphere for the first month after the SSW (September 17 to October
351 15, 2019). The decreased barotropic/baroclinic instability indicates less energy for
352 amplifying the waves passing by, causing perturbations from the troposphere to deflect
353 equatorward close to the tropopause rather than continuing vertically into the
354 stratosphere. Crucially, the deflection of EP fluxes leads to less westward momentum
355 transport into the SH stratosphere, preventing the seasonal weakening of the polar
356 vortex.

357 Under the influence of the downward control and wave-mean interaction, the
358 westward anomalies of the zonal wind propagate downward and reach the troposphere
359 after October 15, 2019. The anomalous zonal wind thus modulates the vertical and
360 meridional wind shear. It increases the atmospheric barotropic/baroclinic instability in
361 the midlatitude troposphere, which provides energy to amplify the PWs passing through
362 and removes the midlatitude propagation barrier for EP fluxes. As a result, more EP
363 flux PWs from the lower troposphere can propagate into the stratosphere (Figure 3d).
364 The refractive index influenced by barotropic/baroclinic instability and static stability
365 anomalies (Jucker et al., 2022) guides the PWs to propagate poleward in the

366 stratosphere. The westward momentum transport by the anomalous PWs decreases the
367 circumpolar eastward wind and benefits the early break of the stratospheric polar vortex.

368 The early reversal of the SH stratospheric zonal wind in November 2019 caused
369 the filtering of westward GWs by westward zonal wind rather than filtering the
370 eastward GWs by the eastward zonal wind in the lower stratosphere. More westward-
371 propagating GWs enhance the mesospheric meridional mean residual circulation,
372 including anomalous upwelling over the polar region and northward flow in the upper
373 mesosphere. The enhanced upwelling in the SH polar region is key to cooling the polar
374 mesosphere and increasing the water vapor mixing ratio. Both contributed to the early
375 onset of PMCs in the SH mesosphere in November 2019.

376 To conclude, our results indicate a mechanism in which the early spring
377 stratospheric perturbation could affect the vertical coupling from the troposphere to
378 the mesosphere in early winter. While we studied this mechanism concerning the
379 2019 SH September SSW, the early break of the SSPV, and the response in the polar
380 SH mesosphere in November 2019, it does not have to be limited to such events. It
381 can be expected to be relevant whenever lower stratospheric and upper tropospheric
382 barotropic/baroclinic instability interacts with the zonal wind anomalies and PW
383 activities. Thus, future work will explore the dynamical coupling during other
384 occurrences of stratospheric perturbation in both the Southern and Northern
385 Hemispheres. In addition to the dynamics process, the interplay between dynamics
386 and radiation heating could influence the long-lasting coupling process induced by the

387 stratospheric perturbation, but further work is required to explore this.

388

389 **Acknowledgment**

390 This work was supported by the National Natural Science Foundation of China
391 grants (42130203, 41874180, 41974175, 41831071); the B-type Strategic Priority
392 Program of the Chinese Academy of Sciences, grant no. XDB41000000; the pre-
393 research project on Civil Aerospace Technologies no. D020105 funded by China's
394 National Space Administration; and the Open Research Project of Large Research
395 Infrastructures of CAS –“Study on the interaction between low/mid-latitude
396 atmosphere and ionosphere based on the Chinese Meridian Project.” XW is supported
397 by the NSF via the NCAR's Advanced Study Program Postdoctoral Fellowship.

398

399 **Data Availability Statement**

400 The Cloud Imaging and Particle Size (CIPS) observed by AIM/aura are available at
401 <https://lasp.colorado.edu/aim/>. The subsets of MERRA-2 tavg3_3d_asm_Nv: 3d,3-
402 Hourly, Time-Averaged, Model-Level, Assimilation, Assimilated Meteorological
403 Fields V5.12.4 data are downloaded at
404 https://disc.gsfc.nasa.gov/datasets/M2T3NVASM_5.12.4/summary?keywords=MERRA-2%20tavg3_3d_asm.

406 The Aura/MLS temperature and water vapor mixing ratio measurements are

407 downloaded at
408 https://acdisc.gesdisc.eosdis.nasa.gov/data/Aura_MLS_Level2/ML2T.005/ and
409 https://acdisc.gesdisc.eosdis.nasa.gov/data/Aura_MLS_Level2/ML2H2O.005/,
410 respectively.

411

412 **Figure captions**

413 **Figure 1.** Anomalous SH polar cap (65° - 90° S) temperature from August 2019 to
414 December 2019 from (a) MLS observations and (b) MERRA2 reanalysis datasets; (c)
415 the mean SH PMC occurrence percentage derived from AIM/aura from 2007 to 2021
416 (thick black line) for the solstice. The blue line indicates the SH PMC occurrence during
417 2019, and the gray shading indicates 1 standard deviation.

418 **Figure 2.** (a) MERRA2 zonal mean temperature anomalies from 70-80 km, averaged
419 over 65° S and poleward (red line and shading), and the vertical component of the
420 residual circulation anomalies in the SH polar mesosphere (80° S and poleward, 65-80
421 km, green line, and shading) from August to December. (b) MERRA2 zonal mean zonal
422 wind at 60° S, 10 hPa from August to December (light gray lines). (c) Anomalous eddy
423 heat flux averaged over 45 - 75° S, 100 hPa from August 2019 to December 2019. (d)
424 The meridional gradient of potential vorticity averaged over 50 - 70° S from August 2019
425 to December 2019. The purple line denotes 2019, the thick black line indicates the mean
426 from 2004 to 2018, and the red and blue shadings indicate positive and negative

427 anomalies compared to the climatological mean.

428 **Figure 3.** (a) Zonal mean zonal wind anomalies (shading) superimposed by the
429 anomalous meridional gradient of the potential vorticity (\bar{q}_θ) multiplied by a (the
430 Earth's radius) at 60°S, 5-50 km (contours, white solid lines indicate positive anomalies,
431 white dashed lines indicate negative anomalies, and the contour interval is 30 m⁻¹) for
432 August-December 2019. The vertical red dashed line indicates the occurrence of the
433 SSW, while the vertical gray dashed line indicates the date of Oct 15. The horizontal
434 red solid line denotes the location of the lowermost stratosphere. (b) The anomalous
435 $\bar{q}_\theta * a$ due to the meridional wind shear (upper) and vertical wind shear at 60°S from
436 5 to 15 km for August-December 2019; (c) The latitude-altitude cross-section for the
437 SH \bar{q}_θ anomalies (shading), EP flux (green vector) and the wavenumber 1 refractive
438 index multiplied by a² (contour lines, the solid and dashed gray lines indicate 10 and -
439 10, respectively) averaged from September 17 to October 15, 2019; (d) is the same as
440 (c) but for the period from October 15 to November 15, 2019.

441 **Figure 4.** (a) Latitude versus altitude cross section of the anomalous meridional residual
442 mean circulation (m/s, streamlines), zonal mean temperature (K, shading) and zonal
443 mean volume mixing ratio of water vapor (ppmv, contours) from November 15 to 30,
444 2019; (b) anomalous vertical residual circulation (cm s⁻¹) averaged over 85°S-70°S
445 from 20 to 80 km from November 15 to 30, 2019, the thick black line associated with
446 the gray shading indicate the total vertical residual circulation anomalies, while the red
447 and blue lines associated with the red and blue shadings indicate the vertical residual

448 circulation anomalies due to none-resolved and resolved waves, respectively.

449

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