

Asymmetric Backward Peaking Radiation Pattern from a Relativistic Particle Accelerated by Lightning Leader Tip Electric Field

Mert Yücemöz¹, Martin Füllekrug²

¹University of Bath

²University of Bath

Key Points:

- A novel formula was developed to calculate bremsstrahlung radiation patterns of a charged non/relativistic particle with the Doppler effect.
- The bremsstrahlung radiation pattern of a charged relativistic particle exhibits forward and backward peaking due to symmetry conservation.
- The two forward and backward peaking lobes are asymmetric with respect to the velocity vector due to the curved particle trajectory.

Corresponding author: Mert Yucemoz, m.yucemoz@bath.ac.uk

This article has been accepted for publication and undergone full peer review but has not been through the copyediting, typesetting, pagination and proofreading process, which may lead to differences between this version and the [Version of Record](#). Please cite this article as doi: [10.1029/2020JD033204](https://doi.org/10.1029/2020JD033204).

This article is protected by copyright. All rights reserved.

Abstract

Terrestrial Gamma-ray Flashes exhibit slopes of ionizing radiation associated with bremsstrahlung. Bremsstrahlung has a continuous spectrum of radiation from radio waves to ionizing radiation. The Poynting vector of the emitted radiation, i.e., the radiation pattern around a single particle under the external lightning electric field during interaction with other particles or atoms, is not quite well known. The overall radiation pattern arises from the combination of radiation of parallel and perpendicular motions of a particle caused by the acceleration from the lightning electric field and the bremsstrahlung. The calculations and displays of radiation patterns are generally limited to a low-frequency approximation for radio waves and separate parallel and perpendicular motions. Here we report the radiation patterns of combined parallel and perpendicular motions from accelerated relativistic particles at low and high frequencies of the bremsstrahlung process with an external lightning electric field. The primary outcome is that radiation patterns have four relative maxima with two forward peaking and two backward peaking lobes. The asymmetry of the radiation pattern, i.e., the different intensities of forward and backward peaking lobes, are caused by the Doppler effect. A novel outcome is that bremsstrahlung has an asymmetry of the four maxima around the velocity vector caused by the curvature of the particle's trajectory as it emits radiation. This mathematical modeling helps to better understand the physical processes of a single particle's radiation pattern, which might assist the interpretation of observations with networks of radio receivers and arrays of γ -ray detectors.

1 Introduction

It was recently suggested that high-frequency radiation emissions observed in the atmosphere could originate from muons interacting with electric fields inside thunderclouds. This novel idea is based on a reduction of the muon detection during thunderstorm occurrences by the ground based telescope GRAPES-3 located in Ooty, India (Hariharan et al., 2019). Gamma-Ray Bursts (GRBs) are commonly thought to result from the interaction of neutron stars in outer space or comet collisions. GRBs emit photons in the energy range from keV to MeV that last ~ 10 seconds. However, a ~ 90 minute long GRB was detected with photon energies ~ 18 GeV (Hurley et al., 1994). When Terrestrial Gamma-Ray Flashes (TGFs) were first observed by detectors of the Compton Gamma Ray Observatory (CGRO) (Fishman et al., 1994), their association with bremsstrahlung was demonstrated by the observation of the characteristic slopes of ionizing radiation (Dwyer et al., 2012a), supported by Monte Carlo simulations that included the bremsstrahlung process (Dwyer, 2007). Another example of bremsstrahlung associated with lightning discharges is the detection of ultra-low frequency (ULF) and very low frequency (VLF) radio emissions of the same electrons that are also responsible for emitting terrestrial gamma-ray flashes (Connaughton et al., 2013). TGFs are associated with low-frequency radio emissions, and these observations were used to identify their source altitude (Pu et al., 2019; Cummer et al., 2014). The source altitude was located to lie between two charged cloud layers in a thunderstorm. All the above discoveries offer experimental evidence for the continuous radiation spectrum of bremsstrahlung to occur. Relativistic runaway electrons are the source of high-frequency X- and γ -ray emissions observed in the upper troposphere at altitudes from ~ 12 -14 km height (Celestin, 2016). High energy relativistic electrons have a larger mean free path such that they can attain larger velocities until they collide with an atom or molecule in the atmosphere. As these electrons are capable of reaching large velocities, they can emit ionizing radiation through the bremsstrahlung process. Low energy electrons are much more likely to collide with atmospheric atoms or molecules, leading to an increase in the number of free electrons in the atmosphere (Celestin, 2016). Another working hypothesis is that bremsstrahlung radiation is emitted by thermal runaway electrons accelerated by intra-cloud lightning leader tips (Xu et al., 2015). Bremsstrahlung has a continuous electromagnetic spectrum. Low-frequency radio and optical emissions could also be due to fluorescence, where high-frequency TGFs

67 are absorbed by air molecules (Xu et al., 2015). Numerical Monte Carlo simulations demon-
68 strated the significance of the bremsstrahlung process as the primary process behind high-
69 frequency emissions (Dwyer et al., 2012b). Bremsstrahlung electrons emit radiation in
70 forward peaking radiation patterns with an angle that scales with the inverse of the Lorentz
71 factor of the relativistic electrons (Koch & Motz, 1959).

72 Asymmetric signal of γ -ray bursts measured by the Gamma-Ray Burst Monitor
73 on the Fermi Gamma-ray Space Telescope reveal the lightning leader charge structure.
74 Asymmetric γ -ray pulses indicate the lightning leader charge flux, which exhibits a fast
75 rise and slow decay of the leader tip electric field (Foley et al., 2014). The asymmetries
76 in γ -ray pulses are thought to be caused by Compton scattering (Xu et al., 2019). The
77 rise to decay time ratio of single γ -ray pulses was measured to be approximately 0.67 (Nemiroff
78 et al., 1994). Data from the Burst and Transient Source Experiment (BATSE) reveals
79 two different types of spectra of γ -ray bursts known as bright and dim GRBs. It was found
80 that dim GRBs have less photon energy than bright GRBs (Norris et al., 1994). It was
81 observed that as time passes, overall γ -ray photons transit from bright to dim photons
82 as a photon bunch due to a time delay of approximately 100 μ s between the peaks aris-
83 ing from hard and soft photons (Grefenstette et al., 2008).

84 Experimental measurements of ionizing radiation and optical emissions by the At-
85 mosphere Space Interactions Monitor (ASIM) on the International Space Station recently
86 reported the detection of 217 TGFs from June 2, 2018, to April 1, 2019 (stgaard et al.,
87 2019), some associated with radio emissions from charged particles that are observed on
88 the ground. All these measurements reveal the properties of γ -ray bursts. After the com-
89 bination of the measurements from ground-based radio receivers and spacecraft, it was
90 found that TGFs are produced at the very beginning of the lightning discharge process.
91 It is well known that the observed γ -rays originate from the bremsstrahlung process (Xu
92 et al., 2015). There are approximately $\sim 10^{17} - 10^{19}$ Gamma-ray bursts emitted during the
93 bremsstrahlung process. It is well known that the initially emitted ionizing radiation is
94 not the same in terms of energy and direction compared to the radiation measured by
95 sensors. This difference is because the emitted radiation loses energy by back-scattering
96 and interacting with other air molecules. The interaction causes an ionization and re-
97 leases more electrons, which can explain why $10^{17} - 10^{19}$ γ -rays are emitted (Dwyer, 2008).
98 Another theory explains γ -ray bursts to originate from the large electric fields of leader
99 tip streamers producing $\sim 10^{12}$ electrons which then increase the number of electrons
100 within the relativistic runaway electron avalanche (RREA) process that emits γ -ray pho-
101 tons (Babich et al., 2014, 2015a; Moss et al., 2006; Chanrion & Neubert, 2010; Celestin
102 & Pasko, 2011; Skeltved et al., 2017).

103 This contribution reports the modeling of an asymmetric forward peaking radia-
104 tion pattern and an asymmetric backward peaking radiation pattern of a single parti-
105 cle bremsstrahlung process. The asymmetry occurs around the horizontal axis parallel
106 to the direction of motion of the charged particle, and it is unique to the bremsstrahlung
107 process as the particle continuously follows a curved trajectory of an anticlockwise ro-
108 tation. Radiation patterns are calculated for both relativistic and non-relativistic veloc-
109 ities. The main asymmetry with four radiation peaks is unique to the bremsstrahlung
110 process and occurs when the particle radiation transits from a dipole towards forward
111 and backward peaking radiation patterns.

112 The particle trajectory is mostly determined by the Coulomb field and modified
113 by the presence of the external lightning leader tip electric field. The change of the particle
114 trajectory caused by the external leader tip electric field is accounted for by only adjusting
115 the radius of the curvature of the particle's trajectory over time. This is possible because,
116 when the leader tip electric field is resolved into its vector components, one vector com-
117 ponent is tangent to the trajectory, whereas, the other vector component is perpendic-
118 ular to it. The perpendicular component points either in the same or opposite direction
119 to the centripetal acceleration caused by the target particle depending on the polarity

120 of the leader tip electric field. The tangent component of the leader tip electric field contributes
 121 to determining the amount of angular rotation that can be covered by the particle during
 122 Coulomb interaction. The perpendicular component of the external leader tip electric
 123 field contributes to the radius of curvature of the curved trajectory caused by Coulomb
 124 interaction.

125 The particle trajectory defined by the Coulomb field and the particle velocity arising
 126 from the leader tip electric field are two independently derived equations that serve
 127 as ingredients that contribute to the bremsstrahlung radiation pattern (Figure 1a). This
 128 independence is important because, under a repulsive or attractive Coulomb force, the
 129 target particle mostly determines the trajectory of the incoming particle. Also, param-
 130 eters defining particle trajectory by the target particles Coulomb field can account for
 131 the trajectory changes caused by the external leader tip electric field. Hence, when the
 132 radius of curvature modification caused by the external leader tip electric field is excluded
 133 from the particle trajectory, the temporal derivative of the particle trajectory is the ve-
 134 locity that an incoming particle experiences from the Coulomb force. The overall par-
 135 ticle velocity is the superposition of the velocity arising from both the leader tip and the
 136 target particle's electric field. The particle velocity is considered to be dominated by the
 137 external lightning leader tip electric field.

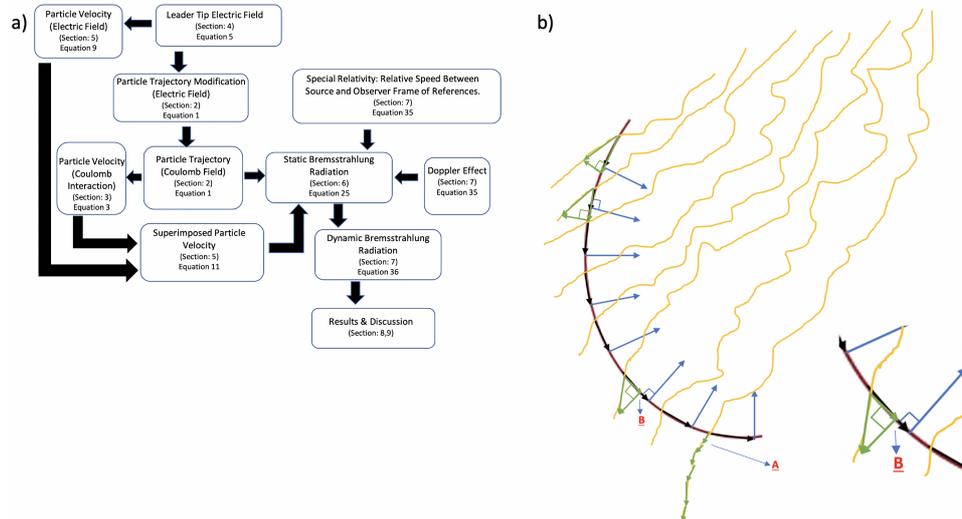


Figure 1. a) The flow chart shows the general structure of the theory as it is developed throughout the text. b) Explains how particle's curved trajectory caused by Coulomb field changes as a result of external non-uniform leader tip electric field and how it is accounted for in the theoretical model. The yellow lines are the external electric field black arrows are the tangential Coulomb velocity, thick blue arrows are centripetal acceleration. Finally, green lines are acceleration caused by the external electric field. As can be seen in region B, one component of the leader tip electric field acceleration is in the direction of tangential Coulomb velocity. Whereas, the other component is in the opposite direction of the centripetal acceleration. Leader tip electric field acceleration in the opposite direction of the centripetal acceleration of the Coulomb field only changes the radius of curvature of the trajectory. Hence, it can be accounted with parameter "b" in the trajectory equation 1, defining the radius of curvature. The acceleration of the leader tip electric field tangent to the trajectory is accounted with the leader tip electric field velocity equation 8. Finally, region A explains how the bremsstrahlung process ends, when the leader tip electric field acceleration component opposite to centripetal acceleration is larger than the centripetal acceleration of the target particle.

138

1.1 Aims & Objectives

139

140

141

142

143

144

145

146

147

148

149

150

151

152

The primary aim of this contribution is to construct a generic mathematical model of electromagnetic radiation emitted by a single relativistic charged particle via the bremsstrahlung process that applies to high and low-frequency radiation. Moreover, a generic mathematical model that is also capable of explaining a particle motion whose velocity and the acceleration vectors have both parallel (particle following linear path) and perpendicular (particle following circular trajectory) orientations with each other, at the same time. The secondary aim is to use the model to investigate the details of the actual particle motion and radiation mechanisms during the bremsstrahlung process. Initial assumptions for the particle position vector and lightning leader tip electric field were made to achieve the stated aims. In addition, all the necessary physical details of the interaction, such as form factors and collision cross-sections, are included and investigated. They both play a crucial role in explaining high-frequency radiation. Moreover, the Doppler effect was added to transform the model from a static to a dynamic model. Transformation enables an explanation of the asymmetry of the radiation in forward and backward direction relative to the axis perpendicular to the particle's direction of motion. Finally, the second-order differential equation (Eq. 13) was solved as an initial step to establish a generic model to explain radiation patterns using algebra, calculus, geometry, and the table of integral transforms.

153

2 Particle Position Vector

154

Starting with defining a curved path for a bremsstrahlung electron.

155

156

The position vector is formulated for a particle trajectory that is an anti-clockwise rotating spiral as a function of the retarded time characteristic for bremsstrahlung radiation.

157

158

159

160

161

162

163

164

165

166

167

168

The position vector in equation 1 defines a spiral trajectory for an incoming particle, i.e., an electron, induced by the Coulomb force of the target particle that causes the emission of bremsstrahlung radiation (Figure 2a). The spiral trajectory in Figure 2a and mathematically defined in equation 1 is realistic even though the mean free path is quite short, e.g., nm- μm in the atmosphere with a high recombination rate. For example, a circle with a radius of 1 m could also have a radius of 2 μm , depending on the medium and the recombination rate. It is still a circle but a scaled microscopic version of the initial macroscopic circle. Preserving geometry at different scales is also true in the formulated spiral trajectory (Eq.1). The decision on a specific particle trajectory considers the ratio of particle size to a curved trajectory radius. If the particle size is larger than the curvature radius, the particle trajectory is approximately a straight line. Therefore, a spiral particle trajectory is realistic because an electron has a size of $< 2.8 \times 10^{-19}$ m as measured by the Hadron-Electron Ring Accelerator (HERA) in Hamburg, Germany at the Deutsches Elektronen Synchrotron (DESY) facility (Bourilkov, 2000).

169

170

171

172

173

174

175

Overall acceleration caused by the Coulomb field can be resolved into its orthogonal components as centripetal and tangent Coulomb accelerations. The centripetal acceleration caused by the target particle defines the curved trajectory of the incoming particle and it is this centripetal acceleration that forces the incoming particle to stay on its curved trajectory. The tangent Coulomb velocity and acceleration components caused by the target particle due to the Coulomb field are perpendicular to the position and centripetal acceleration vectors. Hence, tangent to the particles curved trajectory.

176

177

178

179

180

181

182

Similar to Coulomb field, leader tip electric field can also be resolved into its orthogonal vector components. As shown in figure 1b, tangent to the particle trajectory component of the leader tip electric field increases the incoming particle's tangential Coulomb velocity, hence acceleration. The remaining acceleration component of the external lightning electric field acts in the opposite direction to centripetal acceleration and it only changes the radius of curvature of the incoming particles trajectory which can be accounted with mathematical variable b in equation 1 representing particles radius of curvature.

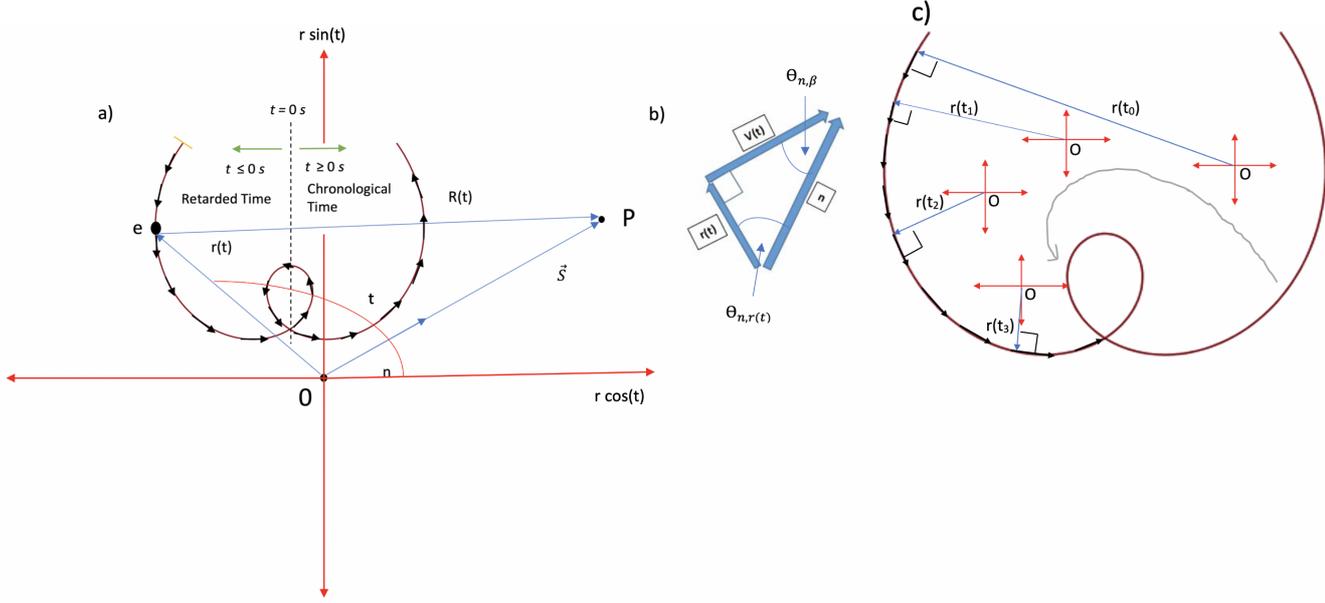


Figure 2. a) Trajectory of the bremsstrahlung electrons given by equation (1) in a polar coordinate system and radiation emissions by the change in velocity over time by a Coulomb force of other charges. O represents the target particle that defines the electron's spiral trajectory due to the Coulomb force (equation 1). The tangential component of the electron acceleration and corresponding velocity vectors are displayed with black arrows and are tangential to the spiral trajectory (red line) and perpendicular to the position vector $r(t)$. $R(t_r)$ is the distance between the accelerated electron and the observer, which is a function of both retarded and chronological time (t_r, t). P is the position of an observer. \vec{S} is the Poynting vector, or radiant energy flux, which determines the direction of the energy flow per area of an emitted electromagnetic wave. The dimensionless unit vector n points in the direction of the Poynting vector \vec{S} . b) The position vector $r(t)$, velocity vector $v(t)$ and the unit radiation vector n form a rectangular triangle. c) Co-ordinate system and all the associated vectors are drawn from the target particle which is positioned at the origin, O . As the target particle is not a rigid body and displaces under the incoming particle's Coulomb field, the moving co-ordinate system is constructed (Brock, 2019). Displacement of the moving co-ordinate system is described with the collision cross-section and form factors (Supplemental information, text S2). Position vector, $r(t)$ connects both target and incoming particle together. Time derivative of this position vector is the tangential velocity of the incoming particle and it is perpendicular to the position vector in the moving co-ordinate system. Without radius of trajectory changes caused by the external leader tip electric field, time derivative of the position vector is the tangential component of the Coulomb velocity. With radius of trajectory changes caused by the external leader tip electric field, time derivative of the position vector is the tangential component of the overall velocity.

$$r(t) = \frac{(t^R)^2 b^R (\omega')^R \cos(\theta_{n,r(t)})^R c}{\tau^2 R c^R \omega' \cos(\theta_{n,r(t)})} - \frac{at}{\tau}, \quad (1)$$

where $r(t)$ is the position vector as a function of time t in s, R is the dimensionless bremsstrahlung asymmetry index, τ is the mean free time in s. Also, ω' is the angular frequency of the emitted electromagnetic wave in the frame of reference of the particle in rads/s, c is the speed of light, $\theta_{n,r(t)}$ is the angle between

the emitted radiation unit vector n and the particle's position vector. The time range of the position vector is $-\infty < t < +\infty$. In addition, the factor b in m describes the interaction distance between the incoming particle and the target particle, which is the radius of the time-dependent position vector. The radius of the position vector is directly proportional to the parameter b . This radius of curvature is related to the mean free path because the curvature increases with time, contributing to the overall arc length, i.e., the mean free path of the accelerated particle. In other words, $b \propto \lambda_{v/c}$, where $\lambda_{v/c}$ is the mean free path of a particle at the velocity v which is given as a percentage of the speed of light v/c . The relation between a and $\lambda_{v/c}$ will be derived in section 3. The factor a in m is an arbitrary adjustment parameter. It is introduced to correct the radius of the curvature of a particle during the bremsstrahlung process. This correction is required because the trajectory of a relativistic particle shrinks in size over time, and a propagation close to the speed of light introduces significant changes in the mean free path. Finally, equation 1 is a function of time t . Hence, the particle will only cover some segment, or arc length, of the spiral, or the complete arc length of a particle's spiral trajectory when $t = \tau$.

3 Interaction Gap

Defining the mean free path, $\lambda_{v/c}$ and the particle velocity arising only from the Coulomb interaction, $\frac{dr}{dt}$.

The arc length, or the mean free path, of a curve in polar coordinates is given by

$$\lambda_{v/c} = \int_0^{\theta_f} \sqrt{r^2 + \left(\frac{dr}{d\theta}\right)^2} d\theta, \quad (2)$$

where $\lambda_{v/c}$ is the mean free path, r is the position vector given in equation 1, and the derivative with respect to the polar angle is $\frac{dr}{d\theta}$. Equation (2) is a geometric equation used to calculate the arc length of any curve in polar coordinates. As the particle covers its mean free path, it follows the curved spiral trajectory in small segments with time t . In this case, equation (2) has to be formulated with the parameter t rather than θ . To satisfy the dimensional accuracy, the Jacobian is used to carry out a variable transform from θ to t to preserve the dimension m^2 in the square root of the integrand. In geometry, the unit of angle in rad is considered to be dimensionless. An example of this is the arc length formula of a circle ($s = r\theta$). Hence, $\theta = \frac{t}{\tau}$. The mean free time τ is a function of time due to its dependence on particle velocity that changes over time. Therefore, the quotient rule $\left(\frac{d\theta}{dt} = \frac{\Lambda \frac{du}{dt} - u \frac{d\Lambda}{dt}}{\Lambda^2}\right)$ has to be used to find the derivative $\frac{d\theta}{dt}$. Input parameters to the quotient rule are, $u = t$, $\Lambda = \tau$, $\frac{du}{dt} = 1$ and $\frac{d\Lambda}{dt} = \tau'$ where τ' is the first derivative of the mean free time with respect to time t . The mean free time τ of the two identical particles is defined as $\tau = \frac{1}{\pi n v d^2}$, where n is the particle or molecule number density per unit volume in m^{-3} and d is the diameter in m of both an incoming and a target particle or molecule. The basic mean free time is defined between two identical particles with same diameter d where effective collision area is defined as a function of the two identical particle's diameter as $A_c = \pi d^2$. Finally, $\frac{d\theta}{dt} = \frac{\tau - t\tau'}{\tau^2}$ and the limits of the integration are $0 \leq \frac{t}{\tau} \leq \theta_f = 1$ rad. The maximum limit of the integral is 1, which means that the ratio of t to τ as time progresses should not exceed allowed mean free time τ for the particle to radiate.

Coulomb velocity of an incoming particle arising from electrostatic interaction between a target particle can be written as

$$\frac{dr}{dt} = \frac{b^R (\omega')^R \cos(\theta_{n,r(t)})^R c}{c^R \omega' \cos(\theta_{n,r(t)})} \left(\frac{\tau^{2R} \left[t^{2R} \frac{2R}{t} + t^{2R} 2R' \ln(t) \right] - t^{2R} \left[\tau^{2R} \frac{2R}{\tau} + \tau^{2R} 2R' \ln(\tau) \right]}{\tau^{4R}} \right) - \frac{a\tau - at\tau'}{\tau^2}. \quad (3)$$

225 Second terms, $t^{2R}2R'\ln(t)$ and $\tau^{2R}2R'\ln(\tau)$ inside the square brackets of the equation 3 represents a
 226 trajectory of varying radius of curvature i.e. spiral. Changing radius of curvature with time is directly
 227 linked to the changing bremsstrahlung asymmetry, R with time, R' . In the case of $R' = 0$, particle
 228 follows a constant radius trajectory where the bremsstrahlung asymmetry, R remains at a constant value.
 229 Such trajectory could be a circle.

230 This leads to

$$\lambda_{v/c} = \int_0^1 \sqrt{r^2 + \left(\frac{dr}{dt} \frac{\tau^2}{\tau - t\tau'} \right)^2} \frac{\tau - t\tau'}{\tau^2} dt. \quad (4)$$

231 Equation 4 means that by knowing the mean free path of an electron in the atmosphere, the parameters
 232 a and b of a position vector $r(t)$ in equation one can be calculated.

233 4 Lightning Leader Tip Electric Field

234 In this section, the lightning leader tip electric field is formulated to define the particle velocity for
 235 the bremsstrahlung radiation. The velocity defined by an integration of the electric field can be equated to the
 236 formulated velocity (equation 3) from the particle position vector to calculate the unknown parameters a and b .
 237 Moreover, this is an important step in defining particle velocity parameters such as the mean free path and time
 238 as the leader tip electric fields can be measured experimentally.

239 The measured lightning discharge electric field has a characteristic behavior of a sharp rise and
 240 a slow decay. This characteristic information was used to mathematically model the lightning leader tip
 241 electric field, E in equation 5. The electric field is mathematically modeled using the ansatz of a sharp rise
 242 and a slow decay

$$E(t) = s_f \left[\frac{An(s_{ft})^{(n-0.489)}}{(t/t_0)^{(n-0.489)}} - \frac{2B(s_{ft})^m}{(t/t_0)^m} \right], \quad (5)$$

243 where $A = 4.083 \times 10^7 \left[\frac{\text{J}}{\text{Cm}} \right]$ or $\frac{\text{V}}{\text{m}}$, $B = 3.840 \times 10^7 \left[\frac{\text{J}}{\text{Cm}} \right]$ or $\left[\frac{\text{V}}{\text{m}} \right]$, $n = 1.95$ is a dimensionless
 244 constant, and $m = 1.5$ is a dimensionless constant. The scaling factor $s_f = 1$ scales the lightning
 245 leader tip electric field magnitude, i.e., it is a dimensionless constant that can be used with different
 246 constant values to adjust the leader tip electric field to the required strength, $s_{ft} = 1$ is a dimensionless
 247 constant scaling factor of time, and the time t/t_0 is the relative time with respect to $t_0 = 1$ s where $t[s]$ is
 248 an independent time variable. The time range of the electric field is $s_{ft} 0.39694 \mu\text{s} \leq t \leq \infty$.
 249 Values closer to $0.397 \mu\text{s}$ describe the initial state and values closer to $72.2 \mu\text{s}$ describe the final state
 250 of the lightning leader tip electric field shown in Figure 3. The minimum value of time was determined such
 251 that it is the minimum value of the electric field before it goes to $-\infty$ at $t = 0$ s. The scaling factor determines
 252 the acceleration of the particle. As a result, the magnitude of the emitted radiation intensity scales in an
 253 indirect way by scaling the electric field which scales the force and hence the acceleration.

254 The lightning leader tip electric field in equation (5) is derived such that it goes to zero at infinite time,
 255 as shown in Figure 3b. Any remaining residual electric field integrated over infinite time would accelerate
 256 a particle to velocities exceeding the speed of light.

257 The strength and duration of the electric field affect the magnitude and pattern
 258 of the emitted radiation. To adjust the electric field's strength and duration conveniently
 259 without affecting the graphical shape of Figure 3a, the scaling factor s_f for magnitude
 260 and s_{ft} for the duration are introduced.

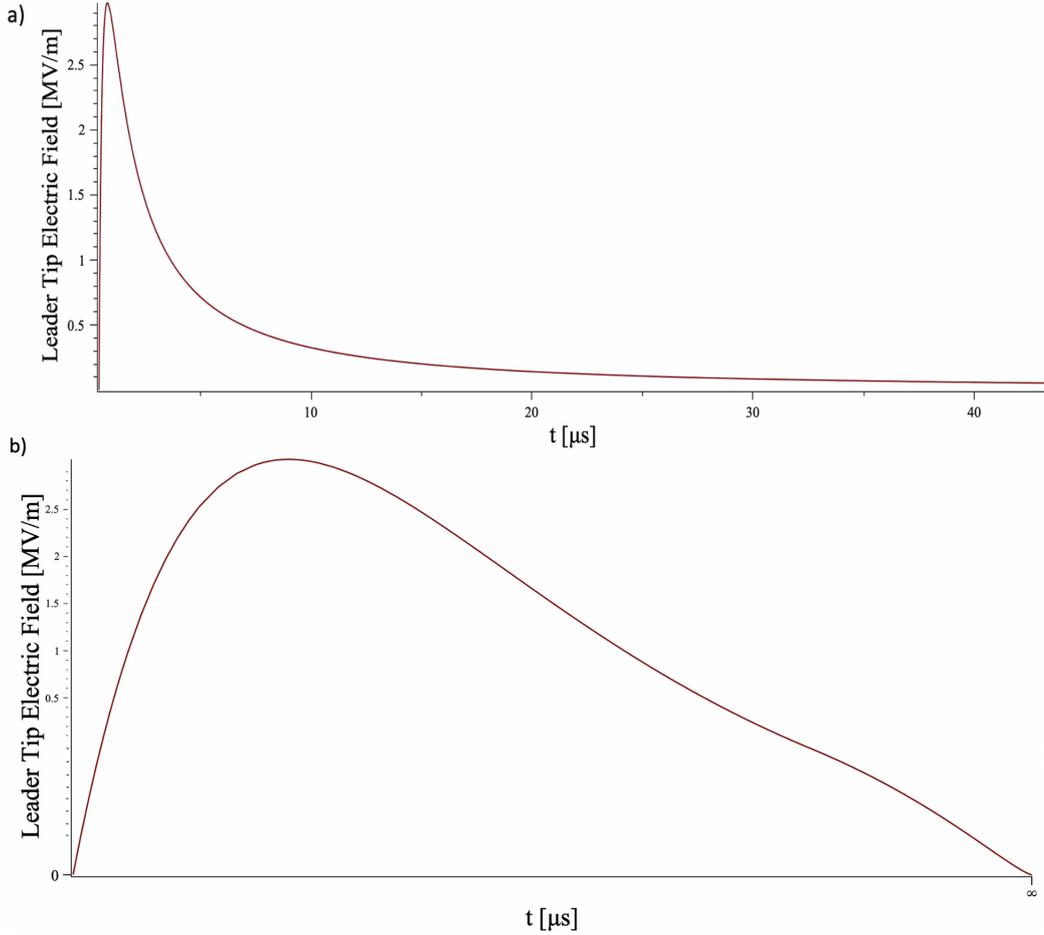


Figure 3. Derived electric field in equation 5. a) The leader tip electric field is modeled to fit observed electric field characteristics of lightning discharges. Lightning discharges exhibit a sharp rise and slowly decaying electric field. The peak value of the electric field was chosen to be ~ 3 MV/m, which is the approximate electric field for the conventional dielectric breakdown of air. For a RREA to develop, the electrons need to be relativistic such that bremsstrahlung occurs. In this case, the leader tip electric field is on the order of ~ 26 MV/m at 101 kPa of atmospheric pressure (Babich et al., 2015b). b) The derived electric field from 0 s to ∞ is plotted on logarithmic time scale. It is important that the leader tip electric field is 0 V/m at infinity. Any function used to describe the leader tip electric field that never goes to zero (i.e., an exponential) integrated up to infinity would cause a particle to exceed the speed of light.

5 Velocity Function

The velocity of the particle accelerated by the leader tip electric field is found by integrating the derived electric field (Figure 3, Eq. 5). The upper and lower limits of the integration are chosen such that the result of the integral gives a velocity equation as a function of time. Next, the solution of integral is scaled to accelerate the particle up to 94% of the speed of light, i.e., 280,000 km/s, and thereby prevents the particle from exceeding the speed of light in vacuum, c . Finally, the particle velocity function should decay to the thermal velocity. Afterward, the particle should either remain constant at a thermal velocity or slowly decay to 0 m/s and remain stationary at an infinite time for the particle to stop radiating. A stationary or constant velocity at infinite time is im-

271 portant because equation 13 requires an integral with limits between $\mp\infty$, and the par-
 272 ticle cannot radiate for an infinite time. The derived velocity function (Eq. 9) presented
 273 in Figure 4 accelerates the particle from a thermal velocity (92.493 km/s) up to 94% of
 274 c within ~ 12.427 ns and then decelerates back to a thermal velocity. The whole pro-
 275 cess takes ~ 24.854 ns.

276 The velocity function is the integral of the electric field function with respect to
 277 time, multiplied with the constant charged particle properties, which are elementary charge
 278 e , charge number z , and the inverse of accelerated electron mass $\frac{1}{m_e}$.

$$v(t) = \frac{ez}{m_e} \int_{s_{ft}0.39694\mu s + \left| \frac{t[s]}{t_o[s]} \right| t_o}^{\infty} E(t) dt. \quad (6)$$

279 Where $t_o = 1$ s . The upper limit was set to be infinite in order to prevent the
 280 lower boundary of the integral from exceeding the upper limit. The lower limit involves
 281 a time parameter to prevent the indefinite integral from being a definite integral in or-
 282 der to create a velocity function. The absolute time is used here to create a symmetric
 283 velocity behavior in retarded time to explain the particle motion and also to satisfy the
 284 integral limits between $\mp\infty$ in equation 13. The factorial was incorporated in order to
 285 prevent the electric field from becoming singular at $t=0$ s.

286 Separating integral (Eq. 6) into the two terms of the electric field function (Eq. 5)
 287 gives

$$v(t) = \frac{ez}{m_e} s_f A n s_{ft}^{(n-1.489+1)} \int_{s_{ft}0.39694\mu s + |t|}^{\infty} \frac{1}{(t)^{(n-1.489)} t} dt - \frac{ez}{m_e} s_f 2B s_{ft}^m \int_{s_{ft}0.39694\mu s + |t|}^{\infty} \frac{1}{t^m} dt \quad (7)$$

and

$$v(t) = \frac{ez}{m_e} s_f A n (s_{ft})^{(n-1.489+1)} \left[\frac{(125)2^{2.511+n} 3^{n-0.489} (9.657 \times 10^6)^{n-0.489} (23s_{ft} + 5.794 \times 10^7 |t|)^{-n+0.489} (3.969 \times 10^7 s_{ft} + |t|)}{1000n - 1489} - \frac{ez}{m_e} s_f 2B (s_{ft})^m \left[\frac{(3.969 \times 10^{-7} s_{ft} + |t|)^{-m+1}}{-1 + m} \right] \right]. \quad (8)$$

288 The factorial of time only works for an integer number of time values. However,
 289 this is only a problem when the velocity function is used with non-integer values. To ap-
 290 proximate velocities with a factorial of non-integer time values, Stirling's approximation
 291 can be used. The velocity function can be written with the substitution of all the con-
 292 stant values that do not require to be changed in order to scale. Also, to preserve the
 293 physical meaning of the characteristic lightning leader tip electric field feature with a sharp
 294 rise and a slow decay taking place in the atmosphere. These constants are A , B , n , and
 295 m . Substituting and simplifying gives

$$v(t) = s_{fv} s_f z (s_{ft})^{1.461} \frac{4.365 \times 10^{26}}{(1.333 \times 10^9 s_{ft} + 3.357 \times 10^{15} |s_{ftv} t|)^{0.461}} - s_{fv} s_f z (s_{ft})^{1.5} \frac{1.565 \times 10^{27}}{((1.333 \times 10^9 s_{ft} + 3.357 \times 10^{15} |s_{ftv} t|)^{0.5}}. \quad (9)$$

296 As the particle has a short mean free path with a short mean free time \sim ns, the time scale in the velocity
 297 function (8) was scaled accordingly to ns with $s_{ftv} = 1 \times 10^9$ to reflect atmospheric electrodynamic conditions.

Previous to the scaling, as the time t is integrated up to ∞ , the time scaling in the velocity function (8) has grown and it took ~ 12.427 s for the electron to reach 94% of the speed of light from its thermal velocity $v_{rms} = \sqrt{\frac{3kT}{m}}$ of 92.493 km/s at an example ambient air temperature of 188.15 K, when considering the mesosphere at an altitude of 80 km. Similarly, as the time t is integrated up to ∞ , the particle attained a large velocity with the multiplication of the constant term $\frac{eZ}{m_e}$ in front of the integral. This is solved with a new velocity scaling factor $s_{fv} = 8.19 \times 10^{-11}$.

By knowing the particle velocity, the particle velocity equation (Eq. 9) can be scaled directly by using the corresponding scaling factors for the required velocity. The mean free path (Eq. 4) of a particle is a function of the total particle velocity. Although the derivative of the position vector (Eq. 3) alone determines the velocity arising from the Coulomb force, equation 3 is the total velocity when used for the mean free path (Eq. 4). Hence, if the particle velocity is known, it can be equated to the derivative of the position vector on a spiral trajectory (Eq. 3) to determine the parameters a (the relativistic mean free path correction parameter) and b (the interaction distance between an incoming and a target particle). This algebraic calculation transforms the velocity arising from the Coulomb field, i.e., the derivative of the position vector in equation 1, to an overall velocity, i.e., the sum of the velocity arising from the Coulomb field and the velocity originating from the external lightning leader tip electric field for a particle on a spiral trajectory.

As can be seen from Figure 4, the above assumptions help to construct a velocity function (Eq. 9). The complete radiation model is based on the fact that the particle reaches approximately 94% of the speed of light. Then the particle decelerates back to a thermal velocity as bremsstrahlung is the braking radiation due to obstacle particles or atoms, and as the source electric field from the leader tip starts to decay.

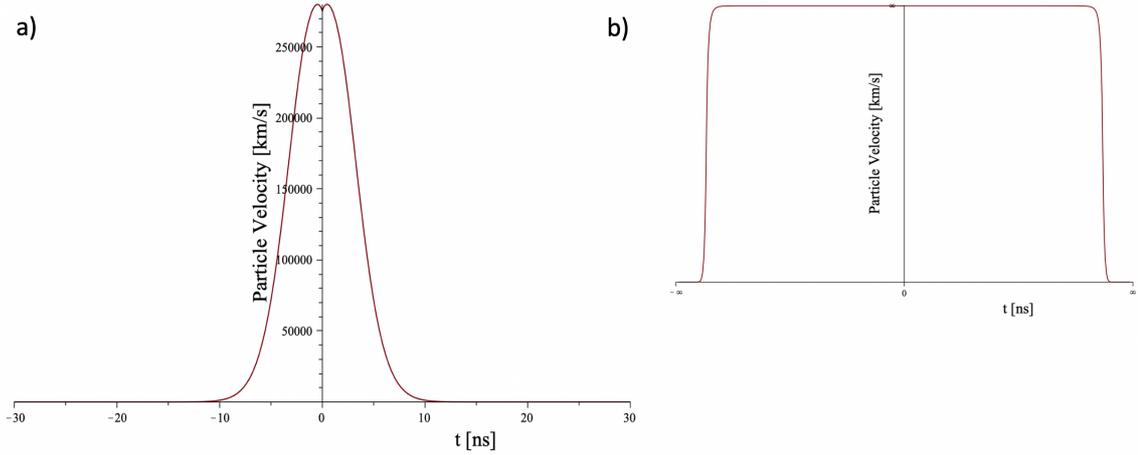


Figure 4. Velocity function (Eq. 9) of the derived electric field. (a) There is no negative chronological time. However, negative time has a meaning in physics that it describes an off the record process of an experimentally observed electromagnetic pulse. Negative time is the retarded time in LinaardWiechert potentials that describes the radiation emission by the acceleration of a single particle. Retarded time in LinaardWiechert potentials is also the reason why equation (13) has integral limits between $-\infty$ and ∞ . Negative time describes the duration of the process until the pulse detection time of the observer. Therefore, starting from ~ -10 ns up to to ~ 10 ns, the particle gains velocity through an acceleration by the external leader tip electric field, and as it accelerates, it radiates electromagnetic radiation. When the particle velocity is peaking around $t = 0$ s, the observer starts receiving a signal and can measure the pulse as the process of particle acceleration, and the propagation of the electromagnetic wave from the source to an observer requires some time - known as the retarded, or negative, time. b) The electron's velocity on an infinite time scale displays the end of a particle acceleration outside the range of $\sim \mp 11$ ns until $\mp \infty$. The peak velocity remains the same at 94% of the speed of light.

The relative particle velocity is

$$\beta(t) = \frac{v(t) + \frac{dr}{dt}}{c}. \quad (10)$$

Hence,

$$\beta(t) = \frac{\frac{s_{fv}s_{fz}(s_{ft})^{1.461}}{c} \frac{4.365 \times 10^{26}}{(1.333 \times 10^9 s_{ft} + 3.357 \times 10^{15} |s_{ftv}t|)^{0.461}} - \frac{s_{fv}s_{fz}(s_{ft})^{1.5}}{c} \frac{1.565 \times 10^{27}}{((1.333 \times 10^9 s_{ft} + 3.357 \times 10^{15} |s_{ftv}t|)^{0.5} + \frac{dr}{dt} \frac{1}{c})}}{c}}{c}, \quad (11)$$

where t is the time in s and the range of the time changes in the velocity function (Eq. 9), hence in β (Eq. 11) function. The time range in velocity function differs from the electric field function given in equation 5 as a result of the integration process. The time range of the velocity function is $-\infty \leq t \leq \infty$. An important information to note about the indicated range of time t values separately in both, the scaled velocity function (Eq. 9) and the electric field function (Eq. 5), is that they are specific for this scaled velocity and non-scaled electric field function separately. There is no time parameter in any other equation nor in the final equation due to the nature of the definite integral of time. However, as indicated in equation 13, the definite integral between the limits of $-\infty$ and ∞ of time results in a time t parameter to vanish in the final equation (Eq. 43). Therefore, the range of time t values presented in sections 4 and 5 can be neglected for the final equation 43.

Before moving to the next section, the concepts explained so far make up the base of the complete theory. The next section uses the information of sections 2, 4, and 5 to construct the core of the theory and to derive the novel generic radiation intensity pattern of a single particle unique to the bremsstrahlung process. Figure 5 shows how this theory fits into practical observations.

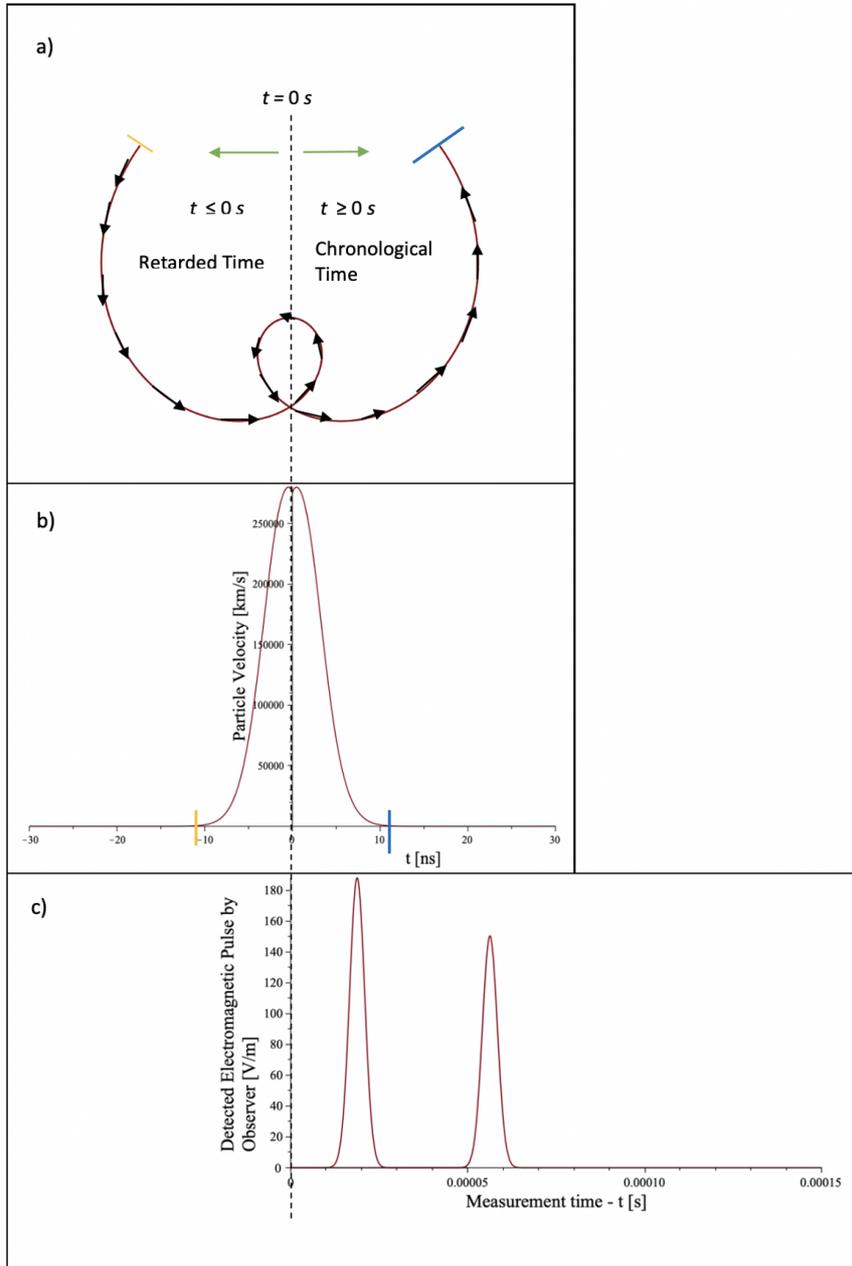


Figure 5. Complete order of events from initial particle acceleration until pulse detection by an observer. a) The yellow line indicates the particle's starting position. It starts accelerating from -10 ns until $t = 0$ s as a result of an external leader tip electric field and Coulomb electric field (Bremsstrahlung) of the target particle in the atmosphere. b) Shows how the particle's velocity changes as a function of time. The particle is still allowed to propagate to positive time values. However, the starting time is from negative time values (retarded time). Negative time is required to reflect the reality that when the signal is measured by an observer at $t = 0$ s, it has to travel some time from the source to the observer. The acceleration process that causes the emission also requires some time to take place. The observer in chronological time does not observe all of these processes. Hence, they occur in negative retarded time. c) Sketches when the observer would receive the signal. It is plotted using the Dirac Delta function to sketch when the observer would receive the signal, which does not reflect the actual timing. The first signal could also be received at $t = 0$ s, however not in retarded time as it would then contradict the observations where there is no negative time.

6 Mathematical Model of Combined High and Low-Frequency Bremsstrahlung Radiation Pattern

Deriving the bremsstrahlung radiation pattern.

This section introduces collision cross-sections as a part of explaining complete radiation pattern. Further information on collision cross-section and mathematical definitions consisting of form factors can be found in supplemental information file in the form of text, S2 (Punjabi & Perdrisat, 2014). Collision cross-section and form factors are very important in defining the final form of the radiation patterns at high frequency (Perdrisat et al., 2007). Form factors are not used in the final equation and in simulations of radiation patterns. For high-frequency emissions, the details of collision and interaction are crucial. They are accounted for by adding a cross-section of a particle of interest.

Hence, it is assumed that

$$\frac{d^3\chi}{d\omega' d\Omega_{rad} d\Omega_{par}} = \frac{d^2I}{d\omega' d\Omega_{rad}} \frac{d\sigma}{d\Omega_{par}}. \quad (12)$$

Dimensionally, $\chi = I \times \sigma$, where, σ in m^2 is the Coulomb scattering cross section, χ in Js^{-1}m^2 is the radiation cross section, and I in Js^{-1} is the radiation intensity. As a result of LinardWiechert potentials, the radiation emitted by a single charged particle (Jackson, 1999, p.675) is given by

$$\frac{d^2I}{d\omega' d\Omega_{rad}} = \frac{z^2 e^2 (\omega')^2}{4\pi^2 c \epsilon_0} \left| \zeta \right|^2, \quad (13)$$

where $\zeta = \int_{-\infty}^{\infty} n \times (n \times \beta) e^{i\omega'(t - \vec{n} \cdot r(t)/c)} dt$. In addition, z is the dimensionless charge number, c is the speed of light, ϵ_0 is the permittivity of free space, ω' is the angular frequency of the emitted wave in rad/s in the particle's frame of reference and t is the time in s.

The radiation unit vector n , the position vector $r(t)$ in equation 1, the particle motion, the position of an observer, and the overall theoretical concept about the bremsstrahlung radiation emission modeling that the equations are built upon are shown in the Figure 2.

As n is a unit vector in the direction of the emitted radiation, the integrand of the integral in equation (13) can be re-written in the sinusoidal form of the cross-product as

$$\zeta = \int_{-\infty}^{\infty} |\beta| |\sin(\theta_{n,\beta})| e^{i\omega'(t - \vec{n} \cdot r(t)/c)} dt, \quad (14)$$

or in the sinusoidal form of dot product as

$$\zeta = \int_{-\infty}^{\infty} |\beta| |\sin(\theta_{n,\beta})| e^{i\omega'(t - r(t)\cos(\theta_{n,r(t)})/c)} dt. \quad (15)$$

Substituting equation (1) and the first term of equation (11) into equation (14) and taking constants out results in

$$\zeta = \frac{s_{fv} s_f z(s_{ft})^{1.461} 4.365 \times 10^{26}}{c} \sin(\theta_{n,\beta}) \int_{-\infty}^{\infty} \frac{1}{(1.333 \times 10^9 s_{ft} + 3.357 \times 10^{15} |s_{ftv} t|!)^{0.461}} e^{i\omega(t - \left[\frac{(t^R)^2 b^R (\omega')^R \cos(\theta_{n,r(t)})^R c}{\tau^2 R_c R \omega' \cos(\theta_{n,r(t)})} - \frac{at}{\tau} \right] \cos(\theta_{n,r(t)})/c} dt. \quad (16)$$

A variable transformation from t to s requires the Jacobian. Assuming that $s = -t$ it follows that $\frac{ds}{dt} = -1$ such that

$$\zeta = -\frac{s_{fv} s_f z(s_{ft})^{1.461} 4.365 \times 10^{26}}{c} \sin(\theta_{n,\beta}) \int_{-\infty}^{\infty} \frac{1}{(1.333 \times 10^9 s_{ft} + 3.357 \times 10^{15} |s_{ftv} (-s)|!)^{0.461}} e^{i\omega(-s - \left[\frac{((-s)^R)^2 b^R (\omega')^R \cos(\theta_{n,r(t)})^R c}{\tau^2 R_c R \omega' \cos(\theta_{n,r(t)})} - \frac{a(-s)}{\tau} \right] \cos(\theta_{n,r(t)})/c} ds. \quad (17)$$

The exponential term can be simplified such that

$$e^{-i\omega s} e^{-i\omega \frac{((-s)^R)^2 b^R (\omega')^R \cos(\theta_{n,r(t)})^R c}{\tau^2 R_c R \omega' \cos(\theta_{n,r(t)})} \cos(\theta_{n,r(t)})/c} e^{-i\omega \frac{as}{\tau} \cos(\theta_{n,r(t)})/c}. \quad (18)$$

Further simplification of the common parameters results in

$$e^{-i\omega s} e^{-i \frac{((-s)^R)^2 b^R (\omega')^R \cos(\theta_{n,r(t)})^R}{\tau^2 R_c R}} e^{-i\omega \frac{as}{\tau} \cos(\theta_{n,r(t)})/c}, \quad (19)$$

where $e^{-i\omega s}$ can be neglected as there is no $\frac{1}{c}$ term to reduce the rate of exponential decay. Therefore, this term goes to zero quickly with time such that

$$e^{-i \frac{((-s)^R)^2 b^R (\omega')^R \cos(\theta_{n,r(t)})^R}{\tau^2 R_c R}} e^{-i\omega \frac{as}{\tau} \cos(\theta_{n,r(t)})/c}. \quad (20)$$

Bringing the whole integral in equation (16) together with the simplified exponential gives

$$\zeta = -\frac{s_{fv} s_f z(s_{ft})^{1.461} 4.365 \times 10^{26}}{c} \sin(\theta_{n,\beta}) \int_{-\infty}^{\infty} \frac{1}{(1.333 \times 10^9 s_{ft} + 3.357 \times 10^{15} |s_{ftv} (-s)|!)^{0.461}} e^{-i \frac{((-s)^R)^2 b^R (\omega')^R \cos(\theta_{n,r(t)})^R}{\tau^2 R_c R}} e^{-i\omega \frac{as}{\tau} \cos(\theta_{n,r(t)})/c} ds. \quad (21)$$

Finally, the velocity equation (9) converging to 0 m/s at $\mp\infty$ (see Figure 4.b) enables the integral in equation 21 to be evaluated. Convergence is due to the nature of derived electric field function, which goes to zero at infinite time and does not accelerate the particle any more. Furthermore, the use of functions such as Planck's radiation curve, Heidler current (Heidler et al., 2013), or Poisson distribution function with a variable transform to represent electric field would not be beneficial in deriving a velocity function as they diverge due to their exponential nature. The solution to the integral

382 is approximated by the use of a formula in the book of Fourier transforms of exponen-
 383 tial functions (Bateman et al., 1954, p.121, eq. 23). The integral can be approximated
 384 by seeking a solution from the table of integral transforms by Harry Bateman for the ex-
 385ponential Fourier transform of the function $f(x) = f(s)$ being equal to

$$f(s) = \frac{1}{(1.333 \times 10^9 s_{ft} + 3.357 \times 10^{15} |s_{ftv}(-s)|!)^{0.461}} e^{i\alpha^2 s^2}. \quad (22)$$

386 The solution of the integral requires index ν in the variable $(ix)^\nu$ that multiplies
 387 the exponential function in Exponential Fourier Transform of function $f(x) = f(s)$ to
 388 be $\nu > -1$ (Bateman et al., 1954, p.121, eq. 23). The derivation of the velocity (Eq.
 389 9) results in the beta function (Eq. 11) satisfying this requirement.

390 As the integrand of the integral in equation 14 multiplies with the β function con-
 391 sisting of 2 terms given in equation 11, the integral can be separated and solved indi-
 392 vidualy for each term such that all terms can be added together to give the final result.

393 Hence, the approximated solution of the integral ζ for the first term of the scaled
 394 velocity function (11) β is $\nu_1 = -0.461 > -1$ such that

$$\frac{d^2 I}{d\omega \Omega_{rad}} = -\frac{s_{fv} s_f z(s_{ft})^{1.461} 4.365 \times 10^{26}}{c} \left[\pi^{1/2} 2^{-(1/2)\nu_1} \alpha^{-\nu_1-1} e^{-\frac{y^2 \alpha^{-2}}{8}} \times D_{\nu_1}(2^{-1/2} \alpha^{-1} y) \right]. \quad (23)$$

395 The approximated solution of the integral ζ for the second term of the scaled ve-
 396locity function (11) β is $\nu_2 = -0.5 > -1$ such that

$$\frac{d^2 I}{d\omega \Omega_{rad}} = \frac{s_{fv} z s_f(s_{ft})^{1.5} 1.565 \times 10^{27}}{c} \left[\pi^{1/2} 2^{-(1/2)\nu_2} \alpha^{-\nu_2-1} e^{-\frac{y^2 \alpha^{-2}}{8}} \times D_{\nu_2}(2^{-1/2} \alpha^{-1} y) \right], \quad (24)$$

397 where $\alpha^2 = \frac{b^R (\omega')^R (\cos(\theta_{n,r(t)}))^R}{(\tau^R)^2 c^R} [s^{-2}]$ and $y = \frac{\omega' \cos(\theta_{n,r(t)}) a}{c\tau} [s^{-1}]$.

398 Therefore, the final radiation pattern for both high and low-frequency emission is
 399 the sum of the two terms

$$\begin{aligned} \frac{d^2 I}{d\omega' \Omega_{rad}} = & \frac{z^2 e^2 (\omega')^2}{4\pi^2 c \epsilon_0} \left| \sin(\theta_{n,\beta}) \right| \left[-\frac{s_{fv} s_f z(s_{ft})^{1.461} 4.365 \times 10^{26}}{c} \right. \\ & \left. \left[\pi^{1/2} 2^{-(1/2)\nu_1} \alpha^{-\nu_1-1} e^{-\frac{y^2 \alpha^{-2}}{8}} \times D_{\nu_1}(2^{-1/2} \alpha^{-1} y) \right] + \right. \\ & \left. \frac{s_{fv} z s_f(s_{ft})^{1.5} 1.565 \times 10^{27}}{c} \left[\pi^{1/2} 2^{-(1/2)\nu_2} \alpha^{-\nu_2-1} e^{-\frac{y^2 \alpha^{-2}}{8}} \times D_{\nu_2}(2^{-1/2} \alpha^{-1} y) \right] \right] \Bigg|^2, \quad (25) \end{aligned}$$

400 where R is the bremsstrahlung asymmetry control parameter, which plays a cru-
 401 cial role in distinguishing the parameter y from α , which is a requirement in order to be
 402 able to approximate the solution to an integral (dimensionless). $D_\nu(z)$ is the parabolic
 403 cylinder function.

404 The parabolic cylinder function is given by (Whittaker & Watson, 1927, p. 347)

$$D_\nu(z) = 2^{v/2+1/4} z^{-1/2} W_{v/2+1/4,1/4}(1/2 z^2), \quad (26)$$

405 where, $W_{v/2+1/4,1/4}(1/2z^2)$ is a Whittaker function (Whittaker & Watson, 1927,
406 p. 346) and

$$W_{\kappa,\mu}(1/2z^2) = \frac{\Gamma(-2\mu)M_{\kappa,\mu}(1/2z^2)}{\Gamma(1/2 - \mu - \kappa)} + \frac{\Gamma(2\mu)M_{\kappa,-\mu}(1/2z^2)}{\Gamma(1/2 + \mu - \kappa)}, \quad (27)$$

407 where, $M_{\kappa,\mu}(1/2z^2)$ is another Whittaker function (Kiyosi Ito and The Mathemat-
408 ical Society of Japan, 1993) & (Whittaker & Watson, 1927, p. 347) and

$$M_{\kappa,\mu}(1/2z^2) = {}_1F_1(1/2 + \mu - \kappa; 2\mu + 1; 1/2z^2)(1/2z^2)^{1/2+\mu} e^{-1/2z^2}. \quad (28)$$

409 For the second term of equation (27) with Whittaker M function of negative μ

$$M_{\kappa,-\mu}(1/2z^2) = {}_1F_1(\mu - \kappa; 2\mu; 1/2z^2)(1/2z^2)^{1/2+\mu} e^{-1/2z^2}, \quad (29)$$

where ${}_1F_1(1/2 + \mu - \kappa; 2\mu + 1; z)$ is a confluent hypergeometric function of the
first kind (Abramowitz & Stegun, 1972) and

$${}_1F_1(1/2 + \mu - \kappa; 2\mu + 1; 1/2z^2) = \frac{U_{\kappa,\mu}(1/2z^2)}{e^{-1/2z^2/2}(1/2z^2)^{\mu+1/2}}, \quad (30)$$

410 where, $U_{\kappa,\mu}(\frac{z^2}{2})$ is the confluent hypergeometric function of the second kind, named Kumer's
411 U function such that

$$U_{\kappa,\mu}\left(\frac{z^2}{2}\right) = \left(\frac{z^2}{2}\right)^{\mu+1/2} e^{-\frac{z^2}{4}} \sum_{n=0}^{\infty} \frac{(m-k+1/2)_n}{n!(2\mu+1)_n} \left(\frac{z^2}{2}\right)^n. \quad (31)$$

412 By direct comparison of equation 26 with 25 and equation 26 with 27, the constant
413 quantities z , κ and μ are defined as $z = 2^{-1/2}\alpha^{-1}y$, $\kappa = v/2 + 1/4$, $\mu = 1/4$.

414 Finally, in order to complete equation 25, the observation angles ($\theta_{n,\beta}$ and $\theta_{n,r(t)}$)
415 need to be reduced from two to one as they are related to each other.

416 Currently, the radiation pattern is observed from $\cos(\theta_{n,r(t)})$ and $\sin(\theta_{n,\beta})$. In ad-
417 dition, $\cos(\theta_{n,r(t)})$ coming from substitute parameter α and y in equation 25 and $\sin(\theta_{n,\beta})$
418 being a variable parameter in equation 25 defining the radiation around the particle, also
419 known as the solid angle. In addition, by looking at the Figure 2b, it can be seen that
420 the velocity vector, which is the derivative of the position vector $r(t)$ is always perpen-
421 dicular to the position vector where radiation unit vector pointing in the direction of the
422 emitted radiation makes up the hypotenuse of the formed rectangular triangle. Hence,
423 the angles of the rectangular triangle can be equated to each other.

424 The velocity function defined in equation 9 provides the particle speed attained by
425 the acceleration due to the external leader tip electric field, and it is scalar. The deriva-
426 tive of the position vector $r(t)$ (Eq. 1) provides the velocity caused by the Coulomb in-
427 teraction of an incoming and target particle, which also determines the curved trajec-
428 tory of the bremsstrahlung given by the position vector (Eq. 1). Hence, the particle's
429 overall velocity moving perpendicular to the position vector $r(t)$ is the sum of the Coulomb
430 velocity arising from the derivative of the position vector $r(t)$ plus the scalar speed at-
431 tained by the external leader tip electric field given in equation 9. Equation 9 is not the
432 velocity, but the speed. It only provides magnitude but gives no information about the
433 direction of the particle's motion. The direction of the particle's motion is already pro-
434 vided by the position vector $r(t)$ given by equation 1. As a result

$$\sin(\theta_{n,\beta}) = \frac{r(t)}{n} \quad (32)$$

and

$$\cos(\theta_{n,r(t)}) = \frac{r(t)}{n}, \quad (33)$$

such that

$$\sin(\theta_{n,\beta}) = \cos(\theta_{n,r(t)}). \quad (34)$$

7 Dynamic Bremsstrahlung - Special Relativity & Doppler effect

Now the Doppler shift is added to the emitted radiation.

Figure 6 displays a stationary laboratory frame ω where the relativistic effects and the Doppler shift is observed, particle's moving frame of reference ω' and the relative velocity $c\beta S_{SpecialR}$ between the observer and the two different frames that determines what an observer would detect. In addition, $S_{SpecialR}$ (dimensionless) is a scaling factor of the relative velocity between the two different frames of references. In particular, $S_{SpecialR} = 1$ means the observer is stationary, and the particle is propagating at its velocity. As the $S_{SpecialR}$ approaches zero, the relative velocity between the observer and the particle falls, meaning that the observer is getting closer to the particle's velocity and would start to observe or detect a non-relativistic dipole radiation pattern.

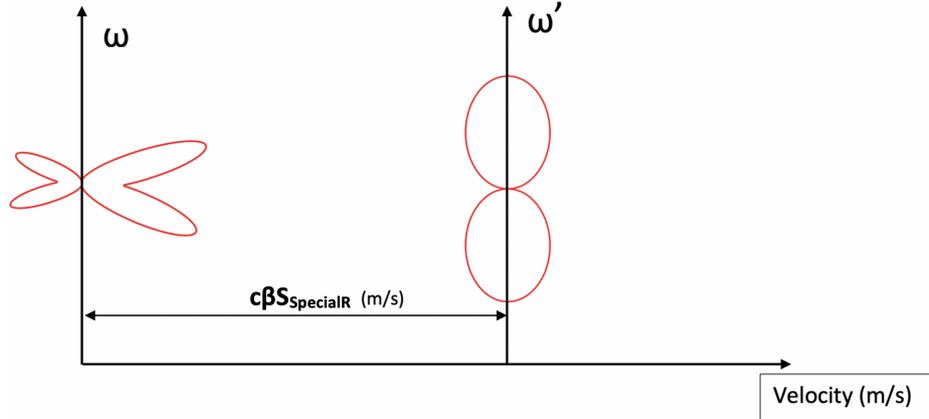


Figure 6. Particle radiation at two different frames of references. Laboratory (stationary) frame of reference ω and moving frame of reference ω' with the velocity $c\frac{v(t)}{c}S_{SpecialR}$. In addition, $S_{SpecialR} = 0$ meaning the observer is moving with the particle at the same velocity at all times, hence would see no radiation emission as the particle is not accelerating. As $S_{SpecialR} \rightarrow 0$, the observer would see dipole radiation pattern. As $S_{SpecialR} \rightarrow 1$, the observer would approach rest and would see the particle propagating at its own velocity defined by equation (9) with the corresponding radiation pattern.

The equation for translating the angular frequency of the emitted wave into the laboratory frame for the Doppler shift is given by (Jackson, 1999, p. 720, eq. 15.40).

$$\omega' = \gamma\omega(S_{SpecialR} - \beta S_{SpecialR}\cos(\theta_{n,\beta})). \quad (35)$$

As shown in Figure 2b, $\theta_{n,\beta}$ is the angle between the emitted radiation unit vector and the particle velocity vector.

Substituting equation 34, and 35 into the equation 25 gives

$$\frac{d^2 I}{d\omega \Omega_{rad}} = \frac{z^2 e^2 (\gamma \omega (S_{SpecialR} - \beta S_{SpecialR} \cos(\theta_{n,\beta})))^2}{4\pi^2 c \epsilon_0} \left| \sin(\theta_{n,\beta}) \left[-\frac{s_{fv} s_f z (s_{ft})^{1.461} 4.365 \times 10^{26}}{c} \right. \right. \\ \left. \left. + \frac{s_{fv} z s_f (s_{ft})^{1.5} 1.565 \times 10^{27}}{c} \left[\pi^{1/2} 2^{-(1/2)\nu_1} \alpha^{-\nu_1-1} e^{-\frac{y^2 \alpha^{-2}}{8}} \times D_{\nu_1}(2^{-1/2} \alpha^{-1} y) \right] + \right. \right. \\ \left. \left. \frac{s_{fv} z s_f (s_{ft})^{1.5} 1.565 \times 10^{27}}{c} \left[\pi^{1/2} 2^{-(1/2)\nu_2} \alpha^{-\nu_2-1} e^{-\frac{y^2 \alpha^{-2}}{8}} \times D_{\nu_2}(2^{-1/2} \alpha^{-1} y) \right] \right] \right|^2, \quad (36)$$

where the Lorentz factor γ is $\gamma = \frac{1}{\sqrt{1-\beta^2}}$ (dimensionless), and ω is the received angular frequency in the laboratory (stationary) frame of reference in rad/s. In addition, the new definition of α is $\alpha^2 = \frac{b^R (\gamma \omega (S_{SpecialR} - \beta S_{SpecialR} \cos(\theta_{n,\beta})))^R (\sin(\theta_{n,\beta}))^R}{(\tau^R)^2 c^R} [s^{-2}]$, and the new definition of y is $y = \frac{\gamma \omega (S_{SpecialR} - \beta S_{SpecialR} \cos(\theta_{n,\beta})) \sin(\theta_{n,\beta}) a}{c \tau} [s^{-1}]$.

The scaling factor s_f in equation 36 is carried over from the definition of the electric field equation 5. As a result of the integration process, s_f scales the magnitude of the emitted radiation, which is directly linked with its previous purpose of scaling the electric field, which defines the magnitude of the emitted radiation. The scaling factor s_f can be approximated at high frequency with a scalar equation that gives the magnitude of the emitted high-frequency radiation from a single particle in an instantaneous circular motion (Jackson, 1999, p. 679, eq. 14.84).

Hence,

$$s_f \simeq \frac{\frac{d^2 I}{d\Omega d\omega} |_{\theta=0}}{\left| \frac{d^2 I}{d\Omega d\omega} \right|} \simeq \frac{\frac{3e^2 \gamma^2 \omega e^{-\frac{\omega}{\omega_c}}}{4\pi c \omega_c}}{\left| \frac{d^2 I}{d\Omega d\omega} \right|} \quad (37)$$

and the critical angular frequency (Jackson, 1999, p. 679, eq. 14.81) is

$$\omega_c = \frac{3\gamma^3}{2} \left(\frac{c}{p} \right), \quad (38)$$

where p is the radius of curvature of instantaneous circular motion in m, ω_c is the critical angular boundary frequency. Beyond critical angular boundary frequency, the emitted radiation would have a minimal value in all directions such that it can be neglected.

8 Results

This section presents predictions of radiation patterns of a single electron accelerated under an external lightning leader tip electric field using the derived equation 36.

Radiation patterns change as a result of changing particle velocity, and acceleration as a particle get relativistic. As the acceleration of a charged particle increases, it emits higher intensity, hence higher frequency radiation. Therefore, the input parameter, emitted radiation frequency represents particle acceleration. To simulate relativistic and non-relativistic radiation patterns, the range of input velocity and frequency values were chosen and tabulated in table 1.

<i>Velocity is shown as ratio with the speed of light β and emitted radiation frequency (Hz)</i>		
Radiation Patterns	Scaled Bremsstrahlung Electron Velocity β (Dimensionless)	Emitted Radiation Frequency (Hz)
Pattern 1	0.02	1 k
Pattern 2	0.36	10 k
Pattern 3	0.67	1 M
Pattern 4	0.69	7 M
Pattern 5	0.77	10 M
Pattern 6	0.85	100 M
Pattern 7	0.87	500 M
Pattern 8	0.90	1 G

Table 1. Radiation pattern input parameters are velocity as a ratio to the speed of light β , and emitted radiation frequency by the particle for eight different radiation patterns. Emitted radiation frequency by a single particle is an input parameter to the final equation (36). It plays a crucial role in delivering valuable information and determining the particle's total acceleration, hence the total velocity is indirect as there is no input acceleration parameter in the final equation 36. In addition, β function determines the particle's total velocity arising significantly from the external lightning leader tip electric field. The remaining contribution to the overall velocity comes from the Coulomb force between an incoming and target particle, where the target particle causes an incoming particle to follow and covers some arc length of the spiral trajectory described in the position vector (Eq. 1) during its mean free time of the collision.

477 The new radiation patterns reveal that as well as forward peaking, there are also backward peaking
 478 lobes as the particle gains speed. The magnitude of the acceleration determines the magnitude of the
 479 radiation intensity. The horizontal axis is the radiation intensity per solid angle per emitted frequency range
 480 in eight different radiation patterns displayed in Figure 7. Although there is overall more energy in the forward
 481 peaking lobes, the peak radiation intensity in the backward direction is slightly higher in patterns 5, 6, and 7
 482 as they are more horizontally projected. However, the total radiation intensity is higher in the forward
 483 direction as the peaking lobes have larger beam areas due to the Doppler effect's presence.

484 The peak intensity of the backward lobes starts to decay as a result of the Doppler shift from pattern
 485 seven onwards. The Doppler shift effect can be observed from the consideration of both radiation length
 486 in the horizontal direction (as the horizontal axis is a measure of the radiation intensity per solid angle,
 487 i.e., more horizontally projected) and the total area of the emitted radiation pattern. The larger the area
 488 of the radiation pattern, the more the beam extends in the horizontal direction or, when it is more projected
 489 towards the horizontal axis, the more energy the emitted wave has. An increasing particle velocity
 490 translates into increasing radiation frequency. Therefore an increase in energy in the forward direction
 491 of the particle lowers the frequency and hence lowers the energy at the backward direction of the particle
 492 due to the Doppler effect. For example, in the eighth radiation pattern, which is at $\sim 94\%$ of the speed of light
 493 and has 1 GHz frequency, increasing the frequency of the observed radiation leads to an increase in the Doppler
 494 effect that reduces the frequency, hence the energy ($E = hf$) of the beam at the back of the particle.
 495 As a consequence, backward radiation lobes start to shrink such that overall radiation
 496 pattern looks like a basic forward peaking radiation pattern (the pattern 8).

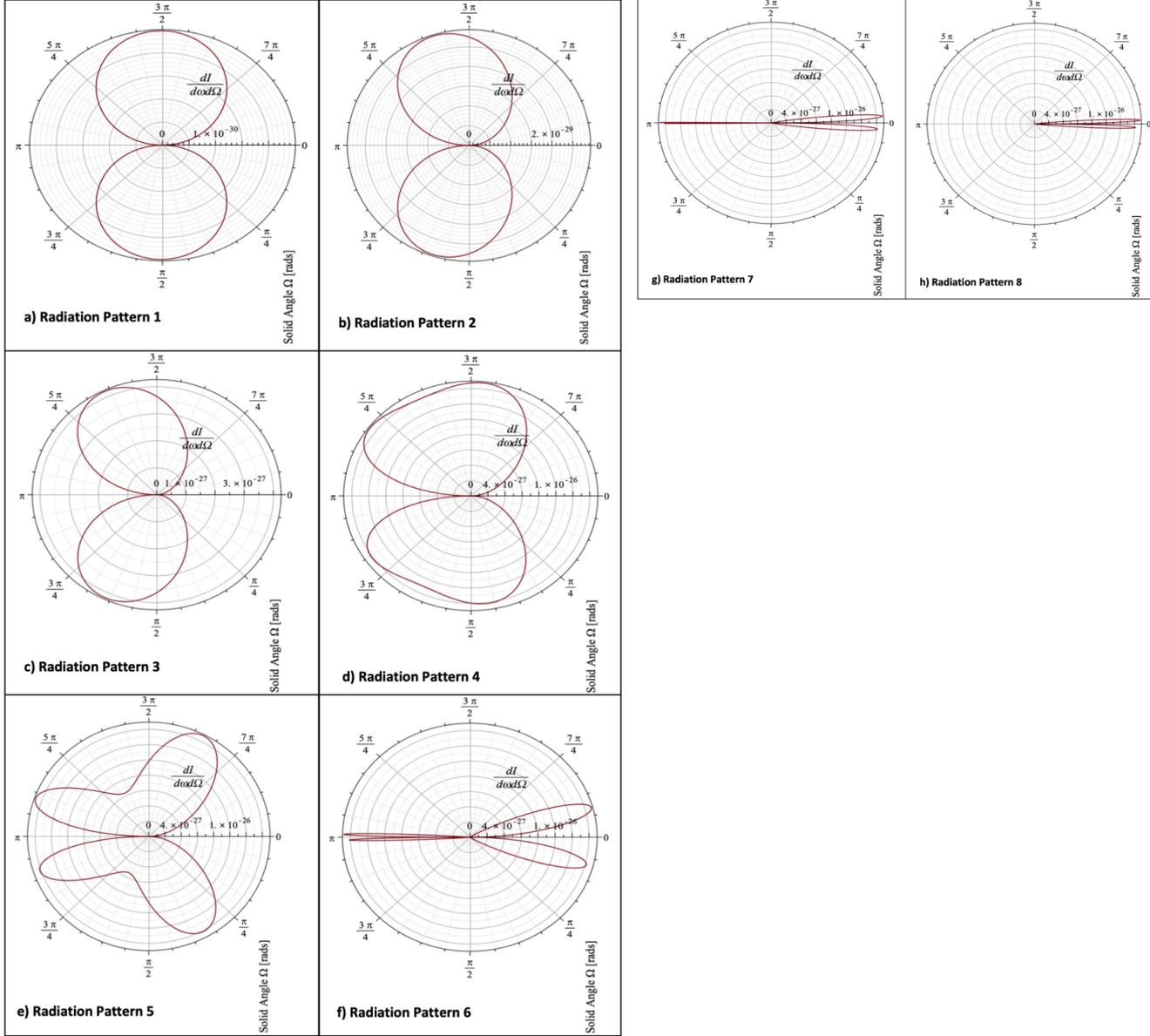


Figure 7. The Radiation patterns emitted by anti-clockwise rotating charged particle - bremsstrahlung process. Plot is in Polar co-ordinates. Horizontal axis gives the radiation intensity per Solid angle, Ω , per emitted angular radiation frequency, ω . In addition, angle of the Polar plot is the Solid angle, Ω . a-f) The radiation pattern connecting dipole with a forward-backward peaking radiation pattern is called the transition pattern. The radiation patterns and formula explaining the radiation patterns is not complete if it doesn't demonstrate the transition patterns. The transition patterns explain how the particle attains relativistic forward-backward peaking lobes. The radiation pattern of a particle during the bremsstrahlung, starting from a low velocity up until a relativistic velocity starts from the dipole radiation pattern and exhibits the forward and backward peaking radiation pattern. The transition is demonstrated in radiation patterns 4 and 5, where the dipole collapses to form four maxima. The values used for plotting are: mean free time $\tau = 30 \mu\text{s}$, number of charges $z = 1$, $a = 100 \mu\text{m}$, $b = 1 \text{ nm}$ (a and b are related to mean free path), $s_{ft} = 1$, $s_f = 1$, $S_{SpecialR} = 1$, velocity-time scaling factor $s_{ftv} = 1 \times 10^9$ and velocity scaling factor $s_{fv} = 8.19 \times 10^{-11}$. Finally, the bremsstrahlung asymmetry is $R = 1/8$. In addition, $\frac{1}{9} \leq R \leq \frac{1}{3}$. g) A particle reaching relativistic speed attains a focused beam in both forward and backward direction. h) A particle reaching to ultra-relativistic speed experiences the take over of the Doppler effect that minimizes the low-frequency radiation in the backward direction and emits almost fully in the forward direction.

497 Radiation patterns on the polar plot have to start from 2π rad until 0 rad to show
 498 the particle's complete radiation. Progression towards the zero is because a particle starts
 499 from the retarded time and radiates until the chronological time $t=0$ s and beyond, as
 500 shown in the Figures 2 and 5. If the polar plot were made with time as a representation
 501 of the solid angle, the plot would have started from ~ -10 ns (when the particle starts
 502 accelerating, see the Figures 2 and 5) until 0 s. In addition, this is the same as starting
 503 from positive time and ending at $t = 0$ s. Therefore, the final result is independent of
 504 positive or negative start time as long as the time flow is towards zero seconds. The in-
 505 dependence of the sign is caused by the absolute time and symmetry of the velocity func-
 506 tion. Progression towards zero seconds is caused by the initial start time being retarded
 507 negative time.

508 The derived mathematical model (Eq. 36) is specifically for the bremsstrahlung proc-
 509 cess. Unlike many of the other radiation emission processes of charged particles (i.e., from
 510 linear acceleration or crossing the boundary between two different dielectric media), the
 511 bremsstrahlung process affects the shape of the emitted radiation. This bremsstrahlung
 512 effect causes an asymmetry of the emitted radiation about the particle's velocity vec-
 513 tor or, in other terms, the direction of motion. To understand this effect, we can compare an electron
 514 to a car that travels in the dark with the headlight turned on. The headlights are the emitted electromagnetic
 515 radiation by the car. When the car gets into the bend, like the bremsstrahlung process of an electron,
 516 an observer outside the car can immediately tell the radiation shape of the headlight would
 517 be asymmetric by looking at the reflections on the road compared to the case when the car follows
 518 a straight path. This novel effect is clearly shown in supporting information in the Fig-
 519 ure S1, which displays a real-life example of visible asymmetry of the car headlights (electromagnetic radiation)
 520 on a bend.

521 Figure 8 displays the effects of the bremsstrahlung radiation asymmetry control
 522 quantity R . The bremsstrahlung asymmetry R depends on the radius of the curvature
 523 of the incoming particle trajectory undergoing the bremsstrahlung process. In addition,
 524 the bremsstrahlung asymmetry R increases as the radius of curvature of the bremsstrahlung
 525 trajectory decreases.

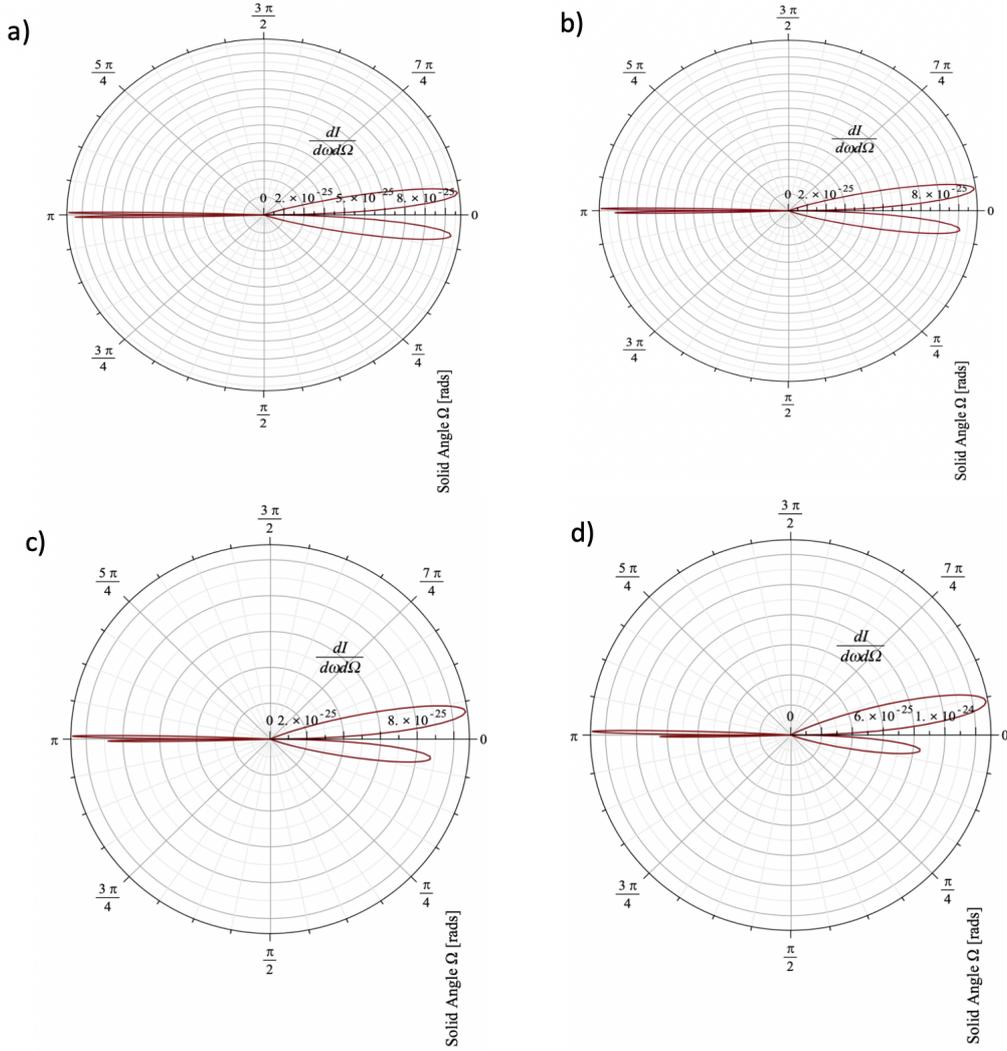


Figure 8. Bremsstrahlung Asymmetry Quantity R with different values. a) $R = \frac{1}{9}$, b) $R = \frac{1}{6}$, c) $R = \frac{1}{4}$, d) $R = \frac{1}{3}$

9 Discussion

Firstly, the created model describing radiation patterns specifically for the bremsstrahlung process is modeled to be generic and applicable to all bremsstrahlung process on earth, outer space, or other planets. As mentioned in the third paragraph of the introduction, TGFs are believed to originate as a result of the bremsstrahlung process of RREA driven by external leader tip electric field. Hence, the theoretical model assumes a single particle accelerated under an external leader tip electric field interacting with a target particle or ion via Coulomb field. In addition, the incoming particle is assumed to follow a spiral particle trajectory. This assumption is based on the in-balance of centripetal (Coulomb field between two particles + Earth's external magnetic field) and centrifugal (external leader tip electric field) forces acting on the incoming particle when it interacts with a target particle via Coulomb field. Any in-balance in centripetal and centrifugal force acting on an incoming particle that follows a curved trajectory would result in a spiral particle trajectory. Hence, the choice of spiral trajectory considers the effects of the Earth's external magnetic field, B on an incoming particle trajectory in an indirect way.

541 Particle speed defined in equation (9) relies on the ansatz of a sharp rise and slow
 542 decay form of leader tip electric field. This ansatz originates and is backed up by the asym-
 543 metric γ -ray bursts (GRBs) measurement that reveals the internal charge structure of
 544 the leader tip described in the introduction section, in the second paragraph.

545 Although acceleration is modeled from the electric field that assumes the same be-
 546 havior as the leader tip electric field observed in the atmosphere, it only provides a mag-
 547 nitude component of the electric field acceleration. Independent input parameters such
 548 as particle trajectory, the bremsstrahlung asymmetry, R , and the Doppler effect are fun-
 549 damental physics where the particle trajectory, and the bremsstrahlung asymmetry, R
 550 are specific to the bremsstrahlung process. These input parameters enable the created
 551 model to apply to any bremsstrahlung processes.

552 The bremsstrahlung process has a continuous electromagnetic spectrum meaning
 553 it is capable of explaining both ionizing and low-frequency radiations. The bremsstrahlung
 554 radiation patterns presented in figure 7 are unique to a bremsstrahlung process. They
 555 are always the same in geometry with forward-backward radiation patterns at relativistic
 556 particle speeds and all the asymmetries. However, when the model is applied to any
 557 other bremsstrahlung processes other than Terrestrial-Gamma ray flashes (TGFs), only
 558 radiation intensities are subject to change but not the radiation shapes.

559 Overall, the eight different radiation patterns plotted in Figure 7 are not the com-
 560 plete radiation patterns, especially for high-frequency radiation. As mentioned in the pre-
 561 vious section, a high-frequency radiation pattern is sensitive to the interaction process
 562 described by scattering cross-sections that an incoming bremsstrahlung electron expe-
 563 riences against target particles. Moreover, plotted radiation patterns are the second-order
 564 derivative term $\frac{d^2 I}{d\omega d\Omega_{rad}}$ of the equation (12), and it misses the information of interac-
 565 tion described by Coulomb scattering cross-section. In future work, a complete radia-
 566 tion pattern will be calculated using the complete equation (12), (36), and equation (1)
 567 provided in the supplemental information, S2 with appropriate form factors, and a di-
 568 mensional analysis to describe high-frequency emissions. Furthermore, the first part of
 569 the derived equation demonstrates a success as Figures 7a (patterns 1) and 7h (pattern
 570 8), display the common expected radiation patterns such as low-frequency dipole radi-
 571 ation pattern and relativistic forward peaking radiation patterns (Jackson, 1999, p. 669,
 572 Fig. 14.4).

573 The dipole radiation pattern 1 in Figure 7a displays perfect symmetry about both
 574 the velocity axis and the axis perpendicular to the velocity vector. In the absence of a
 575 Doppler and bremsstrahlung asymmetric radiation effect, the four maxima of the rela-
 576 tivistic radiation pattern are symmetric about both the velocity axis and the axis per-
 577 pendicular to the velocity. Therefore, following the statement of Noether's theorem, con-
 578 servation of energy arising from the symmetry in the radiation lobes about the two axes
 579 mentioned above, causes the dipole radiation pattern 4 and 5 (in Figure 7d, 7e) to col-
 580 lapse and form 4 symmetric maxima. Initially, there were symmetries about two axes,
 581 and as the particle gains up speed, the number of symmetric axes should still be preserved.
 582 Hence, if there is only a forward peaking radiation with none in the backward direction,
 583 one of the symmetry is broken (about axes perpendicular to the velocity). The idea of
 584 preserving the initial number of symmetric axes is quite powerful in understanding why
 585 there are four maxima (two forward peakings and two backward peakings) and why the
 586 dipole collapses to form 4 maxima of the radiation pattern. All the radiation patterns
 587 in Figure 7 incorporate the Doppler effect and the bremsstrahlung trajectory asymme-
 588 try. These two effects cause the radiation pattern to be asymmetric. However, the emit-
 589 ted total radiation energy is still conserved compared between the presence and absence
 590 of radiation pattern asymmetry causing physical effects (The bremsstrahlung trajectory
 591 and the Doppler effect) at any given particle velocity. In other words, the total energy
 592 radiated at a given velocity is the same regardless of the Doppler and the bremsstrahlung
 593 trajectory effect. While these two effects increase the radiation beam intensity in one di-

594 rection, they also decrease the radiation intensity in other directions by the same amount
595 such that the overall energy radiated by the particle at a given speed remains the same.

596 On the other hand, the reported radiation asymmetry of the bremsstrahlung (R-
597 parameter) about the velocity vector due to the particle's curved trajectory was found
598 to be existing at relativistic speeds when forward and backward peaking radiation pat-
599 terns are present. This bremsstrahlung radiation asymmetry (R - parameter) we get is
600 a novel finding. At the non-relativistic speeds, when radiation has a dipole pattern, asym-
601 metry of the bremsstrahlung was found to be absent. This absence indicates that asym-
602 metry is associated with being a physical effect rather than a mathematical artifact. The
603 absence of the bremsstrahlung radiation asymmetry at a non-relativistic dipole radia-
604 tion pattern of a particle following a curved trajectory can again be compared to a car
605 having its headlights on each side of the door radiating perpendicular with respect to
606 the velocity vector of the car. In this case, one can see that two emitted radiation beams
607 would not be asymmetric as two beams are in the opposite direction to each other and
608 are independent of the curvature of the trajectory as headlights radiate perpendicular
609 to the trajectory.

610 Asymmetries are important because they alter the radiation intensity of the emit-
611 ted radiation. Asymmetries increase the radiation intensity in one direction and decrease
612 it in the other direction. The bremsstrahlung radiation pattern of a single particle with
613 involved asymmetries can be used with a Monte Carlo simulation for multiple particles
614 to predict radiation intensity map across an area where lightning discharge takes place.
615 The radiation intensity map is the expected radiation power at each different observa-
616 tion location around the area of lightning discharge location. The predicted radiation
617 intensity map allows improving the lightning detection and location systems. It helps
618 to find the location of peak radiation intensity, hence contributes to the development of
619 better radiation protection systems.

620 10 Summary

621 In summary, the radiation patterns were found to be peaking in a backward direc-
622 tion as well as an already known forward direction. In addition, four maxima (2 in the
623 forward direction and 2 in the backward direction) were found to be due to the conser-
624 vation of symmetry axes arising from the initial dipole radiation pattern. With the in-
625 troduction of the Doppler effect, peaking lobes of radiation in forward and backward di-
626 rections were found to be asymmetric about an axis perpendicular to the particle's ve-
627 locity vector. Moreover, the novel second asymmetry of peaking lobes with respect to
628 particle's velocity vector was found to be unique to the bremsstrahlung due to the par-
629 ticle following a curved trajectory. Finally, it is also found that low-frequency radiation
630 peaks in the backward direction, whereas high-frequency peaks in forward direction.

631 Acknowledgments

632 I would like to thank my supervisor Dr.Martin Fflekrug for all the opportunities, sug-
633 gestions, guidance, reviews, and support throughout my first and second years of PhD.
634 EPSRC and MetOffice sponsor my PhD project under contract numbers EG-EE1239 and
635 EG-EE1077. MF acknowledge support from the Natural Environment Research Coun-
636 cil (NERC) for under grants NE/L012669/1 and NE/H024921/1. MY wishes to thank
637 Dr. Adrian Hill for mathematical support in helping me solve integral with a divergence
638 problem and my family for their support and good wishes. The Maple worksheets used
639 to simulate the particle trajectory, external lightning leader tip electric field, particle ve-
640 locity, and the radiation patterns are openly available from the University of Bath Re-
641 search Data Archive at <https://doi.org/10.15125/BATH-00810>.

References

- 642
- 643 Abramowitz, M., & Stegun, I. A. (1972). Handbook of Mathematical Functions:
644 with Formulas, Graphs, and Mathematical Tables. In (10th ed., Vol. National
645 Bureau of Standards Applied mathematics series 55, p. 1046). Washington,
646 D.C., USA: U.S. Dept. of Commerce : U.S. G.P.O. (ISBN: 9781591242178)
- 647 Babich, L. P., Bochkov, E. I., & Kutsyk, I. M. (2014). Mechanism of generation of
648 runaway electrons in a lightning leader. *JETP Letters*, 99(7), 386–390. Re-
649 trieved from <https://doi.org/10.1134/S0021364014070029> doi: 10.1134/
650 S0021364014070029
- 651 Babich, L. P., Bochkov, E. I., Kutsyk, I. M., Neubert, T., & Chanrion, O. (2015a).
652 A model for electric field enhancement in lightning leader tips to levels allow-
653 ing X-ray and ray emissions. *Journal of Geophysical Research: Space Physics*,
654 120(6), 5087–5100. Retrieved from [https://agupubs.onlinelibrary.wiley](https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1002/2014JA020923)
655 [.com/doi/abs/10.1002/2014JA020923](https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1002/2014JA020923) doi: 10.1002/2014JA020923
- 656 Babich, L. P., Bochkov, E. I., Kutsyk, I. M., Neubert, T., & Chanrion, O. (2015b).
657 A model for electric field enhancement in lightning leader tips to levels al-
658 lowing xray and ray emissions. *Journal of geophysical research.*, 120(6),
659 5087–5100. doi: 10.1002/2014JA020923
- 660 Bateman, H., W.Magnus, F.Oberhettinger, & F.G.Tricomi. (1954). Table of Integral
661 Transforms. In (1st ed., Vol. 1, p. 121). New York, USA: McGraw-Hill Book
662 Company,Inc. (ISBN:07-019549-8)
- 663 Bourilkov, D. (2000, Sep). Search for tev strings and new phenomena in bhabha
664 scattering at cern lep2. *Phys. Rev. D*, 62, 076005. Retrieved from [https://](https://link.aps.org/doi/10.1103/PhysRevD.62.076005)
665 link.aps.org/doi/10.1103/PhysRevD.62.076005 doi: 10.1103/PhysRevD
666 .62.076005
- 667 Brock, C. P. (2019). Atomic form factors-international tables for crystallography. In
668 (6th ed., Vol. C, p. 554-595). Hoboken, United States: Wiley. (ISBN:978-1-119-
669 46870-7)
- 670 Celestin, S. (2016, December). Electron acceleration mechanisms in thunderstorms.
671 *arXiv*, 1-6. doi: arXiv:1701.00105[astro-ph.HE]
- 672 Celestin, S., & Pasko, V. P. (2011). Energy and fluxes of thermal runaway elec-
673 trons produced by exponential growth of streamers during the stepping
674 of lightning leaders and in transient luminous events. *Journal of Geo-*
675 *physical Research: Space Physics*, 116(A3). Retrieved from [https://](https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/2010JA016260)
676 agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/2010JA016260 doi:
677 10.1029/2010JA016260
- 678 Chanrion, O., & Neubert, T. (2010). Production of runaway electrons by negative
679 streamer discharges. *Journal of Geophysical Research: Space Physics*, 115(A6),
680 1-10. Retrieved from [https://agupubs.onlinelibrary.wiley.com/doi/abs/](https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/2009JA014774)
681 [10.1029/2009JA014774](https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/2009JA014774) doi: 10.1029/2009JA014774
- 682 Connaughton, V., Briggs, M. S., Xiong, S., Dwyer, J. R., Hutchins, M. L., Grove,
683 J. E., ... Wilson-Hodge, C. (2013, May). Radio signals from electron beams in
684 terrestrial gamma ray flashes. *Journal of Geophysical Research: Space Physics*,
685 118(5), 2313-2320. Retrieved from [https://agupubs.onlinelibrary.wiley](https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/2012JA018288)
686 [.com/doi/abs/10.1029/2012JA018288](https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/2012JA018288) doi: 10.1029/2012JA018288
- 687 Cummer, S. A., Briggs, M. S., Dwyer, J. R., Xiong, S., Connaughton, V., Fish-
688 man, G. J., ... Solanki, R. (2014, December). The source altitude, elec-
689 tric current, and intrinsic brightness of terrestrial gamma ray flashes. *Geo-*
690 *physical Research Letters*, 41(23), 8586-8593. Retrieved from [https://](https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1002/2014GL062196)
691 agupubs.onlinelibrary.wiley.com/doi/abs/10.1002/2014GL062196 doi:
692 10.1002/2014GL062196
- 693 Dwyer, J. R. (2007). Relativistic breakdown in planetary atmospheres. *Physics*
694 *of Plasmas*, 14(4), 042901. Retrieved from [https://doi.org/10.1063/](https://doi.org/10.1063/1.2709652)
695 [1.2709652](https://doi.org/10.1063/1.2709652) doi: 10.1063/1.2709652
- 696 Dwyer, J. R. (2008). Source mechanisms of terrestrial gamma-ray flashes. *Journal of*

- 697 *Geophysical Research: Atmospheres*, 113(D10), 1-12. Retrieved from [https://](https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/2007JD009248)
 698 agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/2007JD009248 doi:
 699 10.1029/2007JD009248
- 700 Dwyer, J. R., Smith, D. M., & Cummer, S. A. (2012a). High-energy atmospheric
 701 physics: Terrestrial gamma-ray flashes and related phenomena. *Space Science*
 702 *Reviews*, 173(1), 133–196. Retrieved from [https://doi.org/10.1007/s11214-](https://doi.org/10.1007/s11214-012-9894-0)
 703 [012-9894-0](https://doi.org/10.1007/s11214-012-9894-0) doi: 10.1007/s11214-012-9894-0
- 704 Dwyer, J. R., Smith, D. M., & Cummer, S. A. (2012b, June). High-Energy At-
 705 mospheric Physics: Terrestrial Gamma-Ray Flashes and Related Phenomena.
 706 *Space Science Reviews*, 173(1), 133–196. Retrieved from [https://doi.org/](https://doi.org/10.1007/s11214-012-9894-0)
 707 [10.1007/s11214-012-9894-0](https://doi.org/10.1007/s11214-012-9894-0) doi: 10.1007/s11214-012-9894-0
- 708 Fishman, G. J., Bhat, P. N., Mallozzi, R., Horack, J. M., Koshut, T., Kouve-
 709 liotou, C., ... Christian, H. J. (1994). Discovery of intense gamma-ray
 710 flashes of atmospheric origin. *Science*, 264(5163), 1313–1316. Retrieved
 711 from <https://science.sciencemag.org/content/264/5163/1313> doi:
 712 10.1126/science.264.5163.1313
- 713 Foley, S., Fitzpatrick, G., Briggs, M. S., Connaughton, V., Tierney, D., McBreen, S.,
 714 ... Wilson-Hodge, C. (2014, June). Pulse properties of terrestrial gamma-ray
 715 flashes detected by the Fermi Gamma-Ray Burst Monitor. *Journal of Geo-*
 716 *physical Research: Space Physics*, 119(7), 5931-5942. Retrieved from [https://](https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1002/2014JA019805)
 717 agupubs.onlinelibrary.wiley.com/doi/abs/10.1002/2014JA019805 doi:
 718 10.1002/2014JA019805
- 719 Grefenstette, B. W., Smith, D. M., Dwyer, J. R., & Fishman, G. J. (2008, March).
 720 Time evolution of terrestrial gamma ray flashes. *Geophysical Research Letters*,
 721 35(6), 1-5. Retrieved from [https://agupubs.onlinelibrary.wiley.com/](https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/2007GL032922)
 722 [doi/abs/10.1029/2007GL032922](https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/2007GL032922) doi: 10.1029/2007GL032922
- 723 Hariharan, B., Chandra, A., Dugad, S. R., Gupta, S. K., Jagadeesan, P., Jain, A.,
 724 ... Tanaka, K. (2019, Mar). Measurement of the electrical properties of a
 725 thundercloud through muon imaging by the grapes-3 experiment. *Phys. Rev.*
 726 *Lett.*, 122, 105101. Retrieved from [https://link.aps.org/doi/10.1103/](https://link.aps.org/doi/10.1103/PhysRevLett.122.105101)
 727 [PhysRevLett.122.105101](https://link.aps.org/doi/10.1103/PhysRevLett.122.105101) doi: 10.1103/PhysRevLett.122.105101
- 728 Heidler, F. H., Manhardt, M., & Stimper, K. (2013, April). The Slow-Varying Elec-
 729 tric Field of Negative Upward Lightning Initiated by the Peissenberg Tower,
 730 Germany. *IEEE Transactions on Electromagnetic Compatibility*, 55(2), 353-
 731 361. doi: 10.1109/TEMC.2012.2209121
- 732 Hurley, K., Dingus, B. L., Mukherjee, R., Sreekumar, P., Kouveliotou, C., Meegan,
 733 C., ... Niel, M. (1994, December). Detection of a γ -ray burst of very long
 734 duration and very high energy. *Nature*, 372(6507), 652–654. Retrieved from
 735 <https://doi.org/10.1038/372652a0> doi: 10.1038/372652a0
- 736 Jackson, J. D. (1999). Classical Electrodynamics. In (3rd ed., p. 661-732). New
 737 York, United States: John Wiley & Sons, Inc. (ISBN:9780471309321) doi: 10
 738 .1002/3527600434.eap109
- 739 Kiyosi Ito and The Mathematical Society of Japan. (1993). Encyclopedic Dictionary
 740 of Mathematics. In (2nd ed., Vol. 1, p. 2168). London, England: The MIT
 741 Press. (ISBN: 0262590204)
- 742 Koch, H. W., & Motz, J. W. (1959, Oct). Bremsstrahlung cross-section formulas
 743 and related data. *Rev. Mod. Phys.*, 31, 920–955. Retrieved from [https://](https://link.aps.org/doi/10.1103/RevModPhys.31.920)
 744 link.aps.org/doi/10.1103/RevModPhys.31.920 doi: 10.1103/RevModPhys
 745 .31.920
- 746 Moss, G. D., Pasko, V. P., Liu, N., & Veronis, G. (2006). Monte Carlo model
 747 for analysis of thermal runaway electrons in streamer tips in transient lu-
 748 minous events and streamer zones of lightning leaders. *Journal of Geo-*
 749 *physical Research: Space Physics*, 111(A2), 1-37. Retrieved from [https://](https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/2005JA011350)
 750 agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/2005JA011350 doi:
 751 10.1029/2005JA011350

- 752 Nemiroff, R. J., Norris, J. P., Kouveliotou, C., Fishman, G. J., Meegan, C. A., &
 753 Paciasas, W. S. (1994, March). Gamma-Ray Bursts Are Time-asymmetric.
 754 *Astrophysical Journal*, *423*, 432-435. doi: 10.1086/173819
- 755 Norris, J. P., Nemiroff, R. J., Bonnell, J. T., Wickramasinghe, W. A. D. T., Kou-
 756 veliotou, C., Paciasas, W. S., ... Meegan, C. A. (1994, November). Gross
 757 Spectral Differences between Bright and DIM Gamma-Ray Bursts. *Astrophysi-
 758 cal Journal Letters*, *435*, L133. doi: 10.1086/187612
- 759 Perdrisat, C., Punjabi, V., & Vanderhaeghen, M. (2007, Oct). Nucleon electromag-
 760 netic form factors. *Progress in Particle and Nuclear Physics*, *59*(2), 694764.
 761 Retrieved from <http://dx.doi.org/10.1016/j.pnnp.2007.05.001> doi:
 762 10.1016/j.pnnp.2007.05.001
- 763 Pu, Y., Cummer, S. A., Lyu, F., Briggs, M., Mailyan, B., Stanbro, M., & Roberts,
 764 O. (2019, June). Low Frequency Radio Pulses Produced by Terrestrial
 765 Gamma-Ray Flashes. *Geophysical Research Letters*, *46*(12), 6990-6997.
 766 Retrieved from [https://agupubs.onlinelibrary.wiley.com/doi/abs/
 767 10.1029/2019GL082743](https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/2019GL082743) doi: 10.1029/2019GL082743
- 768 Punjabi, V., & Perdrisat, C. (2014, March). The Proton Form Factor Ratio
 769 Measurements at Jefferson Lab. *EPJ Web of Conferences*, *66*, 1-4. Re-
 770 trieved from <https://doi.org/10.1051/epjconf/20146606019> doi:
 771 10.1051/epjconf/20146606019
- 772 Skeltved, A. B., stgaard, N., Mezentsev, A., Lehtinen, N., & Carlson, B. (2017).
 773 Constraints to do realistic modeling of the electric field ahead of the tip of a
 774 lightning leader. *Journal of Geophysical Research: Atmospheres*, *122*(15),
 775 8120-8134. Retrieved from [https://agupubs.onlinelibrary.wiley.com/
 776 doi/abs/10.1002/2016JD026206](https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1002/2016JD026206) doi: 10.1002/2016JD026206
- 777 Whittaker, E. T., & Watson, G. N. (1927). A course in modern analysis : an intro-
 778 duction to the general theory of infinite processes and of analytical functions,
 779 with an account of the principal transcendental functions. In (4th ed., p. 616).
 780 The Pitt Building, Trumpington Street, Cambridge, CB2 1RP: Cambridge
 781 University Press. (ISBN: 0521091896)
- 782 Xu, W., Celestin, S., & Pasko, V. P. (2015, January). Optical emissions as-
 783 sociated with terrestrial gamma ray flashes. *Journal of Geophysical Re-
 784 search: Space Physics*, *120*(2), 1355-1370. Retrieved from [https://
 785 agupubs.onlinelibrary.wiley.com/doi/abs/10.1002/2014JA020425](https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1002/2014JA020425) doi:
 786 10.1002/2014JA020425
- 787 Xu, W., Celestin, S., Pasko, V. P., & Marshall, R. A. (2019, August). Compton
 788 Scattering Effects on the Spectral and Temporal Properties of Terrestrial
 789 Gamma-Ray Flashes. *Journal of Geophysical Research: Space Physics*, *124*(8),
 790 7220-7230. Retrieved from [https://agupubs.onlinelibrary.wiley.com/
 791 doi/abs/10.1029/2019JA026941](https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/2019JA026941) doi: 10.1029/2019JA026941
- 792 stgaard, N., Neubert, T., Reglero, V., Ullaland, K., Yang, S., Genov, G., ... Al-
 793 nussirat, S. (2019, December). First 10 Months of TGF Observations by
 794 ASIM. *Journal of Geophysical Research: Atmospheres*, *124*(24), 14024-14036.
 795 Retrieved from [https://agupubs.onlinelibrary.wiley.com/doi/abs/
 796 10.1029/2019JD031214](https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/2019JD031214) doi: 10.1029/2019JD031214