

# Probing the southern African lithosphere with magnetotellurics, Part I, model construction

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November 26, 2022

## Abstract

The Southern African Magnetotelluric Experiment (SAMTEX) involved the collection of data at over 700 sites in Archean to Proterozoic southern Africa, spanning features including the Kalahari Craton, Bushveld Complex and voluminous kimberlites. Here, we present the first 3D inversions of the full SAMTEX dataset. In this paper, we focus on assessing the robustness of the 3D models by comparing two different inversion codes, jif3D and ModEM, and two different subsets of the data, one containing all acceptable data and the other containing a smaller selection of undistorted, high-quality data. Results show that the main conductive and resistive features are imaged by all inversions, including deep resistive features in the central Kaapvaal Craton and southern Congo Craton and a lithospheric-scale conductor beneath the Bushveld Complex. Despite this, differences exist between the jif3D and ModEM inverse models that derive mainly from the differences in regularization between the models, with jif3D producing models that are very smooth laterally and with depth, while ModEM produces models with more discrete conductive and resistive features. Analysis of the differences between these two inversions can provide a good indication of the model resolution. More minor differences are apparent between models run with different subsets of data, with the models containing all acceptable data featuring higher wavelength conductivity variations than those run with fewer stations but also demonstrating poorer data fit.

# Probing the southern African lithosphere with magnetotellurics, Part I, model construction

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## Key Points:

- We create the first large-scale conductivity models of southern Africa.
- We compare different strategies to construct continental scale models and investigate the impact on the results.
- The main inversion features are common to all models, with most differences being due to regularization.

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**Abstract**

The Southern African Magnetotelluric Experiment (SAMTEX) involved the collection of data at over 700 sites in Archean to Proterozoic southern Africa, spanning features including the Kalahari Craton, Bushveld Complex and voluminous kimberlites. Here, we present the first 3D inversions of the full SAMTEX dataset. In this paper, we focus on assessing the robustness of the 3D models by comparing two different inversion codes, *jif3D* and *ModEM*, and two different subsets of the data, one containing all acceptable data and the other containing a smaller selection of undistorted, high-quality data. Results show that the main conductive and resistive features are imaged by all inversions, including deep resistive features in the central Kaapvaal Craton and southern Congo Craton and a lithospheric-scale conductor beneath the Bushveld Complex. Despite this, differences exist between the *jif3D* and *ModEM* inverse models that derive mainly from the differences in regularization between the models, with *jif3D* producing models that are very smooth laterally and with depth, while *ModEM* produces models with more discrete conductive and resistive features. Analysis of the differences between these two inversions can provide a good indication of the model resolution. More minor differences are apparent between models run with different subsets of data, with the models containing all acceptable data featuring higher wavelength conductivity variations than those run with fewer stations but also demonstrating poorer data fit.

**Plain Language Summary**

We investigate the structure of the upper 200 km of the Earth beneath southern Africa. To achieve this, we utilize an electromagnetic geophysical technique called magnetotellurics which is sensitive to variations in electrical resistivity within the Earth. To reconstruct electrical resistivity from magnetotelluric measurements, we use so-called inversion algorithms. However, the results are non-unique and a variety of different parameters have to be chosen by the user during the inversion process. In order to better understand the possible variability in our Earth models, we use different inversion algorithms and compare different strategies. This allows us to assess the reliability of our results. Based on our models and their comparison, we infer that the lithosphere, the solid outer shell of the Earth, varies in thickness below our study area and is thickest below central South Africa. In addition, we can detect remnants of past continental col-

49 lisions that have been preserved for hundreds of millions of years since this part of the  
50 world was assembled from the collision of various micro-continents.

## 51 **1 Introduction**

52 The lithosphere of southern Africa is among the most important in the world for  
53 understanding continental evolution (e.g., Lee et al., 2011). It contains extensive, Archean  
54 to Paleoproterozoic cratons, including the Kaapvaal, Zimbabwe and Congo cratons, which  
55 are also sampled by voluminous kimberlite magmatism (e.g., De Wit et al., 1992; Begg  
56 et al., 2009). Investigations of the geological, geochemical and geophysical nature of these  
57 cratons help us understand the formation and amalgamation of the Archean continen-  
58 tal lithosphere and the survival of that lithosphere to the present. The southern African  
59 lithosphere also hosts many of the world’s largest mineral deposits (Clifford, 1966), in-  
60 cluding the world’s largest platinum group element deposits in the Bushveld Complex  
61 (itself the world’s largest layered mafic intrusion (e.g., VanTongeren, 2018)), extensive  
62 kimberlite-hosted diamond deposits including the Kimberley, Venetia and Jwaneng de-  
63 posits (e.g., Field et al., 2008), and giant orogenic and placer gold deposits such as those  
64 in the Barberton Goldfields and Witwatersrand Basin (e.g., de Ronde & de Wit, 1994).  
65 Since the formation of many of these deposits involved lithospheric-scale processes, defin-  
66 ing the lithospheric architecture and composition of southern Africa not only helps our  
67 understanding of continental evolution but also aids mineral exploration.

68 Analysis of the vast mantle xenolith and xenocryst databases has shown spatially  
69 and temporally complex patterns of depletion and metasomatism of the southern African  
70 mantle (e.g., Griffin et al., 2003; Kobussen et al., 2009; Grégoire et al., 2003). In some  
71 cratonic regions, inferences from xenoliths and seismic data can seem contradictory. Many  
72 xenoliths have metasomatised and geochemically fertile compositions, while seismic to-  
73 mography models tend to show fast wave speeds extending to depths  $>200$  km and have  
74 been interpreted to represent deep, geochemically depleted lithospheric keels (e.g., Fouch  
75 et al., 2004; White-Gaynor et al., 2020). This apparent contradiction has led to the sug-  
76 gession that cratonic mantle xenoliths and xenocrysts from southern Africa may be un-  
77 representative of the cratonic mantle more generally (Griffin et al., 2009). In contrast,  
78 seismic receiver function data image several low-velocity anomalies within those south-  
79 ern African lithospheric keels, which indicate that broad metasomatism may be more widespread  
80 than suggested by the tomographic models (Sodoudi et al., 2013; Selway et al., 2015).

81 Here and in the accompanying manuscript, we add new magnetotelluric (MT) con-  
 82 straints to understanding the architecture, composition and evolution of the southern  
 83 African lithosphere. MT data are sensitive to mantle metasomatism, both through the  
 84 hydration of nominally anhydrous mantle minerals and through the precipitation of meta-  
 85 somatic minerals (e.g., Selway, 2014), and MT interpretations are generally consistent  
 86 with mantle xenolith and xenocryst compositions (Özaydın et al., 2021). Therefore, MT  
 87 models of southern Africa can provide new insights into the composition and metaso-  
 88 matism of the lithosphere. To do this, we have analysed and inverted the Southern African  
 89 Magnetotelluric Experiment (SAMTEX) database (Jones et al., 2009). These data were  
 90 collected between 2003–2008 and comprise more than 700 MT stations including broad-  
 91 band MT (BBMT) and long-period MT (LMT) measurements along more than 15,000  
 92 line km crossing South Africa, Botswana and Namibia.

93 In this paper, we describe the first 3D MT models of the entire SAMTEX dataset.  
 94 In contrast to when SAMTEX was collected, 3D MT inversions are now routine. How-  
 95 ever, different inversion codes use different regularizations, model discretizations and for-  
 96 ward modelling approaches, and the impact of these differences is not well understood.  
 97 Therefore, we have inverted the dataset with two different algorithms *jif3D* (Moorkamp  
 98 et al., 2011) and *ModEM* (Kelbert et al., 2014)) to compare results and ensure only the  
 99 most robust features are interpreted. In the accompanying paper, we interpret the mod-  
 100 els in terms of the composition and evolution of the southern African lithospheric man-  
 101 tle.

## 102 **2 The SAMTEX magnetotelluric dataset**

103 Magnetotellurics (MT) is a passive electromagnetic technique to infer the resistiv-  
 104 ity of the subsurface from measurements at the surface (Chave & Jones, 2012). Together  
 105 with seismic tomography and potential field methods (e.g., gravity, magnetics), MT is  
 106 one of the foremost geophysical techniques to image the structure of the lithosphere-asthenosphere  
 107 system. It has been used to investigate continental lithospheric structures in many re-  
 108 gions around the world (e.g., Jones, 1999; Gatzemeier & Moorkamp, 2005; Rao et al.,  
 109 2014; Wannamaker et al., 2017; Selway, 2018) and has been a component of large national  
 110 programs in the United States (Kelbert, 2019), Australia (Kirkby et al., 2020) and China  
 111 (S.-W. Dong et al., 2013). Based on simultaneous measurements of naturally occurring

112 variations of the electric field  $\mathbf{E}$  and the magnetic field  $\mathbf{B}$  we can estimate the frequency-  
 113 dependent, complex-valued magnetotelluric impedance  $\mathbf{Z}$ , viz.

$$\mathbf{E} = \mathbf{Z}\mathbf{B}. \quad (1)$$

114 This estimation process is based on robust statistical methods and thus gives for-  
 115 mal estimates of data uncertainties (e.g., Chave & Thomson, 2003) although these are  
 116 typically increased before use in an inversion algorithm (e.g., Miensopust, 2017, and dis-  
 117 cussion below).

118 The MT response estimates used for our inversions were acquired during the South-  
 119 ern African Magnetotelluric Experiment (SAMTEX) between 2003–2008 (Jones et al.,  
 120 2009). They comprise more than 700 broadband MT (BBMT) and long-period MT (LMT)  
 121 measurements across South Africa, Botswana and Namibia. The primary goal of SAM-  
 122 TEX was to image the lithospheric architecture of the cratons and mobile belts in the  
 123 region, and thus measurements were taken at intervals of roughly 20 km and at periods  
 124 0.01 - 10,000 s for the BBMT sites, and 60 km at 10 s - 10,000 s for the LMT sites. Due  
 125 to long recording times and favourable noise conditions in many parts of the study area,  
 126 the data quality at many sites is excellent throughout the period range. To date there  
 127 have been a number of publications performing modelling and inversions for different sub-  
 128 regions and profiles (Hamilton et al., 2006; Muller et al., 2009; Miensopust et al., 2011;  
 129 Evans et al., 2011; Khoza et al., 2013; Finn et al., 2015; Moorkamp et al., 2019). In ad-  
 130 dition, maps of resistivity directly derived from the data at selected periods have been  
 131 used to investigate the structure and composition of the lithosphere-asthenosphere sys-  
 132 tem (Jones et al., 2012, 2013) and multi-observable petrological-geophysical models have  
 133 been created based on subsets of the data (Fullea et al., 2011). Still, to date, no three-  
 134 dimensional resistivity models of the lithosphere based on all the available data have been  
 135 published.

136 We show representative data from six stations across the array in Figure 2. Their  
 137 locations are shown as yellow stars in Figure 1. We concentrate on the period range of  
 138 1–10,000 s as lithospheric-scale structures are the focus of our modelling efforts. The sites  
 139 show good data quality throughout the plotted period range, although at the longest pe-  
 140 riods ( $> 1,000$  s), the scatter and error estimates increase at some of the sites (e.g., KAP45).  
 141 The off-diagonal apparent resistivity sounding curves highlight a moderately conductive

142 shallow subsurface ( $\approx 10 \Omega\text{m}$  at short periods) at the northern sites RAK011A and ZIM117.  
 143 Further south, at sites KIM428 and KAL014, the short period apparent resistivities are  
 144 generally comparable but slightly higher, while the southernmost sites, KAP045 and KAP019,  
 145 show significantly higher short period apparent resistivities of  $\geq 500 \Omega\text{m}$ . Beneath this  
 146 surface layer, the sounding curves of the four northern sites suggest that resistivity in-  
 147 creases with depth to a maximum of  $\approx 500 \Omega\text{m}$  at  $\approx 500 \text{ s}$ , before either remaining con-  
 148 stant or decreasing slightly at longer periods. In contrast, at the two southern sites (KAP45  
 149 and KAP19), most of the apparent resistivity curves have higher initial values and de-  
 150 crease with period.

### 151 **3 Inversions**

#### 152 **3.1 Inversion Algorithms**

153 We invert the observed MT impedances with two different inversion algorithms (Moorkamp  
 154 et al., 2011; Kelbert et al., 2014) and two different strategies for each code. This helps  
 155 us to address model uncertainties related to algorithm-specific choices, such as regular-  
 156 ization, error floor, model discretization, and precision of the forward modelling engine.  
 157 As a result of these choices, each code will fit different aspects of the data, including those  
 158 affected by noise, to different degrees. As both algorithms used here are well established,  
 159 we only give a brief summary of each algorithm and focus on comparing the differences  
 160 and the potential impact on the results.

161 For two inversions, we use the MT inversion module of the joint inversion frame-  
 162 work *jif3D* (Moorkamp et al., 2011). The numerical basis of the forward modelling en-  
 163 gine and the gradient calculation are presented in Avdeev & Avdeeva (2009) and Avdeeva  
 164 et al. (2015). It utilizes an integral-equation based forward engine *x3d* (Avdeev et al.,  
 165 1997) and includes a correction for galvanic distortion at each site (Avdeeva et al., 2015;  
 166 Moorkamp et al., 2020). Galvanic distortion of magnetotelluric impedances is typically  
 167 caused by charge accumulation at small structures (compared to the induction length  
 168 scale), and it can mathematically be described as a site-specific, frequency-independent  
 169 multiplication of the impedances with a real-valued matrix  $\mathbf{C}$  (Chave & Jones, 2012).  
 170 In *jif3D* distortion correction is achieved by estimating the elements of  $\mathbf{C}$  as part of the  
 171 inversion and multiplying the synthetic impedance  $\mathbf{Z}_{synth}$  at each site with the corre-

172 sponding distortion matrix when calculating the data misfit. Details on this methodol-  
 173 ogy can be found in Avdeeva et al. (2015).

174 The inversion algorithm has been used on a range of MT datasets, both commer-  
 175 cial and academic, including imaging the fault structure for an intra-plate event in Botswana  
 176 (Moorkamp et al., 2019) and hydrothermal fluids in the central Andes (Pearce et al., 2020).  
 177 Due to the integral equation based forward modelling algorithm, the horizontal cell sizes  
 178 need to be constant in both orthogonal horizontal directions. In the vertical direction,  
 179 cell sizes can vary and are typically fine near the surface, increasing by a constant fac-  
 180 tor with depth to match the decreasing resolution of MT data. The discretized region  
 181 is embedded in a layered half-space that is kept constant throughout the inversion. Some  
 182 care must be taken to avoid the strong influence of this background conductivity struc-  
 183 ture on the inversion results. In order to enforce positive conductivity values during op-  
 184 timization and restrict model conductivities to realistic values, conductivities in each model  
 185 cell are transformed using the generalized model parameter scheme described in Moorkamp  
 186 et al. (2011). This allows us to use an unconstrained optimization algorithm based on  
 187 a limited-memory quasi-Newton method (L-BFGS, Avdeeva & Avdeev, 2006) to min-  
 188 imize the objective function. Within *jif3D* we regularize the inversion through a first-  
 189 order approximation of the spatial gradient of the generalized model parameters. This  
 190 approach has the advantage of equalizing the vast range of Earth conductivities ( $\approx 10^{-1}$   
 191 to  $10^6 \Omega m$ ) to a range between approximately -2 and 2, and ensuring that the regular-  
 192 ization operates similarly in all parts of the model. As the regularization is purely smooth-  
 193 ness based, it has the potential disadvantage that structures may horizontally or verti-  
 194 cally smear into regions of low resolution, such as those with poor site coverage in the  
 195 heterogeneous SAMTEX array. However, this could be considered a form of natural in-  
 196 terpolation.

197 The other two inversions were performed using *ModEM*, a well-established and freely  
 198 available 3D MT inversion code (Egbert & Kelbert, 2012; Kelbert et al., 2014). It is widely  
 199 used in the academic community and has seen applications on datasets around the world  
 200 (e.g., Kelbert & Egbert, 2012; Meqbel et al., 2014; H. Dong et al., 2020; Robertson et  
 201 al., 2020). Its forward engine is based on a finite-difference formulation (e.g., Mackie et  
 202 al., 1994; Egbert & Kelbert, 2012) and its modular structure allows for inversion of dif-  
 203 ferent combinations of electromagnetic data (e.g., Campanya et al., 2016). Compared  
 204 to *jif3D* the gridding requirements are less strict with variable-sized rectilinear cells in

205 all three coordinate directions. Furthermore, no background layered half-space is pre-  
 206 scribed; instead, the grid must be extensive enough that secondary electromagnetic fields  
 207 are insignificant at the model boundaries. Thus, a typical strategy for designing inver-  
 208 sion grids in *ModEM* is to use an inner core with constant horizontal cell size and padding  
 209 cells of increasing size around it. In *ModEM* the natural logarithm of conductivity is used  
 210 as a model parameter which enforces positivity of conductivity and has an equalizing ef-  
 211 fect similar to the generalized model parameters in *jif3D*. *ModEM* does not allow the  
 212 range of permitted conductivities to be directly limited, but the regularization limits the  
 213 difference from a prior model, often the starting model, and lateral variations of conduc-  
 214 tivity simultaneously (Egbert & Kelbert, 2012). Compared to a pure smoothing-based  
 215 regularization, this combined approach should reduce smearing but can result in arti-  
 216 ficial changes in conductivity if the prior model is not representative of the average con-  
 217 ductivity in the region. In this case, poorly-resolved regions of the model will be kept  
 218 at prior conductivity values, whereas well-resolved regions will exhibit a different con-  
 219 ductivity.

### 220 3.2 Data selection

221 Within each inversion algorithm, we ran an “all data” inversion of the entire SAM-  
 222 TEX dataset (with only clearly erroneous stations removed) and another “selected” in-  
 223 version of a subset of the data. This approach was designed to test the impacts of het-  
 224 erogeneous station coverage and of noisy and distorted data on the results of each in-  
 225 version algorithm.

226 SAMTEX station coverage is highly heterogeneous compared to other large-scale  
 227 initiatives such as USArray (Kelbert, 2019) or AusLAMP (e.g., Robertson et al., 2016;  
 228 Kirkby et al., 2020; Thiel et al., 2020). Due to logistical constraints and the still preva-  
 229 lent two-dimensional inversion approaches at the time of planning the measurements, data  
 230 were collected in relatively dense transects separated by significant gaps. In a 3D regional  
 231 model, the crustal structure will therefore be strongly represented in the data near those  
 232 profiles and completely absent in regions without coverage. In contrast, deeper features  
 233 (50-200 km) will at least be partially sensed by the data even in regions without direct  
 234 station coverage. For this reason, the focus of our inversions will largely be on the re-  
 235 gional imaging of the mantle lithosphere-asthenosphere system. To image the mantle,  
 236 dense sampling along the profiles could be either beneficial or detrimental. On the one

237 hand, dense coverage should result in redundant information and thus reduce the influ-  
 238 ence of noise for deep imaging, but on the other hand, dense measurements can be highly  
 239 affected by local structures that cannot be represented well in the regional model. This  
 240 issue might be further exacerbated by the need to choose a global regularization param-  
 241 eter for the model, as localized structures in densely covered areas might require a small  
 242 regularization parameter. However, small regularization parameters might be inappro-  
 243 priate for regions without dense coverage.

244 The noise levels of SAMTEX data are also heterogeneous (Figure 3). Some sites  
 245 show significant noise across the whole period range, with either highly scattered or phys-  
 246 ically unrealistic data (e.g., ELG010A). These sites were excluded from all inversions.  
 247 Of the remaining sites, some (e.g., BOT405) display smooth sounding curves with phases  
 248 in quadrant but also demonstrate large offsets between the apparent resistivity curves,  
 249 indicating local static distortion. Others (e.g., KAP047 and WIN011) show similar signs  
 250 of static distortion and additionally display rapidly varying phases that extend out of  
 251 the quadrant, which could indicate local noise or strong resistivity contrasts in the shal-  
 252 low subsurface. Even though *jif3D* can correct for distortion and strategies have been  
 253 devised for *ModEM* to mimic the effects of distortion (Meqbel et al., 2014), it is unclear  
 254 to what degree the information from these sites is useful to constrain deep structures and  
 255 whether fitting distorted sites prevents fitting other data. To investigate the impact of  
 256 such sites, they were therefore retained in the “all data” inversion but excluded from the  
 257 “selected” inversion. To further reduce station density and to assess the impact of het-  
 258 erogeneous station coverage, additional stations with low-quality long-period data ( $> 500$ –  
 259  $10000$  s) were also removed. After the selection procedure, the resulting “selected” dataset  
 260 has a station spacing along the station transects of  $\sim 30$ – $80$  km in most regions, com-  
 261 pared to  $\sim 20$  km for the “all data” dataset (compare blue and red dots in Figure 1).

### 262 **3.3 Inversion Setup**

263 To be able to accommodate the entire region in a single model with acceptable com-  
 264 putational run times, the horizontal discretization for the core region was chosen to be  
 265 15 km in the northing and easting directions for all inversions. Including the padding  
 266 cells for the runs with *ModEM*, the inversion domain comprises  $132 \times 133 \times 53$  cells with  
 267 a vertical discretization of 50 m for the topmost cells increasing up 141 km at the bot-  
 268 tom of the domain. Information on ocean bathymetry was introduced from the ETOPO1

269 global topographic dataset, and seawater was assigned a resistivity of  $0.3 \Omega m$ . On land,  
 270 no topography was considered, and a starting resistivity of  $100 \Omega m$  was assigned to all  
 271 cells.

272 Inversions for *ModEM* were run with error floors of 5% of  $\sqrt{Z_{xy}Z_{yx}}$  on all tensor  
 273 elements. The starting  $\lambda$  was set to 10 and decreased by a factor of 5 when the inver-  
 274 sion when RMS misfit difference is less than 0.002. An isotropic smoothing operator was  
 275 constructed with the covariance matrix set to 0.4 in all directions. For the inversions with  
 276 *jif3D*, we used the same error floor as for the inversions with *ModEM*. We removed the  
 277 outer padding cells from the grid used for modelling in *ModEM* resulting in a mesh with  
 278  $119 \times 120 \times 48$  cells and chose a fixed background resistivity of  $100 \Omega m$ .

279 The inversions for *jif3D* were run with a similar approach to regularization as the  
 280 inversions with *ModEM*. However we used different values for the regularization param-  
 281 eter, starting with  $\lambda = 1,000$  and reducing it to  $\lambda = 1$  in the final iterations, since the  
 282 influence of the regularization on the inversion is different between the two algorithms.  
 283 The initial iterations did not include any distortion correction, but this was enabled af-  
 284 ter the first regularization change as this has been shown to yield stable results (Moorkamp  
 285 et al., 2020).

## 286 4 Data fit

287 For the selected datasets the inversion algorithms reach a final RMS of 1.7 (*jif3D*)  
 288 and 2.3 (*ModEM*) after 200 and 146 inversion iterations, respectively. For the inversion  
 289 of the full datasets, the corresponding RMS values are 2.7 and 5.0, respectively. We show  
 290 the final root-mean-square (RMS) misfit at each site for all frequencies in Figure 5. When  
 291 we only invert the selected data (bottom row) using both *jif3D* and *ModEM* we achieve  
 292 a relatively homogeneous RMS between 1.5 and 2.5 at the majority of sites, and only  
 293 a few sites exceed RMS values of 4.5. While there are some differences in how well sites  
 294 are fit, the overall pattern is comparable and some sites are fit better in one inversion  
 295 or the other. In contrast, when inverting the maximum amount of data, the distribution  
 296 of RMS becomes much more heterogeneous. Many sites are still in the 1.5 to 2.5 range,  
 297 but some sites exceed RMS values of 10. This effect appears to be more pronounced for  
 298 *ModEM* than *jif3D* and some sites that were fit well in the “selected” inversion are now  
 299 fitted significantly worse. These observations confirm that some data that were excluded

300 are, in fact, problematic for the inversion. At least for some of these sites, the distortion  
 301 correction used by *jif3D* helps to achieve a better fit. The question remains, though, to  
 302 which degree this impacts the final models.

303 Figure 4 shows the estimated values of the distortion matrix  $\mathbf{C}$  for the two *jif3D*  
 304 inversions at the central area of the array around the sites BOT405 and KAP047 (Fig-  
 305 ure 3) that were previously identified as distorted (a version with all stations can be found  
 306 in the supplementary material). For the selected data inversion,  $\mathbf{C}$  is close to the iden-  
 307 tity matrix at virtually all sites indicating little to no galvanic distortion. This demon-  
 308 strates that the data selection process successfully removed stations with significant dis-  
 309 tortion and that the inversion algorithm does not introduce artificial distortion, for ex-  
 310 ample, to achieve a low data misfit with a smooth model. When inverting the complete  
 311 data set, some but not all of the additional sites show significant distortion and sites BOT405  
 312 and KAP047 are among the most distorted (Figure 4).

313 In theory, if galvanic distortion is caused by structures that are small compared to  
 314 the typical induction scale length at short periods (Chave & Jones, 2012) and the dis-  
 315 tortion correction only represents this structure-related distortion, the estimates of  $\mathbf{C}$   
 316 at neighbouring sites should show little correlation. Although this is the case in some  
 317 regions, we also see clusters of sites with very similar distortion estimates, e.g., south of  
 318 site BOT405. Here the estimate of  $C_{xx}$  is consistently larger than unity and  $C_{yy}$  smaller  
 319 than unity at most sites. There are two possible explanations for this phenomenon: a)  
 320 It is possible that these sites were all installed in similar geological conditions, for ex-  
 321 ample, when looking for softer ground in an environment dominated by outcropping bedrock.  
 322 b) More likely, the distortion estimates capture variability in structures that can, in prin-  
 323 ciple, be resolved by MT measurements but cannot be represented by the chosen hor-  
 324 izontal discretization of 15 km, i.e. they account for the so-called model discrepancy (Kennedy  
 325 & O’Hagan, 2001). These stations are located at the northern end of the Kaapvaal Cra-  
 326 ton crossing into the Magondi Mobile Belt, and thus it is likely that significant defor-  
 327 mation is recorded in the crust.

328 We compare the data fit for the distorted site BOT405 and and the exemplary sites  
 329 RAK011A, KIM428 and KAP045 for the *ModEM* “all data” inversion (Figure 6) and  
 330 the *jif3D* “all data” inversion (Figure 7). The difference between observed and predicted  
 331 data for site BOT405 clearly shows how the distortion correction in *jif3D* helps to achieve

332 a better fit to the off-diagonal apparent resistivity curves. Whereas the model response  
 333 from *ModEM* converges to a common apparent resistivity value at short periods, *jif3D*  
 334 reproduces the constant offset between the two curves. Interestingly, although the off-  
 335 diagonal phases are fit differently by both inversions, there is no clearly superior fit by  
 336 either of the two models. At sites RAK011A and KIM428, both models produce virtu-  
 337 ally identical responses for the off-diagonal apparent resistivities and phases and match  
 338 the observed data well. At site KIM428, the models reproduce all variations of the curves,  
 339 while at site RAK011A, the overall shape is reproduced well by the models, but the phase  
 340 anomaly in the xy-component at periods between 50-100 s is not fully reproduced by ei-  
 341 ther model. At both stations, the diagonal elements are significantly smaller than the  
 342 off-diagonal elements and are matched better by the *jif3D* inversion than the *ModEM*  
 343 inversion. It is our experience that distortion correction helps to match diagonal elements  
 344 better even when these are small (Moorkamp et al., 2020). The data at site KAP045 are  
 345 matched differently by the two inversions, and again the difference is more pronounced  
 346 in the phase than in apparent resistivity. The response from *ModEM* reproduces the short  
 347 period phases well but shows small but consistent differences in the overall shape. In com-  
 348 parison, *jif3D* appears to reproduce aspects of the general shape better but does match  
 349 the phases exactly in any period range.

350 The observed differences in model fit highlight that different inversion algorithms  
 351 reproduce different aspects of the observed data that go beyond the changes expected  
 352 from simply modifying the regularization in a single inversion algorithm. This contrast-  
 353 ing behaviour illustrates the value of inverting data with multiple inversion algorithms.  
 354 It also shows that distortion correction can help fit certain aspects of the data, as demon-  
 355 strated by the misfit maps, but this does not necessarily imply that all aspects of the  
 356 data are matched more closely. All in all both inversions with all data match the obser-  
 357 vations at the majority of sites well. We therefore expect both models to provide rea-  
 358 sonable representations of electrical resistivity in the vicinity of the measurements sites.

## 359 **5 Resistivity models**

360 We show horizontal cross-sections through the derived inversion models between  
 361 50 and 200 km depth in 50 km intervals (Figures 8 - 11) as well as vertical slices in the  
 362 east-west direction at latitude 22° south (Figure 12) and along the Kaapvaal (Figure 13)  
 363 and Kimberley (Figure 14) profiles (see also Figure 1 for location of these profiles). The

364 different inversions show very similar large-scale structures, e.g., a generally resistive ( $\geq$   
 365  $500 \Omega m$ ) central region below 50 km depth, which is significantly more resistive than the  
 366 starting model ( $100 \Omega m$ ). Embedded in this resistive lithosphere are several large con-  
 367 ductors, typically associated with boundaries of different geological units. Even though  
 368 the large-scale picture is similar for all models, there are significant differences in the de-  
 369 tailed resistivity structures and values between the inversion results. We will therefore  
 370 start with a description of the main features based on the “all data” *jif3D* model and  
 371 discuss how these are expressed in the other models. In the next section, we use the dif-  
 372 ferences and similarities between the models to appraise the robustness and resolution  
 373 of inversion results.

374 We observe the maximum resistivity ( $\geq 5,000 \Omega m$ ) around the south-eastern part  
 375 of the array (labelled the Kaapvaal Resistor (KR) on the horizontal slices) and the north-  
 376 western part of the array, north of the Damara Conductive Belt (DCB). In both cases,  
 377 the maximum resistivity is located at depths between 50–100 km and appears to decrease  
 378 at 150 km depth and below. These observations are compatible with the thick, dry litho-  
 379 spheric mantle associated with the roots of the Kaapvaal Craton and the Congo Cra-  
 380 ton, respectively (e.g., Evans et al., 2011; Jones et al., 2013; Khoza et al., 2013). In the  
 381 central part of the array, around latitude  $24^\circ$  south, is a roughly east-west striking band  
 382 of reduced resistivity ( $\sim 100 \Omega m$ ) in the deeper slices (150 km and below) which becomes  
 383 more resistive in the shallower parts of the model. We term the central structure in this  
 384 band at approximately  $24^\circ$  east the Molopo Farms Conductor (MFC). It can be iden-  
 385 tified as a zone of decreased resistivity ( $< 20 \Omega m$ ) on the 150 km and 200 km depth slices  
 386 from the two *ModEM* inversions. The *jif3D* based inversions only show a weak signa-  
 387 ture at 150 km but show a structure with similarly low resistivity displaced slightly to  
 388 the north-west at 200 km depth. The conductor associated with Bushveld Complex (Bushveld  
 389 Conductor, BC) appears on all different modelling schemes north of the Kaapvaal Re-  
 390 sistor (KR, Figure 13). Even though its conductivity differs from model to model, less  
 391 than  $10 \Omega m$  in the inversions with *ModEM* and  $30\text{--}50 \Omega m$  in the inversions with *jif3D*  
 392 , its spatial extent is consistent between models and it is consistently positioned beneath  
 393 the surface expression of the Bushveld intrusive complex, suggesting that it is a robustly  
 394 modelled feature.

395 Further south, the *ModEM* inversions indicate a low resistivity zone ( $20 \Omega m$ ) at  
 396 a depth of 50 km near the south-western terminus of the Kaapvaal Craton, which we term

397 the Southern Kaapvaal Conductor (SKC). The “selected data” *ModEM* inversion shows  
398 this low resistivity extending to depths  $\geq 150$  km, but this is less clearly visible in the  
399 inversion with all data. Both *jif3D* inversions show decreased resistivities of  $100 \Omega \text{ m}$  com-  
400 pared to the surrounding  $1,000 \Omega \text{ m}$ , but no structures with the low resistivity indicated  
401 by *ModEM*.

402 In the north-western part of the array, the signature of the Damara Conductive Belt  
403 (DCB), previously identified by Khoza et al. (2013), is apparent at a depth of 50 km in  
404 all inversions. It is an east-west striking band of decreased resistivity ( $\sim 10 \Omega \text{ m}$ ), inter-  
405 preted to be associated with the collision between the Congo Craton and the adjacent  
406 mobile belts. At depths of more than 100 km, the inversions with all data also contain  
407 an approximately north-south striking, low resistivity feature. The inversions with se-  
408 lected data also show slightly decreased resistivity in the same region, but it appears that  
409 some information on this feature is contained in the sites excluded in the selection pro-  
410 cess.

411 In addition to these four features discussed above, the model contains a variety of  
412 other structures. We do not go into further detail on all these features here but in the  
413 second part of this study (Özaydin et al., 2021) we investigate the relationships between  
414 the geoelectric lithospheric architecture, composition, tectonic and magmatic history of  
415 the southern Africa in detail.

416 The vertical slices through the model shown in Figures 12 - 14 confirm the infer-  
417 ences made by comparing the horizontal slices, showing similar low and high resistivity  
418 features. However, the exact locations, shapes and resistivity values vary between the  
419 different inversions. In all cases, the “all data” inversions show stronger resistivity con-  
420 trasts and more localized features than the “selected data” inversions for the same in-  
421 version algorithm, particularly in the upper 50 km. Below this depth, the differences be-  
422 tween using all data and selected data are less pronounced, but persist. For example, the  
423 *ModEM* “all data” slice along the KAP line (Figure 13) shows a low resistivity zone at  
424 a depth of 200 km towards the northern (right) end of the profile. This feature is not  
425 clear in the “selected data” inversion, which shows resistivity values comparable with  
426 the starting model, possibly suggesting that there is little resolution in this region. Com-  
427 paring the results from the two different inversion algorithms, *ModEM* appears to favour

428 more concentrated features at depth while in *jif3D* the features are generally more dis-  
 429 tributed with less sharp edges.

## 430 6 Model appraisal

431 To provide a quantitative view on the differences between the models, we plot model  
 432 difference matrices at a depth of 50 km in Figure 15 and 150 km in Figure 16. In both  
 433 cases, we show a horizontal slice through each model at the respective depth on the di-  
 434 agonal. Plots above the diagonal show the difference in logarithmic resistivity for the dif-  
 435 ferent model combinations, while plots below the diagonal show the corresponding re-  
 436 sistivity difference histograms. The model histograms show significant differences in re-  
 437 sistivity between all model combinations of up to 2 orders of magnitude (2 in logarith-  
 438 mic units), even though for the vast majority of model cells, the difference is less than  
 439  $\pm 1$  order of magnitude. The histograms appear to be slightly wider at 50 km depth than  
 440 at 150 km depth. Most histograms are centered around a difference of zero, suggesting  
 441 that there is no significant overall bias in the resistivities retrieved in each inversion, ex-  
 442 cept the histogram for the two inversions with selected data, which is centered around  
 443  $\approx 0.2$ , indicating that the model produced by *jif3D* is consistently more resistive.

444 The histograms clearly illustrate that the largest resistivity differences are produced  
 445 by using different algorithms to invert the same dataset, while smaller differences are pro-  
 446 duced by inverting different subsets of data with the same inversion algorithm. At both  
 447 depths and for both inversion algorithms, the histograms comparing the “selected” and  
 448 “all data” inversions show highly symmetric shapes and a concentrated peak at zero, while  
 449 the other histograms are generally broader and exhibit more structure.

450 The spatial difference plots in Figures 15 and 16 add more detail to the global re-  
 451 sistivity differences displayed in the histograms. Spatial comparisons between the *jif3D*  
 452 and *ModEM* models using both “selected data” and “all data” datasets at both depths  
 453 consistently show that the *jif3D* models have higher average resistivities over much of  
 454 the model space than the *ModEM* models, except for in the south-western part of the  
 455 array where *jif3D* produces a consistently less resistive model. The south-western region  
 456 is the part of the model most poorly constrained by station coverage. These differences  
 457 sum to a resistivity difference histogram that centers on zero. In all difference plots, we  
 458 also see a correlation between locations of large scale tectonic boundaries and changes

459 in sign of the resistivity difference, particularly along the northern margin of the Namaqua-  
460 Natal Belt and the margins of the Damara and Ghanzi-Chobe belts. While the details  
461 vary, this phenomenon is observed in all combinations of models to varying degrees. This  
462 indicates that the differences in the models are not merely due to fitting aspects of the  
463 data differently or the influence of noisy measurements, but each inversion images the  
464 Earth in a different way.

465 Comparisons between the two “selected data” inversions and the two “all data” in-  
466 versions demonstrate that data selection has a significant impact on model differences.  
467 The spatial difference plots for the *jif3D* and *ModEM* “selected data” inversions reveal  
468 broad zones of consistent resistivity differences, while those for the two “all data” inver-  
469 sions show much more inhomogeneous, spatially varying resistivity differences. This re-  
470 sults from the stronger influence of regularization in the “selected data” inversions, lead-  
471 ing to overall smoother models. When adding data, the wavelength of the patterns de-  
472 creases, and we see more fine-scaled differences, together with relatively sharp changes  
473 between positive and negative differences. Some of the largest differences are located in  
474 regions without site coverage, e.g., southeast of the KAP line or in the gaps between mea-  
475 surement lines in the northern part of the array.

476 The most likely candidate for causing many of these differences is the different reg-  
477 ularization schemes. This factor is most clearly seen in the comparison between mod-  
478 els produced with the “selected” dataset as the influence of regularization is strongest  
479 there. Where the Earth is more resistive than the starting model in both inversions, *jif3D*  
480 consistently estimates higher resistivities than *ModEM*. Conversely, where the inversions  
481 indicate lower resistivities than the starting model, *jif3D* underestimates resistivity com-  
482 pared to *ModEM* on the larger scale. Both observations can be explained by the fact that  
483 *ModEM* minimizes the difference to the reference model and smoothness simultaneously,  
484 while *jif3D* only aims at recovering a smooth model. This behaviour can explain the ob-  
485 served correlation between major tectonic boundaries and changes in sign of the resis-  
486 tivity difference. A resistive geological region is likely to be modelled with a higher re-  
487 sistivity in *jif3D* than *ModEM*, and an adjacent conductive geological region is likely to  
488 be modelled with a lower resistivity in *jif3D* than *ModEM*. The model difference plot there-  
489 fore highlights the boundary between these two regions.

490 Without additional information, we cannot say which of the inversions is more rep-  
491 resentative of the true resistivity within the Earth. However, we tried to reduce the ef-  
492 fect of regularization in *ModEM* by running an additional inversion with a starting and  
493 reference model based on laterally smoothed apparent resistivities. To construct this model,  
494 for each measurement site we construct a circle with a 4-degree radius centred on the  
495 site and take the median apparent resistivity value of all sites within the circle at peri-  
496 ods longer than 100 s. The resulting resistivity value is assigned to all cells below this  
497 site. We then perform a linear interpolation of logarithmic resistivity between these val-  
498 ues to determine the resistivity in each model cell. The resulting model (Figure 17) shows  
499 laterally varying resistivities between 50 and 1,000  $\Omega m$  and regions of high resistivity that  
500 correlate with the most resistive regions identified in the previous inversions.

501 The resulting median inversion model (Figures 18 and 19) fits the selected data to  
502 an RMS comparable with the inversion run from a homogeneous half-space. Compared  
503 to the homogeneous inversion run, the average resistivities at 50 km depth (Figure 18)  
504 and 150 km depth (Figure 19) are higher, particularly in regions that are not directly  
505 covered by sites. Conductive anomalies show a very similar pattern to the previous in-  
506 versions, although the shape and location differ slightly in some cases, including some  
507 of the individual conductors that form the Damara Conductive Belt at 50 km depth (com-  
508 pare Figure 8 and Figure 18) or the Southern Kaapvaal Conductor at 150 km depth. These  
509 changes are not significant enough to imply a different geological interpretation of these  
510 structures.

511 The spatial difference plot and difference histogram comparing the median inver-  
512 sion and the inversion of the same data with *jif3D* reveals some interesting changes com-  
513 pared to the inversion with a homogeneous starting model. Visually, the spatial differ-  
514 ence plot for the median model contains a lot less long-wavelength structure and is dom-  
515 inated by more small scale differences. This contrast is particularly visible at 50 km depth  
516 where *jif3D* previously produced consistently higher resistivities in the central model re-  
517 gion. However, the spatial difference plot for the median model displays a much more  
518 variable pattern where the sign of the conductivity difference changes within smaller dis-  
519 tances. At 150 km depth, the effect is less pronounced yet still observable.

520 The difference histogram comparing the median inversion and the *jif3D* inversion  
521 at 50 km depth has a maximum very close to zero, while the histogram comparing the

522 homogeneous inversion and *jif3D* is offset to slightly positive values with a maximum  
 523 at  $\approx 0.2$ . While still not fully symmetric, the maximum and minimum (most negative)  
 524 differences now show a similar magnitude. At 150 km depth, the impact of the starting  
 525 model on the histogram is even more pronounced and the median model histogram is  
 526 more significantly offset to negative values associated with the higher average resistiv-  
 527 ity of the median model. While the average resistivities inverted from the homogeneous  
 528 *ModEM* are downward biased compared to *jif3D*, the median model *ModEM* resistiv-  
 529 ities are upward biased.

## 530 7 Discussion and Conclusions

531 The main goal of this paper is to present a new 3D conductivity model for south-  
 532 ern Africa and use the different inversion methodologies to understand uncertainties in  
 533 the results better. An additional result is that the detailed comparisons of the models  
 534 also reveal some technical aspects of inversions and regularization that are of interest to  
 535 both algorithm developers and practitioners and thus warrant some discussion before de-  
 536 scribing some of the geological interpretations implied by these models.

537 It is our impression that most of the differences between the results from the two  
 538 inversion algorithms stem from the different regularization philosophies. The purely smooth-  
 539 ness based approach followed by *jif3D* spreads out structures to their maximum possi-  
 540 ble extent, most clearly visible in Figure 13. Selecting such a smoothing operator has  
 541 the disadvantage that conductive anomalies can be smeared out, and their boundaries  
 542 can be challenging to identify. In contrast, the mixed regularization approach pursued  
 543 by *ModEM* typically produces more localized structures. On the flip side, the regular-  
 544 ization toward a reference model appears to bias the large scale resistivity toward this  
 545 model, particularly in regions of low sensitivity. This observation is mirrored by the sys-  
 546 tematic study of Robertson et al. (2020). Taken together, we conclude that for large ar-  
 547 rays with heterogeneous coverage such as this, *jif3D* produces models with more repre-  
 548 sentative large-scale resistivity values, while *ModEM* produces more focused and local-  
 549 ized anomalies. To some degree, a more representative large-scale resistivity can be ob-  
 550 tained with *ModEM* with a median-based starting model, as shown by the comparison  
 551 at 50 km depth. Still, the shift in bias at 150 km shows that possibly a more detailed  
 552 starting model with varying resistivity with depth is necessary to obtain good average  
 553 resistivities over large areas. Alternatively, one could design a regularization scheme where

554 the balance between smoothness and damping toward a reference model can be finely  
 555 adjusted. While it seems that such an approach could combine the advantages of both  
 556 regularization approaches, it is questionable how an optimal balance could be objectively  
 557 found and how practical such a scheme would be for routine application.

558 The effect of inverting for only selected, high-quality data or the maximum amount  
 559 of data is similar regardless of the inversion algorithm. In both cases, the inclusion of  
 560 more data increases the misfit of the final models as potentially problematic sites are in-  
 561 troduced. To some degree, this effect is reduced by the distortion correction employed  
 562 in *jif3D* which can deal with problems associated with galvanic distortion and achieve  
 563 a better fit at many sites. At the same time, the models with more data exhibit stronger  
 564 resistivity contrasts and additional structures, e.g., the north-south striking conductor  
 565 in the north-western part of the study area that extends from the Congo Craton into the  
 566 Damara Belt. Given the similarity of these features for both inversion algorithms and  
 567 the acceptable misfit for the inversion with *jif3D*, we conclude that these are not arte-  
 568 facts caused by noisy data, but that these features are due to information about the re-  
 569 sistivity of the Earth contained in the measurements included in the “all data” models.  
 570 Still, the inversions with selected data contain the same general features as the inver-  
 571 sions with all data. Based on the similarity with other models and the data fit, our two  
 572 preferred models are the ones produced by *jif3D* with all data and the *ModEM* inver-  
 573 sion with a median starting/reference model and selected data.

574 The most prominent features of our two preferred models are: (1) A resistive core  
 575 of the Kaapvaal Craton as indicated by the Kaapvaal Resistor. This region of high re-  
 576 sistivity ( $> 1,000\Omega m$ ) extends to depths of 150 km (*ModEM*) to 200 km (*jif3D*) and  
 577 indicates a dry lithospheric mantle in line with previous 2D interpretations (Evans et al.,  
 578 2011) and experimental electrical conductivity of common mantle minerals (e.g., Karato  
 579 & Wang, 2012; Özaydın & Selway, 2020). (2) Other high resistivity regions at depths of  
 580 100 km and greater include the Congo Craton in the north-west (Kamanjab Inlier) and  
 581 northern Botswana in the north-eastern part of the array suggesting the presence of litho-  
 582 sphere with a broadly similar composition in these regions. (3) These resistors are in-  
 583 tersected by several deep-seated conductors that are present to varying extents in all in-  
 584 version models. These include the Molopo Farms conductor, the Bushveld conductor and  
 585 the north-south striking feature below the Congo Craton and Damara Belt. In the lat-  
 586 ter case, a possible interpretation is a shallower lithosphere-asthenosphere boundary com-

587   pared to the thick cratonic roots of the Kaapvaal and Congo Cratons (Celli et al., 2020).  
588   In contrast, the Molopo Farms Conductor and the Bushveld Conductor are likely expres-  
589   sions of emplacement of metasomatic material during episodes of magmatism (Beukes  
590   et al., 2019).

591         A more detailed interpretation of the resistivity structures recovered by these in-  
592   versions requires careful consideration of the geological history of the region and the re-  
593   lationships between resistivity, composition and temperature. These considerations are  
594   beyond the scope of this study and are presented in a companion paper (Özaydin et al.,  
595   2021).

596         We have constructed a set of 3D models for southern Africa based on two subsets  
597   of the SAMTEX magnetotelluric dataset and utilizing two independent inversion algo-  
598   rithms. Despite some differences in the shape of structures and the recovered resistiv-  
599   ities, the models show strong similarities. Previous efforts using these data either used  
600   the whole dataset but did not perform inversions or were concentrated on regional sub-  
601   sets of the data. Thus the models presented here are the first large-scale resistivity mod-  
602   els of the region and can serve as a resource for further investigations and integration  
603   with other observations such as gravity and seismology.

## 604   **Acronyms**

605   **BBMT** Broad-Band Magnetotelluric

606   **BC** Bushveld Conductor

607   **DCB** Damara Conductive Belt

608   **MT** Magnetotelluric

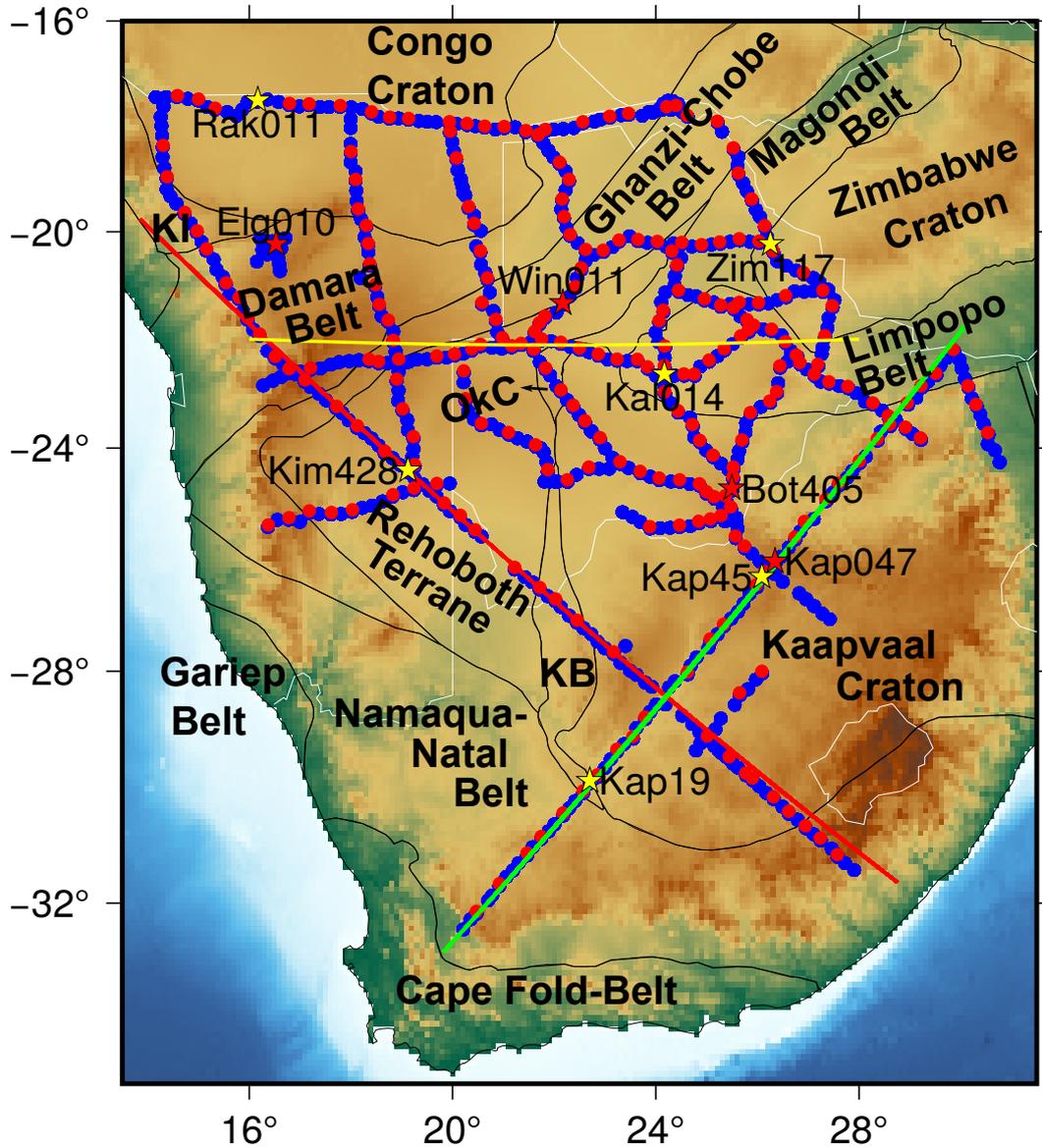
609   **KR** Kaapvaal Resistor

610   **LMT** Long-period Magnetotelluric

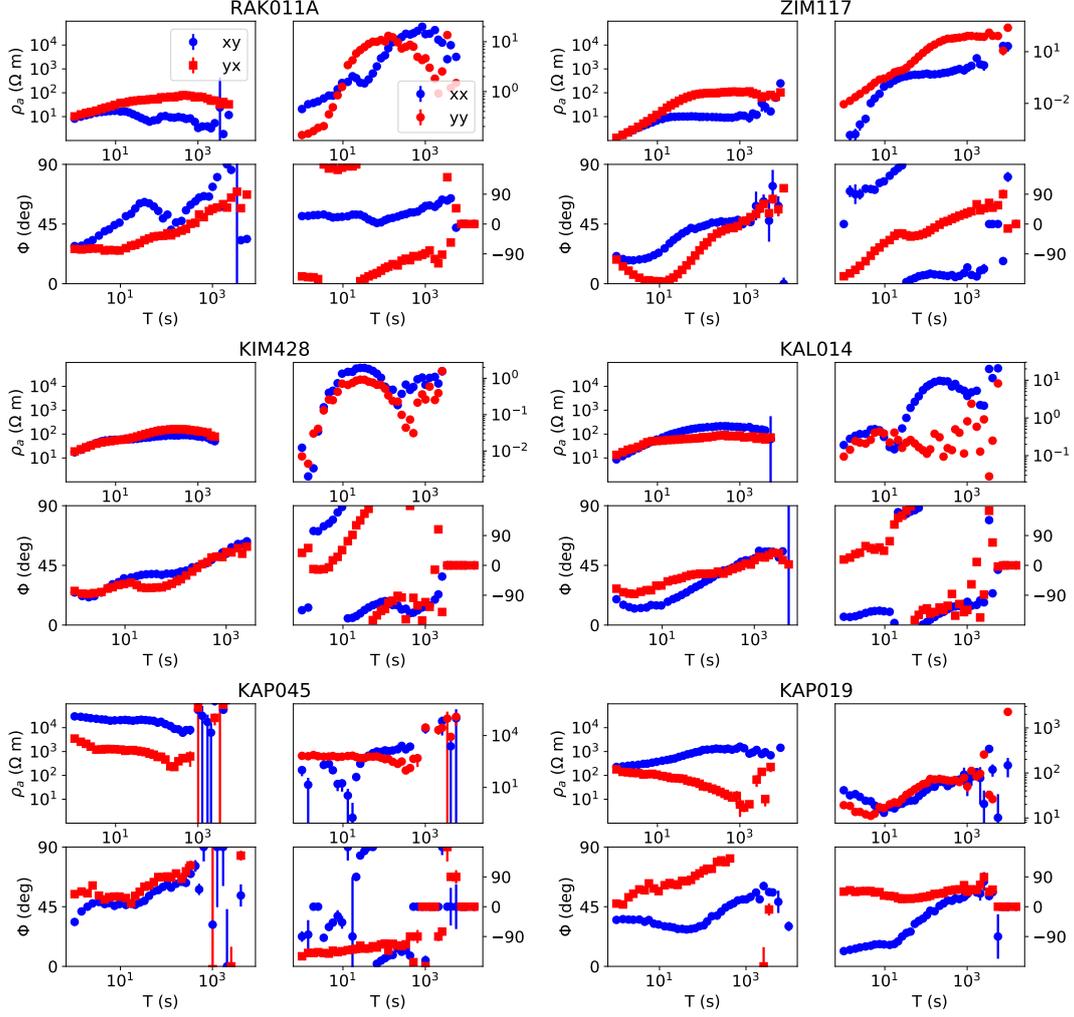
611   **MFC** Molopo Farms Conductor

612   **SAMTEX** South African Magnetotelluric Experiment

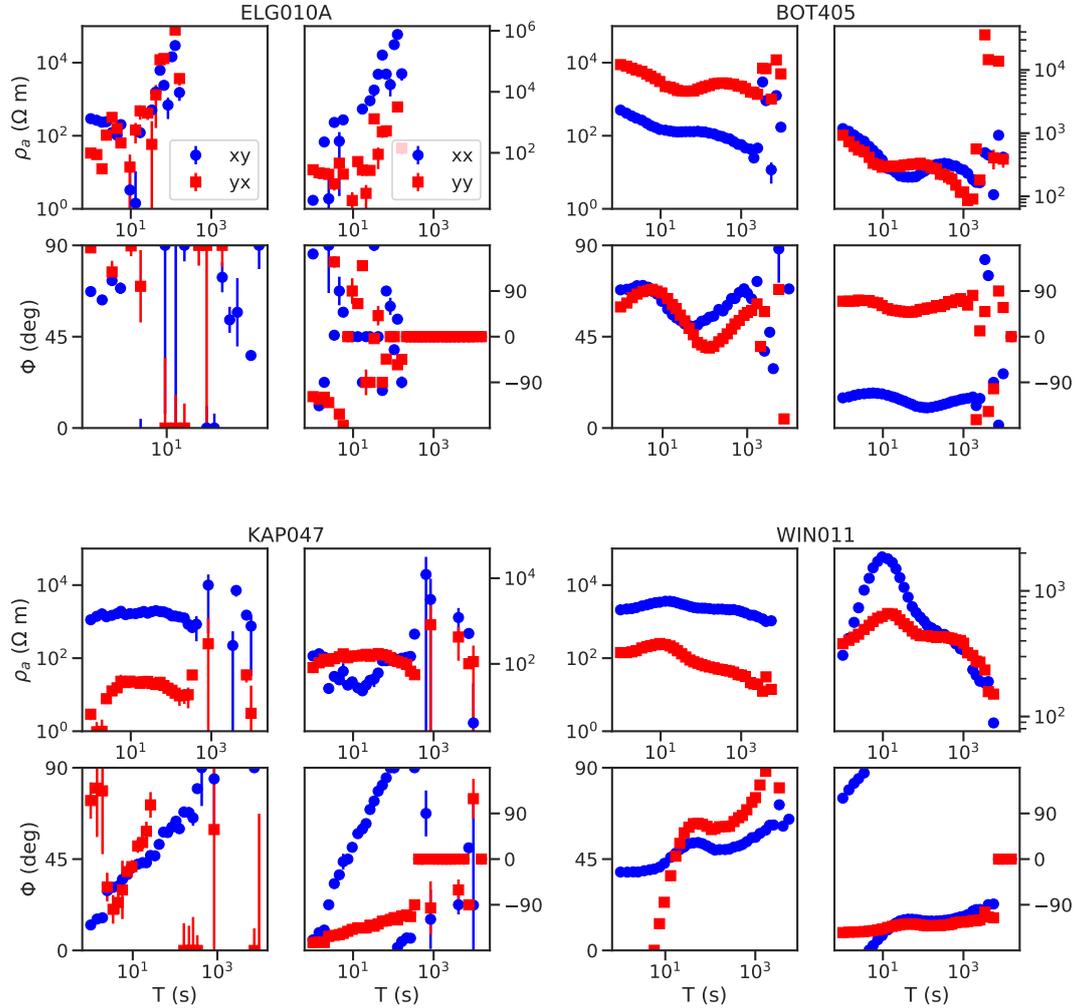
613   **SKC** Southern Kaapvaal Conductor



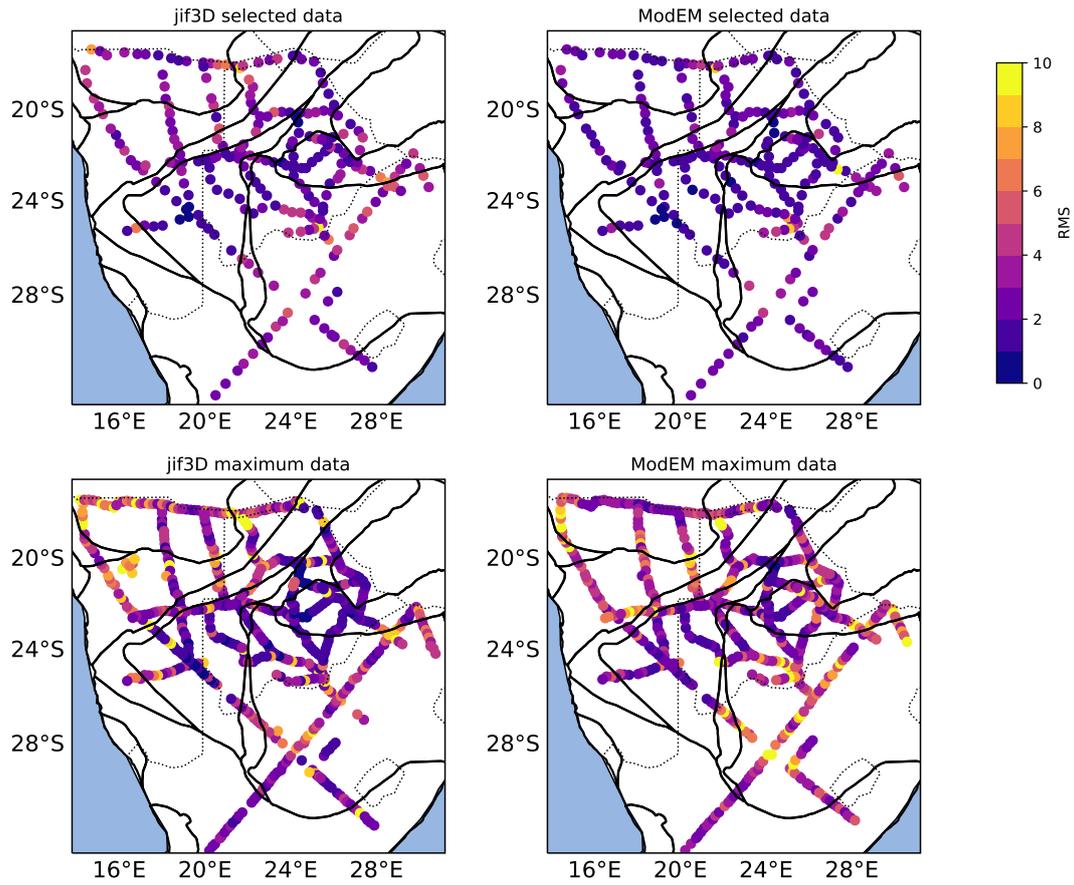
**Figure 1.** Map of the study area. We show the SAMTEX magnetotelluric measurement sites considered in the full data inversions as blue dots and the sites considered in the inversions with selected data as red dots. Yellow stars indicate exemplary stations for different regions shown in Figure 2 and red stars poor quality data excluded from some of the inversions and shown in Figure 3. The red, green and yellow lines mark the locations of vertical model profiles along the Kimberley, Kaapvaal and 22 degree south lines, respectively. Black lines mark the boundaries of tectonic provinces based on McCourt et al. (2013).



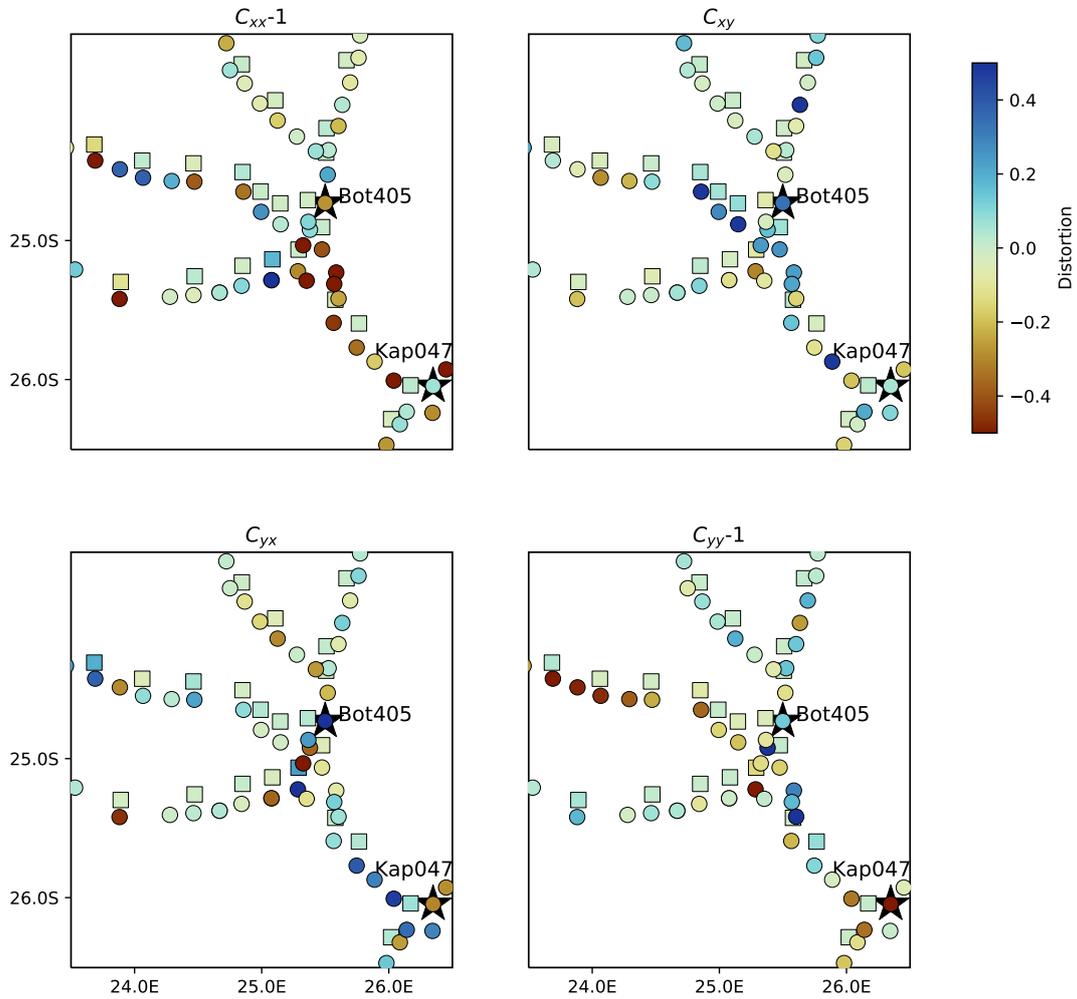
**Figure 2.** Six exemplary sites representing different regions within the inversion domain. For each site we plot apparent resistivity and phase of the four impedance elements. Off-diagonal (xy and yx) apparent resistivities are plotted with a consistent y-axis to highlight differences in average resistivity between different regions, while apparent resistivity for the diagonal components is plotted with a different scale for each site for better readability.



**Figure 3.** Examples of data excluded from the inversions with selected data. Site ELG10A shows overall problematic data and has been excluded from all inversions while the other sites show potentially problematic features as discussed in the text but have been retained for the inversions with maximum data.

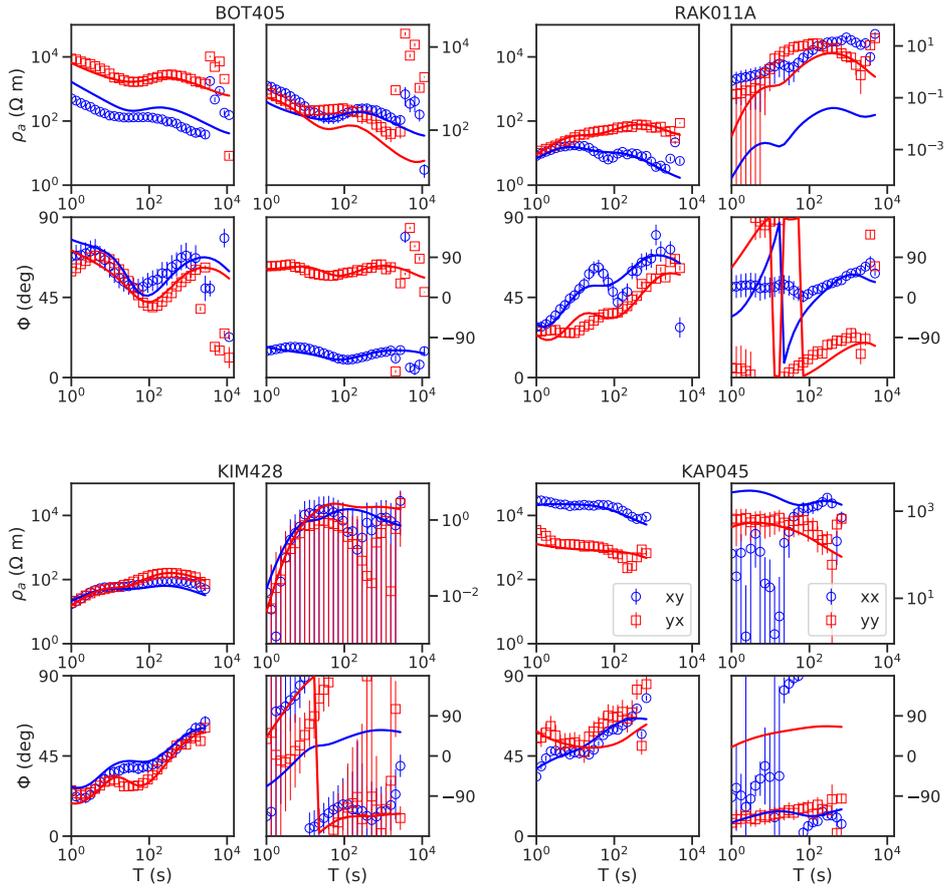


**Figure 4.** Map of misfit for the four inversion runs. We show the error normalized RMS across all frequencies at each site.

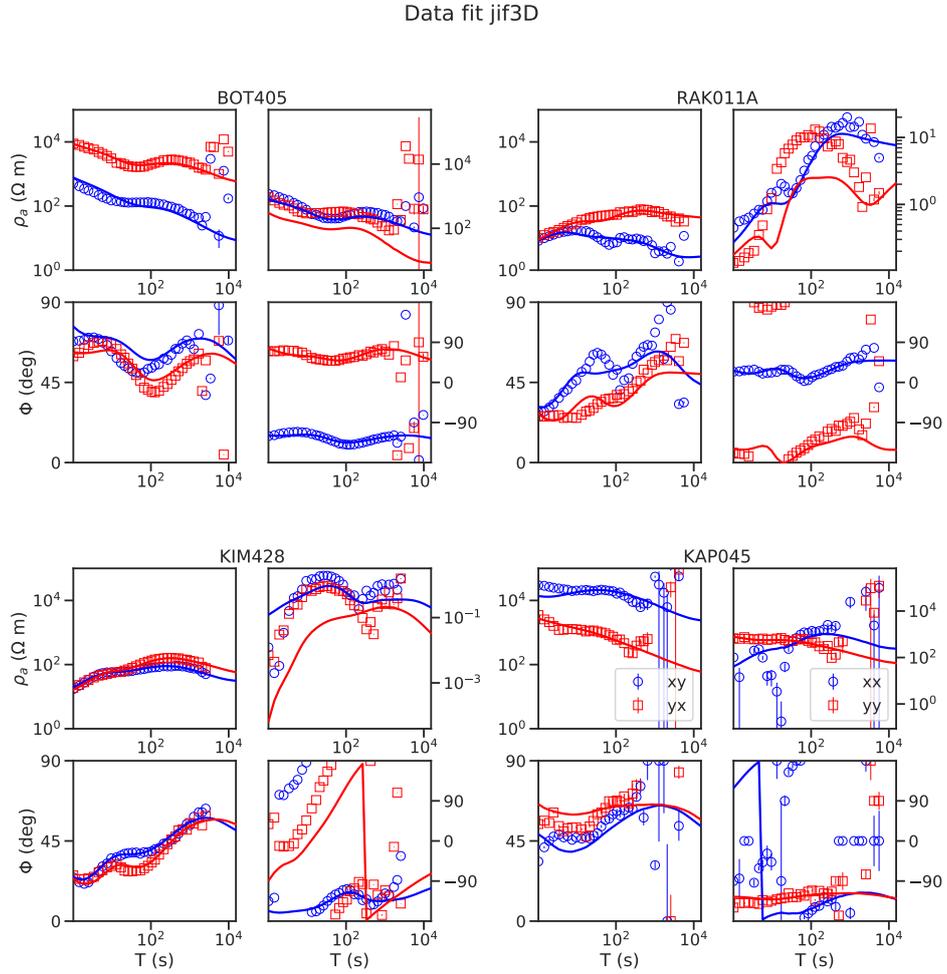


**Figure 5.** Map of distortion estimates for the inversions with *jif3D* in the central region of the array. We show the estimates for the inversion with all data as circles and for the selected data inversion as squares. These have been displaced north from the original locations for better visibility. Colors mark the deviation of the distortion matrix elements from the identity matrix. We highlight sites BOT405 and KAP047 shown in Figure 3 with black stars.

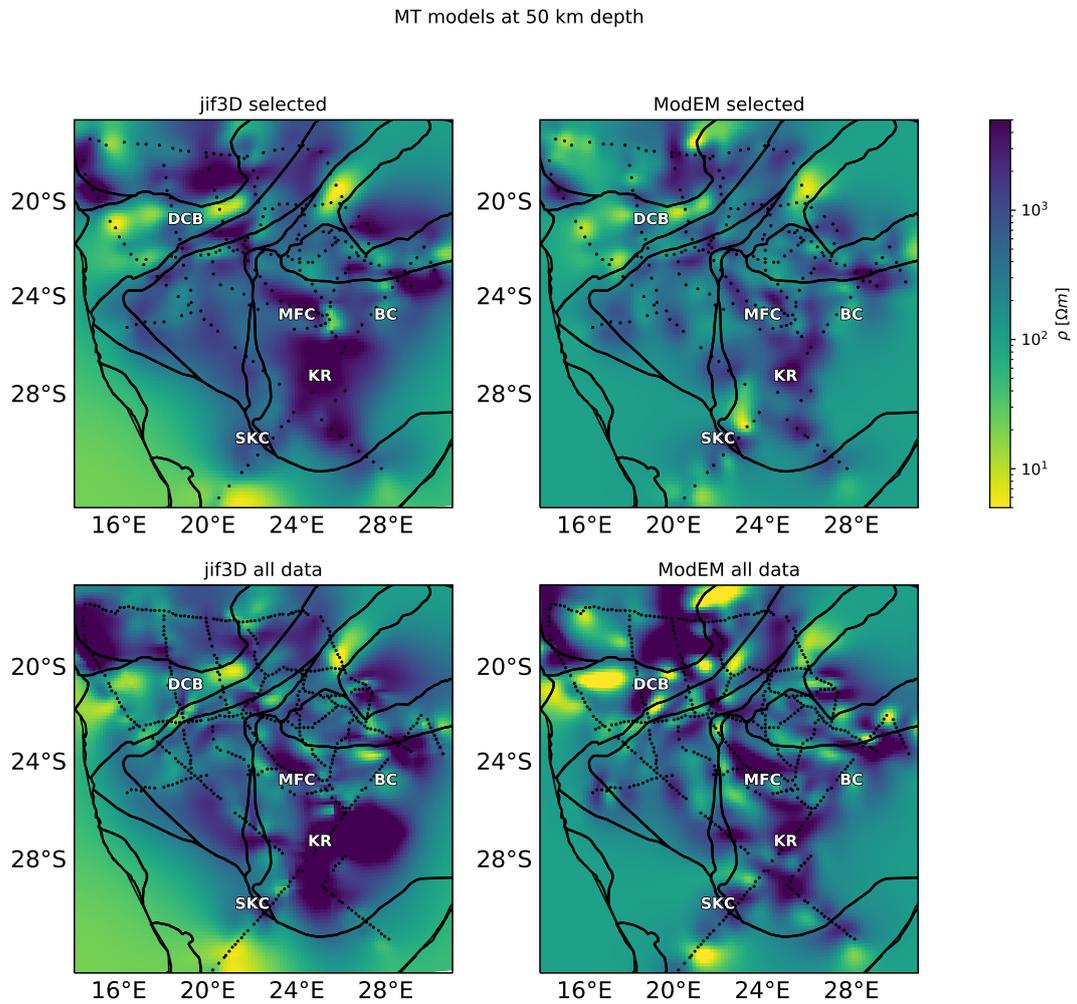
Data fit ModEM



**Figure 6.** Comparison between observed data (symbols) and predicted data (lines) for the inversion run with *ModEM* and all data for four selected sites marked in Figure 1.

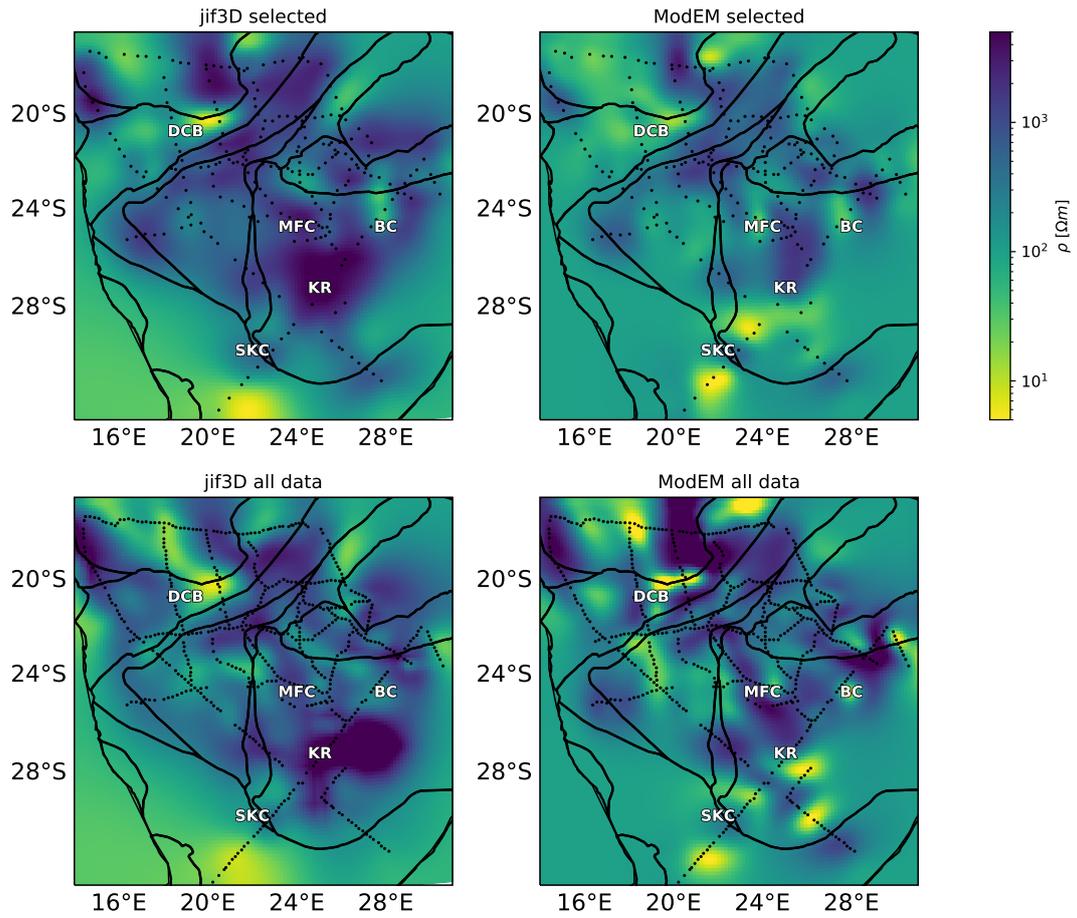


**Figure 7.** Comparison between observed data (symbols) and predicted data (lines) for the inversion run with *jif3D* and all data for four selected sites marked in Figure 1.



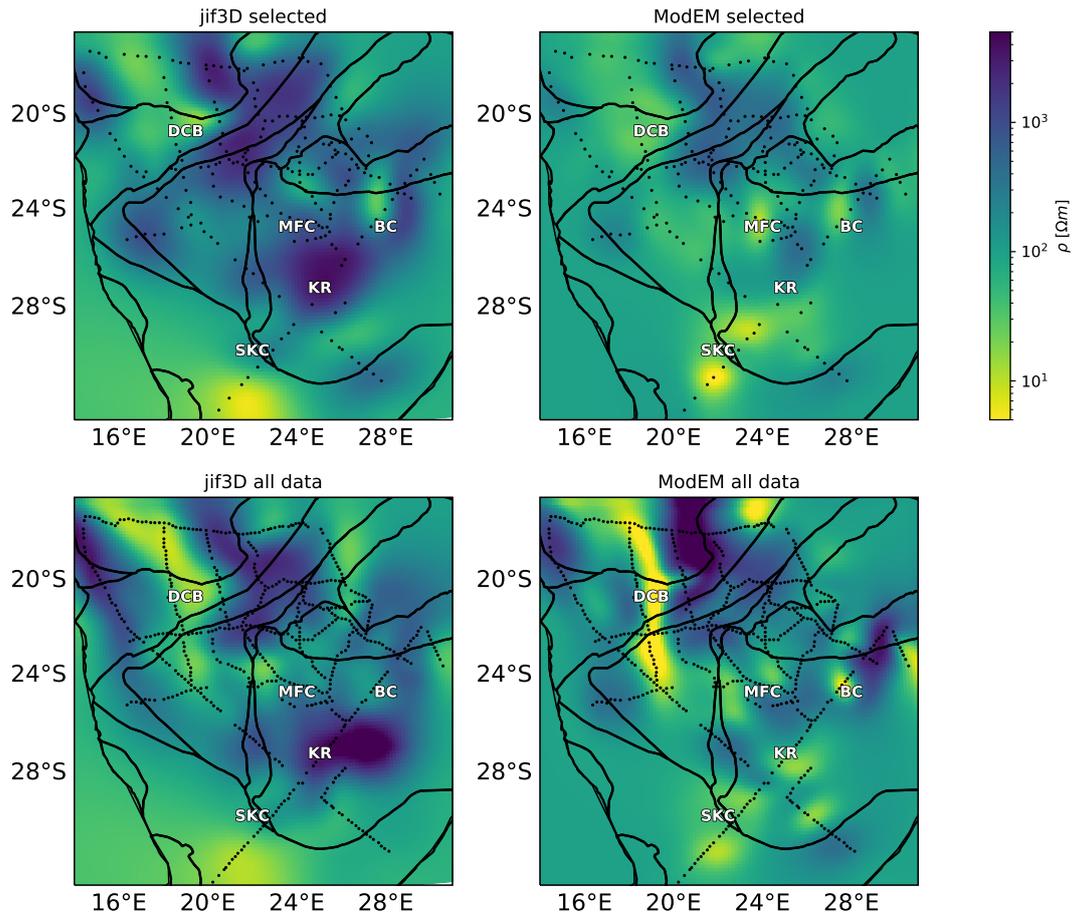
**Figure 8.** Horizontal slices through the inversion models at a depth of 50 km. We mark several notable structures: Bushveldt Conductor (BC), Damara Conductive Belt (DCB), Kaapvaal Resistor (KR), Molopo Farms Conductor (MFC), Southern Kaapvaal Conductor (SKC).

MT models at 100 km depth



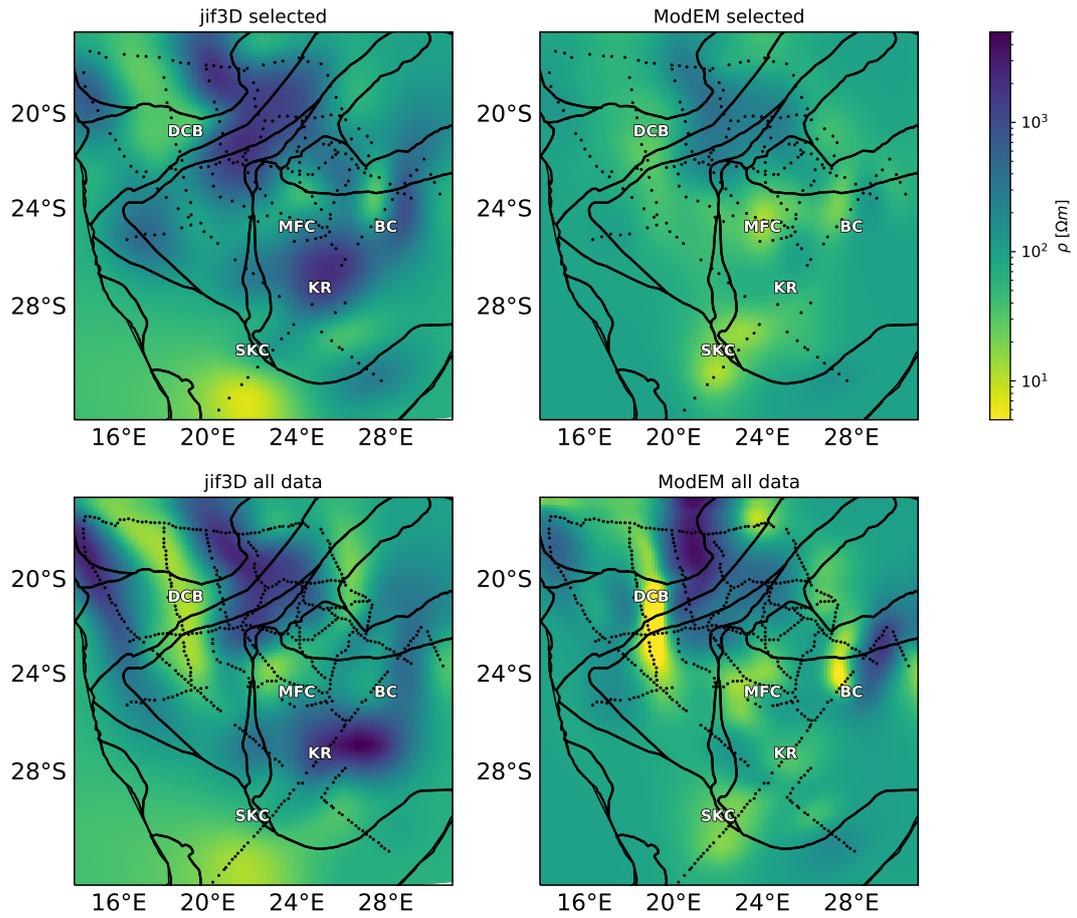
**Figure 9.** Horizontal slices through the inversion models at a depth of 100 km. For an explanation of abbreviations see Figure 8.

MT models at 150 km depth

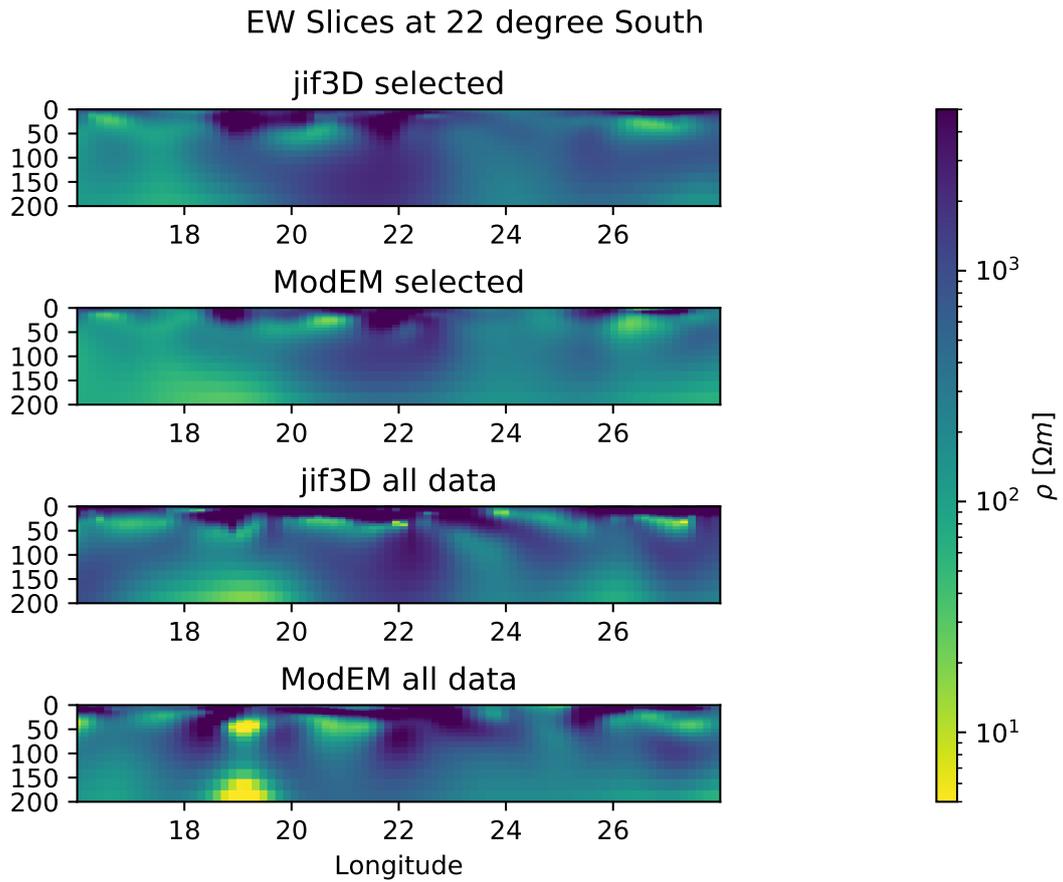


**Figure 10.** Horizontal slices through the inversion models at a depth of 150 km. For an explanation of abbreviations see Figure 8.

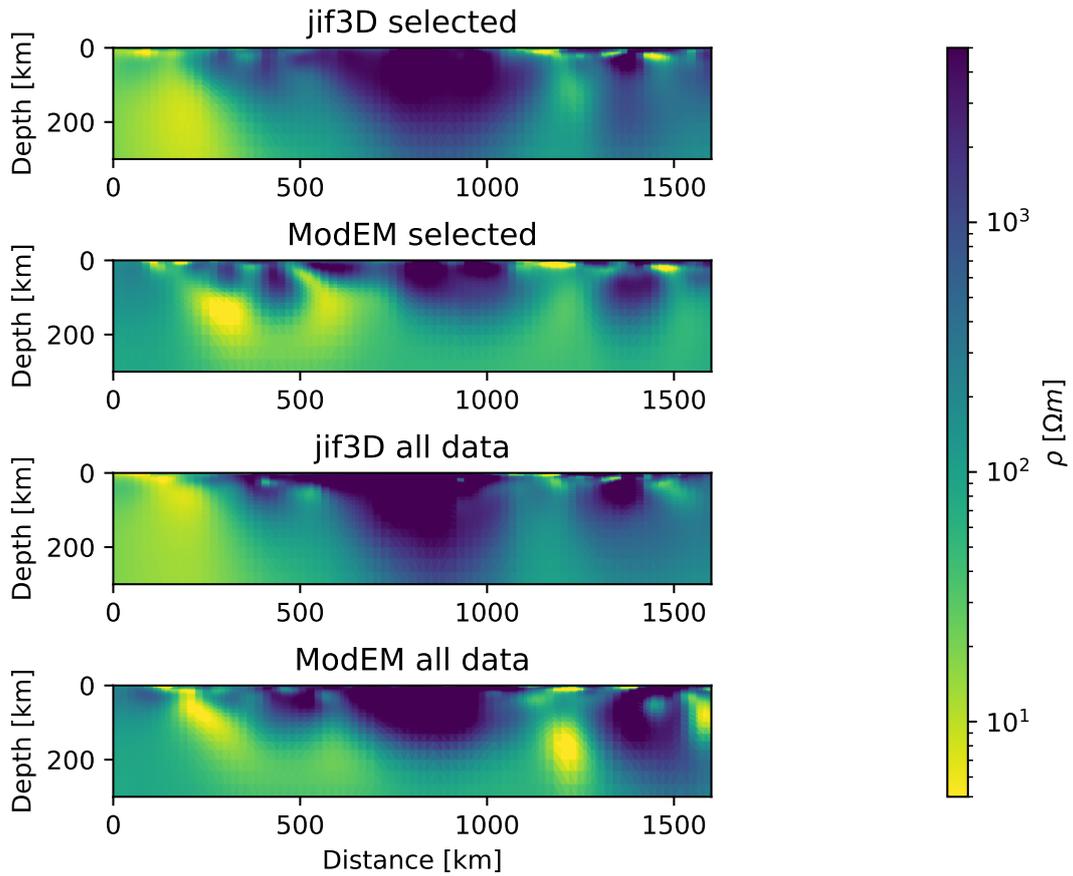
MT models at 200 km depth



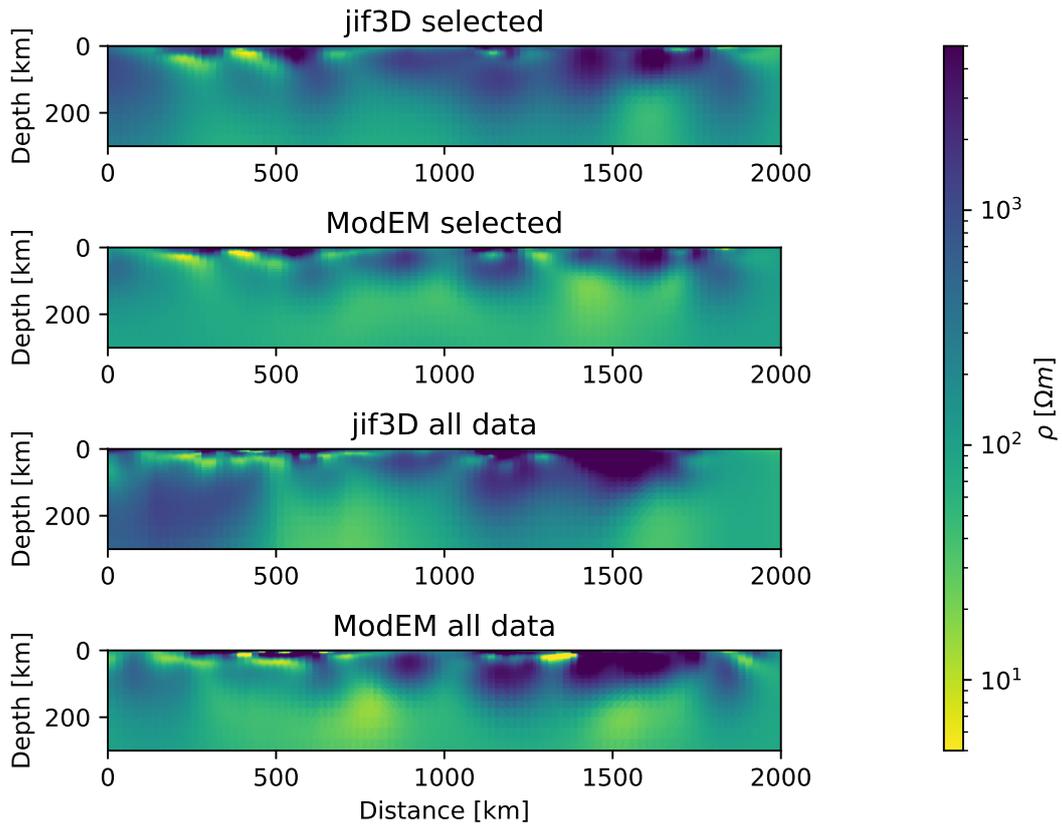
**Figure 11.** Horizontal slices through the inversion models at a depth of 200 km. For an explanation of abbreviations see Figure 8.



**Figure 12.** East-west slice through the four inversion models at 22 degree southern latitude (yellow line in Figure 1).

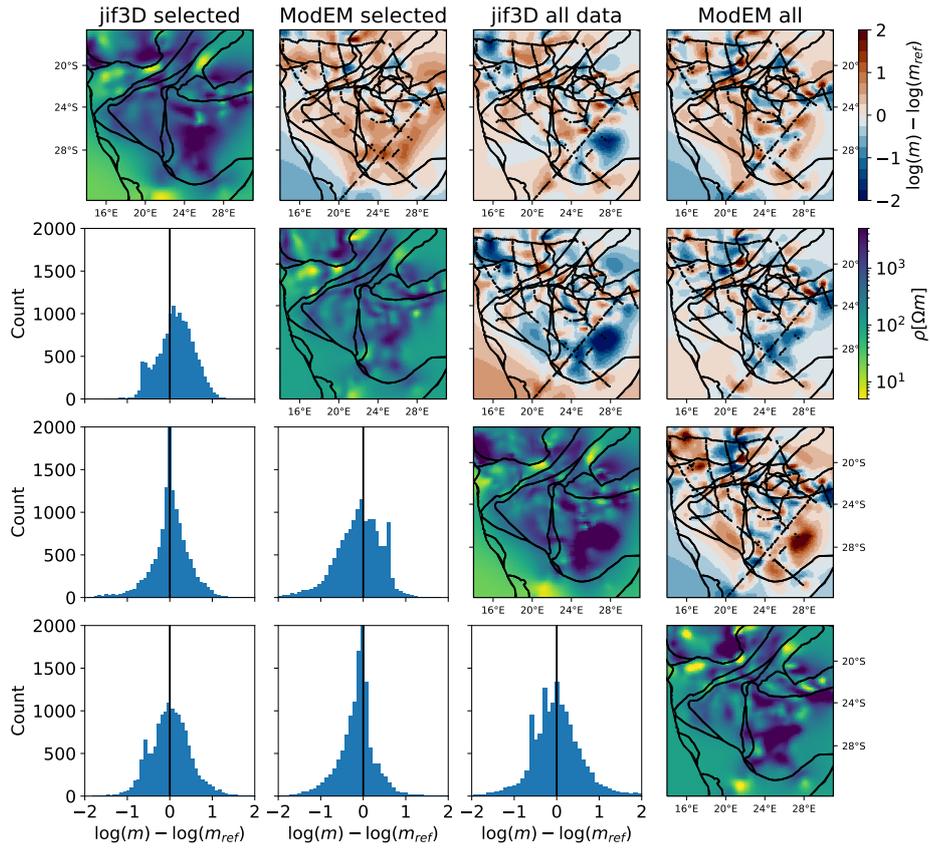


**Figure 13.** Vertical model slices through the four inversion models along the KAP line (green line in Figure 1). From top to bottom the inversion runs are: *jif3D* with selected data, *ModEM* with selected data, *jif3D* with all data, *ModEM* with all data.



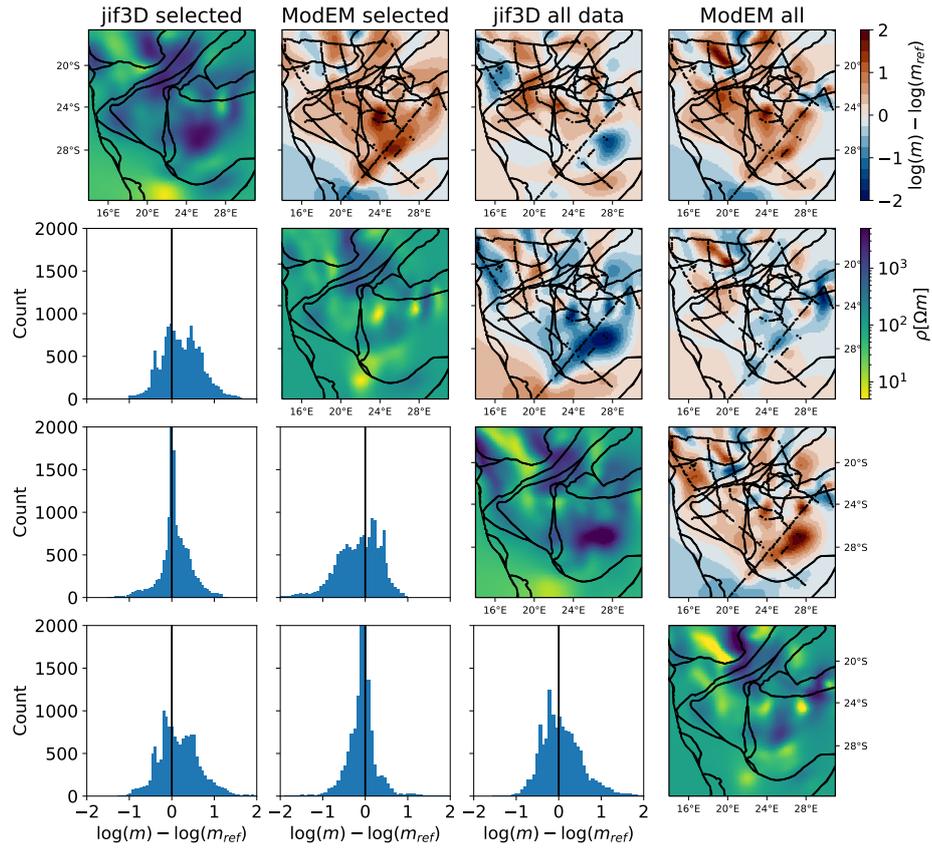
**Figure 14.** Vertical model slices through the four inversion models along the Kimberley line (red line in Figure 1). From top to bottom the inversion runs are: *jif3D* with selected data, *ModEM* with selected data, *jif3D* with all data, *ModEM* with all data.

Model difference at 50 km depth

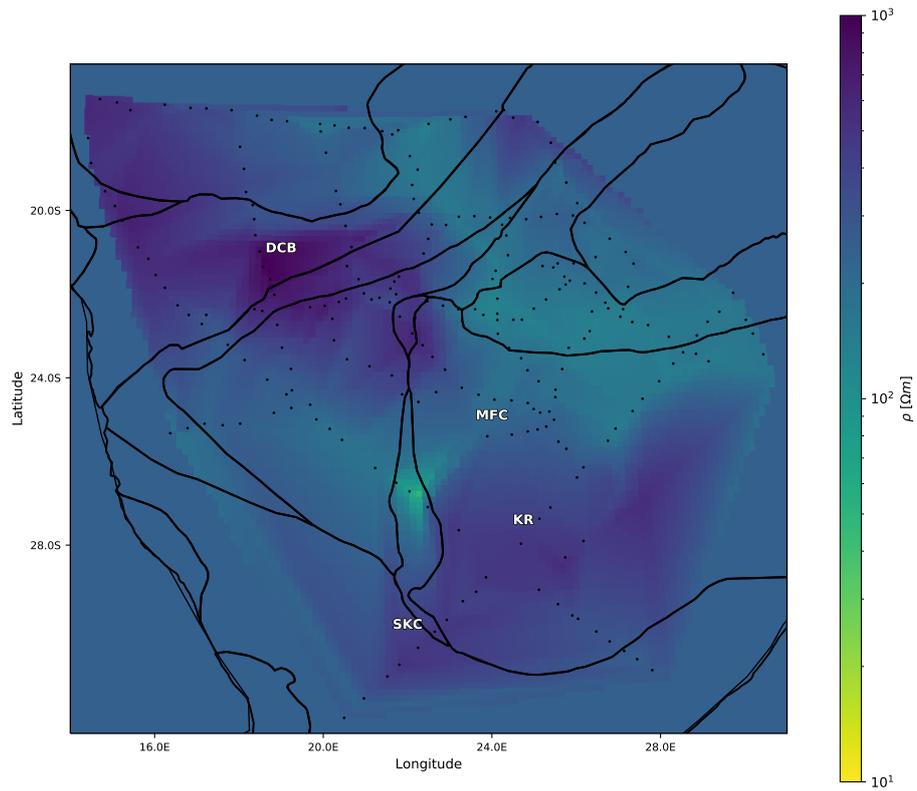


**Figure 15.** Difference matrix for the four inversion runs at a depth of 50 km. We plot the resistivity slice for each model on the diagonal. Plots above the diagonal show the difference in logarithmic resistivity between pairs of models as labelled above each column, Plots below the diagonal show the corresponding histogram.

Model difference at 150 km depth

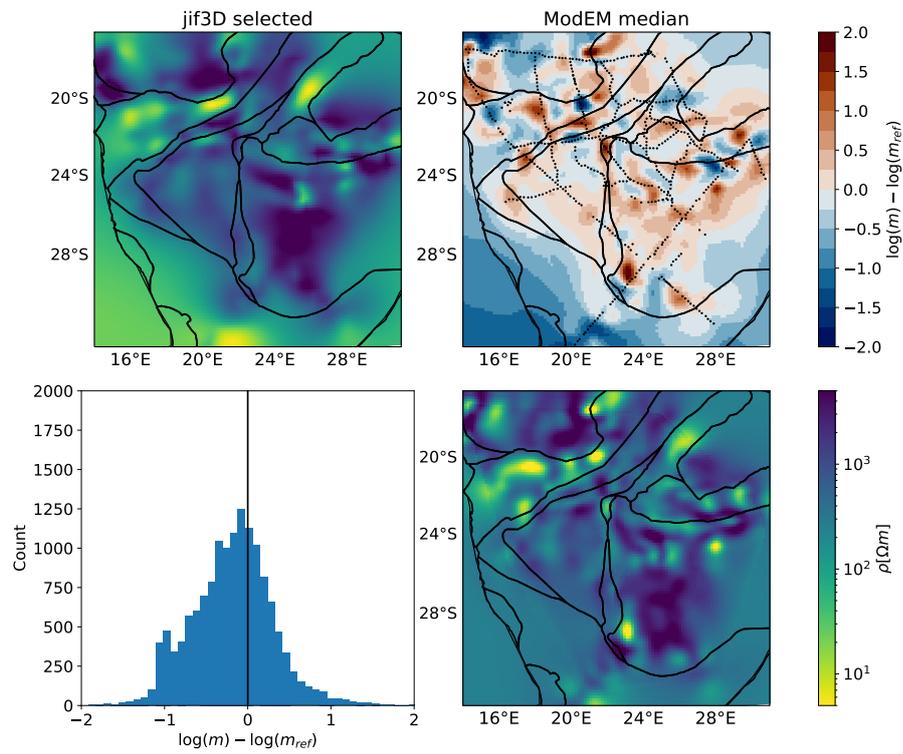


**Figure 16.** Difference matrix for the four inversion runs at a depth of 150 km. We plot the resistivity slice for each model on the diagonal. Plots above the diagonal show the difference in logarithmic resistivity between pairs of models as labelled above each column, Plots below the diagonal show the corresponding histogram.



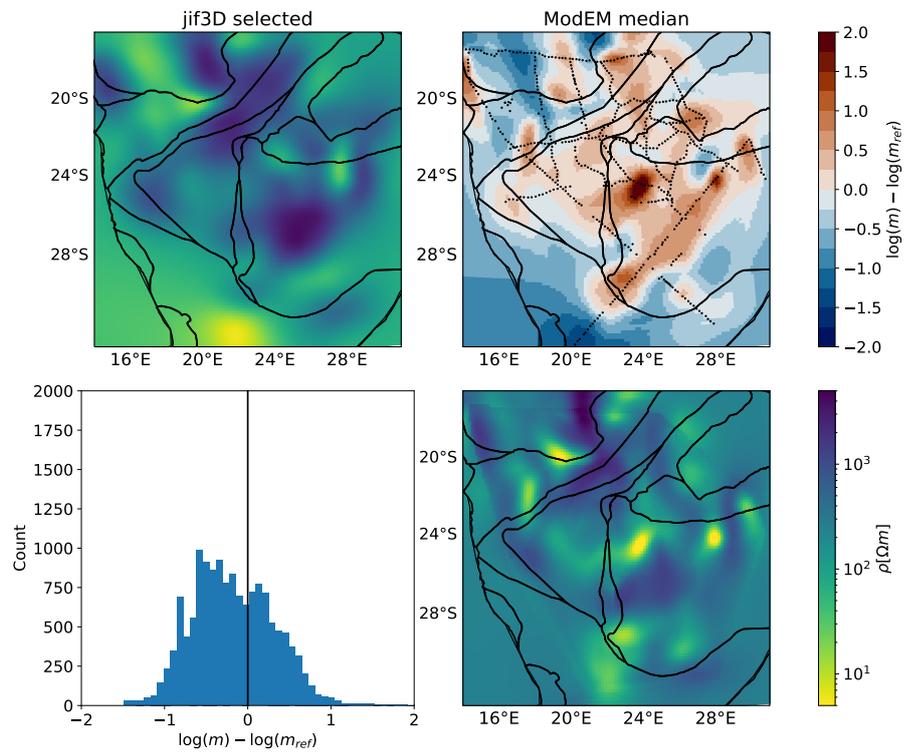
**Figure 17.** Starting model derived from median apparent resistivity for the inversion with ModEM. Note the reduced range of resistivities in the color bar compared to the other model plots.

Model difference at 50 km depth



**Figure 18.** Difference matrix for the ModEM inversion runs with a homogeneous starting model and a median apparent resistivity starting model at a depth of 50 km.

Model difference at 150 km depth



**Figure 19.** Difference matrix for the ModEM inversion runs with a homogeneous starting model and a median apparent resistivity starting model at a depth of 150 km.

614 **Acknowledgments**

615 MM was funded by the German Research Foundation, DFG, under grant MO 2265/4-  
 616 1. SÖ and KS are supported by the Australian Research Council grant FT150100541.  
 617 The SAMTEX magnetotelluric impedance estimates used here are available through the  
 618 IRIS SPUD repository at <https://doi.org/10.17611/DP/EMTF/SAMTEX>. Download in-  
 619 structions for the ModEM inversion software can be found at [https://sites.google](https://sites.google.com/site/modularem/download)  
 620 [.com/site/modularem/download](https://sites.google.com/site/modularem/download). jif3D is available via subversion at [https://svn.code](https://svn.code.sf.net/p/jif3d/jif3dsvn/trunk/jif3d)  
 621 [.sf.net/p/jif3d/jif3dsvn/trunk/jif3D](https://svn.code.sf.net/p/jif3d/jif3dsvn/trunk/jif3d). Figure 1 was prepared using GMT 6.1 (Wes-  
 622 sel et al., 2019), all other Figures with Python. We thank Fabio Cramereri for designing  
 623 perceptually uniform color scales used in some of the plots (Cramereri et al., 2020).

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