



# Artificial Small-scale Field-aligned Irregularities in the High Latitude Ionospheric F-region

*N.F. Blagoveshchenskaya<sup>1</sup>,*

*T.D. Borisova<sup>1</sup>, A.S. Kalishin<sup>1</sup>, T.K. Yeoman<sup>2</sup>*

*<sup>1</sup> Arctic and Antarctic Research Institute, St. Petersburg, Russia,*

*<sup>2</sup> Department of Physics and Astronomy, University of Leicester, UK,*

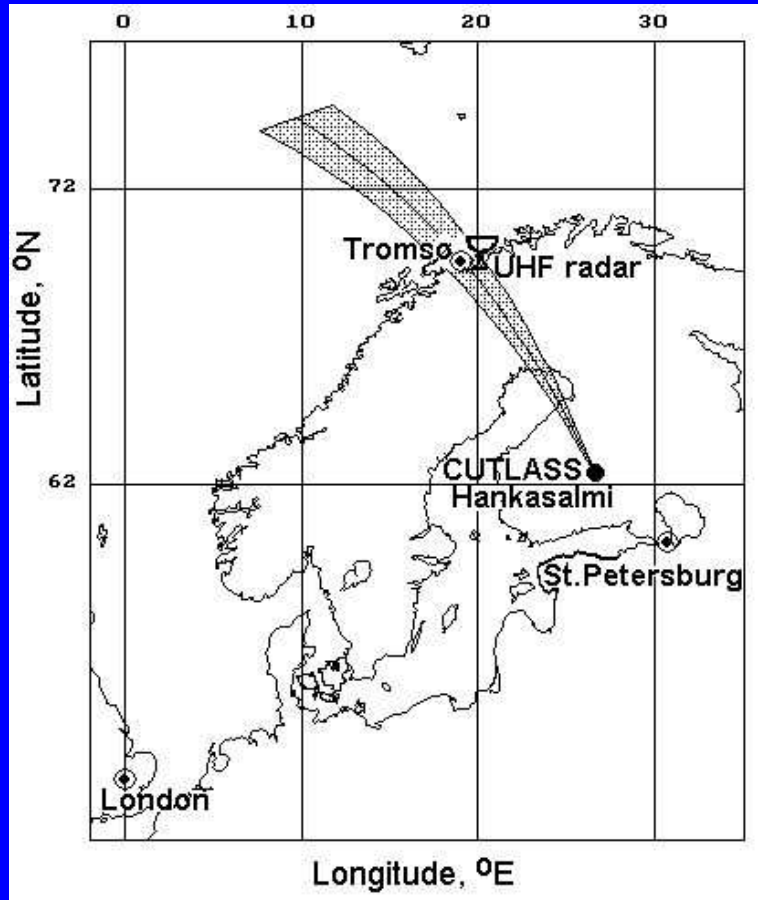
**AGU Fall Meeting**

**10 – 14 December, 2018, Washington, DC, USA**

**SA11C-2745**

The ionospheric modification experiments are providing the means for understanding mechanisms and physical processes leading to the generation and evolution of ionospheric irregularities. We present experimental results concentrating on the features and evolution, generation conditions and mechanisms of small-scale field-aligned irregularities (FAIs), with spatial scale across the geomagnetic field of 7.5 – 15 m, in the high latitude ionosphere F region induced by the controlled injection of the powerful HF radio waves from the ground into the ionosphere. The behavior and properties of FAIs excited by the ordinary polarized (O-mode) HF pump waves are well known (see, for example, Robinson, 1989; Gurevich, 2007 and references therein). The main attention is paid to the recently discovered FAIs induced by the extraordinary polarized (X-mode) HF pump wave (Blagoveshchenskaya et al., 2011; 2014; 2015) and their comparison with the O-mode FAIs.

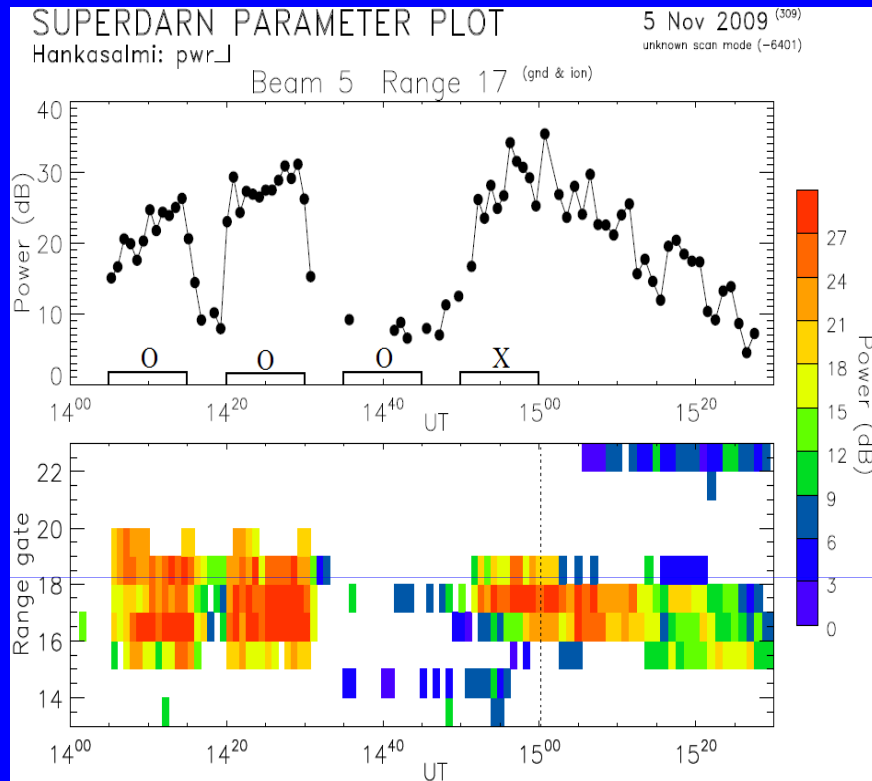
# Instrumentation



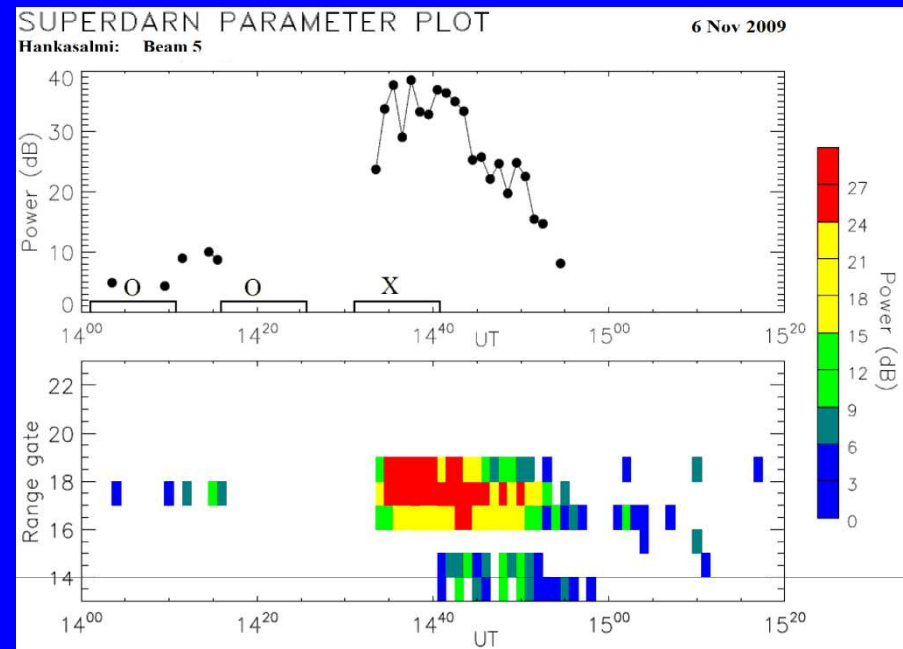
A map showing the experiment geometry. The EISCAT/Heating facility at Tromsø was used for HF ionospheric modification of F-region. HF heating facility was operating at heater frequencies of 4.0 – 8.0 MHz with an effective radiated power of 100 – 750 MW. HF pump wave with ordinary (O-mode) or extraordinary (X-mode) polarization was injected along the magnetic field line.

Instrument diagnostics included the CUTLASS (Co-operative UK Twin Located Auroral Sounding System) radar, the European Incoherent Scatter (EISCAT) UHF radar at 931 MHz near Tromsø and the EISCAT ionosonde (dynasonde).

# Artificial small-scale field-aligned irregularities ( $f_H > f_oF2$ )

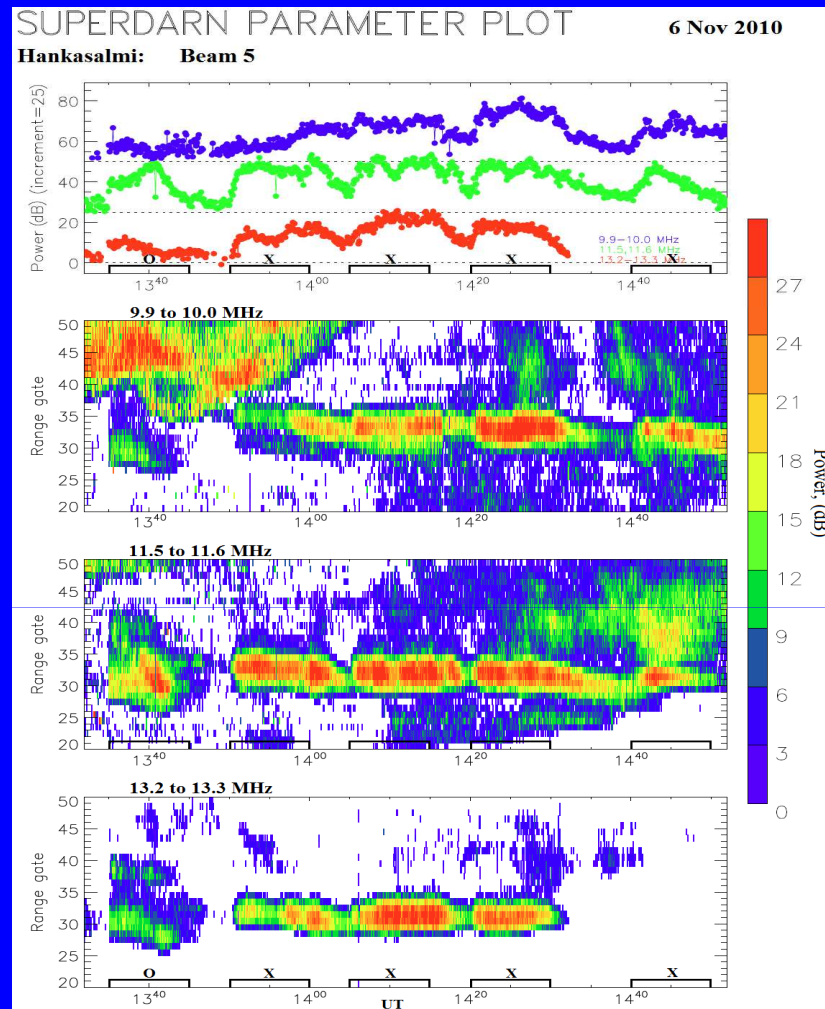


Backscattered power from the Hankasalmi (Finland) CUTLASS radar (beam 5) at  $f \approx 10$  MHz for alternating O/X-mode heating at  $f_H = 4.040$  MHz (ERP= 123 MW) on 5 November, 2009.

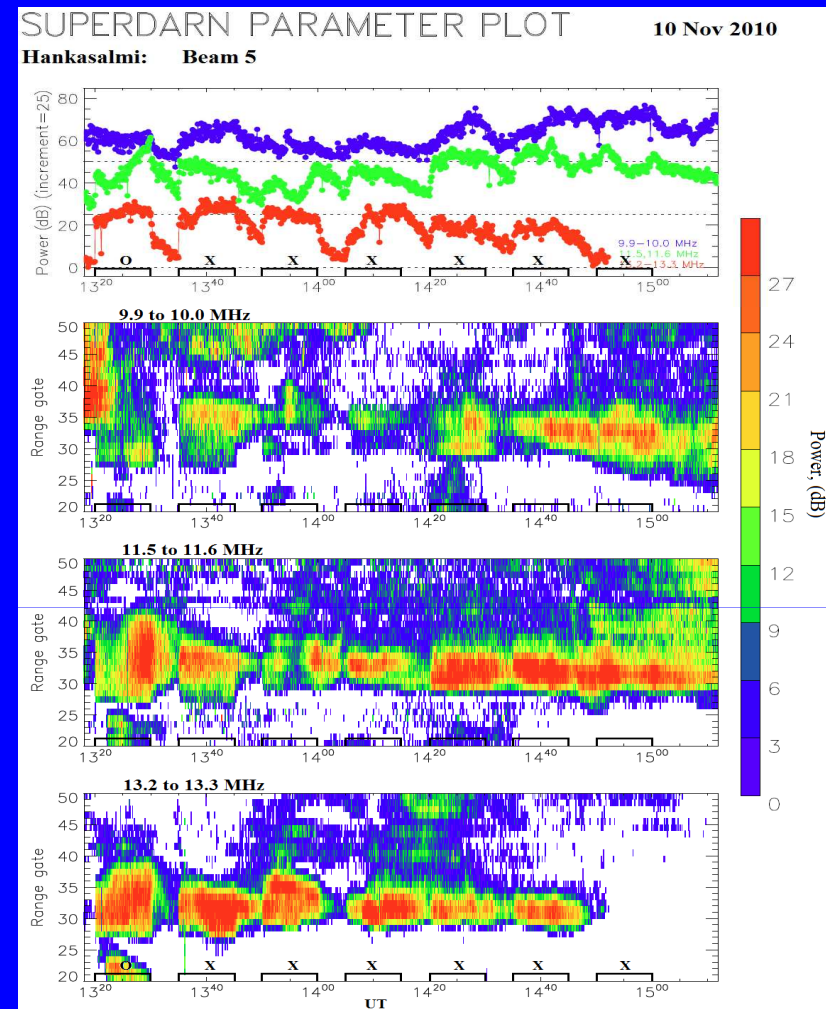


Backscattered power from the Hankasalmi (Finland) CUTLASS radar (beam 5) at  $f \approx 10$  MHz for alternating O/X-mode heating on 6 November, 2009. Power stepping mode of the Tromsø HF heating facility, utilizing an orderly sequence of 20%, 50%, 70%, 85%, 100%, 100%, 85%, 70%, 50%, 20% (from 30 to 157 MW and back) was used. The duration of each step is 1 min. The HF pump wave was radiated at  $f_H = 4.040$  MHz towards magnetic zenith.

# Artificial small-scale field-aligned irregularities ( $f_H > f_oF2$ )



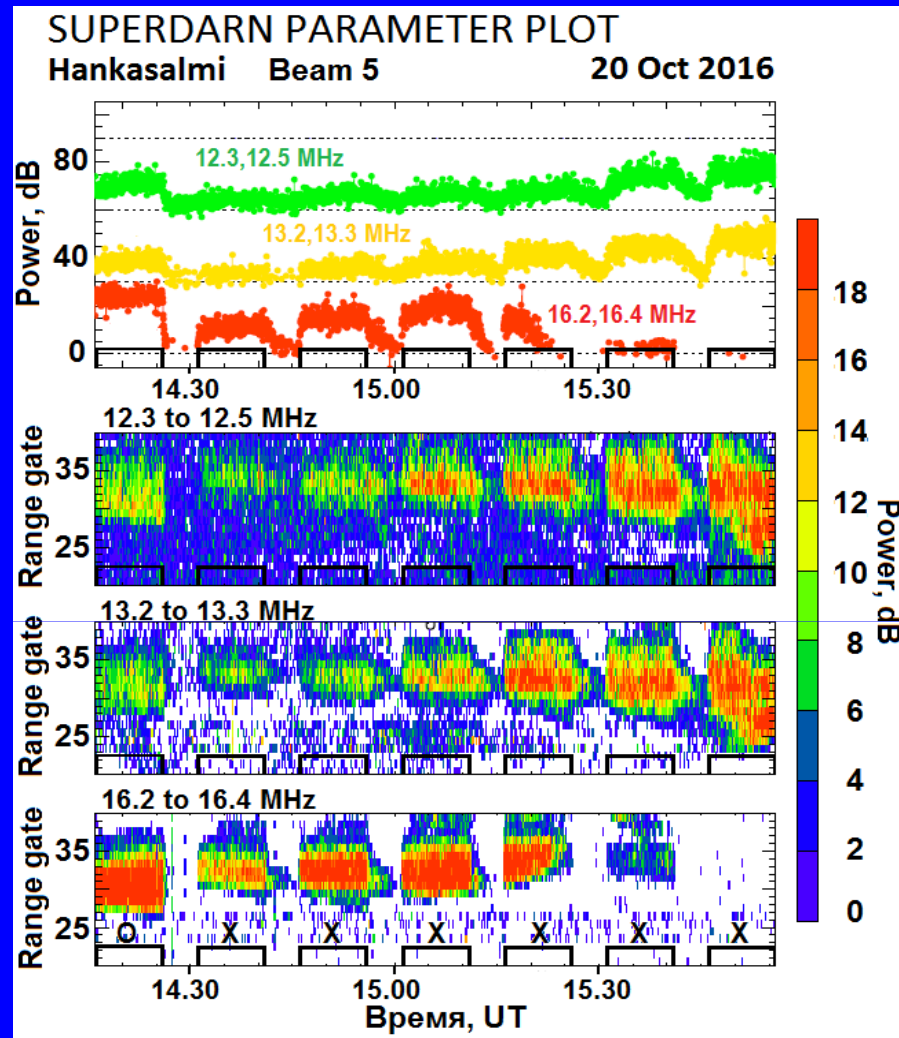
Backscattered power from the Hankasalmi (Finland) CUTLASS radar (beam 5) at  $f \approx 10$ , 11.5, and 13 MHz for contrasting O/X-mode heating at  $f_H = 4.544$  MHz (ERP=95 MW) on 6 November, 2010.



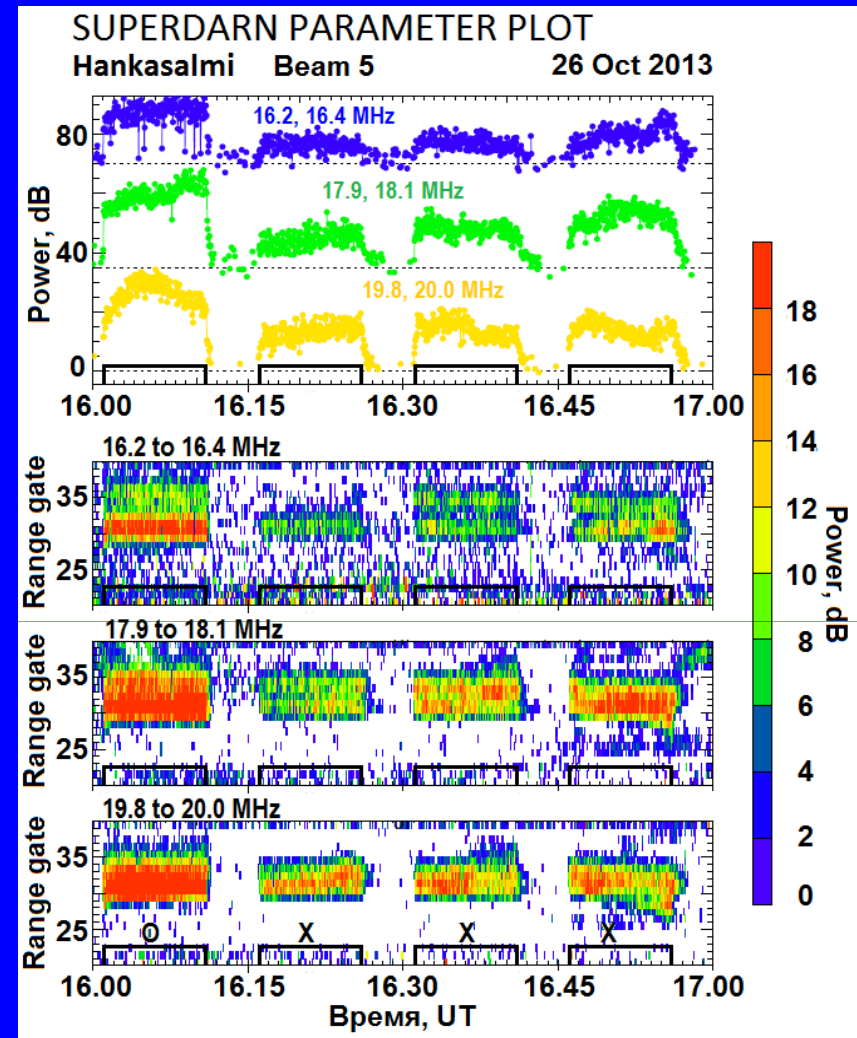
Backscattered power from the Hankasalmi (Finland) CUTLASS radar (beam 5) at  $f \approx 10$ , 11.5, and 13 MHz for contrasting O/X-mode heating at  $f_H = 4.544$  MHz (ERP=113 MW) on 10 November, 2010.



# Artificial small-scale field-aligned irregularities ( $f_H \leq f_oF2$ )

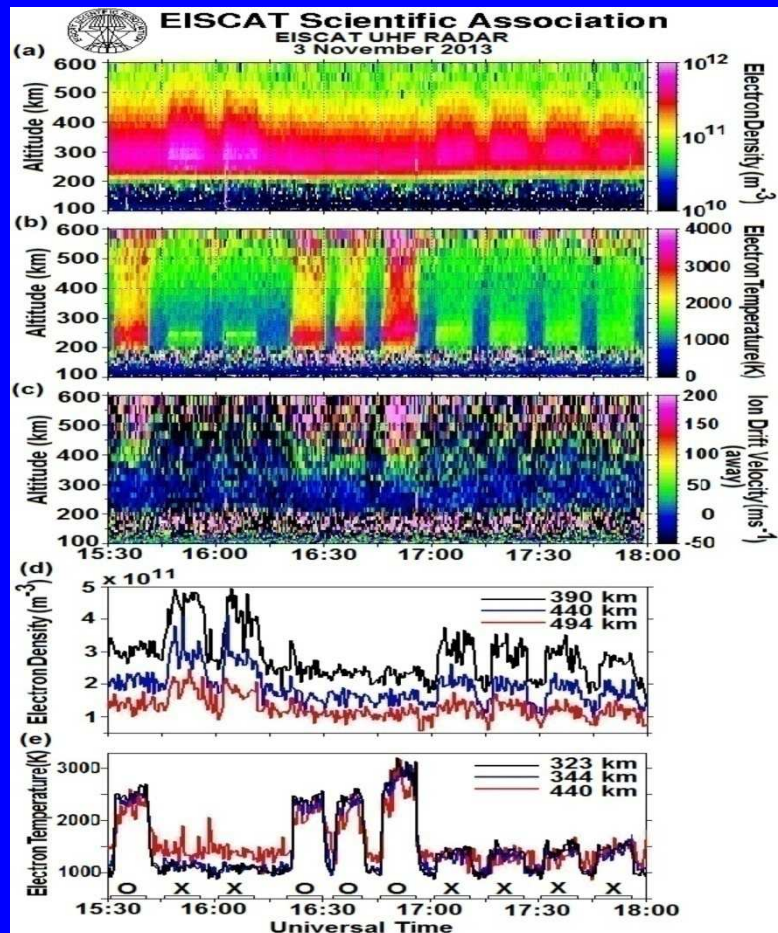


Backscattered power from the Hankasalmi (Finland) CUTLASS radar (beam 5) at  $f \approx 12, 13$ , and  $16$  MHz for O/X-mode heating at  $f_H = 4.544$  MHz (ERP=130 MW) on 20 October, 2016. Heater-on cycles and polarization is shown on the time axis.

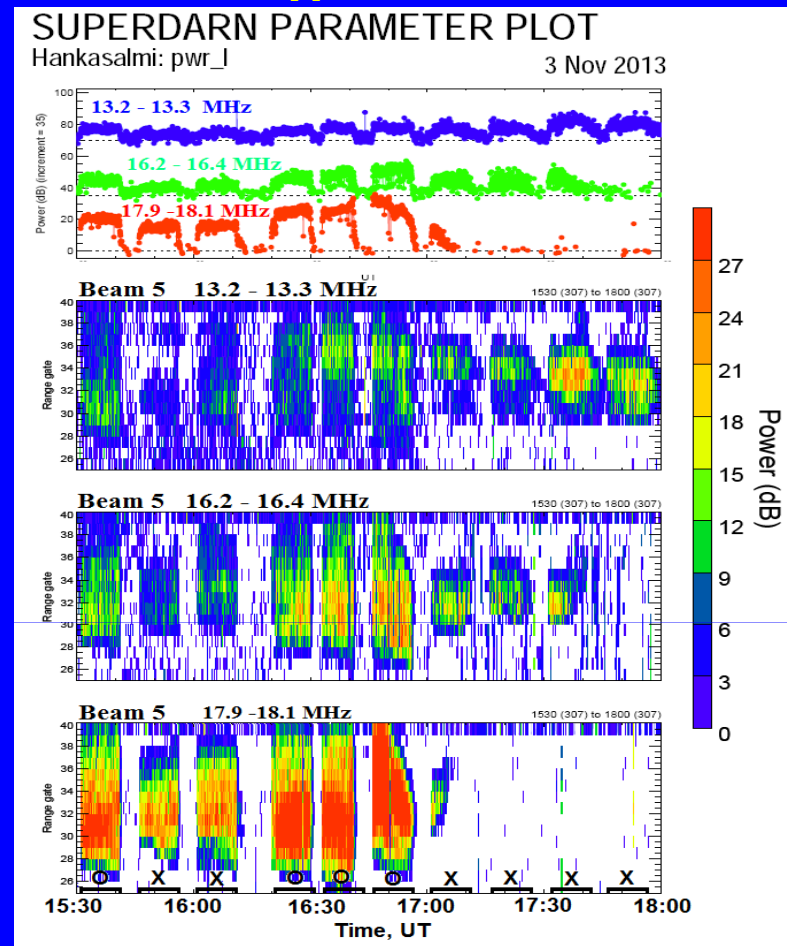


Backscattered power from the Hankasalmi (Finland) CUTLASS radar (beam 5) at frequencies of about  $16, 18$  and  $20$  MHz on 26 October 2013. HF pump wave was radiated towards MZ at  $f_H = 6.96$  MHz (ERP = 550 MW).

# Alternating O/X-mode HF pumping ( $f_H = 6.2$ MHz)

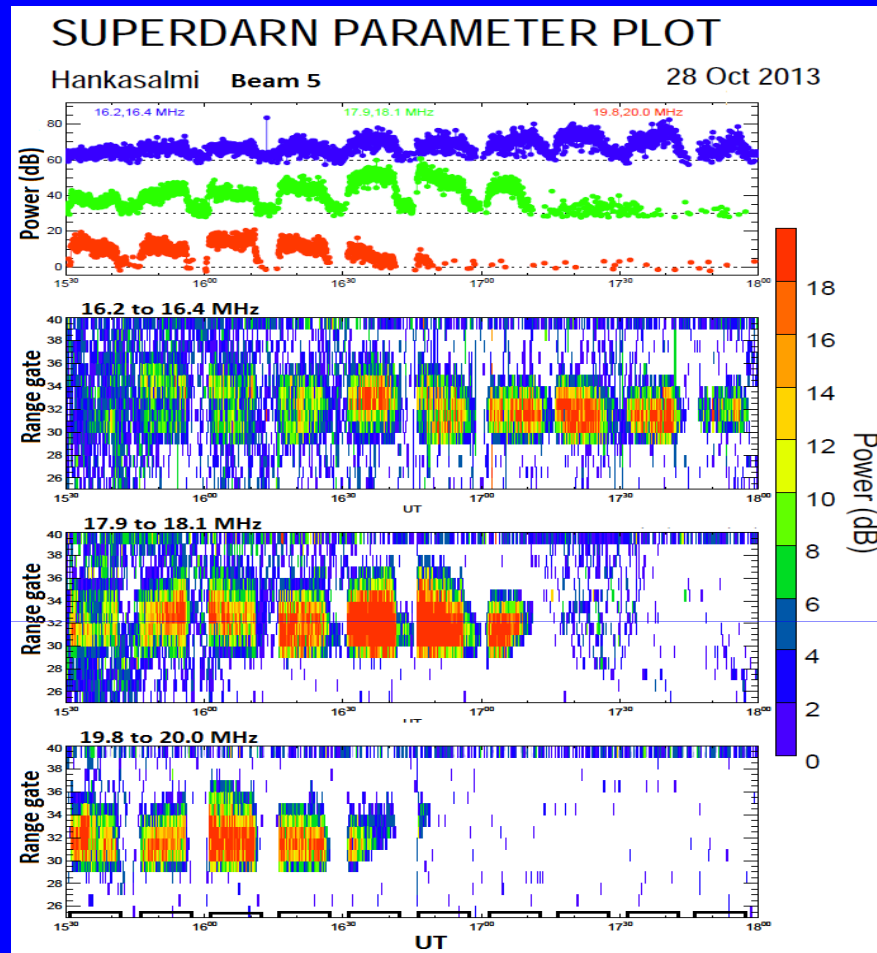


The behaviour of the Ne, Te, Vi from the EISCAT UHF radar observations at Tromso for O/X-mode heating on 3 November 2013. The O/X-mode HF pump wave at  $f_H = 6.2$  MHz (ERP = 450 MW) was radiated towards MZ by 10 min on, 5 min off cycles. The critical frequency foF2 dropped from 6.7 MHz at 15.30 UT to 5.2 MHz at 18 UT ( $f_H / \text{foF2} = 0.92 - 1.2$ ).

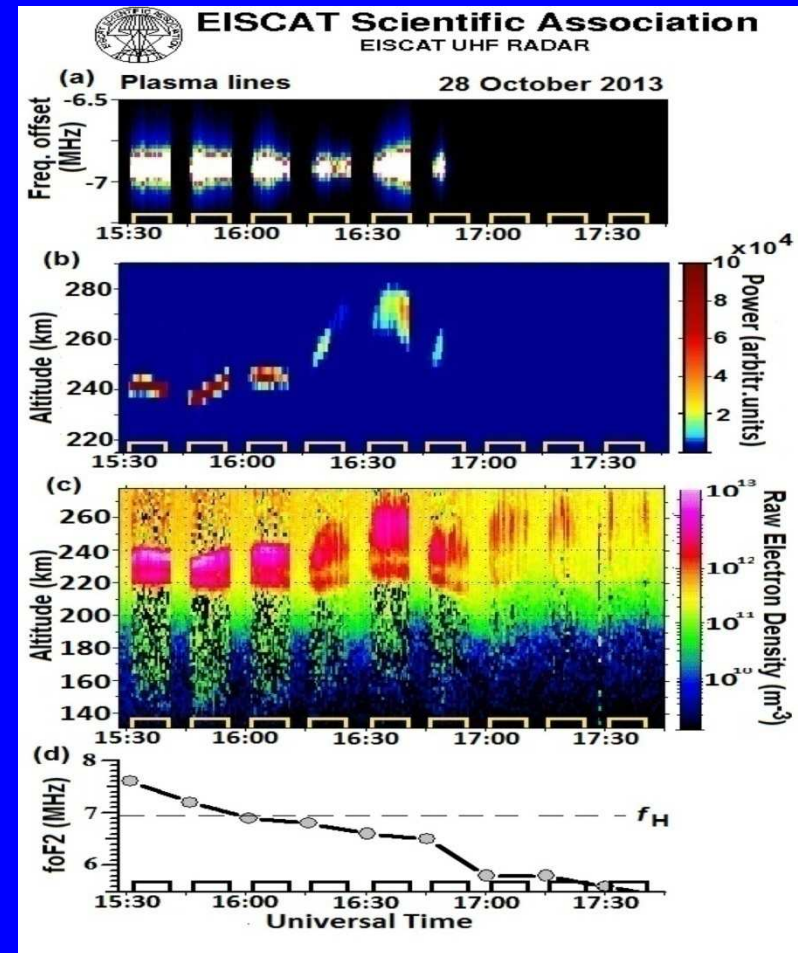


Backscattered power from the Hankasalmi (Finland) SuperDARN radar (beam 5) at operational frequencies of about 13, 16 and 18 MHz on 3 November 2013. HF pump wave was radiated towards MZ at  $f_H = 6.2$  MHz (ERP = 450 MW).

# X-mode FAls and HF-induced plasma turbulences



Backscattered power from the Hankasalmi (Finland) SuperDARN radar (beam 5) at operational frequencies of about 16, 18 and 20 MHz on 28 October 2013. The X-mode HF pump wave was radiated towards MZ at  $f_H = 6.96$  MHz (ERP = 550 MW). The foF2 values dropped from 7.6 to 5.3 MHz.



The behaviour of HF-induced plasma lines and raw electron densities obtained from the EISCAT UHF radar on 28 October, 2013. The X-mode HF pump wave was radiated towards MZ at  $f_H = 6.96$  MHz (ERP = 550 MW). The foF2 values dropped from 7.6 to 5.3 MHz ( $f_H / \text{foF2} = 0.91 - 1.35$ ).



# Distinctive features of FAIs

- |   | X-mode                         | O-mode                         |
|---|--------------------------------|--------------------------------|
| • generation                                    | $f_H \leq foF2$ ;              | $f_H \leq foF2$                |
| •   | $f_H > foF2$ (up to 2 MHz)     |                                |
| • spatial scale                                 | $l_{\perp} \approx 7.5 - 15$ m | $l_{\perp} \approx 7.5 - 15$ m |
| • growth time                                   | 12 – 120 s                     | 3 – 9 s                        |
| • decay time                                    | 40 – more than 300 s           | 12 – 30 s                      |
| • threshold of excitation from the “cold start” |                                |                                |
| • low $f_H = 4$ MHz                             | 75 MW                          | 8 – 10 MW                      |
| • high $f_H = 8$ MHz                            | 160 MW                         | 26 MW                          |
- Generation mechanisms of the FAIs under O-mode HF pumping at  $f_H \leq foF2$  is the thermal parametric (resonance) instability at the upper hybrid resonance altitude (Grach and Trachtengertz, 1975; Gurevich and Vas'kov 1978).
  - Generation mechanisms of the FAIs under X-mode HF pumping excited at  $f_H \leq foF2$  and  $f_H > foF2$  both is not completely understood. It is suggested that X-mode FAIs are generated via two-step process (Borisov et al., 2018; Blagoveshchenskaya et al., 2011). In the first step the generation of elongated large-scale irregularities (with the spatial scale across the geomagnetic field of the order of 1 -10 km ) due to the self-focusing instability is occurred. Excitation and behavior of small-scale FAIs is driven by large-scale irregularities. Their possible generation mechanisms can be the temperature gradient-drift instability (Borisov et al., 2018) or the filamentation instability (Kuo, 2015).

# Summary

- It was found that the features and physical driving mechanisms of FAIs with the spatial scale across the geomagnetic field of 7.5 – 15 m are significantly different for O- and X-mode HF pumping, presenting challenges for understanding the relevant processes.
- X-mode FAIs are excited in the regular high-latitude ionosphere F-region under quiet magnetic conditions when the heater frequency is below as well as much above the critical frequency of the F2 layer (up to 2 MHz ), whereas the O-mode FAIs can not be fundamentally generated when the pump frequency  $f_H$  is above  $f_{oF2}$ .
- The X- and O-mode FAI behaviors exhibit the different grow and decay times, and thresholds of excitation that is indicative of different physical mechanisms of their generation.
- As is found from EISCAT UHF radar measurements FAIs greatly impact on the development of strong artificial turbulence such as Langmuir and ion-acoustic plasma waves.

# References

Blagoveshchenskaya, N. F., T. D. Borisova, T. Yeoman et al. (2011). Artificial field-aligned irregularities in the high-latitude F region of the ionosphere induced by an X-mode HF heater wave, *Geophys. Res. Lett.*, 38, L08802, doi: 10.1029/2011GL046724.

Blagoveshchenskaya, N. F., T. D. Borisova, M. Kosch et al. (2014). Optical and Ionospheric Phenomena at EISCAT under Continuous X-mode HF Pumping, *J. Geophys. Res. Space Physics*, 119, 10,483–10,498, doi:10.1002/2014JA020658.

Blagoveshchenskaya, N.F., T.D. Borisova, T.K. Yeoman, I. Häggström, and A.S. Kalishin (2015), Modification of the high latitude ionosphere F region by X-mode powerful HF radio waves: Experimental results from multi-instrument diagnostics, *J. Atmos. Sol.-Terr. Phys.*, 135, 50–63.

Borisov, N., F. Honary, and H. Li (2018). Excitation of plasma irregularities in the F-region of the ionosphere by powerful HF radio waves of X-polarization . *J. Geophys. Res.: Space Physics*, 123, 5246–5260, [doi.org/10.1029/2018JA025530](https://doi.org/10.1029/2018JA025530).

Grach, S. M., and V. Y. Trakhtengerts (1975), Parametric excitation of ionospheric irregularities extended along the magnetic field, *Radiophys. Quant. Electron.*, 18, 951–957.

Gurevich, A. V. (2007), Nonlinear effects in the ionosphere, *Physics-Uspekhi*, 50, 1091–1121.

Gurevich, A. V., 1978. Nonlinear Phenomena in the Ionosphere. Springer-Verlag, New York.

Kuo, S. (2015), Ionospheric modifications in high frequency heating experiments, *Phys. Plasmas*, 22, 012901, doi: 10.1063/1.4905519.

Robinson, T.R. (1989), The heating of the high latitude ionosphere by high power radio waves, *Phys. Rep.*, 179, 79 - 209.

## Acknowledgements

EISCAT is an international scientific association supported by research organizations in China (CRIRP), Finland (SA), Japan (NIPR and STEL), Norway (NFR), Sweden (VR), and the United Kingdom (NERC). We would like to thank the Staff of the EISCAT Scientific Association for help in providing the Tromsø heating experiments in conjunction with the EISCAT UHF radar observations in the course of Russian campaigns. We are also thankful to the Department of Physics and Astronomy of University of Leicester for providing CUTLASS observations during heating campaigns. ). TKY is supported by Science and Technology Facilities Council Grant ST/H002480/1.