

# Different formation histories between the African and Pacific large low shear-wave velocity provinces as revealed by their water contents

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## Key points:

- H<sub>2</sub>O distribution in a single large low shear-wave velocity province is heterogeneous.
- The African large low shear-wave velocity province has higher H<sub>2</sub>O content and H<sub>2</sub>O/Ce than the Pacific large low shear-wave velocity province.
- The two large low shear-wave velocity provinces may have different formation and evolution histories.

## **Abstract**

Global tomographic models have revealed the existence of two large low shear-wave velocity provinces (LLSVPs) underlying Africa and the Pacific, which are regarded as sources of most typical mantle plumes. Plume-induced basalts have the potential to imply the formation mechanisms and evolutionary histories of the LLSVPs. In this study, we measured H<sub>2</sub>O contents in clinopyroxene and olivine phenocrysts from Cenozoic basalts produced by the Kerguelen and Crozet mantle plumes, which are deeply rooted in the African LLSVP. The results were used to constrain the H<sub>2</sub>O content in the source of basalts, yielding  $1805 \pm 579$  ppm for the Kerguelen plume and  $2144 \pm 690$  ppm for the Crozet plume. H<sub>2</sub>O contents in the mantle sources of basalts fed by other plumes rooted in these two LLSVPs were calculated from literature data. Combining these results together, we show that the African LLSVP seems to have higher H<sub>2</sub>O content and H<sub>2</sub>O/Ce (620-2144 ppm and 184-592, respectively) than the Pacific LLSVP (262-671 ppm and 89-306, respectively). These features could be ascribed to incorporation of subducted material, which had experienced variable degrees of dehydration during its downwelling, into the LLSVPs. Our results imply that the continuous incorporation of subducted oceanic crust modifies the compositions of LLSVPs and induces heterogeneous distribution of H<sub>2</sub>O within individual LLSVPs and distinct H<sub>2</sub>O contents between the African and Pacific LLSVPs. This suggests that the African and Pacific LLSVPs might have different formation and evolution histories.

## **Plain language summary**

Seismic tomography has revealed that large-scale heterogeneity exists at the base of the mantle, dominantly manifested by two large low shear-wave velocity provinces (LLSVPs) beneath Africa

and the Pacific. The LLSVP may play a key role in the global mantle convection and the core-mantle interaction, but their compositions and causes remain enigmatic up to present. Mantle plume-induced ocean island basalts (OIBs), originally related to the LLSVPs, can potentially supply some clues for this point. The presence of H<sub>2</sub>O can remarkably influence the physical and chemical properties of mantle rocks. Moreover, H<sub>2</sub>O contents and H<sub>2</sub>O/Ce ratios in basalts are efficient tracers to characterize the basalt sources. We compiled and inversely estimated the source H<sub>2</sub>O contents for global OIBs to track H<sub>2</sub>O distribution in the LLSVPs. The distribution of H<sub>2</sub>O shows significant heterogeneity both inside a single LLSVP and between the two LLSVPs. The incorporation of subducted oceanic crusts with variable dehydration extents during downwelling may be responsible for the heterogeneous H<sub>2</sub>O distribution within individual LLSVPs and the distinct H<sub>2</sub>O contents and H<sub>2</sub>O/Ce ratios between the African and Pacific LLSVPs. It may also imply the different formation and evolutionary histories for the two LLSVPs.

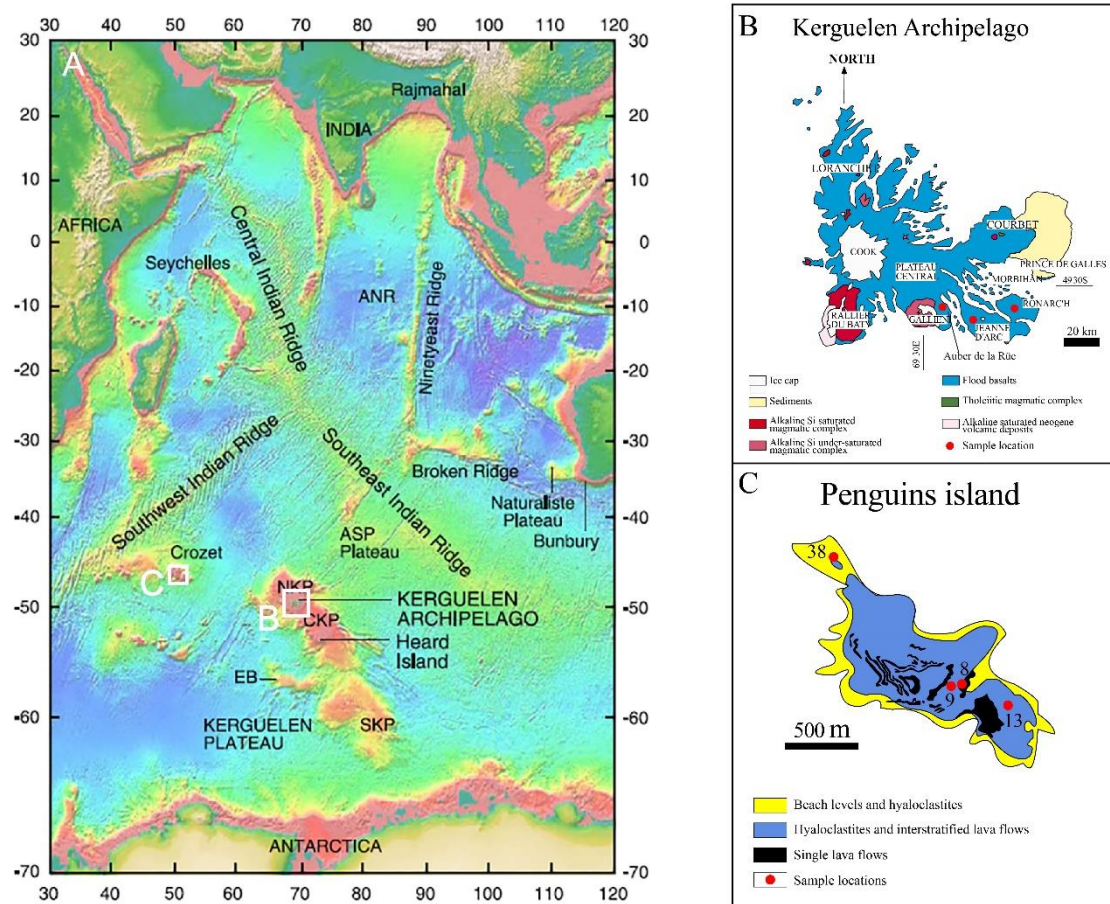
## **1. Introduction**

Seismic tomography revealed large-scale heterogeneity at the base of the mantle, dominated by two large low shear-wave velocity provinces (LLSVPs), one beneath Africa and the other beneath the Pacific, with regions around the LLSVPs corresponding to an assembly of cooler subducted slabs (Cottaar & Lekic, 2016; Garnero et al., 2016). Up to the present, the formation mechanism and geochemical nature of the LLSVPs have remained controversial. Several lines of geophysical evidence support the view that LLSVPs have distinct compositional properties compared with the ambient mantle and can be regarded as thermochemical piles (e.g., Garnero et al., 2016). Two different hypotheses have been put forward to explain the origin of the piles: they are primordial features generated during an ancient differentiation event (e.g., Labrosse et al., 2007; Lee et al.,

2010) or they represent an assembly of subducted oceanic crust (e.g., Brandenburg & van Keken, 2007; Mulyukova et al., 2015; Thomson et al., 2019). Diverse seismic studies have delineated typical mantle plumes as fluxes of material and energy originating from the LLSVPs (e.g., French & Romanowicz, 2015; Marignier et al., 2020; Montelli et al., 2004, 2006). Hotspots have long been considered as surficial expressions of mantle plumes (Hofmann, 1997; White, 2015). Thus, hotspot-related basalts could potentially serve as windows to geochemical characteristics of the LLSVPs, further providing information on their formation and evolution (e.g., Doucet et al., 2020; Homrighausen et al., 2020; Huang et al., 2011; Jackson et al., 2018). Doucet et al. (2020) compiled Sr-Nd-Pb isotope data on plume-generated basalts from ocean islands and oceanic plateaus overlying geographically both the African and Pacific LLSVPs. They proposed that these two LLSVPs have distinct geochemical compositions and evolutionary histories.

Garnero et al. (2016) highlighted the significance of volatiles in the formation of LLSVPs and demonstrated that these two LLSVPs may have compositional differences induced by different influxes of volatiles and recycled materials. Indeed, the presence of H<sub>2</sub>O can significantly affect the physical and chemical properties of mantle rocks under conditions of the lowermost mantle, including, for instance, melting processes (Nomura et al., 2014), viscosity (Nakagawa et al., 2015), and thermal conductivity (Hsieh et al., 2020). Different subducted slabs could undergo variable extents of dehydration during their subduction or residence in the deep mantle, suggesting that the continuous incorporation of subducted oceanic crust can potentially induce heterogeneous distribution of H<sub>2</sub>O in the deep mantle (Dixon et al., 2002; Workman et al., 2006). Since diverse mantle components have distinct H<sub>2</sub>O contents and H<sub>2</sub>O behaves incompatibly during partial melting with a compatibility similar to that of Ce, H<sub>2</sub>O contents and H<sub>2</sub>O/Ce ratios of basalts have

83 been proposed as efficient tracers of compositional features for basalt sources (e.g., Dixon et al.,  
84 2002; Kendrick et al., 2017). Therefore, H<sub>2</sub>O contents and H<sub>2</sub>O/Ce ratios of LLSVP-derived plumes  
85 can give new insights into the formation and evolution of LLSVPs. However, so far, no comparison  
86 of the H<sub>2</sub>O contents between the Africa and Pacific LLSVPs has been undertaken.



87 **Figure 1** Bathymetric map of the Indian Ocean (A) modified from Doucet et al. (2005), with the locations of  
88 the studied Kerguelen basalts from the Kerguelen Archipelago (B) and Crozet basalts from the Penguins  
89 Island (C). Kerguelen basalts with an age of less than 25 Ma were sampled during several field missions in the  
90 southeast Kerguelen Archipelago.

91 In this study, we provide new H<sub>2</sub>O abundance data in primary melts of Cenozoic basalts from the  
92 Kerguelen and Crozet Archipelagos (Figure 1), which have been suggested to originate from the  
93 Africa LLSVP. And we also compile H<sub>2</sub>O contents reported previously for diverse plume-related  
94 oceanic island basalts (OIBs) associated with the two LLSVPs. The large variations in the estimated

H<sub>2</sub>O contents and H<sub>2</sub>O/Ce ratios in the mantle sources of OIBs imply that recycled materials advected into the LLSVPs might have experienced dehydration to various degrees during their subduction or storage in the deep mantle. Higher H<sub>2</sub>O contents and H<sub>2</sub>O/Ce ratios in the African LLSVP than in the Pacific LLSVP are suggestive of distinct formation or evolution histories for these LLSVPs.

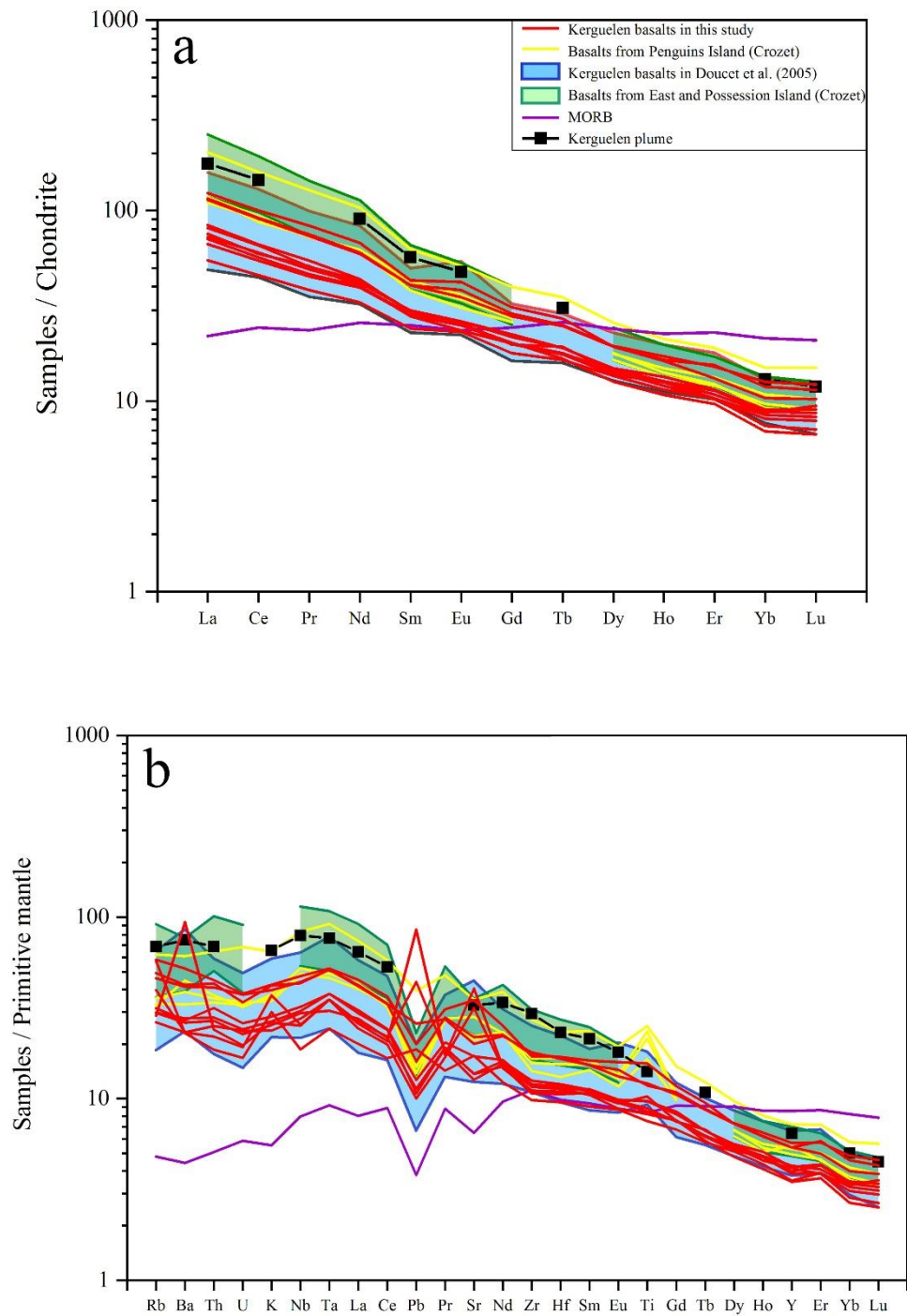
## **2. Geological background**

The Kerguelen mantle plume has produced magmatic activity since ~130 Ma, generating the Kerguelen-Broken Ridge large igneous province and the Ninetyeast Ridge volcanic track with a length of around 5000 km (Coffin et al., 2002; Frey et al., 2000a). Doucet et al. (2006) analyzed He and Ne isotopic compositions of olivine phenocrysts from picrites and high-MgO basalts collected from the same regions and having the same eruption ages as the basaltic samples investigated in this study. They confirmed that the Kerguelen mantle plume entrained primitive material from the deep mantle ( $R/R_a = 8-18$ ; average  $^{21}\text{Ne}/^{22}\text{Ne} = 0.044$  after correction for air contamination). Basalts of the Crozet mantle plume have resulted in a series of islands located in the Indian Ocean. Breton et al. (2013) reported major and trace element data and Sr-Nd-Pb-He isotopic compositions for Crozet plume-related basalts and demonstrated that the plume source is heterogeneous in composition and represents a mixture of two or three endmembers. Four samples from the Penguins Island with primitive He isotopic compositions ( $R/R_a = 13.19-13.98$ ; Breton et al., 2013) were selected for this study. Geophysical observations indicated that the Kerguelen and Crozet mantle plumes are rooted in the African LLSVP (French & Romanowicz, 2015; Montelli et al., 2006). More detailed description of the geological background and petrological features of the studied samples are given in the supporting information.

### 3. Analytical methods

Chemical compositions of bulk-rock samples were obtained using X-Ray fluorescence spectrometry and inductively coupled plasma mass spectrometry at ALS Chemex Co., Ltd. (Guangzhou, China). Fourier transform infrared spectroscopy (FTIR) was applied to measure H<sub>2</sub>O contents of clinopyroxene (Cpx) phenocrysts under unpolarized radiation by a Nicolet iS50 FTIR attached to a Continuum microscope at the School of Earth Sciences, Zhejiang University, China. The compositions of olivine and Cpx phenocrysts were analyzed by a Shimadzu electron probe microanalyzer (EPMA 1720) at the School of Earth Sciences, Zhejiang University, China. The concentrations of trace elements in Cpx phenocrysts were analyzed by the laser ablation inductively coupled plasma mass spectrometry (LA-ICPMS) at the School of Earth Sciences, Zhejiang University, China. Sr-Nd-Pb isotopic analyses were conducted using a multi-collector inductively coupled plasma mass spectrometer (MC-ICP-MS; Thermo Fisher Scientific Neptune Plus) at Hokkaido University, Japan. More details of the used analytical methods are reported in the supporting information.

### 4. Results



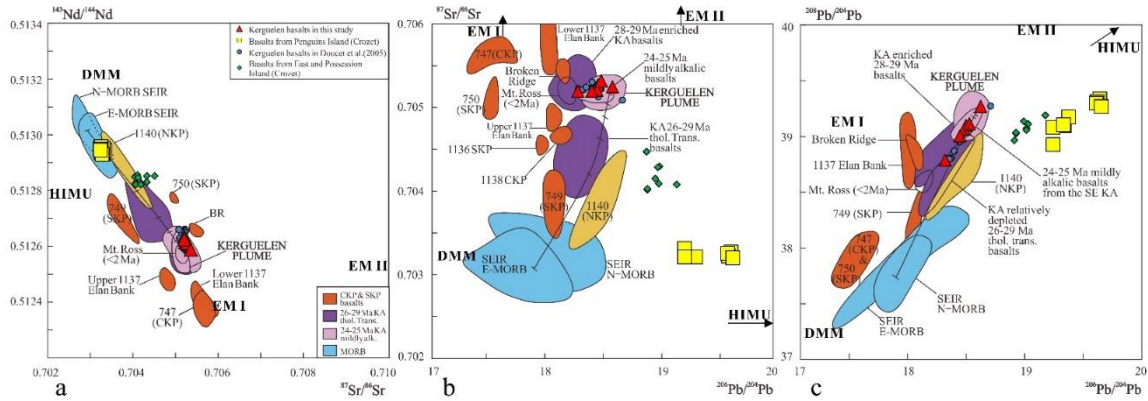
133 **Figure 2 Chondrite-normalized rare earth element (REE) and primitive mantle-normalized trace element**  
134 **patterns for Crozet and Kerguelen basalts.** The trace element compositions of the Crozet basalts are from Breton  
135 et al. (2013). Trace element compositions of high-MgO basalts and picrites studied by Doucet et al. (2005) are also

shown for comparison. The average trace element compositions of global MORB and basalts representing direct melting products of the Kerguelen plume are from Gale et al. (2013) and Frey et al. (2000b), respectively. Normalization values after McDonough and Sun (1995).

Major and minor element and Sr-Nd-Pb isotopic compositions of representative bulk-rock samples are shown in Tables S1-3. The selected samples display a porphyritic texture, with the phenocrysts being mainly olivine and Cpx. All the samples have loss-on-ignition values lower than 2.4 wt.% and MgO contents ranging from 7.1 to 20.0 wt.% (Table S1). The samples with high MgO contents have experienced olivine and/or Cpx accumulation as evidenced by the relationship between other major elements and MgO (Figure S2). All the samples show significant light rare earth elements (LREE) and large ion lithophile elements (LILE) enrichment and no negative Nb-Ta anomalies, as revealed by primitive mantle-normalized trace element patterns (Figures 2a, b). Their Nb/U ratios fall in the range of worldwide mid-ocean ridge basalts (MORBs) and OIBs, whereas several Kerguelen samples have lower Ce/Pb ratios, probably due to surficial alteration influencing the Pb concentrations in these samples (Figures S3a, b).

The Sr-Nd isotopic compositions of the studied Kerguelen basalts resemble those reported in previous studies, showing little variations and falling close to the EMI endmember (Figure 3a; White, 2015). For the Pb isotopic compositions, the studied samples form a trend mostly overlapping with earlier Kerguelen samples, but extending towards lower Pb isotopic ratios (Figure 3c). Similar trends have been ascribed by Doucet et al. (2005) to intrinsic compositional heterogeneity in the Kerguelen mantle plume. The Sr-Nd-Pb isotopic compositions of the Crozet basalts form two groups plotting away from the Kerguelen field towards the MORB fields in the Sr-Nd isotope diagram (Figure 3a) whereas in the Pb-Pb isotope diagram, they plot away from other fields due to their high

$^{206}\text{Pb}/^{204}\text{Pb}$  ratios. The samples selected for  $\text{H}_2\text{O}$  analysis from Crozet follow a trend extending towards the HIMU endmember (Figures 3a, b, c; Breton et al., 2013).



**Figure 3 Sr-Nd-Pb isotopic compositions of basalts from the Crozet and Kerguelen Archipelagos.** The data of Crozet basalts (including from Penguins, East and Possession Islands) are adopted from Breton et al. (2013). The Crozet basalts analyzed for  $\text{H}_2\text{O}$  in this study were sampled from the Penguins Island, and their sources have a contribution from the HIMU component as indicated by the trend to high  $^{206}\text{Pb}/^{204}\text{Pb}$  ratios (b and c). The fields of SEIR MORBs and different basalts related to the Kerguelen plume, including those from the Southern Kerguelen plateau (SKP), Central Kerguelen Plateau (CKP), Broken Ridge (BR), Elan Bank, NKP, etc., are based on data from Doucet et al. (2005) and references therein. The basalts in the Kerguelen Archipelago display temporally compositional variations, with basalts with an age higher than 25 Ma being more depleted due to contamination with SEIR MORB. The fields representing isotopic compositions of the Kerguelen mantle plume are defined by the compositions of basalts from Mt. Crozier (Doucet et al., 2005; Weis et al., 2002).

#### 4.2. $\text{H}_2\text{O}$ contents in Cpx phenocrysts and equilibrated melts for Kerguelen and Crozet basalts

Measured  $\text{H}_2\text{O}$  contents in Cpx phenocrysts and calculated  $\text{H}_2\text{O}$  contents in melts equilibrated with individual Cpx phenocrysts, together with major element compositions of Cpx phenocrysts, are reported in Dataset S1.

##### 4.2.1. Petrographic and geochemical characteristics of Cpx phenocrysts

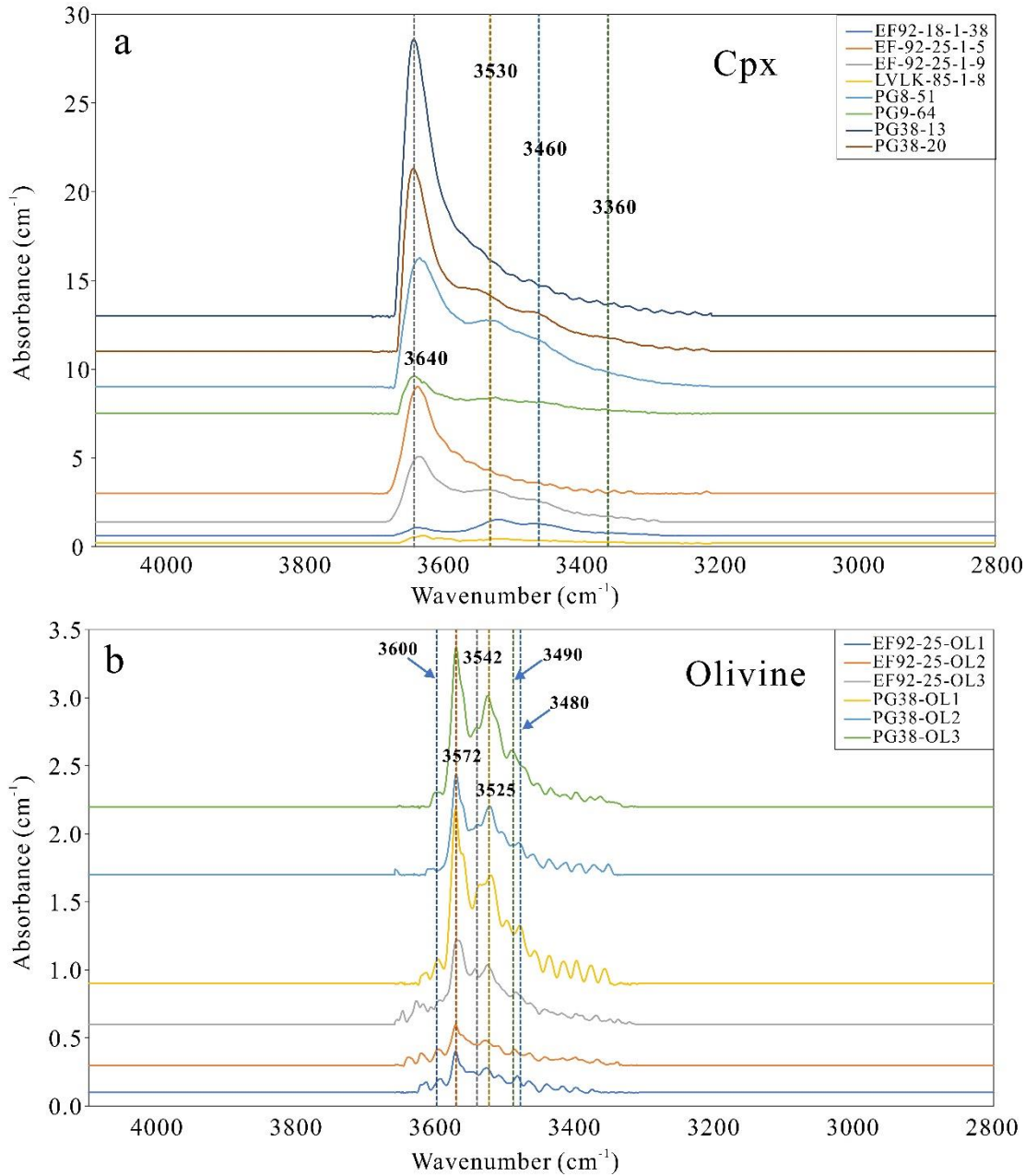
Most of the Cpx phenocrysts are euhedral and no major element diffusion zonation has been observed in the backscattered electron images (Figure S4). The analyzed Cpx phenocrysts have  $\text{Mg}\#$

(=molar  $100 \times \text{Mg}/(\text{Mg}+\text{Fe})$ ) values ranging from 74.1 to 88.7 and from 74.3 to 88.3 for the Kerguelen and Crozet basalts, respectively (Dataset S1). As mentioned in Section 4.1 and the precious studies (Doucet et al., 2005, 2006; Breton et al., 2013), some samples with high MgO contents should have experienced the olivine and/or Cpx accumulation, and the Cpx phenocrysts in several samples display Fe-Mg disequilibrium with the bulk rocks (Table S1 and Dataset S1). However, petrographic and geochemical characteristics suggest that these Cpx phenocrysts should have crystallized from the early batches of melts with similar compositions to the host magmas. Firstly,  $\text{SiO}_2$ ,  $\text{TiO}_2$  and  $\text{Al}_2\text{O}_3$  contents display single co-relationship with Mg# (Figure S5). Secondly, different Cpx phenocrysts in a single sample show parallel REE patterns (Figure S6), which could be explained by the fractionation of magma. Thirdly, most of the Cpx are euhedral, without rounded or resorbed texture and obvious compositional zonation. We selected suitable Cpx phenocrysts that have crystallized from melts with similar characteristics with the host magmas, to retrieve the water content information of primary melts and sources for the Kerguelen and Crozet basalts.

#### **4.2.2. Water contents in Cpx phenocrysts and their equilibrated melts**

The infrared spectra of Cpx phenocrysts display typical structural OH absorption bands similar to mantle Cpx (Figure 4a).  $\text{H}_2\text{O}$  contents of Cpx phenocrysts were determined ranging from 74 to 89 and from 7 to 858 ppm, respectively (Dataset S1). The uncertainty for a single analysis by unpolarized FTIR was estimated to be less than 30% (Liu et al., 2015; Xia et al., 2013).  $\text{H}_2\text{O}$  partition coefficients between Cpx and melt were calculated from major element compositions of Cpx. These were combined with the obtained  $\text{H}_2\text{O}$  contents in Cpx to estimate  $\text{H}_2\text{O}$  contents in the melts in equilibrium with Cpx, yielding values ranging from 0.07 to 7.74 wt.% (Dataset S1). Taking into

account the errors in the H<sub>2</sub>O content measurements and calculation of the H<sub>2</sub>O partition coefficient values, we estimated that the uncertainty is about 32% for the estimation of the H<sub>2</sub>O content in the melts equilibrated with Cpx phenocrysts (Liu et al., 2015).



**Figure 4 Representative infrared spectra for Cpx (a) and olivine (b) phenocrysts in Kerguelen and Crozet basalts.** The spectra display main peaks of structural OH absorption at wavenumbers similar to those previously observed for mantle olivine and Cpx (marked by the dotted lines; e.g., Bell et al., 1995, 2003). The spectra are shown after baseline correction and absorbance normalization to the thickness of 1 cm.

#### 4.3. Estimation of H<sub>2</sub>O contents in the primary melts of Kerguelen and Crozet basalts

The H<sub>2</sub>O contents in Cpx phenocrysts from the Kerguelen and Crozet basalts vary largely, which do not show any co-variation with Mg# of Cpx phenocrysts (Dataset S1; Figures S8a, b). This indicates that many Cpx phenocrysts in Kerguelen and Crozet basalts have more or less lost some of their initial H<sub>2</sub>O due to decompression-induced degassing or diffusive loss during their ascent to the surface, which have been proposed to occur in Cpx from some island arc basalts and OIBs (Lloyd et al., 2016; Weis et al., 2015). Nevertheless, we still observed signatures suggesting that those Cpx phenocrysts with high water content likely did not experience significant overprinting, which can be applied to retrieve the lower bound of the H<sub>2</sub>O contents in corresponding primary melts.

In the Kerguelen samples, some Cpx phenocrysts from sample EF92-25 deviate from the overall trend and have outstandingly high H<sub>2</sub>O contents (e.g., >150 ppm; Figure S8a; Dataset S1). No observable differences in the OH band positions, band widths or peak heights are present in the spectra obtained by FTIR profile analyses for cores and rims of these Cpx phenocrysts. When the H<sub>2</sub>O partition coefficient ( $D_{\text{Cpx/olivine}}$ ) between Cpx and olivine in the basaltic system is considered to be 12.5 (Aubaud et al., 2004), olivine phenocrysts with 12.3-32.7 ppm H<sub>2</sub>O in this sample (Figure 4b; Dataset S1) is close to equilibrium with those Cpx phenocrysts with high water contents (>150 ppm). Because hydrogen diffusivity in olivine is one to two orders of magnitude faster than that in Cpx under high temperatures (e.g., >1250 °C; Tian et al., 2017; Xu et al., 2019), the equilibrium distribution of water in Cpx and olivine indicates that these H<sub>2</sub>O-enriched Cpx phenocrysts should have preserved their initial H<sub>2</sub>O contents during the eruption and cooling process.

The similar observations to the Kerguelen case are also present for Crozet samples. Cpx phenocrysts in sample OVP38 have high H<sub>2</sub>O contents at given high Mg# (311-858 ppm for Cpx phenocrysts with Mg# from 81.9 to 88.3, respectively; Dataset S1; Figure S8b). Several lines of evidence support

that the Cpx phenocrysts in this sample have preserved their initial H<sub>2</sub>O contents: the profile analyses unraveled no core-to-rim H<sub>2</sub>O content variations; H<sub>2</sub>O contents correlates positively with Al<sub>2</sub>O<sub>3</sub> and TiO<sub>2</sub> contents in Cpx phenocrysts (Figures S9a, b). Moreover, olivine phenocrysts in this sample have H<sub>2</sub>O concentrations of 22.5-62.3 ppm (Fig. 4b; Dataset S1), rightly in equilibrium with the coexisting Cpx phenocrysts. In several olivine phenocrysts, the FTIR profile analyses revealed that the H<sub>2</sub>O contents display homogeneous plateaus and steep decrease when approaching grain boundaries (Figure S10).

The high water contents coexisting in the cores of Cpx and olivine phenocrysts would be explained by the assimilation of some altered/hydrated crusts in the deep magma chamber before the eruption. However, the geochemical characteristics of the Cpx phenocrysts do not support this possibility. As shown in Figure S5, the Cpx phenocrysts with both high and low water contents show parallel REE patterns which is consistent with the simple fractional crystallization trend from batches of magmas originating from the same primary melt. In addition, if the measured high water content had been brought by the assimilated altered crust, Ce/Pb ratios should decrease in accompany with the increase of water contents in the Cpx phenocrysts. There is no clear trend between the Ce/Pb ratio and water content in Cpx phenocrysts (Figure S9c). Finally, the calculated water content in the equilibrated melt is as high as ~7.7 wt.%, requiring the exotic contaminated materials have much higher water contents, which is hard to meet for the usual alteration process in the oceanic crust. Overall, we suggest that the high water contents of those Cpx phenocrysts could be used to infer water content of the primary melt.

The average H<sub>2</sub>O contents in the Cpx phenocrysts that have preserved their initial H<sub>2</sub>O are  $220.8 \pm 33.6$  ppm and  $519.6 \pm 117.2$  ppm for the Kerguelen and Crozet samples, respectively, while the

average H<sub>2</sub>O contents in equilibrated melts are 2.02 wt.% and 3.97 wt.%. To decrease the errors of the estimated H<sub>2</sub>O contents in primary melts, only Cpx phenocrysts with Mg# higher than 85 were taken into consideration for the estimation of H<sub>2</sub>O contents in primary melts. The compositions of primary melts were established by incremental addition of olivine to the bulk-rock compositions until Mg# of equilibrated olivine equaled 90.0. Based on the average amount of olivine addition, the H<sub>2</sub>O contents in the primary melts of the Kerguelen and Crozet basalts can be estimated to be  $1.68 \pm 0.54$  wt.% and  $3.69 \pm 1.18$  wt.%, respectively. The detailed calculation procedure is described in the supporting information.

## **5. Discussion**

### **5.1. Potential contamination by depleted upper mantle material?**

During ascending to the surface, basalts might undergo magma mixing and contamination with adjacent rocks. These processes could cause their compositions to deviate from the primary melts and thus make the estimation of the compositional nature of their sources more complex. To access the geochemical characteristics of mantle plumes, basalts representing direct melting products of plumes should be explored. The previous study has demonstrated that the Crozet basalts selected for this study have not experienced magma mixing and contamination by the Indian oceanic lithosphere (Breton et al., 2013) and can be used to evaluate the H<sub>2</sub>O content in the Crozet mantle plume.

No evidence has been found to support the presence of a continental component in Cenozoic basalts from the Northern Kerguelen Plateau (NKP), including the Kerguelen Archipelago (e.g., Doucet et al., 2005; Weis & Frey, 2002). However, the compositions of submarine basalts from the NKP and

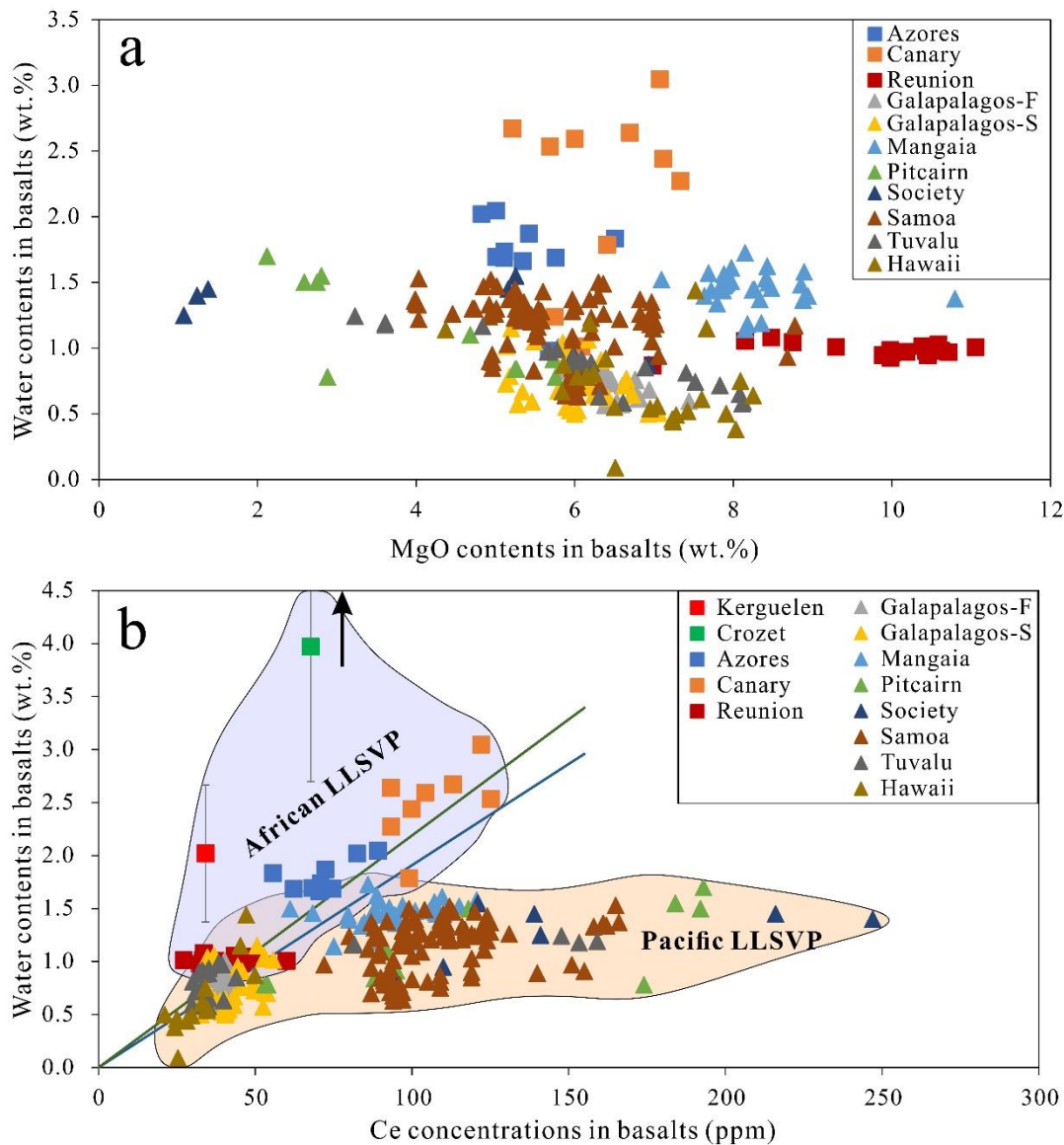
older (>25 Ma) basalts in the Kerguelen Archipelago display a depletion trend towards the Southeast Indian Ridge (SEIR) MORB. This has been interpreted as the result of the interaction between the Kerguelen mantle plume and the SEIR MORB source mantle (Doucet et al., 2005; Frey et al., 2000a; Weis & Frey, 2002). As the SEIR migrated far away from the plume (>400 km) after 25 Ma, contamination with depleted upper mantle material is not reflected in the compositions of Kerguelen basalts, but they exclusively represent geochemical characteristics of the Kerguelen mantle plume itself (Doucet et al., 2005). For the studied samples, the primitive mantle-normalized trace element patterns are similar to those of the inferred enriched component of the Kerguelen plume (Figure 2). Doucet et al. (2005) observed that the Nb/Zr ratio can be used as a good indicator to distinguish basalts contaminated with depleted upper mantle ( $\leq 0.1$ ) from those produced by direct melting of the Kerguelen mantle plume ( $\approx 0.15$ ; Figure S11). The Nb/Zr ratios of  $\sim 0.12$ - $0.17$  in our samples suggest that the compositions of these basalts have not been influenced by materials from the depleted upper mantle. Moreover, the Sr-Nd isotopic compositions of the studied samples exhibit little variability and fall in the range of basalts from Mt. Crozier, which have been proposed to be the best representatives of the compositions of the Kerguelen mantle plume (Figure 3a; Weis et al., 2002). In summary, trace element and Sr-Nd-Pb isotopic compositions indicate that, excluding fractional crystallization, the samples studied here can be regarded as unmodified melting products of the Kerguelen mantle plume.

## **5.2. Comparison with previous estimations of H<sub>2</sub>O contents in the Kerguelen basalts**

H<sub>2</sub>O contents in basalts originally related to the Kerguelen mantle plume have previously been estimated using unaltered glasses recovered from submarine pillow basalts from the NKP (Wallace, 2002). These estimates (0.25-0.69 wt.%) are much lower than that ( $1.68 \pm 0.54$  wt.%) obtained in

294 this study for the primary melts of the Kerguelen mantle plume. It is worth noting that trace element  
295 and Sr-Nd-Pb isotopic compositions of glasses indicate mixing of components from the depleted  
296 SEIR MORB source mantle into the Kerguelen mantle plume (Weis & Frey, 2002). Moreover, the  
297 measured H<sub>2</sub>O contents increase from 0.25-0.27 wt.% in glasses with 90-99% of the depleted  
298 component to 0.69 wt.% in enriched glasses with 63-76% of the depleted component (Wallace, 2002;  
299 Weis & Frey, 2002). This means that the Kerguelen mantle plume, as an enriched endmember,  
300 should contain more H<sub>2</sub>O and produce primary melts with higher H<sub>2</sub>O contents than the enriched  
301 glasses analyzed by Wallace (2002). Based on our estimation, a binary mixing model can simulate  
302 the H<sub>2</sub>O contents in the enriched glasses from the NKP. This model involves a SEIR MORB-like  
303 melt and a primary melt from the Kerguelen mantle plume as the two end-members and yields  
304 mixing proportions similar to those obtained from Sr-Nd isotopic compositions (Figure S12; Weis  
305 & Frey, 2002). Thus, our estimation from Cpx phenocrysts should be reliable in assessing the H<sub>2</sub>O  
306 content in the primary melt of the Kerguelen mantle plume.

307 No previous studies have been conducted to probe the H<sub>2</sub>O contents of the Crozet basalts. Otherwise,  
308 the interaction of the Crozet plume with the segment of the Southwest Indian Ridge between the  
309 Indomed and Gallieni transform faults has been confirmed by geophysical and geochemical studies  
310 (Breton et al., 2013; Sauter et al., 2009). H<sub>2</sub>O contents in the primary melts of basalts sampled from  
311 that segment were estimated to be  $1.3 \pm 0.3$  wt.% and the H<sub>2</sub>O/Ce ratios are higher than 1670, being  
312 much higher than H<sub>2</sub>O/Ce in normal MORBs (Li et al., 2017). This observation indicates that the  
313 mantle source of the Crozet plume is enriched in H<sub>2</sub>O, which is consistent with our results.



315 **Figure 5 H<sub>2</sub>O abundances of plumed-related OIBs reported in previous studies, plotted against MgO contents**  
316 **(a) and Ce concentrations (b).** H<sub>2</sub>O data of basalts affected by later processes (e.g., decompression-induced  
317 degassing, diffusive H gain, shallow alteration) were omitted. While basalts from the Azores, Canary and Mangaia  
318 hotspots have higher H<sub>2</sub>O for the same MgO content, basalts from other hotspots form an overall trend with negative  
319 correlation between H<sub>2</sub>O and MgO produced by fractional crystallization (a). In panel b, H<sub>2</sub>O contents vary with Ce  
320 concentrations along different trends when the hotspots originally related to the Pacific (labelled by triangles) and  
321 African (labelled by squares) LLSVPs are compared. The DMM-FOZO trend defined by the H<sub>2</sub>O-Ce relations in  
322 global MORBs and FOZO OIBs is shown as black lines (Shimizu et al., 2019). For the Galapalago hotspot, only  
323 H<sub>2</sub>O contents in basalts with <sup>3</sup>He/<sup>4</sup>He >8 were taken into consideration (two groups Fernandina and Sierra Negra,

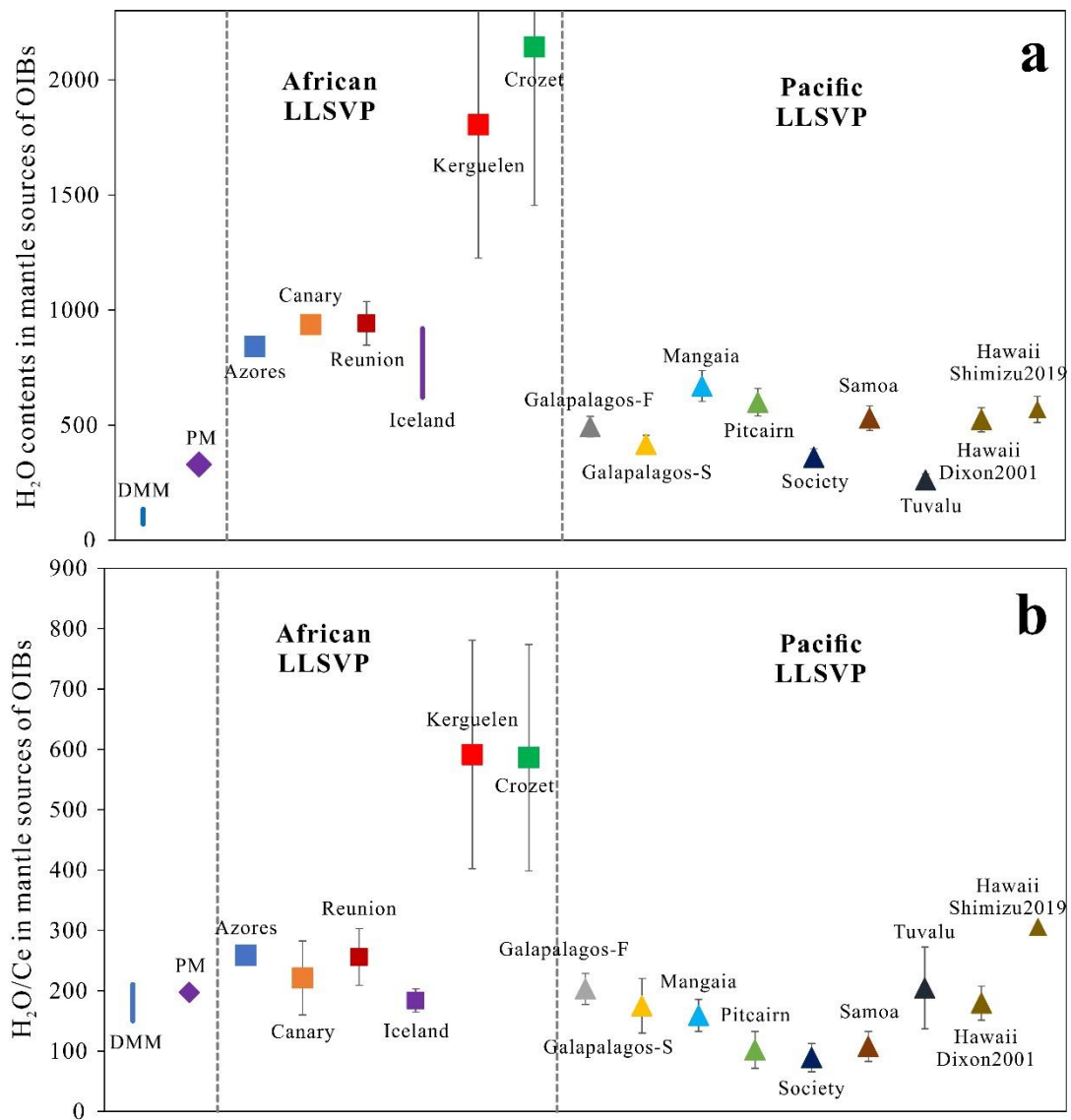
suffixed by 'F' and 'S'; Peterson et al., 2017). The estimated H<sub>2</sub>O and Ce contents in primary melts of the Kerguelen and Crozet basalts are also shown in panel b. Data sources are: Azores (Métrich et al., 2014), Canary (Longpré et al., 2017), Reunion (Walowski et al., 2019), Galapagos (Peterson et al., 2017), Mangaia (Cabral et al., 2014), Pitcairn (Kendrick et al., 2014), Society (Kendrick et al., 2014), Samoa (Workman et al., 2006; Kendrick et al., 2015), Tuvalu (Jackson et al., 2015) and Hawaii (Dixon & Clague, 2001; Shimizu et al., 2019).

We compiled H<sub>2</sub>O content data previously reported for many hotspot-related OIBs. These hotspots have been explicitly evidenced by geophysical observations to overlie mantle plumes and are geographically associated with the two LLSVPs (Dataset S2; Figure 5a; Jackson et al., 2018). Results from some melt inclusions or glasses, disturbed by later processes (decompression-induced degassing, diffusive H gain, surficial alteration, etc.), were not taken into consideration (the detail of data infiltration discussed in the supporting information). It is noteworthy that the melt inclusions and basaltic glasses used in previous studies to explore the H<sub>2</sub>O contents of diverse OIBs have mostly relatively low MgO contents (many lower than 6 wt.%; Dataset S2; Figure 5a). Additionally, the broad negative correlation between H<sub>2</sub>O and MgO contents indicates that the H<sub>2</sub>O inventory in the parental melts of these OIBs have been modified by fractional crystallization processes in variable extents (Figure 5a). To estimate the H<sub>2</sub>O distribution in the LLSVPs, H<sub>2</sub>O contents in primary melts of individual OIBs were roughly retrieved by incremental addition of equilibrium olivine to the bulk-rock compositions. As fractional crystallization in variable extents occurred to melts of OIBs, the compositions of melt inclusions and glasses, with highest MgO contents studied for H<sub>2</sub>O contents of individual OIBs, were applied for the initial compositions for olivine addition. Assuming that H<sub>2</sub>O behaves completely incompatibly during olivine fractionation, the primary H<sub>2</sub>O contents were back calculated (cf. Le Voyer et al., 2015). With the assumption of melting degrees of 5% for alkali OIBs and 10% for tholeiites, we estimated the H<sub>2</sub>O contents in the mantle sources

using the batch and fractional partial melting models (Shaw, 1970). A detailed calculation procedure and results are shown in the supporting information and Table 1. The calculated H<sub>2</sub>O contents in the primary melts and mantle sources display great variations for OIBs associated with a single LLSVP and large differences between OIBs associated with the African and Pacific LLSVPs (Table 1; Figure 6a).

As the transition zone is a significant H<sub>2</sub>O reservoir along the ascent route of deep mantle-rooted plumes, variable amounts of transition zone materials captured in ascending plumes may be a plausible mechanism to induce heterogeneous H<sub>2</sub>O contents in distinct plume-related OIBs. When traversing the transition zone, Archean plumes had higher mantle potential temperatures, causing them to partially melt and potentially take in H<sub>2</sub>O from the transition zone, whereas Phanerozoic plumes may stay solid hampering assimilation of much H<sub>2</sub>O within a short period of time (Sobolev et al., 2016). As most Phanerozoic plumes have similar mantle potential temperatures (consistently lower than Archean plumes; Herzberg & Gazel, 2009; Sobolev et al., 2016), it means that the different H<sub>2</sub>O contents in plume-related OIBs are not caused by entraining variable amounts of materials from the transition zone during plume upwelling.

There is a wide consensus that variable chemical and isotopic compositions of plume-related OIBs reflect notable compositional heterogeneity of the deep mantle (Hofmann, 1997; White, 2015). As accumulative geophysical observations and numerical models propose that most mantle plumes feeding OIBs are rooted in the LLSVPs, the significant variations in H<sub>2</sub>O contents and H<sub>2</sub>O/Ce ratios in primary melts and mantle sources for diverse plume-induced OIBs could indicate heterogeneous distribution of H<sub>2</sub>O in the LLSVPs (Figures 6a, b). Additionally, geophysical observations have also revealed heterogeneity inside individual LLSVPs (e.g., Garnero et al., 2016).



370 **Figure 6** Estimated  $H_2O$  contents (a) and  $H_2O/Ce$  ratios (b) in the mantle sources of OIBs originally related  
371 **to the African and Pacific LLSVPs.**  $H_2O$  contents were estimated from the calculated  $H_2O$  contents of primary  
372 melts using the batch melting model and  $H_2O$  partition coefficient of 0.0085 (Table 1; the calculation details are  
373 shown in the supporting information).  $H_2O/Ce$  ratios in melts can represent those of their mantle sources. The  $H_2O$   
374 contents and  $H_2O/Ce$  ratios in the depleted MORB source mantle (DMM) and primitive mantle (PM) were adopted  
375 from Bizimis and Peslier (2015). The  $H_2O$  content and  $H_2O/Ce$  ratio in the source of basalts originating from the  
376 Iceland plume are adopted from Nichols et al. (2002) and Hartley et al. (2015), respectively. The  $H_2O$  content and  
377  $H_2O/Ce$  ratio in the Hawaii plume estimated from basaltic glasses by Dixon and Clague (2001) and Shimizu et al.  
378 (2019) were shown separately, named as ‘Hawaii Dixon2001’ and ‘Hawaii Shimizu2019’, respectively. The  $H_2O$

content and H<sub>2</sub>O/Ce ratio of the mantle source for ‘Hawaii shimizu2019’ were estimated from the glass with the highest H<sub>2</sub>O content and H<sub>2</sub>O/Ce ratio reported by Shimizu et al. (2019). Shimizu et al. (2019) only analyzed three glasses and their compositions and H<sub>2</sub>O contents display significant variations.

The existence of the Kerguelen and Crozet mantle plumes has been repeatedly demonstrated by diverse geophysical observations (e.g., French & Romanowicz, 2015; Marignier et al., 2020; Montelli et al., 2004, 2006) and is consistent with primitive He isotopic compositions of Cenozoic basalts (Breton et al., 2013; Doucet et al., 2006). Both mantle plumes are deeply rooted at the base of the mantle and the induced hotspots are geographically associated with the African LLSVP (e.g., French & Romanowicz, 2015; Jackson et al., 2018; Montelli et al., 2006). Together with basalts generated by other plumes originating from the LLSVPs, we can catch a glimpse into the H<sub>2</sub>O distribution in the African and Pacific LLSVPs.

As shown in Section 5.3, we have compiled the data of the plume-induced OIBs representing different mantle end-members as many as possible (Dataset S2). For the African LLSVP, the data for St. Helena basalts, compositionally close to the HIMU endmember component, is absent. Only one H<sub>2</sub>O content value for a glass from St. Helena was reported (Kendrick et al., 2017), while its bulk rock MgO content is 2.8 wt.%, indicating that it has too strongly evolved to recover the H<sub>2</sub>O content in the mantle source. The data shows that irrespective of the geochemical affinity to different mantle endmembers, OIBs associated with the African LLSVP form a trend extending to higher H<sub>2</sub>O contents compared with those associated with the Pacific LLSVP (Figure 5b). For the latter, H<sub>2</sub>O contents do not increase with Ce (especially in samples with Ce concentrations higher than 80 ppm; Figure 5b). In comparison, the H<sub>2</sub>O contents in the mantle sources of OIBs from African LLSVP were estimated to range from 620 to 2144 ppm, which are systematically higher than those

estimated for OIBs from the Pacific LLSVP (262-671 ppm; Table 1; Figure 6a; the calculation details are shown in the supporting information). The H<sub>2</sub>O/Ce ratios of the Crozet and Kerguelen basalts are also higher than those of the Pacific LLSVP, the depleted MORB source mantle, and the primitive mantle (Figure 6b). Overall, the H<sub>2</sub>O/Ce ratios of OIBs from the African LLSVP are close to or higher than the upper boundary of the H<sub>2</sub>O/Ce ratio range for the OIBs from the Pacific LLSVP (184-592 versus 89-306; Figure 6b).

Furthermore, remarkably high H<sub>2</sub>O/Ce ratios have been observed in the North (253 ± 33, Dixon et al., 2002) and equatorial (205-405, mostly higher than 250, Le Voyer et al., 2015) Mid-Atlantic MORBs imprinted by mantle plumes, which are rooted in the African LLSVP. The Afar mantle plume, another ‘primary plume’ detected by French and Romanowicz (2015), is also deeply associated with the African LLSVP and underlies the African continent. However, the H<sub>2</sub>O contents of Afar plume-related basalts have never been studied so far, but geophysical observations indicate that the lower mantle source of the Afar plume is enriched in volatiles (Thompson et al., 2015). All these observations indicate that the African LLSVP is more hydrous than the antipodal Pacific LLSVP.

## **5.5. Implications for different formation histories for the LLSVPs**

Geophysical observations and numerical modeling support the view that the LLSVPs are piles of accumulative dense mantle materials, the origin of which has been proposed to be Fe-rich remnants of ancient differentiation in the mantle or continuous accumulation of subducted oceanic crust (Garnero et al., 2016). Lee et al. (2010) suggested that the primordial Fe-rich by-products of Archean differentiation have FOZO-like geochemical characteristics. However, most OIBs geographically

associated with both LLSVPs deviate from the DMM-FOZO trend in the H<sub>2</sub>O versus Ce plot (Figure 5b; Shimizu et al., 2019). Compared with the primitive mantle, the H<sub>2</sub>O content in the mantle sources of all the OIBs and H<sub>2</sub>O/Ce ratios of most OIBs from the African LLSVP are significantly higher, whereas those of the Pacific are mostly lower (Figures 6a, b). The observed features of H<sub>2</sub>O distribution in LLSVPs may not have resulted from ancient differentiation.

High-P-T experiments have shown that hydrous minerals have the ability to transport large amounts of H<sub>2</sub>O by subduction into the bottom of the lower mantle (e.g., Nishi et al., 2014). Subducted slabs may experience variable extents of dehydration during their downwelling into and residence in the lower mantle (e.g., Kendrick et al., 2017; Workman et al., 2006), which was applied to explain the contrasting H<sub>2</sub>O contents in diverse OIBs (e.g., Cabral et al., 2014; Dixon et al., 2002; Kendrick et al., 2017). The distinct H<sub>2</sub>O contents in these two LLSVPs might be indicative of different extents of dehydration in subducted materials incorporated into individual LLSVPs. The dehydration extent of subducting slabs depends on the conditions of the slabs and the subduction environment (e.g., van Keken et al., 2011). Huge amounts of subducted oceanic crust underwent dehydration to different degrees and accumulated at the base of the lower mantle to form the LLSVPs, indicating that the massive subduction responsible for the generation of the African and Pacific LLSVPs may have occurred in different geological periods. White (2015) has proposed that the two LLSVPs may be compositionally distinct, as indicated by the presence of ubiquitous isotopic DUPAL anomalies in basalts from the south Atlantic and Indian Oceans, but not in basalts from the south Pacific Ocean. In addition, the cluster analysis of seismic tomographic models reveals that the African and Pacific LLSVPs have different shapes and vertical extents, which strongly depend on the viscosity of deep mantle materials (Cottaar & Lekic, 2016; Deschamps & Tackley, 2008). Because H<sub>2</sub>O can

significantly affect the rheological properties of mantle minerals, the differences in the H<sub>2</sub>O contents in the two LLSVPs may be an important factor in accounting for their seismic differences.

Our observations, that the two LLSVPs have distinct H<sub>2</sub>O contents and H<sub>2</sub>O/Ce ratios, may imply that the formation and evolution should be different for the African and Pacific LLSVPs. Numerical models show that the formation of LLSVPs can be related to tectonic processes in the last 120 Ma (e.g., McNamara & Zhong, 2005) or super-continental cycles (e.g., Doucet et al., 2020; Li & Zhong, 2009; Zhang et al., 2010). It has been suggested that the structure of the lower mantle beneath Africa may have been generated as the result of mantle upwelling corresponding to the subduction around the supercontinent Pangea (Zhang et al., 2010). In contrast, the thermochemical piles responsible for the Pacific LLSVP could have been emplaced before the assembly of Pangea, perhaps during the lifetime of the supercontinent Rodinia or after the breakup of Rodinia as the upwelling system driving the assembly of Pangea (Doucet et al. 2020; Li & Zhong, 2009).

## **6. Conclusions**

H<sub>2</sub>O contents in hotspots-related basalts fed by mantle plumes, deeply rooted in the African and Pacific LLSVPs, were used to obtain information on the H<sub>2</sub>O distribution in these two LLSVPs. The H<sub>2</sub>O concentrations in the calculated primary melts of Cenozoic basalts, unmodified melting products of the Kerguelen and Crozet mantle plumes, were estimated to be  $1.68 \pm 0.54$  wt.% and  $3.69 \pm 1.18$  wt.%, respectively. Combined with melting models, the H<sub>2</sub>O contents of the Kerguelen and Crozet mantle plumes were estimated to be  $1805 \pm 579$  ppm and  $2144 \pm 690$  ppm for the Kerguelen and Crozet mantle plumes, respectively, giving an indication of the H<sub>2</sub>O distribution inside the African LLSVP. We compiled H<sub>2</sub>O content data from worldwide plume-fed OIBs and

465 back-calculated the H<sub>2</sub>O contents in their primary melts and mantle sources, with the results  
466 displaying large variations among plume-fed OIBs, even for basalts geographically associated with  
467 a single LLSVP. We conclude that the African LLSVP is more enriched in H<sub>2</sub>O and has higher  
468 H<sub>2</sub>O/Ce ratios compared with the Pacific LLSVP. Incorporation of subducted oceanic crust with  
469 variable extents of dehydration during its downwelling can be responsible for the heterogeneous  
470 distribution of H<sub>2</sub>O within individual LLSVPs and the distinct H<sub>2</sub>O contents and H<sub>2</sub>O/Ce ratios  
471 between the African and Pacific LLSVPs. The present observations suggest that these two LLSVPs  
472 have different formation and evolution histories.

|                  |                             | Compositions of OIBs       |         |                |                       |                                  |                      | Primary water<br>content (wt.%)      Errors |      | Water content in mantle source |        |               |              | Water content in mantle source |               |              |
|------------------|-----------------------------|----------------------------|---------|----------------|-----------------------|----------------------------------|----------------------|---|------|--------------------------------|--------|---------------|--------------|--------------------------------|---------------|--------------|
|                  |                             | H <sub>2</sub> O<br>(wt.%) | Errors* | MgO<br>(wt.%)& | Olivine<br>addition\$ | H <sub>2</sub> O/Ce <sup>£</sup> | Errors <sup>\$</sup> |   |      | Batch melting                  |        |               |              | Fractional melting             |               |              |
|                  | OIBs                        |                            |         |                |                       |                                  |                      |   |      | D =<br>0.0085 <sup>#</sup>     | Errors | D =<br>0.0046 | D =<br>0.011 | D =<br>0.0085                  | D =<br>0.0046 | D =<br>0.011 |
| African<br>LLSVP | Azores                      | 1.83                       | 0.18    | 6.51           | 0.21                  | 259.12                           | 33.51                | 1.45  | 0.14 | 841.20                         | 84.53  | 787.53        | 875.60       | 725.97                         | 724.24        | 724.23       |
|                  | Canary                      | 2.27                       | 0.09    | 7.33           | 0.29                  | 221.30                           | 61.20                | 1.61  | 0.06 | 936.70                         | 37.29  | 876.94        | 975.01       | 808.39                         | 806.47        | 806.46       |
|                  | Reunion                     | 1.01                       | 0.10    | 11.06          | 0.13                  | 256.16                           | 47.02                | 0.88  | 0.09 | 942.71                         | 94.51  | 911.97        | 962.41       | 875.72                         | 875.71        | 875.71       |
|                  | Iceland **                  |                            |         |                |                       | 184.00                           | 19.00                |   |      | 620-920                        |        |               |              |                                |               |              |
|                  | Kerguelen &&                | 2.02                       | 0.65    |                | 0.17                  | 591.57                           | 189.30               | 1.68  | 0.54 | 1804.86                        | 579.00 | 1746.01       | 1842.58      | 1676.61                        | 1676.60       | 1676.60      |
|                  | Crozet                      | 3.97                       | 1.27    |                | 0.07                  | 586.22                           | 187.59               | 3.69  | 1.18 | 2144.19                        | 689.51 | 2007.39       | 2231.87      | 1850.48                        | 1846.08       | 1846.05      |
| Pacific LLSVP    | Galapagos_F <sup>\$\$</sup> | 0.60                       | 0.05    | 7.44           | 0.23                  | 203.06                           | 25.62                | 0.46  | 0.04 | 493.20                         | 44.50  | 477.12        | 503.51       | 458.15                         | 458.15        | 458.15       |
|                  | Galapagos_S <sup>\$\$</sup> | 0.50                       | 0.05    | 7.06           | 0.23                  | 175.19                           | 45.26                | 0.39  | 0.03 | 417.77                         | 37.69  | 404.15        | 426.50       | 388.08                         | 388.08        | 388.08       |
|                  | Mangaia                     | 1.37                       | 0.14    | 10.80          | 0.16                  | 159.07                           | 26.41                | 1.15  | 0.12 | 670.61                         | 67.39  | 627.83        | 698.04       | 578.75                         | 577.37        | 577.37       |
|                  | Pitcairn                    | 0.79                       | 0.08    | 5.94           | 0.29                  | 101.98                           | 30.55                | 0.56  | 0.06 | 599.99                         | 60.15  | 580.42        | 612.53       | 557.35                         | 557.35        | 557.35       |
|                  | Society                     | 0.88                       | 0.09    | 6.93           | 0.29                  | 89.19                            | 23.49                | 0.62  | 0.06 | 360.79                         | 36.26  | 337.77        | 375.55       | 311.37                         | 310.63        | 310.63       |
|                  | Samoa                       | 1.17                       | 0.12    | 8.78           | 0.22                  | 107.71                           | 24.72                | 0.91  | 0.09 | 529.99                         | 53.26  | 496.18        | 551.67       | 457.40                         | 456.31        | 456.30       |
|                  | Tuvalu                      | 0.59                       | 0.06    | 8.12           | 0.24                  | 204.64                           | 67.62                | 0.45  | 0.05 | 261.63                         | 26.29  | 244.93        | 272.32       | 225.79                         | 225.25        | 225.25       |
|                  | Hawaii <sup>££</sup>        | 0.64                       | 0.06    | 8.25           | 0.24                  | 179.54                           | 28.16                | 0.49  | 0.05 | 523.61                         | 52.49  | 506.54        | 534.55       | 486.40                         | 486.40        | 486.40       |
|                  | Hawaii <sup>\$\$</sup>      | 1.44                       | 0.14    | 7.52           | 0.32                  | 305.67                           |                      | 0.98  | 0.10 | 567.97                         | 57.08  | 531.74        | 591.20       | 490.17                         | 489.01        | 489.00       |

**Table 1 Estimated water contents in primary melts and mantle sources of global OIBs.** The details about the calculation procedure are shown in the supporting information. The references for the compositions and water contents of individual OIBs are listed in the caption of Figure 5. The data of melt inclusions or glasses with the highest MgO contents reported by literatures for individual hotspot-related OIBs were applied to retrieve the compositions and water contents of corresponding primary melts and sources. \* The errors of water contents for OIBs are estimated

476 from the analytical errors reported in the literatures. <sup>&</sup> The highest MgO contents of melt inclusions or glasses adopted from referenced literatures for individual OIBs. <sup>‡</sup> We use the average H<sub>2</sub>O/Ce  
477 ratios for all the melt inclusions or glasses (preserving initial information) in our referenced literatures to represent the H<sub>2</sub>O/Ce ratios of corresponding primary melts. <sup>§</sup> The proportions of olivine  
478 addition to the highest-MgO OIB compositions for primary melt compositions. <sup>§</sup> The standard deviations for averaging H<sub>2</sub>O/Ce ratios of all the melt inclusions or glasses. <sup>#</sup> D is the partition  
479 coefficient of H<sub>2</sub>O between the source and primary melt. The choose of partition coefficients is discussed in detail in the supporting information. <sup>\*\*</sup> The H<sub>2</sub>O content and H<sub>2</sub>O/Ce ratio in the source  
480 of the Iceland hotspot are not calculated in this study, and are just adopted from Nichols et al. (2002) and Hartley et al. (2015). <sup>&&</sup> Water contents in Kerguelen and Crozet basalts are estimated  
481 for melts equilibrated with Cpx phenocrysts (Mg# higher than 85). The errors are estimated from the uncertainty of 32% for the whole calculation procedure. Ce concentrations are from the bulk  
482 rock measurements. <sup>§§</sup> For the Galapalago hotspot, only H<sub>2</sub>O contents in basalts with <sup>3</sup>He/<sup>4</sup>He >8 were taken into consideration (two groups Fernandina and Sierra Negra, suffixed by ‘F’ and ‘S’;  
483 Peterson et al., 2017). <sup>‡‡</sup> The H<sub>2</sub>O content and H<sub>2</sub>O/Ce ratio in the Hawaii plume were estimated from basaltic glasses studied by Dixon and Clague (2001). <sup>§§</sup> The H<sub>2</sub>O content and H<sub>2</sub>O/Ce ratio  
484 in the Hawaii plume were estimated from basaltic glasses studied by Shimizu et al. (2019).

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