

20 **Abstract**

21 We develop an automatic workflow for denoising the fundamental mode surface wave
22 from ambient noise cross correlations (ANCs) calculated for a dense linear array. The
23 surface wave signal traveling between each station pair is first enhanced through three-
24 station interferometry. Then, phase travel times at different periods are determined in the
25 frequency domain. The proposed array-based method is applied to a 1.6-km-long dense
26 linear nodal array crossing surface traces of the San Jacinto fault near Anza, California.
27 Surface wave signals in ANCs of the nodal array are significantly enhanced after
28 denoising, particularly at high frequencies (> 2 Hz). Phase travel times are extracted
29 reliably in the period ranges of 0.3-1.3 s and 0.3-1.6 s for Rayleigh and Love waves,
30 respectively. The corresponding period-dependent phase velocity profiles derived from
31 the eikonal equation reveal high-resolution details of fault zone internal structures
32 beneath the array. A broad (500-1000 m) low-velocity zone that narrows with increasing
33 period is observed, illuminating a flower-shaped structure of the San Jacinto fault damage
34 zone.

35 **Plain Language Summary**

36 Properties of fault damage zone (width of 100-1000's meters), such as its geometry and
37 velocity reduction compared to the surrounding host rock, can have a profound impact on
38 our understandings of earthquake ruptures and the long-term behavior of the fault.
39 Several dense nodal arrays with 10-100 m spacing and aperture of a few kilometers were
40 deployed crossing surface traces of major faults, to provide high-resolution images of the
41 fault zone internal structures. Surface waves propagate between every two sensors with
42 frequency-dependent speeds are extracted from ambient noise cross correlations. By
43 measuring the relation between velocity and frequency, we can infer structures at depth
44 as surface waves are more sensitive to shallow structures at a higher frequency. However,
45 surface waves extracted from ambient noise at high frequencies (> 2 Hz) that are essential
46 to image fault zone in the top 100's meters are often very noisy. Here, we develop a new
47 method that utilizes three-station interferometry to suppress signals that are not traveling
48 on the surface. The quality of surface waves is significantly improved after the denoising,

49 especially at high frequencies (> 2 Hz), providing more reliable measurements and better
50 constraints on fault zone internal structures at shallow depth.

51 **1. Introduction**

52 Noise-based surface wave tomography has been widely used to resolve crustal
53 structures at various scales (e.g., Lin et al., 2009; Qiu et al., 2019; Wang et al., 2019;
54 Zigone et al., 2019). Analysis of high frequency (e.g., > 1 Hz) surface waves provides
55 crucial information on the shallow (top 10s to 100s of meters) materials with
56 unprecedented spatial resolution and thus improves our understanding of the local
57 seismic hazard. In contrast to high-quality signals at long periods (e.g., > 2 s), extraction
58 of surface waves from ambient noise cross correlations (ANCs) calculated at high
59 frequencies (e.g., > 1 Hz) remains a challenging topic due to its low signal to noise ratio
60 (SNR). Previous studies that utilize ANCs at high frequencies must first enhance the
61 surface wave signals by performing preprocessing and/or postprocessing steps that are
62 often ad hoc and may only work well for a specific dataset.

63 In the present paper, we develop a simple workflow, based on the idea of three-station
64 interferometry proposed by Zhang et al. (2020), that effectively enhances surface waves,
65 particularly at high frequencies, for ANCs of a 1-D linear array. We apply this method to
66 data recorded by a dense linear array deployed at the Ramona Reservation (RR) site
67 across surface traces of the San Jacinto fault, near Anza (Fig. 1). Seismic waveforms
68 from the RR array have been analyzed for fault zone internal structures in Qin et al.
69 (2020). ANCs were computed for each station pair of the RR array in Wang et al. (2019).
70 To enhance surface waves with low SNR in ANCs at high frequencies, Wang et al.
71 (2019) first applied a period-dependent tapering window and then applied double-
72 beamforming tomography to derive Rayleigh wave phase velocities for periods from 0.3
73 s to 0.8 s beneath the array.

74 Here, we first describe station configuration and ANC data of the RR array in section
75 2. Then, in section 3, following the flow chart illustrated in Figure 2, we present the
76 theoretical formulation for three-station interferometry using ANCs of a 1-D linear array
77 and illustrate the surface wave denoising process with a subset of ANC data computed
78 from the linear segment of the RR array. In section 4, surface wave phase travel times are

79 first extracted from the denoised wavefield and then inverted for phase velocity
80 dispersion models via the eikonal equation. Discussion of the denoising method and
81 comparison between the resulting phase velocity profiles and fault zone images from
82 previous studies (Qin et al., 2020; Wang et al., 2019) are presented in section 5.

83 **2. Data**

84 The RR array (red triangle in Fig. 1b) is located at north of Anza (blue square in Fig.
85 1b), California, and crosses surface traces of the Clark segment of the San Jacinto fault
86 (Fig. 1a). The array consists of 94 three-component 5-Hz Fairfield geophones (balloons
87 in Fig. 1a) that were set to record continuously for a month with a sampling rate of 500
88 Hz. ANC is obtained by first computing cross correlations of ambient noise data in 5-min
89 windows, and then stacking them over the entire recording period for each station pair
90 (Wang et al., 2019). The positive and negative time lags of the monthlong stacked ANC
91 are fold and averaged to suppress the effects of the asymmetric noise source distribution.
92 We use ANCs of a sub-array RR01-RR47 (yellow, blue, and red balloons in Fig. 1a) to
93 demonstrate the surface wave denoising process (Fig. 2a) developed in this study. The
94 sub-array has 47 stations with an average station spacing of ~ 30 m and an aperture of
95 ~ 1.6 km.

96 We project stations in the sub-array to the straight line connecting RR01 and RR47
97 (cyan dashed line in Fig. 1a) and compare interstation distances calculated using station
98 locations before and after the projection. The comparison yields negligible differences ($<$
99 1%) suggesting that the sub-array RR01-RR47 is in a 1-D linear configuration (later
100 referred to as “the linear RR array”). In Wang et al. (2019), a period-dependent velocity
101 threshold is applied to taper off the contamination of body waves or potential higher-
102 mode surface waves. In this study, however, we use a tapering window, between a
103 moveout velocity range of 2 km/s and 0.1 km/s (white dashed lines in Figs. 1c-d), to
104 ANCs of Transverse-Transverse (TT) and Vertical-Vertical (ZZ) components. The
105 tapered ANCs are then filtered between 0.2 Hz and 10 Hz and depicted as colormaps in
106 Figures 1c-d. Before performing surface wave denoising (Section 3) and tomography
107 (Section 4) illustrated in Figure 2, a series of Gaussian narrow bandpass filters centered
108 on consecutive periods between 0.3 s and 1.6 s are applied to the ANCs shown in Figures

109 1c-d and then each filtered waveform is normalized by its corresponding maximum. The
 110 Gaussian narrow bandpass filters are generated following section 3.1 of Qiu et al. (2019).

111 3. Surface Wave Denoising

112 Let $G_{i_j}(t)$ be the positive lag of ANC for the station pair of i -th (virtual source) and j -
 113 th (virtual receiver) sensors in the linear RR array (yellow, green, and red triangles in Fig.
 114 1a), we can expand it as

$$G_{i_j}(t) = S_{i_j}(t) + B_{i_j}(t) + N_{i_j}(t), \quad (1)$$

115 where $S_{i_j}(t)$ and $B_{i_j}(t)$ represent signals traveling on the surface (i.e. surface waves) and
 116 at depth (i.e. body waves) between the source i and receiver j , respectively. $N_{i_j}(t)$ is the
 117 residual (later referred to as “background noise”). This section aims to develop a
 118 denoising process that preserves $S_{i_j}(t)$ while suppressing $B_{i_j}(t)$ and $N_{i_j}(t)$ in equation 1.

119 3.1 Three-station interferometry for a 1-D linear Array

120 Since surface waves are dispersive, let $\tilde{G}_{i_j}(\omega)$ be the Fourier transform of $G_{i_j}(t)$ at
 121 the angular frequency ω , we can rewrite equation 1 in the frequency domain

$$\begin{aligned} \tilde{G}_{i_j}(\omega) &= A_{G_{ij}} \cdot e^{i\varphi_{G_{ij}}} = \tilde{S}_{i_j}(\omega) + \tilde{B}_{i_j}(\omega) + \tilde{N}_{i_j}(\omega) \\ &= \sum_S A_{S_{ij}} \cdot e^{-i(\omega \cdot T_{ij}^S + \varphi_S)} + \sum_B A_{B_{ij}} \cdot e^{-i(\omega \cdot T_{ij}^B + \varphi_B)} + \tilde{N}_{i_j}(\omega), \end{aligned} \quad (2a)$$

122 where $A_{S_{ij}}$ and T_{ij}^S are amplitude spectrum and phase travel time of surface wave signals
 123 in ANC at the angular frequency ω that propagate between the i -th and j -th stations,
 124 while $A_{B_{ij}}$ and T_{ij}^B represent those of body wave signals. φ_S and φ_B are initial phases of
 125 surface- and body-wave signals in the ANC, respectively, and dependent on the
 126 distribution of ambient noise sources (e.g., Lin et al., 2008). Since the fundamental mode
 127 surface wave, $\tilde{F}_{i_j}(\omega)$, is often the dominant signal in ANC, by assuming the higher-
 128 mode surface waves are negligible, we, therefore, can simplify equation 2a as:

$$\tilde{G}_{i_j}(\omega) = \tilde{F}_{i_j}(\omega) + \tilde{O}_{i_j}(\omega) = A_{F_{ij}} \cdot e^{-i(\omega \cdot T_{ij}^F + \varphi_F)} + \tilde{O}_{i_j}(\omega), \quad (2b)$$

129 where the symbol or subscript F stands for the fundamental mode surface wave.
 130 $\tilde{O}_{i_j}(\omega) = \tilde{B}_{i_j}(\omega) + \tilde{N}_{i_j}(\omega)$, that consists of signals from body waves and background
 131 noise, is the term we want to suppress in the denoising process. It is interesting to note

132 that $\varphi_F = \pi/4$ for an azimuthally homogenous ambient noise source distribution
 133 (Snieder, 2004), whereas $\varphi_F = 0$ when noise sources are only present in line with the
 134 station pair i and j (Lin et al., 2008).

135 For surface waves of a certain (e.g., fundamental) mode traveling between three
 136 stations $i < j < k$ in a 1-D linear array, the travel times satisfy the following relation

$$T_{ik}^S = T_{ij}^S + T_{jk}^S, \quad (3a)$$

137 whereas

$$T_{ik}^B < T_{ij}^B + T_{jk}^B, \quad (3b)$$

138 for body waves. Therefore, we introduce a third station k and perform three-station
 139 interferometry following Zhang et al. (2020):

$$\tilde{I}_{i_j}(\omega; k) = \begin{cases} \tilde{G}_{i_k}^*(\omega) \cdot \tilde{G}_{j_k}(\omega), & k < i \\ \tilde{G}_{i_k}(\omega) \cdot \tilde{G}_{j_k}(\omega), & i < k < j. \\ \tilde{G}_{i_k}(\omega) \cdot \tilde{G}_{j_k}^*(\omega), & k > j \end{cases} \quad (4a)$$

140 In equation 4a, we cross correlate $G_{i_k}(t)$ and $G_{j_k}(t)$ in the time domain, when $k < i$ or $k >$
 141 j (later referred to as “outer-source zone”). The interferometry becomes equivalent to the
 142 convolution of $G_{i_k}(t)$ and $G_{j_k}(t)$ in the time domain for station k located within the two
 143 virtual sources (i.e., $i < k < j$; later referred to as “inter-source zone”). For the case $k = i$ or
 144 j , we define $\tilde{I}_{i_j}(\omega; k) = A_{G_{ij}}^2 \cdot e^{i\varphi_{G_{ij}}}$ that approximates the convolution of $G_{i_j}(t)$ and
 145 $G_{i_j}(t)$ or $G_{j_j}(t)$, by assuming the amplitude spectrum of the auto-correlation $G_{i_j}(t)$ or
 146 $G_{j_j}(t)$ is similar to that of $G_{i_j}(t)$, i.e., $A_{G_{ii}} \approx A_{G_{jj}} \approx A_{G_{ij}}$.

147 Combining equations 2b, 3a, and 4a, if the fundamental mode surface wave is the
 148 dominate signal in ANC (i.e., \tilde{O}_{i_j} in Equation 2b is negligible), phase term of the
 149 interferogram $\tilde{I}_{i_j}(\omega; k) = A_{ij_k} \cdot e^{i\varphi_{ij_k}}$ is given by

$$\varphi_{ij_k}(\omega) = \begin{cases} -\omega \cdot T_{ij}^F - 2\varphi_F, & i < k < j \\ -\omega \cdot T_{ij}^F - \varphi_F, & k = i \text{ or } j. \\ -\omega \cdot T_{ij}^F, & k > i \text{ or } k < j \end{cases} \quad (4b)$$

150 φ_F and T_{ij}^F denote the initial phase and phase travel time of the fundamental mode surface
 151 wave signal (Equation 2b) extracted from the ANC of station pair i and j . Equation 4b
 152 suggests that the interferograms within either the inter- or outer-source zones share the
 153 same phase, whereas interferograms from different zones are only aligned in phase when

154 φ_F is zero. In cases when the term $\tilde{O}_{i,j}$ is significant, we can divide the interferogram
 155 $\tilde{I}_{i,j}(\omega; k)$ into two components: $\tilde{I}_{i,j}^F(\omega; k)$ and $\tilde{I}_{i,j}^O(\omega; k)$. $\tilde{I}_{i,j}^F(\omega; k)$ represents the
 156 interferogram that only involves the fundamental mode surface wave signal, i.e., when
 157 $\tilde{O}_{i,j}$ is set to zero. The phase of $\tilde{I}_{i,j}^F(\omega; k)$, given by equation 4b, is independent of k
 158 when φ_F is zero. In contrast, $\tilde{I}_{i,j}^O(\omega; k)$ engages contributions from body waves and
 159 background noise, and thus has a phase that varies significantly with k .

160 **3.2 Surface wave denoising of the linear RR Array**

161 As described in section 2, we first apply a series of Gaussian narrow bandpass filters
 162 centered on periods between 0.3 s and 1.6 s to the ANC's of the linear RR array (Figs. 1c-
 163 d). The period range is determined based on the station spacing and array aperture. Then
 164 the denoising process (Fig. 2a) is performed on the filtered ANC's for each period and
 165 component separately. Figure 3 shows results of the three-station interferometry applied
 166 to ANC's of TT component filtered at 0.3 s (Figs. S1a-c for 0.8 s and Figs. S2a-c & S3a-c
 167 for ZZ component) for an example station pair RR10 ($i = 10$) and RR40 ($j = 40$). The
 168 ANC's with RR10 and RR40 as the virtual source, $G_{i,k}(t)$ and $G_{j,k}(t)$ filtered at 0.3 s, are
 169 shown in Figures 3a and 3b, respectively. Black and blue waveforms denote ANC's with
 170 station k inside the outer-source ($k < 10$ or $k > 40$) and inter-source ($10 < k < 40$) zones,
 171 respectively, whereas the ANC of station pair RR10 and RR40 is depicted in red (i.e., k is
 172 the y-axis of Fig. 3).

173 Figure 3c demonstrates the resulting interferograms in the time domain computed
 174 following equation 4a (blue for convolution, black for cross correlation). Both the phase
 175 and envelope functions of the interferograms vary significantly with station k . Such k -
 176 value-dependent interferograms are also observed using ZZ component ANC's filtered at
 177 0.3 s (e.g., Fig. S2c). This is because of the low SNR for the fundamental mode surface
 178 wave (i.e., $\tilde{O}_{i,j}$ in Equation 2b is non-negligible) in the ANC's filtered at 0.3 s (e.g., Figs.
 179 3a-b and S2a-b). On the other hand, the interferograms are coherent and aligned well in
 180 phase for all k values at 0.8 s (e.g., Figs. S1c and S3c), when high-quality signals of the
 181 fundamental mode surface wave are observed in the filtered ANC's (e.g., Figs. S1a-b and
 182 S3a-b). This suggests that the initial phase φ_F is zero, i.e., $\varphi_{ij,k}(\omega) = -\omega \cdot T_{ij}^F$ (Equation
 183 4b), for ANC's of the linear RR array filtered at 0.8 s. We also verified that the initial

184 phase φ_F is zero for all periods (from 0.3 s to 1.6 s) analyzed in this study, by estimating
 185 the systematic phase difference between interferograms in the inter- and outer-source
 186 zones computed for the example station pair RR10 and RR40 (not shown here).

187 Therefore, we can simply stack the interferograms $\tilde{I}_{i_j}(\omega; k) = A_{ij_k} \cdot e^{i\varphi_{ij_k}}$, defined
 188 in equation 4a, over all available station k to enhance the contribution from the
 189 fundamental mode surface wave, $\tilde{I}_{i_j}^F(\omega; k)$, which has a phase that is independent of k
 190 (Equation 4b with $\varphi_F = 0$). The other component, $\tilde{I}_{i_j}^O(\omega; k)$, that involves contributions
 191 from body waves and background noise is suppressed through stacking, as its phase
 192 varies significantly with k (Section 3.1). The denoised waveform obtained through linear
 193 stacking is given by:

$$\tilde{C}_{i_j}^3(\omega) = \sum_{k=1}^N e^{i\varphi_{ij_k}} \sqrt{A_{ij_k}} / N. \quad (5a)$$

194 Here, N is the number of stations in the 1-D linear array. We take the square root of the
 195 amplitude spectrum A_{ij_k} in Equation 5a to suppress the effect of source spectra
 196 multiplication introduced in the three-station interferometry (Equation 4a). This is based
 197 on the assumption that amplitude spectra of the filtered ANC's are similar for all station
 198 pairs, i.e., $A_{ij_k} = A_{G_{ik}} \cdot A_{G_{jk}} \approx \hat{A}_G^2$.

199 In this study, we perform phase weighted stacking rather than linear stacking, as
 200 phase weighted stacking is more efficient in suppressing incoherent patterns and has
 201 negligible effects on phase measurements (e.g., Fig. S4; Schimmel and Paulssen, 1997).
 202 Let $\hat{C}_{i_j}^3(t; \omega_c)$ be the waveform in time domain denoised through linear stacking using
 203 ANC's filtered at the angular frequency ω_c , the corresponding waveform denoised
 204 through phase weighted stacking in the time domain is given by

$$C_{i_j}^3(t; \omega_c) = W_{i_j}(t; \omega_c) \cdot \hat{C}_{i_j}^3(t; \omega_c), \quad (5b)$$

205 where $W_{i_j}(t; \omega_c)$ is a weighting function that indicates the phase coherence of
 206 interferograms in the time domain (e.g., Fig. 3c) averaged over all k values.

207 3.3 Results

208 Figure 4a shows the comparison between the TT component ANC (black) of the
 209 example station pair RR10 and RR40 filtered at 0.3 s and the corresponding denoised

210 waveforms, computed through linear stacking (blue; Equation 5a) and phase weighted
 211 stacking (bottom red curve; Equation 5b). Although coda waves in the denoised
 212 waveform $C_{i_j}^3(t; \omega_c)$ are greatly suppressed, the fundamental mode surface wave is still
 213 not the dominant signal (e.g., large-amplitude wavelets before the surface wave). This is
 214 likely due to the poor SNR of surface waves in the ANC's filtered at 0.3 s (Figs. 3a-b).
 215 Therefore, we further enhance the surface wave signal by repeating the denoising process
 216 (Fig. 2a): first self-normalize the output wavefield of the current iteration, and then use
 217 the normalized wavefield as the input for the next iteration. The number of iterations is
 218 determined so that the difference between input and output wavefields of the last iteration
 219 is negligible (Fig. 2a).

220 We use symbol $C_{i_j}^{2+n}(t; \omega_c)$ to represent the waveform of station pair i and j , after
 221 applying n (≥ 1) iterations of the denoising process described in section 3.2 to ANC's
 222 filtered at angular frequency ω_c . Figure 4a suggests that four more iterations (red
 223 waveforms; $n = 5$) are needed to obtain surface waves with sufficient quality from the
 224 ANC's filtered at 0.3 s, i.e., the difference between the input, $C_{i_j}^6(t; \omega_c)$, and output,
 225 $C_{i_j}^7(t; \omega_c)$, waveforms is visually negligible. Although the SNR gradually increases in
 226 waveforms from bottom to top (Fig. 4a), the surface wave signal is always coherent and
 227 aligned in phase. As the SNR is much higher for surface waves in ANC's filtered at low
 228 frequencies (> 0.6 s), the number of iterations used to extract good quality surface waves
 229 is smaller ($n = 3$; e.g., 0.8 s in Fig. S1d).

230 Comparison between the TT component ANC data $G_{i_j}(t; \omega_c)$ filtered at 0.3 s and the
 231 denoised waveforms $C_{i_j}^7(t; \omega_c)$ is illustrated in Figures 4b-c for all station pairs.
 232 Although coherent fundamental mode surface waves are seen propagating at a group
 233 velocity slightly slower than 0.5 km/s in the filtered ANC data (Fig. 4b), wavelets with
 234 large amplitudes are observed before and after the surface wave signals (e.g., black
 235 waveform in Fig. 4b). The large amplitude waves arriving prior to the surface wave are
 236 likely related to the tapering window at short interstation distances (e.g., < 1 km), as the
 237 apparent moveout velocity is 2 km/s (the upper limit velocity of our tapering window;
 238 Section 2). At long interstation distances, the precursor wavelets have an apparent
 239 moveout velocity of ~ 1 km/s and may represent the contribution from body wave energy.

240 These waves that we want to suppress in the denoising process are sometimes even larger
241 than the surface wave signal in the filtered ANC (e.g., black waveform in Fig. 4b).

242 After five iterations of the denoising process for ANCs filtered at 0.3 s, surface wave
243 signals are well preserved whereas the other signals are greatly suppressed (Fig. 4c). The
244 amplitude spectra averaged over all station pairs for data before (in black) and after (in
245 red) denoising are demonstrated in Figure 4d. The observation of a smoother mean
246 amplitude spectrum after denoising (red curve in Fig. 4d) suggests that the difference
247 between amplitude spectra of every two station pairs is much smaller. Since SNR of the
248 fundamental mode surface wave is much larger in ANCs filtered at 0.8 s (Figs. S1a-b and
249 S3a-b), the difference between waveforms before and after denoising is still noticeable
250 but much smaller (Figs. S1e-g and S3e-g). As the fundamental mode surface wave is the
251 dominant signal after denoising (e.g., Figs. 4c and S1f-S3f), all results in section 4 refer
252 to the fundamental mode surface wave.

253 **4. Surface Wave Tomography**

254 In this section, we use waveforms of TT and ZZ components denoised at each period
255 (e.g., Figs. 4c and S1f-S3f) to infer phase velocity structures of Love and Rayleigh waves
256 beneath the array, respectively. Following the flow chart shown in Fig. 2b, we first
257 determine cycle-skipped phase travel times of surface waves propagating between all
258 available station pairs at each period (e.g., Fig. 5a) in the frequency domain, which is
259 much simpler than measuring in the time domain but requires high SNR (Section 4.1).
260 Second, we infer phase velocity structures beneath the linear RR array, using travel time
261 measurements after cycle-skipping correction from section 4.1, via the eikonal equation
262 in section 4.2 (e.g., Fig. 5b). The aim of this section is to demonstrate that robust surface
263 wave phase velocity models can be resolved from the denoised waveforms.

264 **4.1 Determination of phase travel time**

265 Frequency time analysis (FTA) is widely used in previous studies to determine phase
266 travel time of surface wave in ANC (e.g., Bensen et al., 2007; Lin et al., 2008; Qiu et al.,
267 2019). First, Gaussian narrow bandpass filters centered on a series of consecutive
268 frequencies are applied to the ANC, then the phase travel time dispersion is measured
269 using the envelope and phase functions of the filtered ANC in the time domain. The

270 advantage of FTA is that reliable phase travel times can still be extracted when SNR is
 271 low at high frequencies. However, ad hoc criteria and thresholds are required to automate
 272 the FTA. Additional details on the FTA method can be found in section 3 of Qiu et al.
 273 (2019). Since our goal is to verify that the signals after denoising are representative of
 274 surface waves and high SNR is achieved for all frequencies, we thus measure phase
 275 travel times from the denoised waveforms in the frequency domain, which is much
 276 simpler than the FTA method.

277 Although surface wave is the dominant signal in the denoised waveform, we still
 278 observe waves with small amplitudes before and after the surface wave (e.g., black
 279 waveform in Fig. 4c). This is because we can only suppress rather than remove signals
 280 that are not surface waves. Here, we apply a tapering window (e.g., black dashed lines in
 281 Figs. S5d-S8d) centered on the surface wave to further remove these background signals.
 282 Width of the tapering window is set to four times the dominant period of the array-mean
 283 amplitude spectrum (e.g., red curve in Fig. 4d), whereas the center is determined by the
 284 average phase and group velocities of the array (e.g., red and cyan stars in Figs. S5b).
 285 Details of the tapering window can be found in Text S1.

286 Assuming these background signals are negligible after denoising and tapering (e.g.,
 287 blue waveform in Fig. 5a), combining equation 2b and $\varphi_F = 0$, we have

$$\tilde{C}_{i-j}^{DT}(\omega; \omega_c) = A_{F,ij}(\omega; \omega_c) e^{-i\omega T_{ij}^F(\omega)}, \quad (6)$$

288 where $\tilde{C}_{i-j}^{DT}(\omega; \omega_c)$ is the spectrum of the tapered waveform that is denoised at the angular
 289 frequency ω_c for station pair i and j . Equation 6 suggests that we can extract cycle-
 290 skipped phase travel time from the phase spectrum of the tapered waveform. Therefore,
 291 we measure the wrapped phase (i.e., between -2π and 0) of the spectrum $\tilde{C}_{i-j}^{DT}(\omega; \omega_c)$ at
 292 the peak frequency f_{\max} (3.2 Hz; e.g., dashed line in Fig. 4d) of the array-mean amplitude
 293 spectrum, where the array-mean SNR of the surface wave is the highest. Then, the cycle-
 294 skipped phase travel time is computed as the wrapped phase divided by $-2\pi f_{\max}$.

295 Figure 5a shows the cycle-skipped phase travel time (white circles) measured for
 296 surface waves filtered at 0.3 s from station pairs associated with a common virtual source
 297 RR10 (y-axis of 0 km). To obtain the actual phase travel time, we perform a simple
 298 cycle-skipping correction as follows:

- 299 (1) Similar to Figure 5a, we first extract all the cycle-skipped phase travel times for
 300 surface waves of a virtual shot gather and arrange them as a function of the location
 301 to the virtual source.
- 302 (2) We perform cycle-skipping correction for surface waves traveling NE (toward RR47)
 303 and SW (toward RR01) separately.
- 304 (3) For surface waves traveling in the same direction, the principle of the cycle-skipping
 305 correction is to ensure that the travel time of any virtual receiver is larger than those
 306 of receivers that are closer to the virtual source after the correction.
- 307 (4) In practice, we examine measurements T_i and T_{i+1} of every two adjacent virtual
 308 receivers with the i -th station being closer to the virtual source. If $T_i \geq T_{i+1}$, we use T_i
 309 as the reference and add N/f_{\max} (N is an integer) to T_{i+1} so that $T_{i+1} + N/f_{\max} > T_i \geq$
 310 $T_{i+1} + (N-1)/f_{\max}$. The correction is performed for closer-to-source pairs first.
- 311 Travel times, for the virtual shot gather of RR10, after the correction are illustrated as red
 312 stars in Figure 5a. We note that a more sophisticated cycle-skipping correction (e.g.,
 313 using phase velocity structure inferred at a longer period as the reference) is needed when
 314 station spacing is larger than one wavelength.

315 4.2 1-D eikonal tomography

316 We use the eikonal equation to derive phase velocity structures using travel time
 317 measurements of all station pairs in the linear RR array (Section 4.1). First, we project all
 318 stations to the straight line connecting RR01 and RR47 (cyan dashed line in Fig. 1a).
 319 Second, travel time measurements associated with each virtual source i at the target
 320 frequency f_{\max} are extracted and interpolated (e.g., black curve in Fig. 5a) with a regular
 321 grid size of $\Delta=50$ m. Since variations in topography (Fig. 2b of Qin et al., 2020) have a
 322 negligible effect ($< 0.5\%$) on the results, the eikonal tomography can be simplified as:

$$\tilde{v}_i(x; f_{\max}) = 2 \cdot \Delta / [T_i(x + \Delta; f_{\max}) - T_i(x - \Delta; f_{\max})], \quad (7)$$

323 where $\tilde{v}_i(x; f_{\max})$ and $T_i(x; f_{\max})$ are the local phase velocity and interpolated phase
 324 travel time, respectively, of the grid cell at location x . Since the local phase velocity \tilde{v}_i
 325 only varies with the grid cell location, it is independent of virtual source i . Thus, we can
 326 average the 1-D phase velocity profiles resolved from all available virtual sources at the
 327 same frequency f_{\max} to achieve a more reliable phase velocity model:

$$\bar{v}(x; f_{\max}) = \sum_{i=1}^{N_x} \tilde{v}_i(x; f_{\max}) / N_x, \quad (8a)$$

328 and estimate the corresponding uncertainty as the standard deviation:

$$\delta(x; f_{\max}) = \sqrt{\sum_{i=1}^{N_x} [\tilde{v}_i(x; f_{\max}) - \bar{v}(x; f_{\max})]^2 / N_x}, \quad (8b)$$

329 where N_x is the number of virtual sources available for stacking at location x .

330 In surface wave studies, phase velocities derived at near-virtual-source grid cells are
 331 often excluded to satisfy the far-field approximation (e.g., Bensen et al., 2007). The size
 332 of the exclusion zone is usually multiples of the analyzed wavelength (e.g., one
 333 wavelength in Wang et al., 2019). Here, however, we set an exclusion zone with a fixed
 334 size of 100 m, i.e., discard phase velocities derived at the four grid cells closest to the
 335 virtual source. Figure 5b shows the 1-D phase velocity profile, in white dots, derived
 336 using measurements associated with the virtual source RR10 (black curve in Fig. 5a) for
 337 Love waves at 3.2 Hz (~ 0.3 s), whereas phase velocity profiles resolved from all virtual
 338 sources are illustrated in gray curves and as the colormap. The average phase velocity
 339 and uncertainty profiles are calculated via equation 8 and demonstrated as red stars and
 340 error bars, respectively, in Figure 5b. Results for Love waves at a lower frequency (~ 0.8
 341 s) and those of Rayleigh waves are shown in Figures S9 and S10-S11, respectively.

342 Figure 6 shows phase velocity models resolved at periods ranging from 0.3 s to 1.6 s
 343 for Love waves (Fig. 6a) and 0.3 s to 1.3 s for Rayleigh waves (Fig. 6b), together with the
 344 corresponding uncertainty estimations (Figs. 6c-d). The period range in the plot is
 345 determined so that the resolved maximum uncertainty is smaller than 0.1 km/s. In
 346 general, the uncertainties are smaller than 0.03 km/s for both Rayleigh and Love waves at
 347 all analyzed periods, indicating the resolved phase velocity structures are robust and
 348 reliable. This also justifies our choice of an exclusion zone with a 100-m-radius, as one
 349 wavelength at low frequency (e.g., ~ 900 m for Rayleigh wave at ~ 0.8 s; Fig. S11b) is
 350 much larger than 100 m. We note that results of Love waves at ~ 0.4 s are excluded here
 351 due to anomalous large uncertainties (Fig. S12b). This is due to the observation of non-
 352 negligible signals, likely representing higher-mode surface waves, traveling at a different
 353 phase speed compared to that of the fundamental mode Love waves (Figs. S13-S14). We

354 note that reliable phase travel times can still be measured for the fundamental mode Love
355 waves at ~ 0.4 s if FTA is used.

356 Phase velocity models of both Love and Rayleigh waves show a ~ 500 - to 1000 -m-
357 wide low-velocity zone at low frequencies (e.g., > 0.8 s) that gradually narrows with the
358 period. Combined with the fact that phase velocity at lower frequency is more sensitive to
359 structures at greater depth, this observation likely indicates a flower-shaped (i.e., width
360 decreases with depth) fault damage zone beneath the linear RR array. We also see several
361 ~ 100 -m-wide narrow zones, that are close to the mapped fault surface traces (black
362 dashed lines in Figs. 6a-b), with extremely low phase velocities (< 500 m/s) at high
363 frequencies (e.g., 0.3 - 0.6 s). However, the shape and location of these low-velocity zones
364 are different between Figures 6a and 6b. In addition, complicated patterns of low-velocity
365 anomalies at various scales in both phase velocity models (Figs. 6a and 6b) suggest high
366 degrees of heterogeneity in shear wave velocity structures beneath the linear RR array.
367 Structure patterns that are inconsistent between models of Love and Rayleigh waves may
368 indicate the existence of radial anisotropy or complicated structures of V_p/V_s ratio.

369 **5. Discussion**

370 We compare the Rayleigh wave phase velocity models derived from this study and
371 Wang et al. (2019) in the overlapping period (0.3 - 0.8 s) and spatial (RR01-RR47) ranges
372 (Fig. 7). The same ANC dataset of ZZ component is utilized in Wang et al. (2019) to
373 derive the phase velocity model (Fig. 7b) for Rayleigh waves. In their study, double
374 beamforming technique is used, where the local phase velocity is obtained through grid
375 search: first sum all ANCs of two sub-arrays (three nearby stations) via slant-stacking
376 with different velocity values, then determine the local phase velocity of each sub-array
377 based on the maximum amplitude of the envelope function for the stacked waveform.
378 Through slant-stacking, Wang et al. (2019) was able to enhance the low SNR surface
379 waves in ANCs at high frequencies and generate a robust phase velocity model.

380 In general, consistent velocity values and structural patterns are observed in both
381 phase velocity models, such as an ultra-low velocity (< 0.4 km/s) zone on the NE side of
382 the middle fault surface trace (F2 in Figs. 7a-b). This indicates that the enhanced signals
383 in ANCs after denoising are surface waves. It is also important to note that our Rayleigh

384 wave phase velocity model covers a wider period range (0.3-1.3 s; Fig. 6b) compared to
385 that of Wang et al. (2019) (0.3-0.8 s). There are two reasons for that: first, we only
386 exclude data associated with station pairs that are less than 100 m apart, whereas the size
387 of exclusion zone is one wavelength in Wang et al. (2019); second, phase velocities
388 inferred from slant-stacking have much larger uncertainties than those derived from
389 travel times, as the peak of envelope function is much sensitive to noise than phase
390 function (e.g., Section 3.3 of Qiu et al., 2019), at longer periods.

391 The uncertainties estimated from this study (Fig. 7c) are significantly smaller than
392 those of Wang et al. (2019) (Fig. 7d), suggesting the phase velocity structure is better
393 constrained in this study, particularly near the fault surface traces (e.g., NE of F3 and
394 between F1 and F2 in Fig. 7) and at high frequencies (e.g., < 0.6 s). The larger uncertainty
395 values observed in Wang et al. (2019) are likely due to the low SNR of surface waves in
396 ANCs at high frequencies and should be significantly reduced if the denoised waveforms
397 (e.g., Fig. 4c) are used. By comparing phase velocity models resolved from this study and
398 Wang et al. (2019), we demonstrate that fundamental mode surface waves in ANCs are
399 successfully enhanced through the proposed denoising method. Besides, phase velocity
400 structures are better constrained (i.e., much smaller uncertainties) using the denoised
401 waveforms in this study than those inferred from the raw ANCs in Wang et al. (2019).

402 The derivations presented in section 3.1 are based on two assumptions: higher-mode
403 surface waves are negligible, and stations are aligned in a straight line. After performing
404 the denoising process on ANCs of the linear RR array filtered at different periods
405 between 0.3 s and 1.6 s for TT and ZZ components, we do not observe higher-mode
406 surface waves except for TT component data filtered at 0.4 s (Figs. S13-S14). Even when
407 higher-mode surface waves are present, we further demonstrate that the denoising process
408 based on three-station interferometry will only enhance surface waves of each mode
409 separately in Text S2. Regarding the effect of array geometry, we do not investigate in
410 detail the bias in surface wave travel times measured from the denoised waveforms when
411 stations are not perfectly aligned in a straight line. Assuming the velocity structure is
412 homogenous perpendicular to the strike direction of the array, the station-configuration
413 error is proportional to the difference between interstation distances calculated using
414 station locations before and after projecting the array to a straight line (e.g., Fig. 1a).

415 For a rough estimation, the mean and maximum differences between interstation
416 distances of the linear RR array calculated before and after the linear projection (Section
417 2) are $\sim 0.1\%$ and 1% , respectively. The uncertainties estimated from eikonal tomography
418 (Figs. 6c-d) suggest $\sim 0.5\text{-}1\%$ and $\sim 3\%$ for the mean and maximum perturbations in the
419 resolved phase velocities, which is comparable and larger than the bias in array geometry.
420 Therefore, we conclude that this denoising method is still robust when the station-
421 configuration error is less than the allowable uncertainty of the resulting phase velocity
422 model (e.g., mean and maximum of 1% and 3% in this study). Surface wave denoising of
423 ANC_s computed from arrays in a 2-D configuration (e.g., include all stations in the RR
424 array) will be the subject of a future study.

425 Although we only illustrate the denoising method using ANC_s computed following
426 the preprocessing steps described in Wang et al. (2019), it can be applied to any noisy
427 wavefield that consists of surface waves propagating within a linear array. Therefore, this
428 method can be used as a routine postprocessing procedure for improving the quality of
429 surface waves in ANC_s of a linear array with any type of preprocessing steps (e.g., coda
430 wave C^3 or C^N ; Froment et al., 2011). After denoising, we may be able to perform surface
431 wave studies on data from linear array deployments that were previously considered to be
432 too short in recording duration or too “noisy” (i.e., deployed during high seismic activity
433 period) to retrieve good quality surface wave signals from ANC_s.

434 We note that Lin et al. (2009) used the term “apparent phase velocity” for the phase
435 velocity resolved from the eikonal equation (Equation 8a). They concluded that the
436 apparent phase velocity is accurate when the wave frequency is sufficiently high or the
437 wave amplitude is varying smoothly in space. Otherwise, a correction term must be
438 added (Lin & Ritzwoller, 2011). Considering the linear RR array is dense (average
439 station spacing of ~ 30 m) and deployed crossing fault damage zones that can
440 significantly amplify seismic motions (Qin et al., 2020), both the “high frequency” and
441 “smooth amplitude variation” approximations are likely violated for phase velocity
442 models derived in section 4.2. Therefore, we propose performing travel-time based full
443 waveform tomography (e.g., Zhang et al., 2018) on the denoised surface waves, rather
444 than inverting the “apparent” phase velocity models (Figs. 6a-b), to infer shear wave
445 velocity structures beneath the linear RR array as the subject of a future study.

446 In the present study, we only show the denoised ANC's filtered at frequencies up to ~ 3
447 Hz (Figs. 4 and S2d-g) by requiring a minimum wavelength of 150 m (i.e., the maximum
448 station spacing) for eikonal tomography (Section 4). However, this denoising method can
449 be applied to ANC's at even higher frequencies when surface wave signals are present
450 (e.g., ~ 5 Hz in Figs. S15-S16). Using high-quality surface waves denoised from ANC's of
451 linear arrays crossing major fault zones for a wide range of frequencies (e.g., 0.5-5 Hz in
452 this study; 2-40 Hz in Zigone et al., 2019), shear wave velocity model that extends to
453 both shallow (top few tens of meters) and deep (top few kilometers) structures can be
454 derived. Such fault zone velocity model with unprecedented high-resolution will
455 complement the qualitative model inferred from traditional fault zone analyses (e.g., Qin
456 et al., 2018, 2020; Qiu et al., 2017, 2020; Share et al., 2017, 2019). An integration of both
457 quantitative and qualitative fault zone models can have significant implications for
458 seismic hazard evaluations (e.g., Ben-Zion & Shi, 2005; Spudich & Olsen, 2001) and
459 long-term behavior of the fault (e.g., Thakur et al., 2020).

460 **6. Conclusions**

461 We demonstrate the effectiveness and robustness of the three-station-interferometry-
462 based surface wave denoising method using ANC's of the linear RR array, particularly at
463 high frequencies (e.g., > 2 Hz). The proposed surface wave denoising method can be
464 applied to a wide range of topics in the future using data recorded by 1-D linear arrays:

- 465 1. Reduce the minimum duration of ambient noise recording and preprocessing steps
466 needed to achieve high-quality surface waves from ANC's.
- 467 2. Obtain good quality surface waves from ANC's computed for fault zone linear arrays
468 deployed right after a major earthquake (e.g., fault zone arrays in Ridgecrest area;
469 Catchings et al., 2020) without applying any sophisticated preprocessing steps.
- 470 3. Provide high-quality surface wave signals both at high (> 2 Hz) and low frequencies
471 (< 1 Hz) for better constraints of shallow (top 10s to 100s of meters) materials
472 through full-waveform surface wave tomography.
- 473 4. Investigate the initial phase φ_F (Equation 2b) for a wider frequency band and arrays
474 at different locations, and its relation to the ambient noise source distribution

475

476

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571

572

573 **Figure captions**

574

575 Figure 1. (a) Google map for the RR array (colored balloons) deployment that crosses
576 surface traces of the San Jacinto fault (colored lines). The stations colored in white are
577 not analyzed in this study, whereas the green balloons denote three sensors closest to
578 each corresponding fault surface trace. Surface wave denoising procedure is
579 demonstrated for an example station pair (red balloons). (b) Zoom out map of the San
580 Jacinto fault zone. The background colors indicate topography. The red star and blue
581 square denote locations of the RR array and the town of Anza. The black lines illustrate
582 surface traces of major faults in this area. EF – Elsinore Fault; SAF – San Andreas Fault;
583 SJF – San Jacinto Fault. (c) Ambient noise cross correlations at TT component of all
584 station pairs for the sub-array RR01-RR47. The cross correlations are arranged according
585 to interstation distance with red and blue colors representing positive and negative values.
586 All the waveforms are first tapered using a velocity range of 2 km/s and 0.1 km/s (dashed
587 lines), and then bandpass filtered between 0.2 and 10 Hz. (d) Same as (c) for the ZZ
588 component.

589 Figure 2. (a) Flow chart of the surface wave denoising and imaging procedure
590 developed in this study. The dashed box outlines the part of the diagram that performs
591 surface wave tomography. The workflow adopted in this study for surface wave
592 tomography is shown in (b). ANC – Ambient Noise Cross-correlation.

593 Figure 3. (a) Ambient noise cross correlations (ANCs) of TT component narrow
594 bandpass filtered at 0.3 s associated with the virtual source RR10 (red star). Waveforms
595 are arranged by the station number of the virtual receiver. (b) Same as (a) for virtual
596 source RR40. Waveforms in black and blue represent ANCs of virtual receivers in the
597 outer- and inter-source zones, respectively, while the red waveform denotes the ANCs of
598 the station pair of the two virtual sources RR10 and RR40 (red balloons in Fig. 1a). (c)
599 Interferograms calculated via three-station interferometry (Equation 4a).

600 Figure 4. (a) Comparison between the ANC (red waveforms in Figs. 3a-b), linear
 601 stacked (LS) C3 (Equation 5a), and phase weighted stacked (PWS) C3 – C7 (Equation
 602 5b) of the example station pair RR10 and RR40 at TT component. (b) ANCs of TT
 603 component (Fig. 1c) narrow bandpass filtered at 0.3 s with red and blue colors
 604 representing positive and negative values. The three white dashed lines illustrate moveout
 605 velocities of 2 km/s, 0.5 km/s, and 0.1 km/s. The black waveform denotes ANC of the
 606 example pair RR10 and RR40, i.e. bottom black waveform in (a). (c) PWS C7 of all
 607 station pairs at 0.3 s. The white dashed line denotes a moveout velocity of 0.5 km/s. The
 608 black waveform denotes PWS C7 of the example station pair RR10 and RR40, i.e. top
 609 red waveform in (a). (d) Array-mean amplitude spectra of waveforms (b) before and (c)
 610 after denoising are depicted in black and red, respectively. The peak frequency of the red
 611 amplitude spectrum is illustrated by the red dashed line and labeled in the top right of the
 612 panel (c).

613 Figure 5. (a) Love waves associated with the virtual source RR10 after denoising at
 614 0.3 s and tapering. Red and blue colors represent positive and negative values,
 615 respectively. White circles denote cycle-skipped phase travel time measurements,
 616 whereas red stars indicate travel times after cycle skipping correction (Section 4.1). The
 617 black curve illustrates the corrected phase travel time after interpolation using a grid size
 618 of 50 m. (b) Phase velocity profiles resolved for Love waves at 0.3 s underneath the 1-D
 619 linear array. White circles depict the 1-D phase velocity profile derived via eikonal
 620 equation using travel time measurements shown as the black curve in (a). The colormap
 621 illustrates phase velocity profiles obtained using different stations as virtual sources (x-
 622 axis), with white space representing the near-source exclusion zone. The phase velocity
 623 profile averaged over results of all virtual sources are depicted as red stars, with the error
 624 bar representing the corresponding standard deviation. The black vertical dashed line
 625 denotes the array-mean phase velocity estimated at 0.3 s (red star in Fig. S4b).

626 Figure 6. Phase velocity dispersion profiles for (a) Love and (b) Rayleigh waves
 627 beneath the RR array. The vertical dashed lines denote locations of the mapped fault
 628 surface traces (Fig. 1a), while the horizontal arrow outlines the group of stations that
 629 record fault zone trapped waves (Qin et al., 2020). Uncertainties of the resolved phase

630 velocity profiles are shown in (c) and (d) for Love and Rayleigh waves, respectively.
631 Phase velocity and uncertainty profiles for Love wave at 0.4 s are excluded here due to
632 large uncertainties (Figs. S12).

633 Figure 7. Comparison of (a) phase velocity and (b) uncertainty profiles of this study
634 and those, (b) and (d), of Wang et al. (2019) in the overlapping period range. The white
635 space indicates the area not covered by the final model. Note, the definition of phase
636 velocity uncertainty is different in this study (standard deviation) compared to that of
637 Wang et al. (2019) (standard deviation of the mean). A conversion is made to transform
638 the uncertainties shown in Wang et al. (2019) into the uncertainty profiles of (d).

Figure 1.

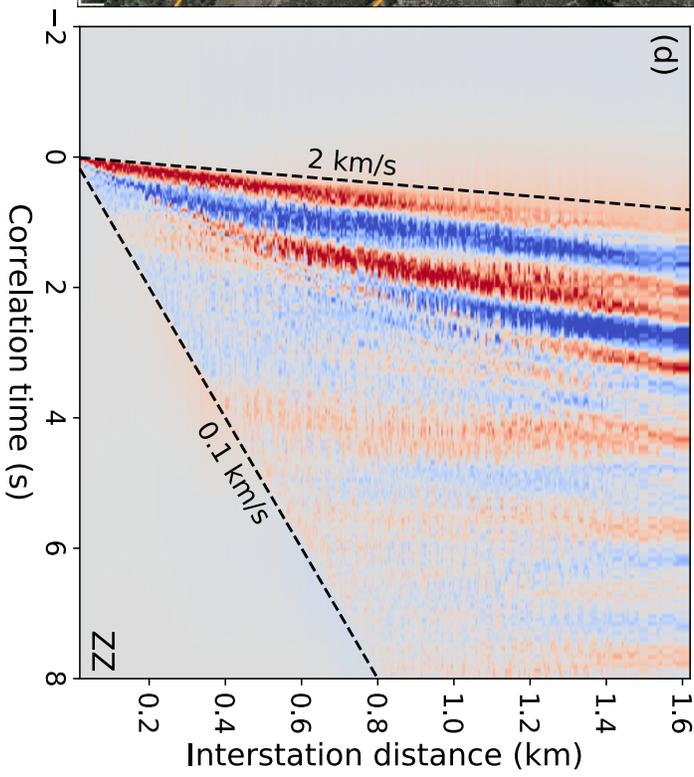
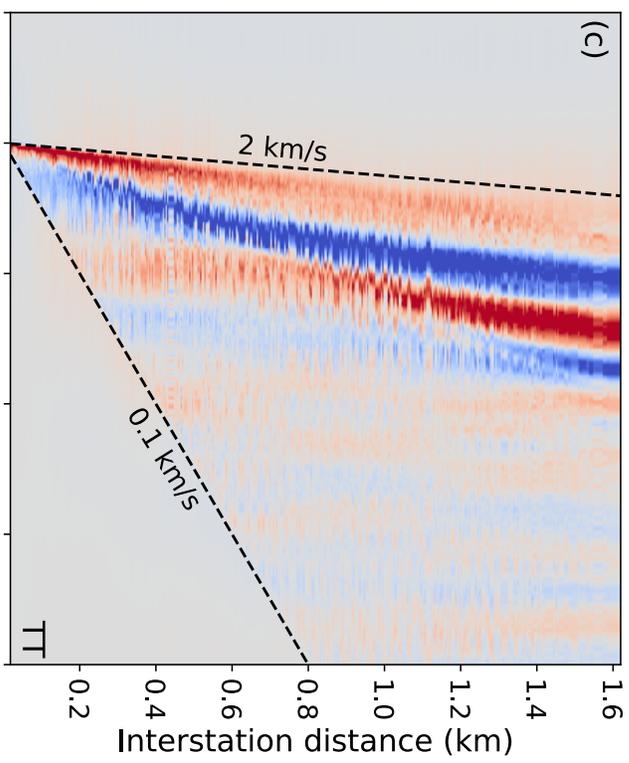
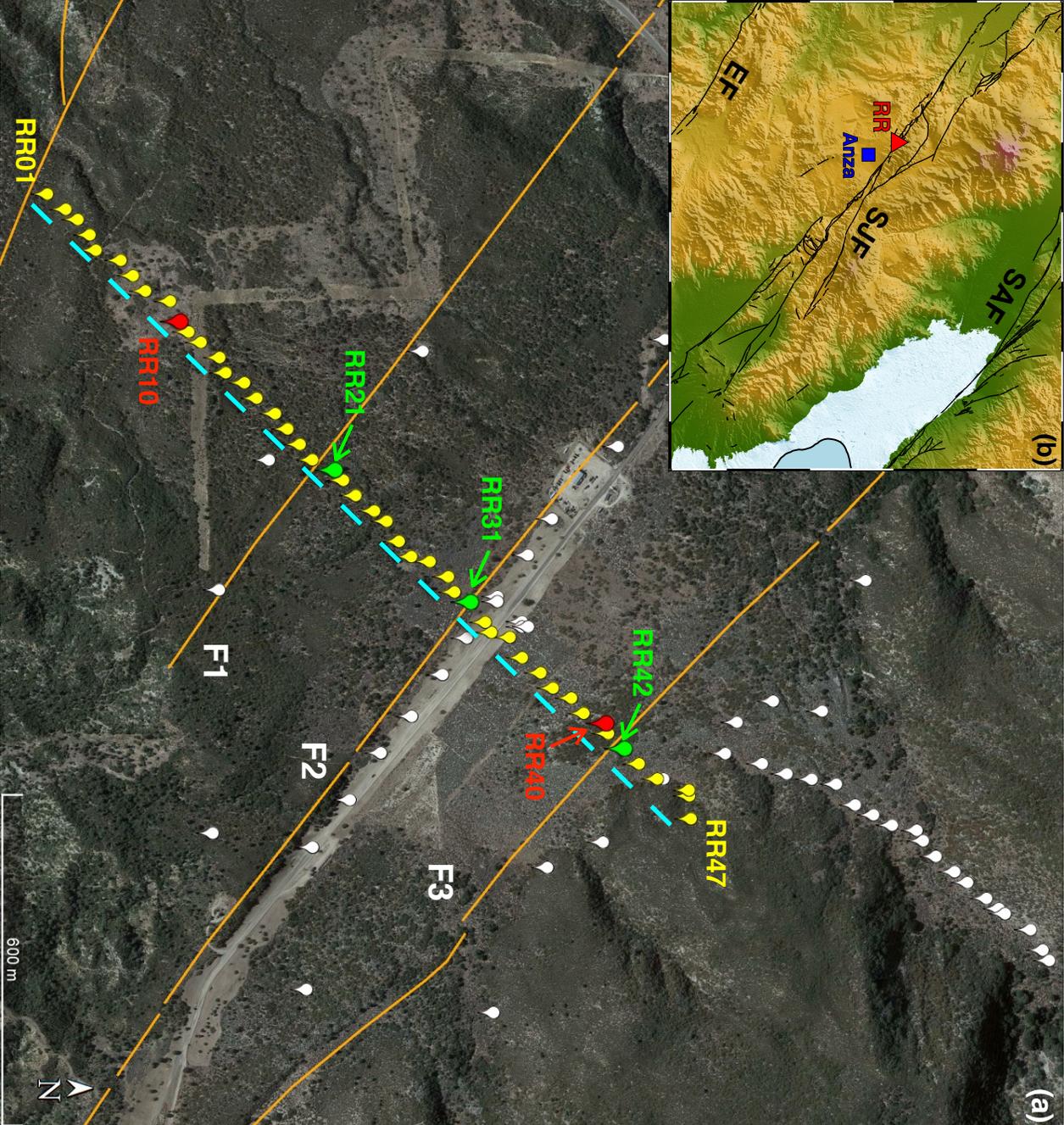
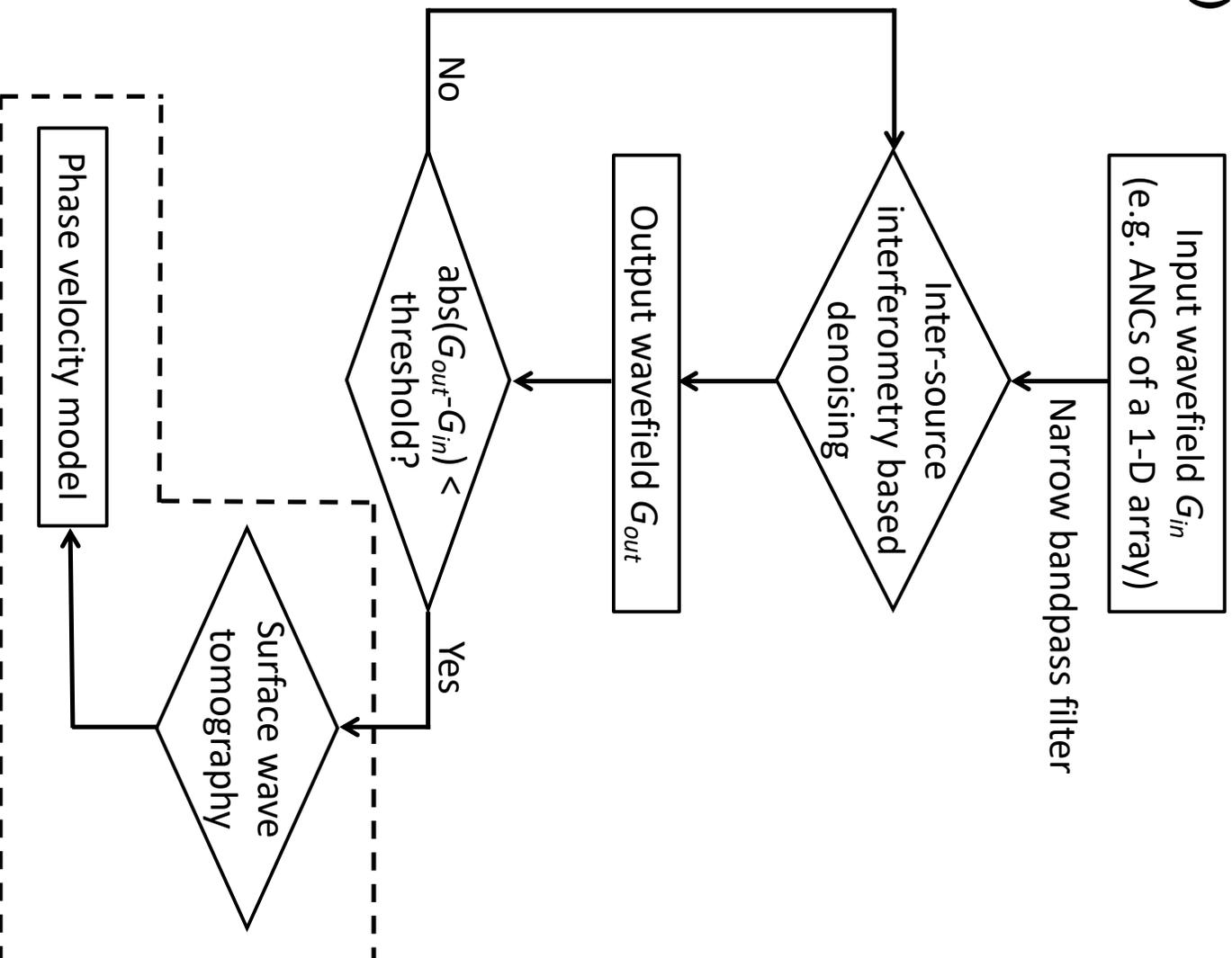


Figure 2.

(a)



(b)

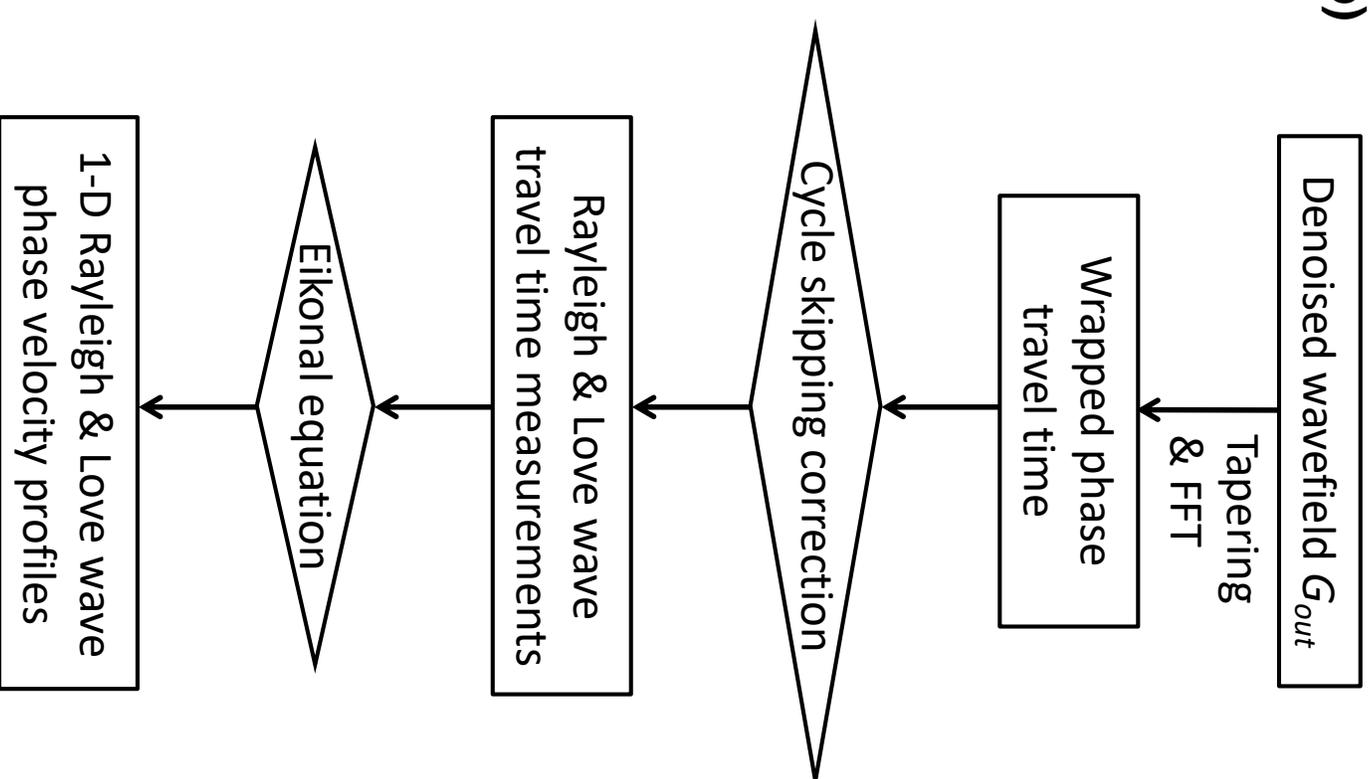


Figure 3.

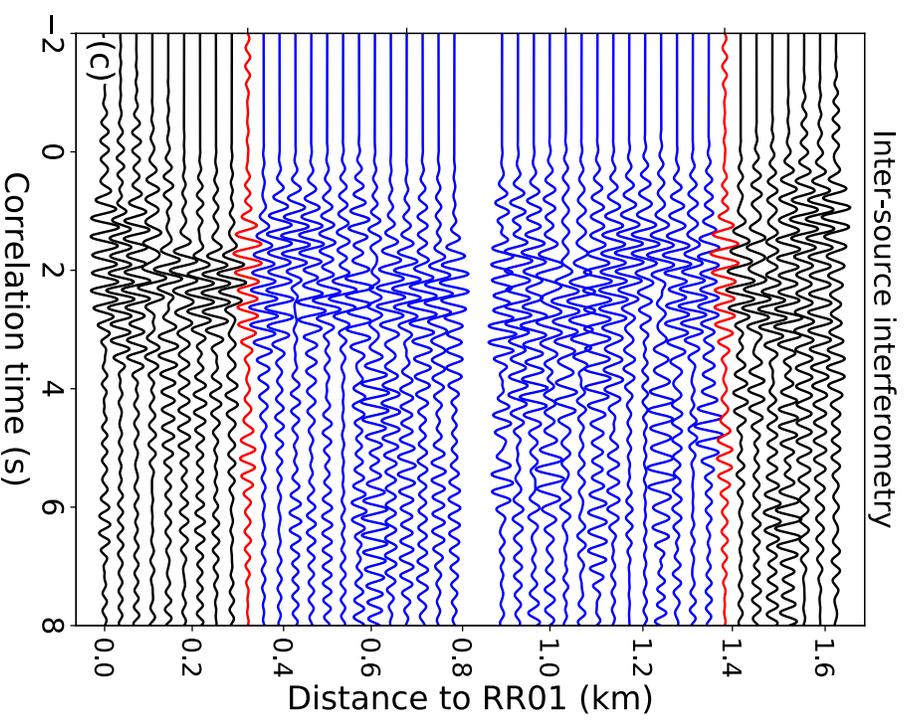
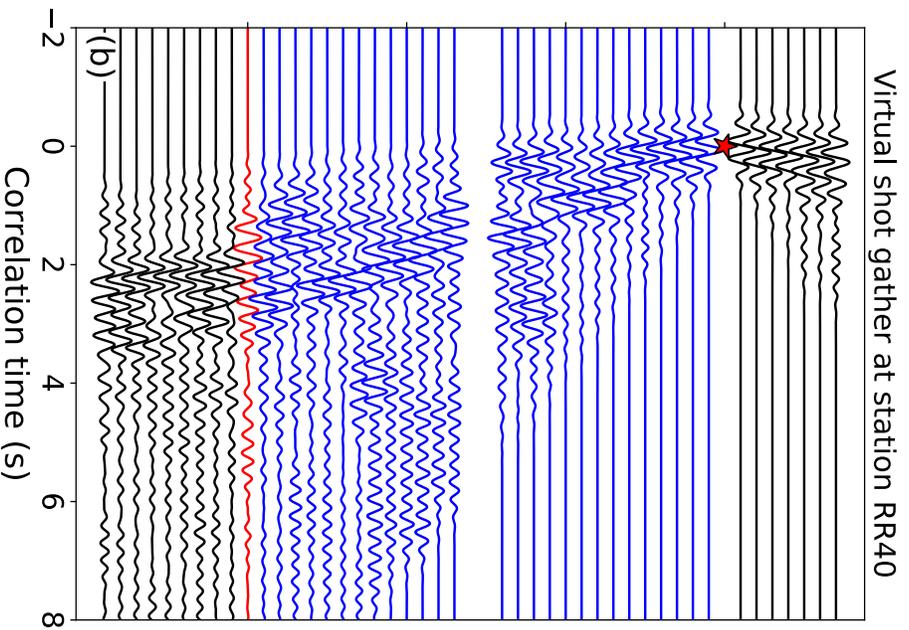
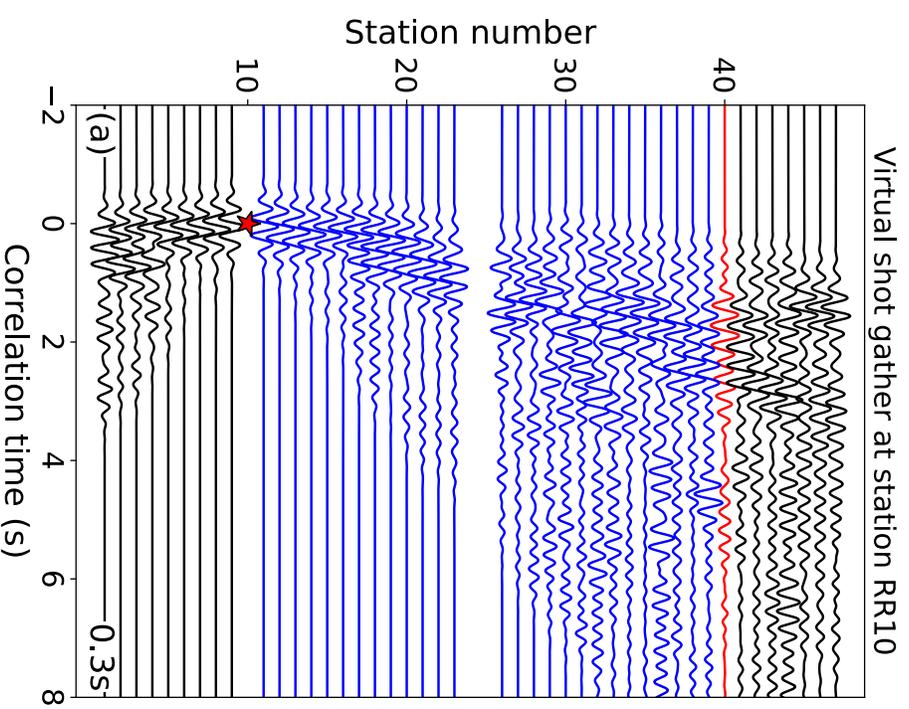


Figure 4.

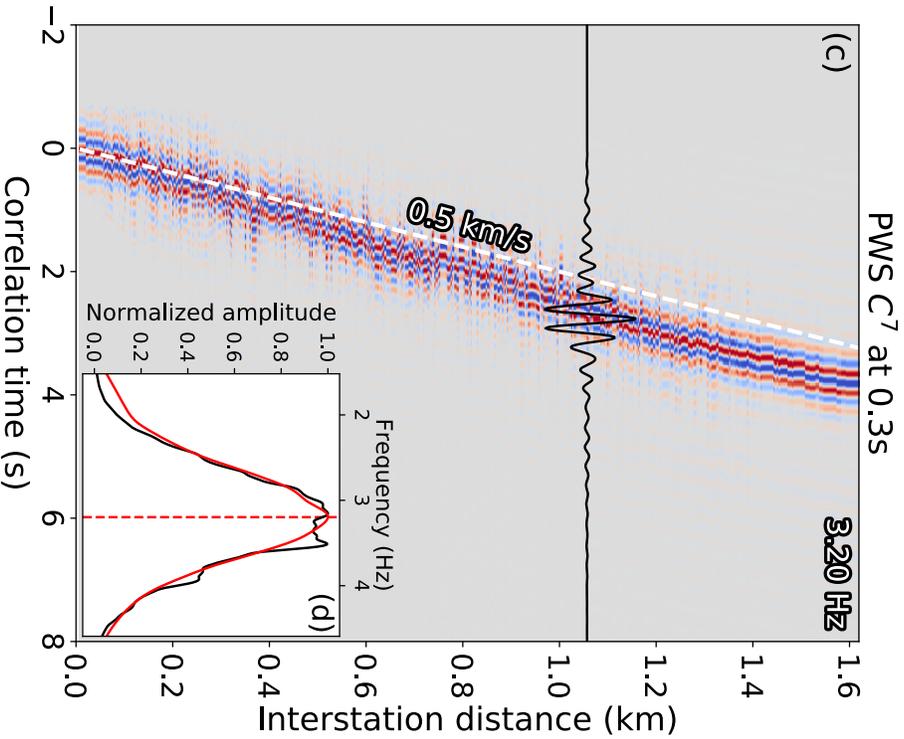
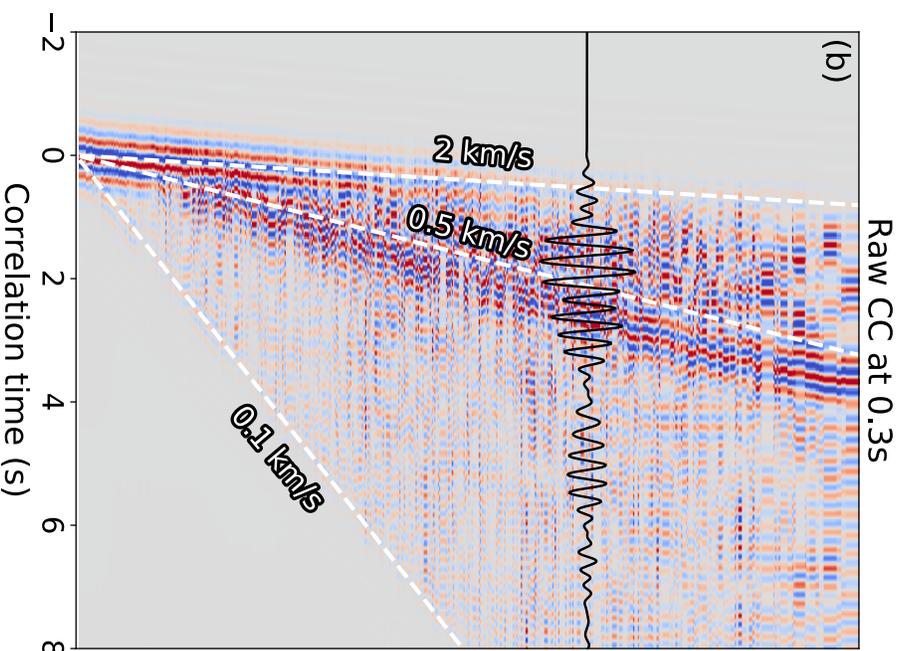
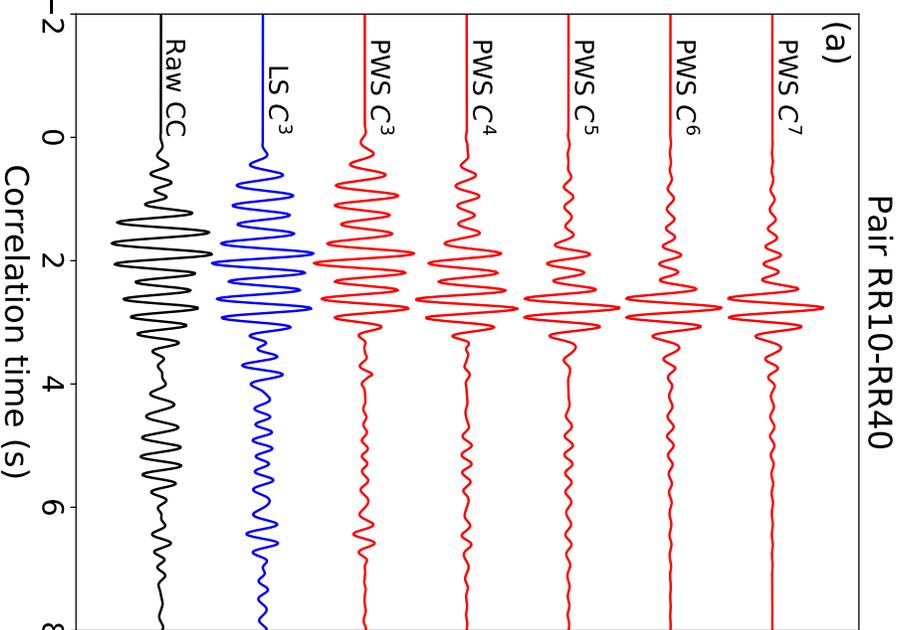


Figure 5.

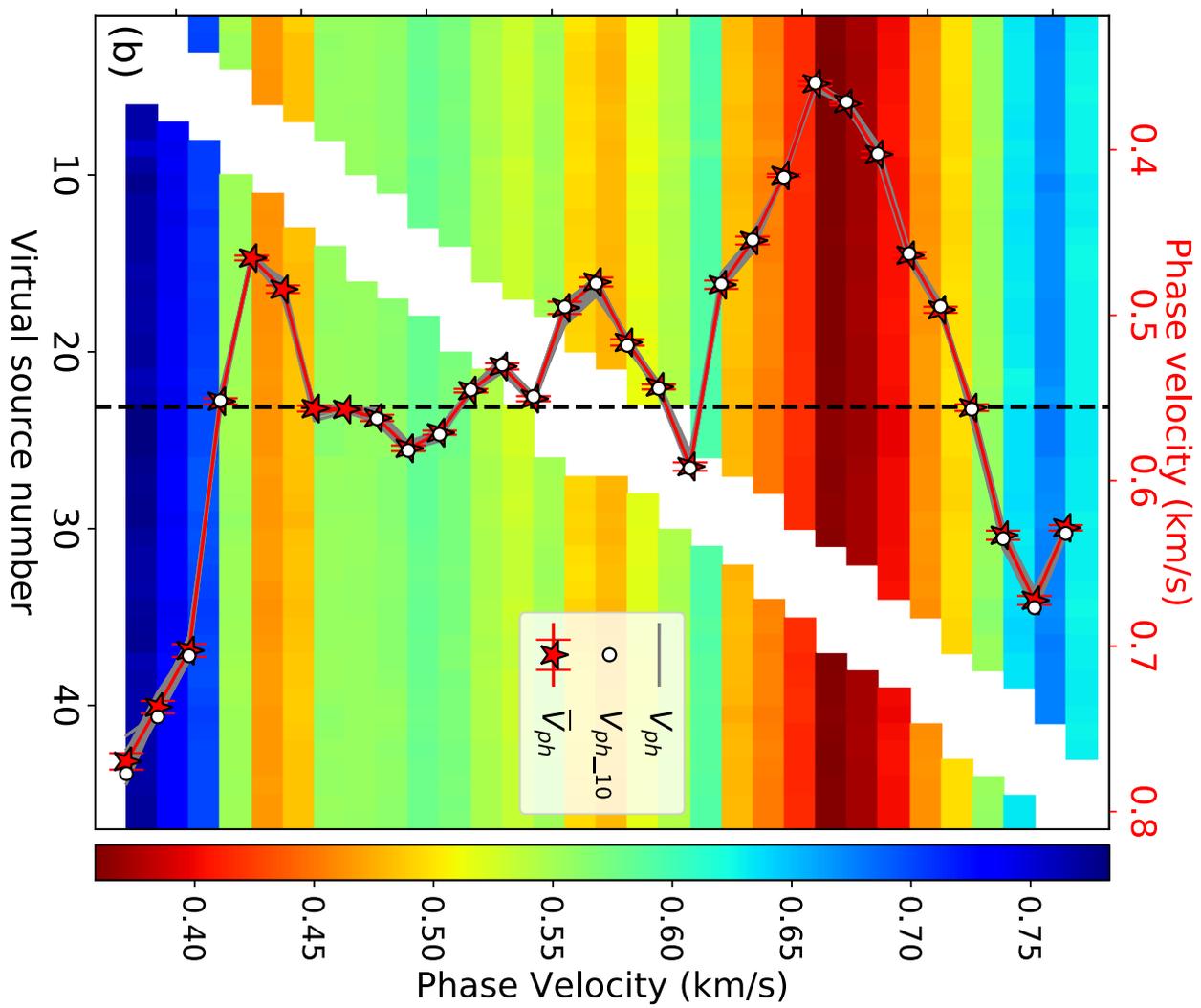
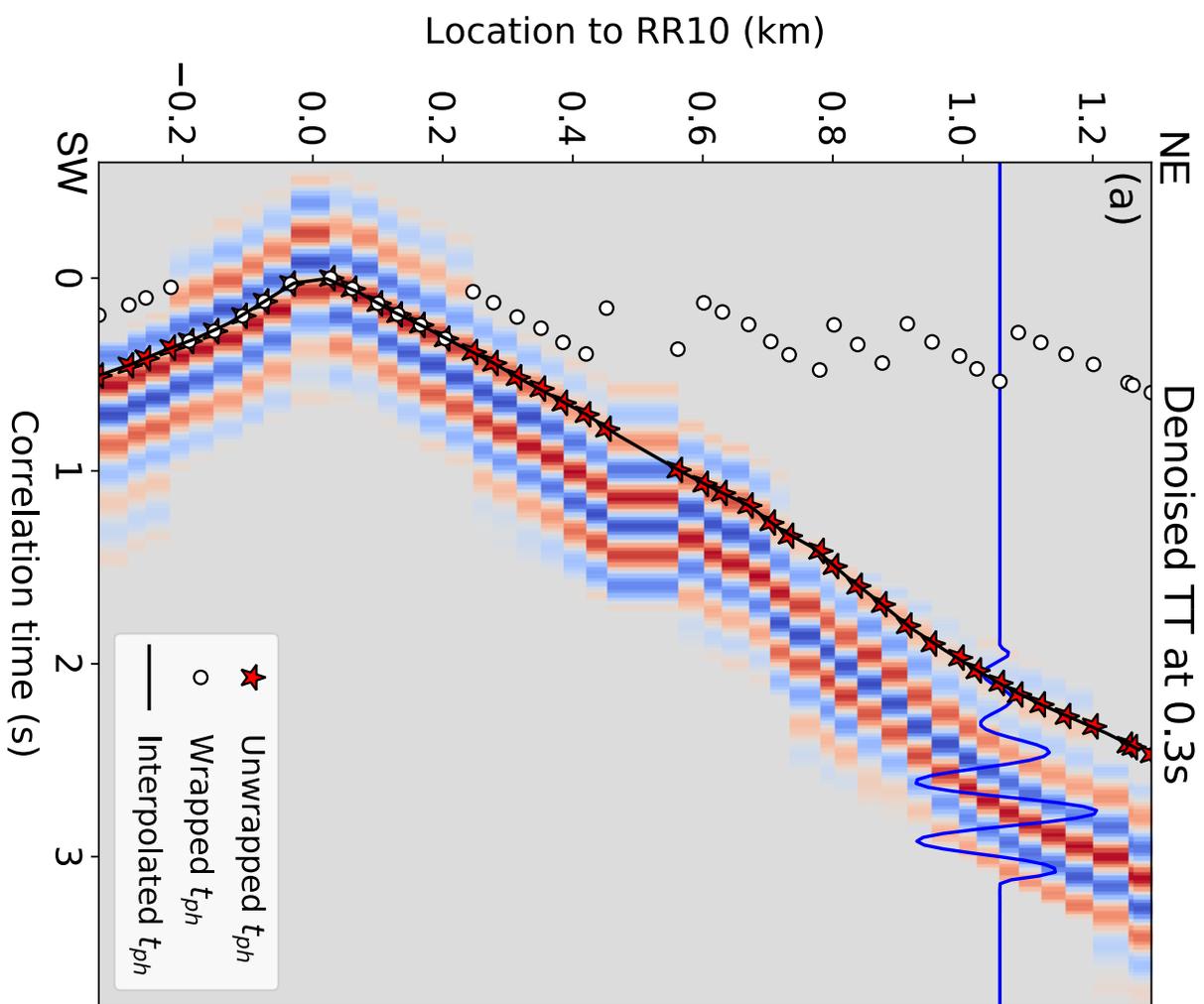


Figure 6.

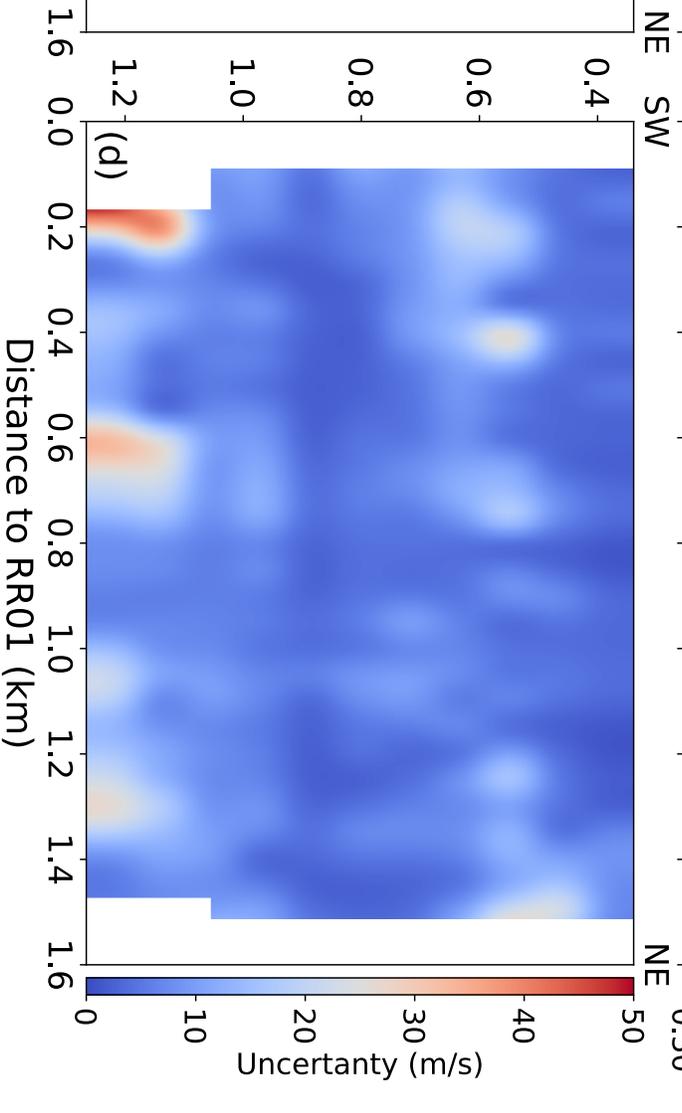
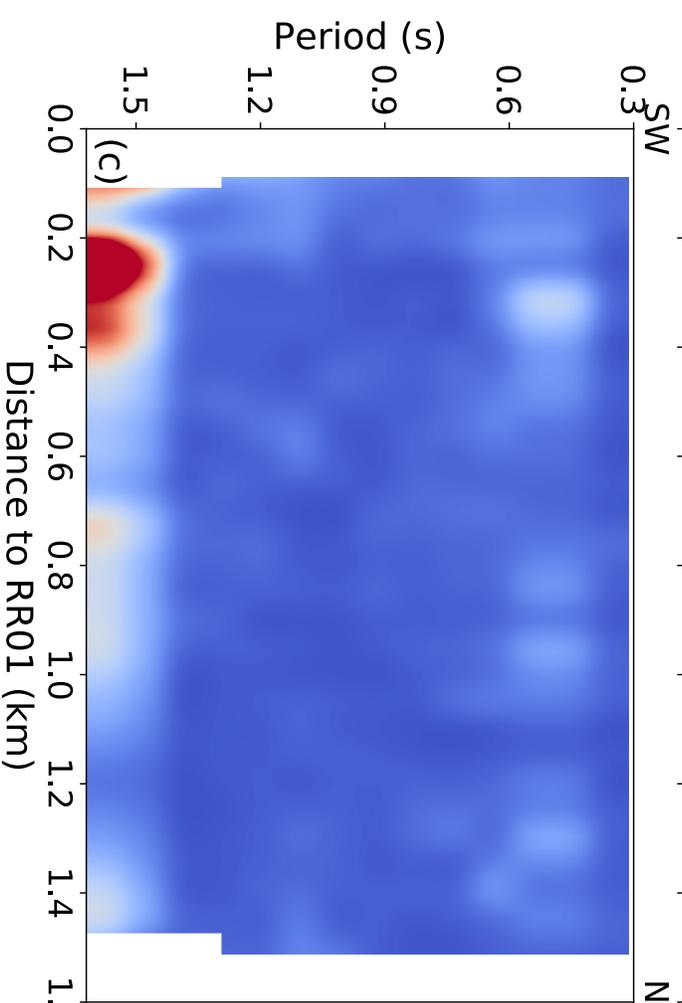
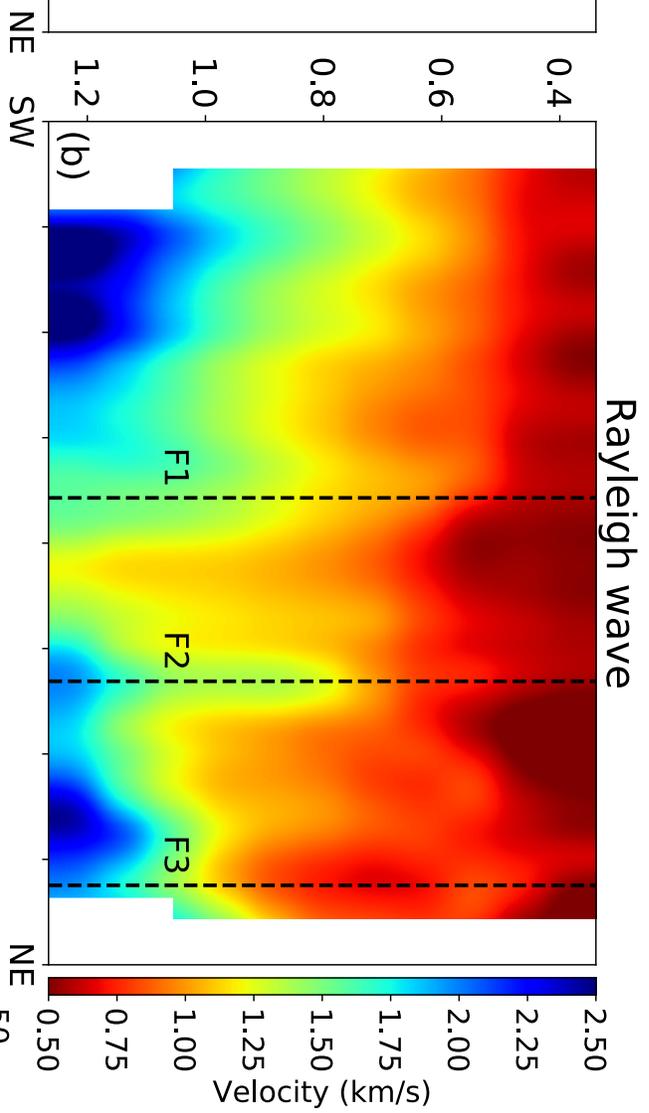
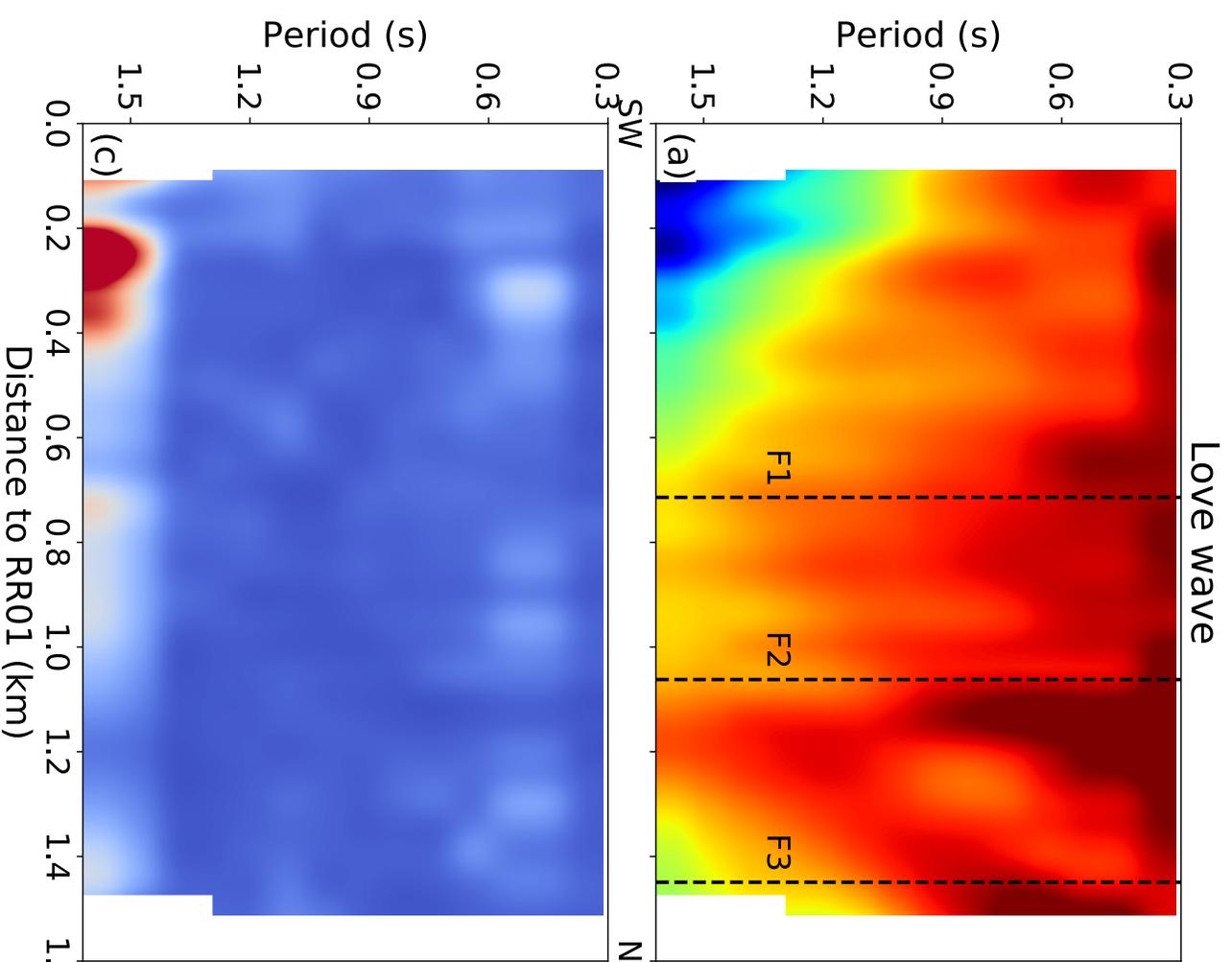


Figure 7.

