

Non-Volcanic Earthquake Swarm Near the Harrat Lunayyir Volcanic Field, Saudi Arabia

Mohammad Youssof¹, P. Martin Mai¹, Adriano Nobile¹, and Sigurjón Jónsson¹

¹King Abdullah University of Science and Technology (KAUST)

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Abstract

Understanding the origin of seismic swarms can be controversial, especially when they occur near volcanic areas. Here, we investigate a seismic sequence which is steadily active in a non-volcanic area close by the volcanic field of Harrat Lunayyir in the western shield of Arabia. Our results unveil a planar zone of seismicity with ~ 5 km long E-W, sub-vertically ~ 9 km south-dipping structure, which is characterized by a dominant tensional focal mechanism. Independent evidence for the tectonic style dominance came from assessing the ground deformation images using the InSAR technique. This local seismicity might be attributed to a reactivated structure along a regional weakness zone of the Najd Fault System, which dominates the Precambrian structure of our area. Comparing the effects of high- and low-frequency datasets for the moment tensor inversion conclude a consistency of our solution. The frequency index analysis for P- and S- waves spectral datasets, does not suggest fluid-driven processes. We observe average stress drop of ~ 5.40 MPa with corner frequency of ~ 2.75 Hz. Our study confirms a localized reactivation of a brittle crustal seismogenic zone in the area of interest. This interpretation relies on the integration of several analysis methods, including spatial and magnitude-frequency distributions statistics.

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3 M. Youssef, P. M. Mai, A. Nobile, and S. Jónsson
King Abdullah University of Science and Technology (KAUST),
Thuwal 23955-6900, Saudi Arabia

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23 seismogenic zone in the area of interest. This interpretation relies
24 on the integration of several analysis methods, including spatial and
25 magnitude-frequency distributions statistics.

26 **Keywords**— Non-volcanic swarm, full waveform moment tensor inver-
27 sion, double-difference relocation, spectral index and stress drop analysis,
28 flow-chart of data processing.

29 Plain Language Summary

30 In this study, we investigate a seismic sequence (started in February 2017),
31 which is steadily active in a non-volcanic area close by (~ 50 km NW) the
32 volcanic field of Harrat Lunayyir in the western shield of Arabia. We conclude
33 that the occurrence of this earthquake swarm is not directly associated with
34 a magmatic cause.

35 In this analysis, we implement various integrated approaches in a sequen-
36 tial workflow that provides a road-map for source parameter estimates. By
37 applying these techniques, ruling out particular causes of seismicity gradu-
38 ally in a step by step process by a comprehensive, integrated data analysis
39 approach is followed. This series of analyses determine whether the seismic
40 sequences are caused by fluid-driven processes as they may occur in any area
41 susceptible to volcano-earthquakes interactions.

1 Introduction

Earthquake swarms can be defined as sequences of events clustered in time and space, lacking a clear main shock (e.g. Hill 1977). They are different from the standard main-shock/aftershock scaling laws (Roland & McGuire 2009). Swarms might be related to the occurrence and migration of fluids which may reduce normal stress along existent faults. In this case, physical processes modulate elastic strain energy released by such frequent events, which often characterized by a hypocentre migration in time and space (e.g. Waite & Smith 2002; Hayashi & Morita 2003; Hainzl 2004).

Swarm-like earthquakes are observed in a diverse range of geological settings including volcanic (Bianco et al. 2004; Guglielmino et al. 2011, Passarelli et al. 2015, White & McCausland 2019), geothermal regions (Dziak et al. 2003), along transform plate boundaries, as well as active rift zones (Baer et al. 2008; Pallister et al. 2010), where earthquake swarms are mainly associated with shallow extensional fractures (Pollard et al. 1983; Rubin & Pollard 1988, Vidale & Shearer 2006).

Swarms' seismic signals can be related to both contributions of fluid- and tectonic-driven processes that may coexist in the same interactive system. Nevertheless, the manifestations of seismicity related to volcanic activity can be spatially and temporally ambiguous, especially when the sequences are close by volcanic areas (Legrand et al. 2002; Hill et al., 2002; Manga and Brodsky, 2006).

The seismotectonic Cenozoic activity in the Arabian Shield is considered to be, at least partially, associated with rifting of the Red Sea that has led to

66 uplift and volcanism throughout the shield, resulting in extensive lava fields
67 (called harrats, locally) that cover an area of $\sim 180,000 \text{ km}^2$ (Coleman et al.,
68 1983) (Figure 1 and supplementary Figure S1). Intraplate volcanism resulted
69 in at least 21 eruptions in Arabia during the past 1500 years (Camp et al.,
70 1987). Some of these eruptions fields display geothermal features such as
71 elevated groundwater temperatures and fumarole emissions (Roobol et al.,
72 1994). It has also been suggested that the entire area of harrats is underlain
73 by asthenospheric flow channelized northward from Afar (Camp and Roobol,
74 1992; Hansen et al., 2006; Chang et al., 2011).

75 Earthquake swarms in Arabia are taking place at different locations where
76 they are recorded by dense seismic arrays of ~ 300 stations (Soliman et al.
77 2019). The seismically active regions around the Red Sea flanks are gener-
78 ally similar in terms of formation age, and dominant geological settings. The
79 main geological exception represents a prominent volcanism at the eastern
80 margin of the Red Sea, in contrast to the western side in Egypt and Sudan
81 where a few flood basaltic fields exist (Pallister 1987). Regardless, the origin
82 of several seismic activities in this region are enigmatic. As shown in the
83 supplementary material (Table S1 and supplementary Figure S1), we assem-
84 bled some data about these swarms. Different studies suggests that some of
85 them are triggered by magmatic processes underneath or close by the surface
86 exposure of the harrats areas.

87 This western part of the Arabian Shield, where Harrat Lunayyir is lo-
88 cated, comprises of amalgamated belts of sedimentary and metamorphic rock
89 units that are penetrated by the regional Precambrian Najd Fault System
90 and intervened by numerous dikes (e.g., Blasband et al. 2000; Johnson 2003).

91 Harrat Lunayyir is a small volcanic field in the Arabian Shield, covering an
92 area of $\sim 3500 \text{ km}^2$, situated within the passive margin of the northern Red
93 Sea region. This region contains large number of volcanic cones (>50) that
94 follow the NW-SE trending normal faults (Baer and Hamiel, 2010; Al-Amri
95 et al., 2012, Jónsson, 2012, Trippanera et al. 2019).

96 Here, we mainly focus our investigation on a sequence of relatively small-
97 magnitude events located NW of Harrat Lunayyir, which started in February
98 2017 and it is still taking place with a daily rate. The Saudi Geological Survey
99 (SGS) provided us with ~ 10 months waveform dataset of events started from
100 February 8th, 2017 (characterized by a magnitude range of M_L -0.75 to 3.73),
101 registered in this locality (Figure 1). This swarm attracted the interests of
102 local authorities due to previous intense seismic activity in the vicinity in
103 2009, related to volcanic unrest at Harrat Lunayyir. The local network of
104 SGS registered this unrest with more than 30k events between April to July
105 2009 with magnitudes ranged from M_L -0.7 to 5.4 (Baer and Hamiel, 2010).
106 This activity also experienced a dike intrusion, including ~ 8 km surface
107 rupturing M_W 5.7 earthquake (Pallister et al. 2010). In the post diking
108 phase, micro-seismicity has been continuously registered in the dike-induced
109 graben up to present (Nobile et al. 2020).

110 In this study, we integrate several seismic analysis methods to mainly in-
111 vestigate the properties of the 2017 earthquake sequence. We inspect the
112 spatiotemporal statistics from the catalog information by examining the
113 magnitude-frequency distribution. We apply the double-difference earth-
114 quake relocation algorithm (Waldhauser et al. 2000) to relocate the swarm
115 events. We constrain source depth and focal mechanism for selected events

116 using full waveform moment tensor inversion constrained with a grid-search
117 over source depth (Ichinose et al. 2003). Furthermore, we estimate static
118 stress drop, corner frequency, and seismic moment from displacement am-
119 plitude spectra for the largest events. We then apply the frequency index
120 method to characterize the spectra of the swarm signals (Buurman & West
121 2010). Finally, to further constrain our results, we evaluated the ground
122 deformations in the area through InSAR imaging. To compare the results
123 of analyzing the 2017 swarm with some swarms around, we study two other
124 seismic sequences which occurred in 2009 and 2018. For analyzing these
125 two cases, we use only two methods because of data limitation. Our inte-
126 grated data analyses lead us to conclude that the ongoing seismic sequence
127 (started in 2017) is of a tectonic origin, and not directly linked to the nearby
128 continuous activity of the volcanic field of Harrat Lunayyir.

129 **2 Earthquake Data : Statistics and Locations**

130 In the study area, we use seismic records from 44 stations operated by SGS.
131 The network geometry forms a polygonal area of $\sim 375 \times 200 \text{ km}^2$ with a cente-
132 riod point of $25.5^\circ, 37.5^\circ$ bounded between Harrat Lunayyir to the south and
133 the northern 2017 swarm location (Figure 1 and supplementary figures S1
134 and S2). Despite 33 existing stations in the area before this 2017 swarms be-
135 gun, SGS densified the local network by adding 11 more permanent stations
136 during the first three months of the swarm. All instruments are broadband
137 three-component sensors with the ability to record static displacements down
138 to the DC offsets and up to frequencies limited by the sampling rate of 100

139 Hz. SGS provided the data in a compacted full-SEED format with a time
140 duration of 565 s, for a period spanning from February 2017 to November
141 2017.

142 We filtered the seismic data between 1–25 s, depending on the study
143 purposes and method. We categorize the high frequency range between 1–10
144 s, and the low frequency range between 5–25 s. The advantages of using
145 different frequency bands is multifold. For instance, the relatively long-period
146 signals (0.04 Hz – 0.15 Hz) improve the estimation of earthquake source
147 parameters because they are relatively insensitive to the effects of lateral
148 velocity and density heterogeneities (e.g., Ritsema and Lay, 1995, Ichinose et
149 al. 2003). Using relatively short-period signals (0.1 Hz – 1.0 Hz) help refining
150 the sensitivity for structural details at a given depth and help verifying the
151 velocity model used in the inversions.

152 The first-order approach to identify whether seismic events accommo-
153 dated by some fault-like characters or not is to estimate the magnitude-
154 frequency distribution (MFD) in addition to applying a relative relocation
155 technique. The following two subsections will help shaping an initial rough
156 understanding of the temporal and spatial evaluations of the 2017 seismic
157 swarm, north of Harrat Lunayyir region.

158 **2.1 Magnitude Frequency Distribution**

159 Statistical properties of a given seismicity can be analyzed using MFD. This
160 method describes the rate of events occurrences across all magnitudes. We
161 determine the MFD following the Gutenberg-Richter relationship: ($\log N = a - bM$),

162 where N is a number of events having a magnitude $\geq M$, while a and b are
163 constants. The b -value indicates the ratio of small to large events, the con-
164 stant a is the logarithm of the events number with $M \geq 0$, which quantifies
165 the events productivity of a sequence. In this context, the magnitude of com-
166 pleteness (M_C), represents the lowest magnitudes that is reliably recorded
167 by the seismic network. We estimated the M_C by the maximum curvature
168 approach (Wiemer and Wyss, 2000; Woessner and Wiemer, 2005), which
169 defines it by the largest value of the second derivative of the MFD curve.

170 For calculating the MFD, a complete catalog should be used containing
171 magnitudes $M_L \geq M_C$, and M_L ranges at least over 2.0 magnitudes. Note
172 that only the MFDs derived from similar definition of local magnitude are
173 comparable. The network-based standard magnitudes produced by SGS are
174 based on two definitions for M_L (Soliman et al. 2019), depending on the area
175 distances and tectonics (supplementary material: Appendix (A)).

176 For a total number of 390 events for the 2017 swarm within the local
177 area (NNW Harrat Lunayyir, Figure 1), we obtain a b -value of 0.73, M_C of
178 0.73, and $a= 3.10$. For comparison, we also use a complete catalog ($\sim 15k$
179 events during 2015-2018) within the entire area of Lunayyir volcanic field.
180 Calculating b , M_C , and a -values for the whole region (during these four years,
181 excluding the events of 2017 swarm) results in b -value=1.27, $M_C=0.25$, and
182 $a=4.02$ respectively (Figure 2). These results of the two M_C values agree
183 with Soliman et al. 2019.

184 Figure (2) shows the main difference between the background larger-scale
185 seismicity and the local 2017 swarm. Figure 2a represents the seismicity
186 peak during the 2017 unrest within the temporal and spatial boundaries

187 of this seismic activity, while the cumulative seismic moment inset curve
188 shows the jump in the moment release versus time for the 2017 and 2018
189 main events. Figure 2b confirms the varying statistical relation between the
190 MFD curves for the two different seismicity. This statistics result suggests
191 an interpretation of the 2017 activity to be more of a tectonic origin than a
192 magmatic one.

193 **2.2 Relative Relocation**

194 We compute relative earthquake locations using the double-difference tech-
195 nique (Waldhauser and Ellsworth, 2000), based on an enhanced HypoDD
196 code that includes 3D ray tracing to calculate travel times within a volu-
197 metric velocity model. The standard process of this algorithm iteratively
198 minimizes arrival time residuals using weighted least squares methods, with
199 either a singular value decomposition (SVD) or a conjugate gradient (LSQR)
200 approaches. SVD performs well for up to few hundreds of events to pro-
201 duce more accurate error estimates than the computationally efficient LSQR
202 method (Waldhauser and Ellsworth, 2000). As background models, we use
203 the P-wave velocity model of Tang et al. (2016) along with a calculated
204 S-wave model (Figure 3) using a constant V_P/V_S of 1.76.

205 We apply this algorithm to obtain precise hypocenters of a total of 390
206 earthquakes, as reported in the SGS catalog for ~ 10 months in 2017. Earth-
207 quakes locations before and after applying the double-difference technique
208 are shown in Figure 3 and supplementary Figure S3.

209 Relative relocations results show interesting space-time pattern for the

210 events with larger magnitudes clustered in the swarm beginning (during the
211 first few months), at the deepest level of ~ 12 km. Events with smaller
212 magnitudes progressively migrated upward. Shallower events cluster between
213 5 to 8 km depth, forming an E-W narrow corridor of 5 Km length (as shown
214 in the surface projection in Figure 3 and supplementary Figure S3). The
215 bulk spatial shift in the horizontal E-W plane between the initially located
216 events by SGS and the new relocated events is 0.03° while the difference in
217 depths represents clustering the initially located scattered events into deeper
218 depths for the new relocated ones (Figure 3). Overall, the original locations
219 show a diffuse spatial pattern whereas the relocation solution represents a
220 sense of fault-like structure.

221 The hypoDD errors depend on the array geometry, data quality, and
222 maximum separation between any pair of events, where this offset has to be
223 at least 10 times smaller than event-station distance. The available stations
224 here are sparsely distributed (uneven but dense array, Figure 1) but of a
225 good data quality (supplementary Figure S4). We obtain a total number of
226 21138 P- and S- wave differential travel-times using the 44 stations around
227 the events. The average offset between linked events is 2.2 km, while the
228 maximum offset is 14.8 km. This offset ranges are within the average station
229 separation of ~ 40 km of the array, and events within the region are on average
230 linked by at least 10 arrivals. We choose the SVD method as it produces
231 reliable error estimates in this case of small dataset. Note that the double-
232 difference relocations have much smaller errors than the network locations.
233 While the aim is to relocate the swarm by combining all the P- and S-wave
234 available datasets, relocating events using each data type independently was

235 useful to assess the solution consistency and quality of both datasets.

236 Utilizing this dense array and its high-quality data (Figures 1 and 4a, re-
237 spectively), we apply a waveform-based sensitivity test for constraining the
238 location uncertainty using the largest event of the 2017 swarm as a refer-
239 ence. Backprojecting the incident rays into the source via a beamforming
240 technique amplifies phases with the appropriate slowness, while suppressing
241 incoherent noise and phases with different slowness (supplementary material:
242 Appendix (B)). The frequency–wave number analysis (fk-analysis) measures
243 the complete slowness vector (i.e., back azimuth and horizontal slowness si-
244 multaneously), and allows to calculate the power distributed among different
245 slownesses and directions of approach (Aki and Richards, 1980).

246 In the current case, for the coherent incident waves with a frequency of 1
247 Hz, the maximum power spectral density (PSD) for a P-wave signal arrives
248 with a slowness of 15.65 s/deg and a back azimuth of 324° (Figure 4c). For
249 comparison, we calculate the expected phase travel times and ray parameters
250 given the coordinates of the stations and relocated source for the same event.
251 The average theoretical ray parameters are 16.1, and 28.7 s/° for P- and S-
252 wave, respectively (Table S2). These predicted values agree well with the
253 fk-analysis, confirming the observed and relocated hypocenter of the event
254 of interest.

255 **3 Earthquake Source Characteristic**

256 Full waveforms techniques for investigating focal mechanism and spectral
257 content can help reflecting some features of the fault plane. For instance,

258 seismic moment tensors provide a useful tool for distinguishing between tec-
259 tonic earthquakes and events associated with volcanic processes (e.g. Dreger
260 et al., 2000), as well as other man-made sources of seismic radiation such as
261 explosions or mining activity (e.g. Ford et al., 2009). Additionally, analyz-
262 ing the spectral content can identify the radiated seismic energy and hence
263 predicts the stress-changes. In this section, we apply a couple of techniques
264 utilizing full-waveform data of some selected events of interest (taken place
265 in 2009, 2017, and 2018), to delineate the frequency contents and the focal
266 mechanisms. Furthermore, we apply InSAR imaging to assess whether any
267 discernible ground deformation was associated with the 2017 seismic swarm.

268 **3.1 Seismic Moment Tensor Inversion**

269 We use both first-motion fault mechanisms and full-waveform moment tensor
270 inversion, following Ichinose et al. (2003). We compute Green’s functions for
271 2 km depth increments, using a fast reflectivity and fk-summation (Zeng &
272 Anderson 1995). We then iteratively solve for the source depth using a grid
273 search scheme.

274 The sensitivity of the moment tensor solutions was tested by using dif-
275 ferent local and regional velocity models as implemented in the relocation
276 method, as well as by using different frequency bands in the inversions. Note
277 that using long-period energy avoids the need for modeling complex crustal
278 structure, while large epicentral distances allow for using simple 1D velocity
279 models (Jost and Herrmann 1989).

280 The quality of waveform fits for different stations are shown in Figure

281 5 (for the M_W 3.6 event of March 10, 2017) and in supplementary figures
282 S5 and S6 (for some more selected stations recorded the same event). The
283 mismatches in phase show an average variance reduction of 86% and 83% for
284 the long- and short-period data, respectively, where the misfit in amplitude
285 between observed and synthetic waveforms provides useful information about
286 the accuracy of the available velocity model.

287 We compare both solutions of high- and low-frequency moment tensors.
288 Both results (Figure 5) provide same fault-plane solution, indicating that
289 the local velocity structure model used in the inversion is accurate to predict
290 both high- and low-frequencies signals. Furthermore, this solution confirms
291 the exact same geometrical trend of E-W fracture zone implied by the relative
292 relocation analysis (Figure 3). The depth of the largest events are reasonably
293 in agreement in both methods of double-difference and full-waveform moment
294 tensor inversions. More details about the moment tensor inversion output is
295 shown in Figure 5, which is also presented in the supplementary materials
296 with all stations used in the inversion (supplementary figures S5 and S6).

297 The same inversion procedure is additionally applied to eight more events,
298 ranging in magnitude from M_W 2.8 to 3.6 from the 2017 swarm. The results
299 are summarized using the fundamental lune of Riedesel & Jordan (1989) and
300 Tape & Tape (2018) (Figure 6a). This plot visualize the geometry of a point
301 source moment-rate tensor estimates. It also demonstrates the decomposition
302 of moment tensors into isotropic (ISO), double-couple (DC) and compensated
303 linear vector dipole (CLVD) components. This result (Figure 6a) reveals that
304 the majority of the events are double-couple components with a small CLVD
305 contribution.

306 As we acquired knowledge of which nodal plane is the main fault, a stress
307 inversion from the focal mechanism can be conducted. Note that the main
308 stress regime is a function of the orientation of the principal stress axes and
309 the shape of the stress ellipsoid, meaning it results in extensional mechanism
310 when σ_1 is vertical. We therefore compute the stress axes following Vavryčuk
311 (2014) where the input data for the inversion are the strike, dip, and rake
312 angles obtained from the moment tensor solutions. The stress tensor inver-
313 sion results in a sub-vertical σ_1 axis and sub-horizontal σ_2 and σ_3 axes as in
314 Figure 6b.

315 **3.2 Stress Drop and Spectral Index Analysis**

316 One of the important earthquake source parameters is stress drop $\Delta\sigma$, i.e.,
317 the difference between the average shear stress on the fault plane before and
318 after an earthquake. The main consideration about this method is the results
319 non-uniqueness because $\Delta\sigma$ uncertainty quantification is not often helping
320 to interpret results with confidence (e.g., Abercrombie 2015). For instance,
321 some stress-drop studies show higher stress drops for both normal (Shearer et
322 al., 2006) and strike-slip events (Allmann and Shearer, 2009), whereas others
323 report no dependence on focal mechanisms (e.g., Oth, 2013). Other studies
324 suggest that stress drop depends on tectonic setting, depth, or both (e.g.,
325 Boyd et al., 2017).

326 In this work, using the source model of Brune (1970) and Madariaga
327 (1976), we estimate stress drop from the Fourier source spectra (computed
328 for the displacement time-series), which include the corner frequency f_c (e.g.,

329 Boatwright, 1984). We calculate ($\Delta\sigma$) using the Eshelby (1957) relationship
330 (supplementary material: Appendix (C)).

331 Figure 7 shows few findings of our spectral-fitting procedure for the largest
332 earthquake at different stations. These examples represent results of the four
333 main azimuths, which surround the M_L 3.73 earthquake. From this example,
334 the best fitting theoretical model (dashed blue line) has corner frequencies
335 between 2.14 Hz and 4.95 Hz, with a stress drop of 4.56 MPa and 11.32
336 MPa, respectively. Most values for the stress drop and seismic moment fall
337 in the ranges (0.95 - 17 MPa) and ($0.58 \times 10^{13} - 1.74 \times 10^{14} Nm$), respectively
338 (Figure 8). Additionally, we notice that some parts of the amplitude spectra
339 can not be fitted using the predicted models. At low-frequency (≤ 0.7 Hz), the
340 misfit might be attributed to the static and permanent displacements where
341 the background seismicity can be a reason for such low-frequency noises.
342 While at the other end of the spectrum, a high-frequency range (≥ 35 Hz)
343 contaminates the signal with less contribution than the low frequency (Figure
344 7).

345 Furthermore, our estimates of $\Delta\sigma$, f_c , and M_0 indicate azimuthal vari-
346 ations around event hypocenters. The azimuthal variations for the median
347 $\Delta\sigma$ range from 4.7 to 6.9 MPa over different epicentral distances (Table S3).
348 This variation needs further investigation, which is out of this paper's scope,
349 but this directional variability is probably due to directivity effects. Our
350 results also indicate that the individual event stress drops are heterogeneous
351 and span from 0.95 to over 17 MPa (for the largest three events, as shown in
352 Figure 8). Note that the upper limit is not reliably well-determined because
353 resolution decreases for corner frequencies. Therefore, we estimate the $\Delta\sigma$

354 uncertainty using the spectra of P- and S-waves for a comparison calculations
355 of f_c and M_0 . The results are shown in Figure 8.

356 Another method to discriminate between different source processes can
357 be deduced from the ratio between separated frequency bands within a given
358 seismic signal. The dominant frequency, can be also used as a general proxy
359 for spectral content and to characterize waveform types (e.g. Latter, 1980;
360 McNutt, 2002). However, shortcomings arise when using it as a measure of
361 the overall frequency content, for instance, in case of low signal-to-noise ratio
362 (SNR) recordings or for events with bimodal frequency distributions, because
363 the dominant frequency measures only the highest peak in the spectra and
364 therefore grouping it with other single-peaked events (a particular issue for
365 hybrid-type earthquakes).

366 These limitations associated with dominant frequency led Buurman and
367 West (2010) to develop a measure to discriminate between different types of
368 seismic events, defines the frequency index (FI) based on the ratio of energy
369 in low and high frequency windows (supplementary material: Appendix (D)).
370 For instance, waveforms with equal amounts of high and low energy (as
371 subjectively defined) will have a frequency index around zero. Whereas,
372 a smaller FI than this average means the waveform is dominated by low-
373 frequency energy, while otherwise FI demonstrates a majority of energy in
374 the high-frequency band.

375 To calculate the FI in a consistent manner, we first pick the P- and
376 S-onsets, minimizing the time window to approximately the P-S duration,
377 followed by removing the average amplitude from the selected waveforms
378 signals, with a fixed time series duration of 40 seconds: 10 seconds prior

379 to the earthquake P-onset and 30 seconds after it, ensuring that the high
380 frequency signal is fully captured in the Fourier analysis. Examples of the
381 *FI* analysis for the largest events in the 2009, 2017, and 2018 sequences are
382 shown in Figure 9 with the results values listed in Table S4. Here, this index
383 classified the main event of 2017 as an exclusively high-frequency event, which
384 is contrasted the other known magmatic case of 2009 in Harrat Lunayyir.

385 **3.3 Ground Deformation using InSAR**

386 As magma moves to shallower levels below the surface, it usually produces
387 characteristic ground deformation, seismicity, and gas emissions (e.g., Dzurisin
388 2007; Biggs and Pritchard, 2017; Sigmundsson et al. 2018). During mag-
389 matic intrusions, the seismic moment could be a small fraction of the total
390 geodetic moment (e.g., Nobile et al. 2012). In our area of study, the 2009 Har-
391 rat Lunayyir swarm, which occurred ≥ 50 km southeast of the 2017 swarm,
392 was caused by an ascending magma intrusion that, using InSAR data, was
393 estimated to be ≥ 10 km long, with a volume of 0.13 km^3 , and stops at ~ 1
394 km below the surface (Pallister et al., 2010). Furthermore, the dike intrusion
395 produced over ~ 1.5 m of SW-NE extension as well as 60 cm of graben sub-
396 sidence above the intrusion (Jónsson 2012). Pallister et al. (2010) reported
397 that $\geq 93\%$ of the deformation observed during the 2009 dike intrusion was
398 aseismic. Therefore, the amplitude and pattern of the ground deformation
399 could give valuable information about the origin of the seismicity.

400 Geodetic remote sensing techniques, such as InSAR, allow measuring
401 ground deformation even in areas where ground-based networks are not

402 present, as the case of the area affected by the 2017 swarm. We, there-
403 fore, used InSAR to detect any ground deformation in the area to constrain
404 the results obtained by the analysis of the seismic data. We selected SAR
405 scenes from the Sentinel-1 A/B satellites acquired between January 2017 and
406 January 2019, a total of 51 images from ascending track 87 and 89 images
407 from descending track 123. We processed 100 ascending and 266 descending
408 orbit interferograms with spatial baselines smaller than 200 m and temporal
409 baselines up to 36 days (supplementary Figure S7).

410 Due to high coherence, the resulting interferograms could be easily un-
411 wrapped and used to calculate deformation rate maps in the line of sight
412 (LOS) of the satellites with the Small Baseline Subset (SBAS) technique
413 (e.g., Samsonov 2017). The initial rate maps showed deformation correlated
414 to topography, indicating significant elevation-related atmospheric delays.
415 We reduced these signals by estimating linear correlation coefficients between
416 elevation and the signal and subtracted the results from the rate maps (e.g.,
417 Neelmeijer et al. 2018). However, we were not able to remove completely the
418 signal-topography correlation as evident in the southern part of the ascend-
419 ing deformation rate map (Figure 10a). The final deformation rate maps
420 mostly show smooth variations of ± 0.5 cm/yr (Figure 10), which are due to
421 the noise of the interferograms that could not be fully removed in the time-
422 series analysis. The only clear deformation signal is located ~ 10 km north
423 of the swarm location (Figure 10b), in a narrow WNW-ESE elongated area
424 that corresponds to an ephemeral riverbed (Wadi). This area shows up to 1
425 cm/yr of displacement toward the satellite for both viewing geometries. This
426 corresponds to an uplift of 1.2 cm/yr, which might be attributed to water

427 level changes of the shallow aquifer. No clear deformation is observed in or
428 around the area of the 2017 seismic swarm in these rate maps.

429 We use analytical models to quantify the expected ground deformation
430 due to the seismic swarm. The relocated events are distributed over a 5
431 km \times 9 km planar-like volume that dips $\sim 15^\circ$ SSW with its upper edge at
432 ~ 5 km depth. Given this geometry, a normal focal mechanism of the main
433 events and the total seismic moment of $\sim 12.5 \times 10^{14}$ Nm, less than half
434 a mm of surface displacements would be expected, i.e., less than what is
435 detectable by the InSAR technique. Using the spatial extent of the current
436 swarm as dimensions for a possible dike intrusion (5 km long, 9 km wide at
437 5 km depth) and assuming an opening of 0.5 m, which corresponds to $\sim 1/6$
438 of the volume of the 2009 intrusion, the predicted surface deformation is ~ 2
439 cm that would have been detected by InSAR. However, there is no evident
440 ground deformation in the two InSAR rate maps (Figure 10). Therefore, the
441 InSAR data analysis suggests that the seismic swarm was not accompanied
442 by a magmatic intrusion.

443 4 Discussion

444 The current seismic analysis focuses on one of the most recent earthquake ac-
445 tivity nearby Harrat Lunayyir area. Since February 2017, a swarm located to
446 the north of Harrat Lunayyir is being recorded continuously, with a maximum
447 magnitude of M_W 3.60. We study the source properties using the available
448 seismic records, applying double-difference algorithm, full-waveform moment
449 tensors inversion, frequency index analysis, and stress drop estimations.

450 To identify the activity source-type, we propose a well-defined workflow
451 (supplementary Figure S8), applying a suite of seismological tools. Addition-
452 ally, we advocate analyzing InSAR images to complement the seismological
453 data and results. This flow-chart proposed in this study may serve as guid-
454 ance for future studies on seismic swarms, to characterize and quantify their
455 properties using multiple datasets and analysis techniques to help discrimi-
456 nate volcanic from non-volcanic events.

457 In a regional geographic context, the shield area of the Red Sea flanks
458 is active with a continuous background seismicity. Different kinds of seismic
459 events have been observed in this area as reported in table S1 with some of
460 their main characteristics (supplementary Figure S1).

461 Harratt Lunayyir volcanic field (~ 50 km SE of our study area) hit by a
462 seismic swarm, with intense rate in the first four months between April and
463 July 2009. In this period, more than 30k recorded events struck the area with
464 many events of $M_L > 4$ (e.g., Pallister et al. 2010, Baer & Hamiel 2010, Al-
465 Amri et al. 2012). Several seismic and geodetic studies have confirmed the
466 magmatic intrusion origin as the primary cause of this activity (e.g., Jónsson
467 2012, Duncan & Al-Amri 2013, Koulakov et al. 2014 and 2015, Xu et al.
468 2016). It is worth mentioning that Harrat Lunayyir region is still under a
469 steady background seismicity (Figure 2a).

470 The ongoing activity of our main focus here started in February 8th, 2017,
471 with seven largest events between M_L 3.0 to 3.73, where all of these relatively
472 large events occurred during the first four months since the swarm started.
473 Additionally, another swarm started in October 2018 around Umm-Lujj area,
474 ~ 25 km SW of Harrat Lunayyir, with a maximum magnitude of M_L 3.70.

475 To compare between three swarms in the study region, we start with ap-
476 plying the first step in our flow-chart (statistics with MFD). This comparison
477 study was not conclusive because of the lack of complete datasets (mainly
478 limited catalog for the located events using the standard network approach,
479 in addition to very few available waveforms). Nevertheless, its results turn
480 out to conduct a first-order comparison as it indicates that swarms located
481 to the SW (2018) and SE (2007) of Lunayyir are more associated with rel-
482 ative high b-values, analogous to the background seismicity. Our analysis
483 reveals the b-value varies between 0.85 and 1.3. This high b-value may be
484 associated with transporting fluids out of the deep volcanic system in the
485 region, as interpreted in previous work of Blanchette et al. (2018). Farrell
486 et al. (2009) also concluded that high b-value (up to 1.3 ± 0.1) is attributed
487 to the presence of a high thermal gradient due to fluids emplacement, while
488 the low b-value (as low as 0.6 ± 0.1) might be caused by crustal stress from
489 regional loading.

490 Generally, the b-value could be also connected to the rock physical prop-
491 erties. For instance, Wyss et al. (1997) and Wiemer et al. (1998) pointed out
492 that low b-values could correspond to breaking asperities while the high b-
493 values correspond to creeping sections of faults or due to magmatic processes,
494 where seismicity may also be dominated by the creation of new fractures un-
495 der stress build-up. According to Urbancic et al. (1992) and Wyss et al.
496 (1997), an increase in applied shear stress will be decreasing the b-value.

497 The high b-value could also be indicative of a relative low stress regime
498 resulting from the energy releases by continuous earthquake activities in the
499 vicinity (e.g., Farrell et al. 2009). Another scenario, specifically valid for

500 the SW 2018 swarm, comes from being close to the sea which may cause the
501 presence of fluids in the fault system. In contrast, we found the northern area
502 of the 2017 swarm is characterized by low b-values (0.73 ± 0.03 , Figure 2).
503 This relatively low b-value can be interpreted as a hint of evidence for a high
504 stress regime associated with a dominant extension, which is expected to be
505 found in such intraplate tectonic settings (e.g. Wolfe et al. 2003, Keir et al.
506 2009). In this study, the observed b-values difference between the northern
507 and southern swarms tend to attribute them to different origins.

508 The relative relocation for the 2017 swarm show clustering of the largest-
509 magnitude and earliest events at deeper mid-crustal levels different from the
510 shallow, small-magnitudes events which taken place later in time. This shows
511 an upward time migration of the large early events to form the later (long-
512 lasting) small-size upper crustal events (Figure 3). These results highlight
513 the presence of a fault zone that is accommodating an active strain within
514 the regional Najd Fault System. This observation may imply an evidence for
515 a potentially reactivation mechanism within this Precambrian shear zone.

516 To estimate the relocation errors, we applied a sensitivity analysis by
517 backprojecting the incident rays of the main event. We calculate the expected
518 phase travel times and ray parameters. The predicted values agree well with
519 the fk-analysis, confirming the relocated hypocenter of this event of interest.
520 The uncertainty in slowness values is small ($0.45 \text{ s}/^\circ$), where the backazimuth
521 values have $\sim 1^\circ$ difference. A source of such shift is attributed to the use of
522 only one pair of event-station for the synthetics while using several stations
523 in the fk measurements, however, also the lack of an accurate 3D velocity
524 model contributes to the location uncertainty.

525 The full waveform moment tensor inversion using the largest event of
526 the 2017 swarm shows a typical quality of waveform fits from the traces pre-
527 sented in Figure 5. We used all available stations (supplementary Figure S2),
528 thereby minimizing the effect of the model uncertainty along any given ray-
529 path on the moment tensor solution. Despite relying on this M_W 3.6 event
530 in the interpretation, we also applied the inversions on eight more events of
531 this swarm. We plotted all the inversions results using the fundamental lune
532 plot (Figure 6a). For some waveforms, the amplitude mismatch between ob-
533 served and synthetic low-frequency signals may contain information about
534 the large-scale, structural-related, corrections needed to better calibrate the
535 velocity models.

536 We point out that a reliable velocity model is vital to pursue the full
537 waveform inversion of moment tensor as well as for an accurate relative relo-
538 cation. The two velocity models examined in this study belong to the SGS
539 regional model as reported in Soliman et al. (2019), in addition to the local
540 model of Harrat Lunayyir developed by Tang et al. (2016) (Figure 3). We
541 selected the model of Tang et al. (2016), which is constructed using both P-
542 and S-waves receiver functions and surface waves dispersion measurements
543 to constrain the structure underneath the study area. This velocity model
544 has also some finer details, as the imaged low-velocity seismic perturbations
545 of the mid crust, which might helped for a better relocation and signified the
546 waveforms fits for the high-frequency moment tensor solution.

547 We examine earthquake source properties in terms of stress drop ($\Delta\sigma$)
548 which is proportional to the total seismic moment and rupture size, and could
549 help defining the tectonic environment (e.g., large stress drops are related to

550 more high-frequency energy release). In our analysis for the largest three
551 events of the 2017 swarm, we perform a grid-search to find the parameters
552 that best model the spectrum characteristic; ω_0 , and f_c . Our estimates of
553 these parameters indicate azimuthal variations around event hypocenters.
554 The median of (f_c) and ($\Delta\sigma$) around 2π circumference ranges from ~ 2.3
555 - 3.2 Hz and ~ 4.4 - 6.9 MPa, respectively. This result confirms a similar
556 finding for intraplate events, by Kanamori & Anderson (1975), and Allmann
557 & Shearer (2009).

558 Additionally, we apply *FI* analysis which calculates the mean amplitude
559 of two spectral bands (high (A_{hf}) and low (A_{lf}) ranges) to help describing the
560 relative spectral content of a single event. This is a useful quantity to analyze
561 spectral properties and trends. For instance, Buurman and West (2010) used
562 A_{hf} of 10 - 20 Hz, and A_{lf} of 1 - 2 Hz, finding that low *FI* values are a good
563 indicator of impending eruption at Augustine Volcano in 2006. Our analysis
564 shows that a majority of spectral energy is limited between 10 and 20 Hz at
565 the 2017 swarm, and thus the spectral bands were extended to 10 - 30 Hz
566 for A_{hf} , and ~ 0.015 - 0.045 Hz for A_{lf} . Inspection on the example shown in
567 Figure 9 indicates the spectral difference between the three swarms of 2009,
568 2017 and 2018. Checking this figure, the upper panel (the 2017 swarm) has
569 a relatively higher frequency spectrum and thus higher *FI* than the lower
570 panel (the 2009 swarm, Harrat Lunayyir intrusion event). This method and
571 its result provide further evidence for the tectonic origin of the 2017 swarm.
572 The main event of the 2018 shows a hybrid behaviour. The spectra shows
573 no significant low-frequency signal but it tends to have considerable amount
574 of energy between the two separated windows of high-low frequency ranges.

575 Using InSAR imaging, the deformation rate maps show smooth variations
576 between $\sim \pm 0.5$ cm/yr (Figure 10), which are due to the noise of the inter-
577 ferograms that could not be fully eliminated during the processing. Magma
578 intrusion at shallow depth, generally causes ground deformations that can be
579 observed by geodetic techniques such as InSAR (e.g. Biggs and Pritchard,
580 2017). As indicated from the rate maps, no significant signal of deformation
581 is observed on the surface above the events' focal points, suggesting that the
582 swarm is not associated to magmatic processes.

583 To summarize, the results clearly indicate crustal seismicity with a narrow
584 E-W (~ 500 m wide), and steeply south dipping structure beneath this area.
585 Independent evidence from Szymanski et al. (2016) confirms the existence
586 of fault trace exposure at the relocated events E-W surface corridor. All
587 results out of the above analyses may imply an opening mechanism due
588 to the regional stress fields of the Red Sea tectonic regime which is also
589 evident in the stress inversion result (Figure 6b). Thus, our interpretation
590 suggests a mechanism of extensional faulting on a pre-existing weakness zone
591 (Najd Fault System), proposing that the 2017 swarm activity is mostly a
592 tectonic deformation with overprinting a dipping structural fabric, resulting
593 in a fracture developed in this old transform system.

594 **5 Conclusions**

595 We conclude that the occurrence of the 2017 earthquake swarm (~ 50 km
596 NW of Harrat Lunayyir boundaries) is not directly associated with a mag-
597 matic cause. In this analysis, we implement various integrated approaches

598 in sequential workflow (supplementary Figure S9) that provides a road-map
599 for source parameter estimates. By applying these techniques, ruling out
600 particular causes of seismicity gradually in a step by step process by a com-
601 prehensive, integrated data analysis approach is followed.

602 This series of analyses determines whether, or not, the seismic sequences
603 are caused by fluid-driven processes as they may occur in any area susceptible
604 to volcano-earthquakes interactions.

605 For the swarm analyzed here, our results confirm a localized crustal tec-
606 tonic deformation. Our conclusion comes mainly from the high-frequency
607 content of the events, the fault-like structure from the relative relocation
608 which confirms an upward migration, and finally from the focal mechanism
609 solutions.

610 We conclude that the current extensional-mechanism seismicity of north
611 of Harrat Lunayyir might be attributed to the regional stress fields of the
612 Red Sea stretching continental crust.

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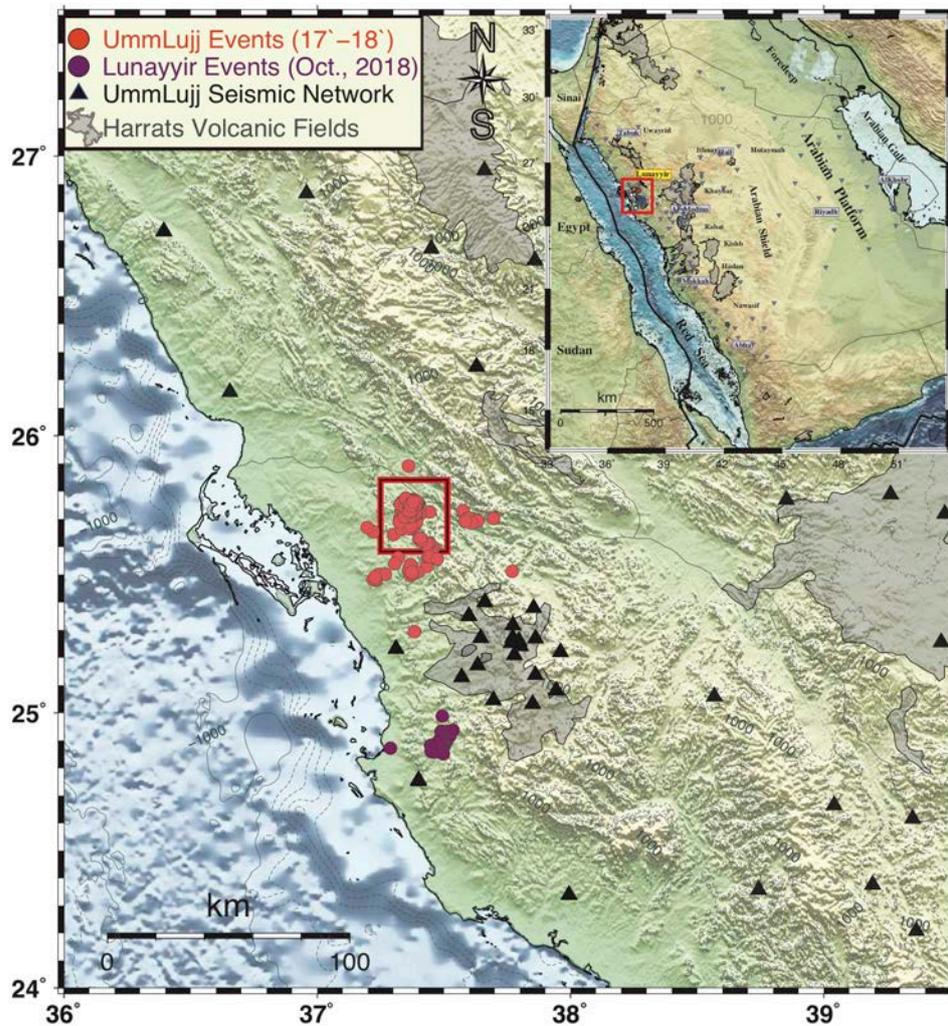


Figure 1: Location map for Lunayyir swarms and local seismic stations. This seismicity is shown by the preliminary location solution using the standard network (hypoinverse) algorithm (circles in red for the northern swarm which started in Feb. 2017, and in purple for the southern swarm which started in Oct. 2018). The inset in the upper right corner shows the Arabian Peninsula with the entire Saudi National Seismic Network (SNSN, details in supplementary Figure S1). Color-coded symbols are shown in the legends.

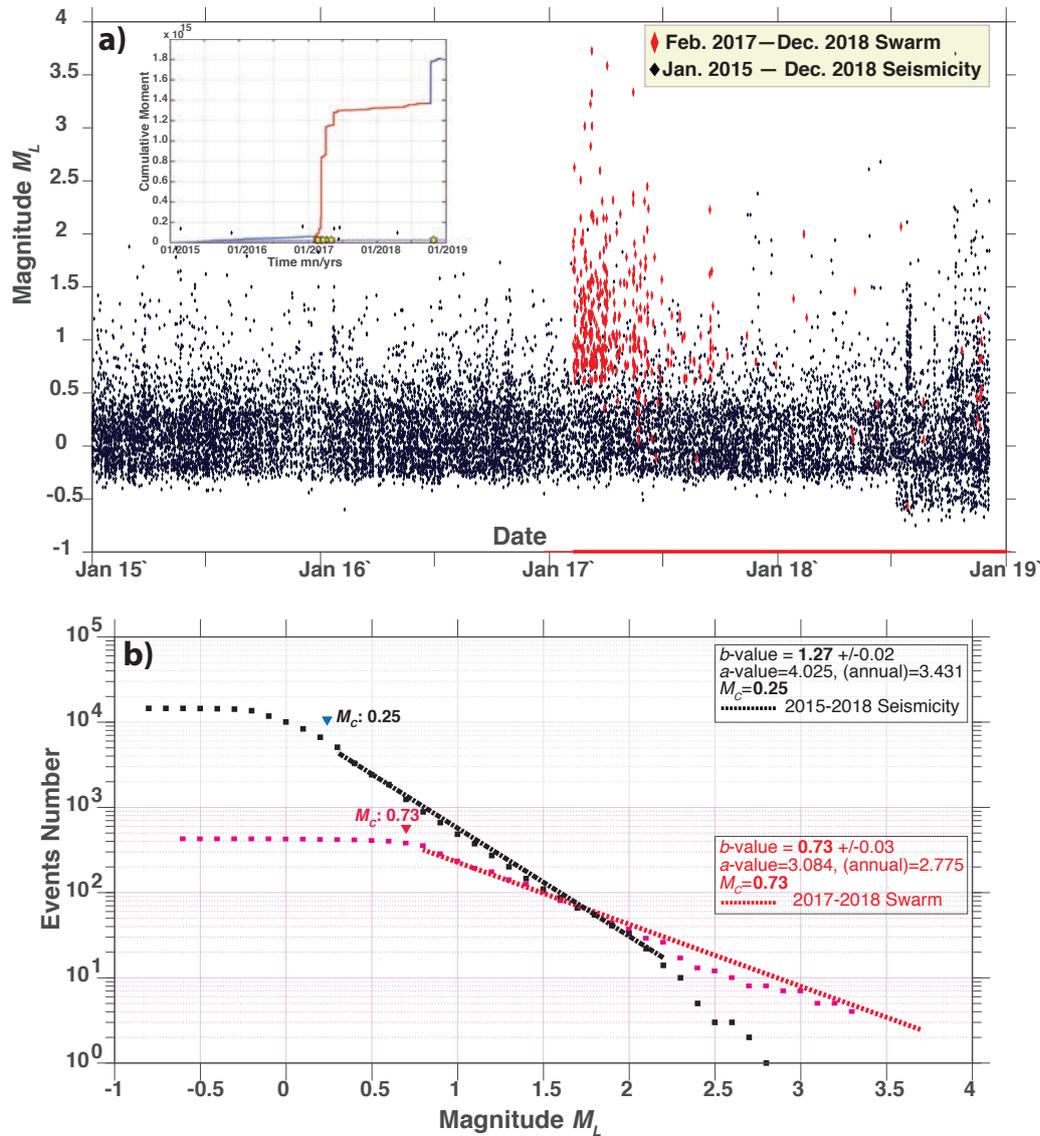


Figure 2: A) Seismicity data of $\sim 15k$ events recorded in the region of interest, where the red symbols represent the 2017 swarm, while the black dots show the background seismicity of Jan. 2015 to Dec. 2018. The inset represents a cumulative seismic moment release during the 2017 and 2018 swarms. B) Magnitude-frequency distributions (MFD) of all events in (A). The black symbols denote the background seismicity for the entire region of Harrat Lunayyir, while red color represents the 2017 swarm in the northern Lunayyir region. The squares denote number of earthquakes within each magnitude range. The dashed lines show the FMD fitted to the observed data. The inverted triangles (at $M_C = 0.25$ and $M_C = 0.73$, respectively) indicate the magnitude of completeness.

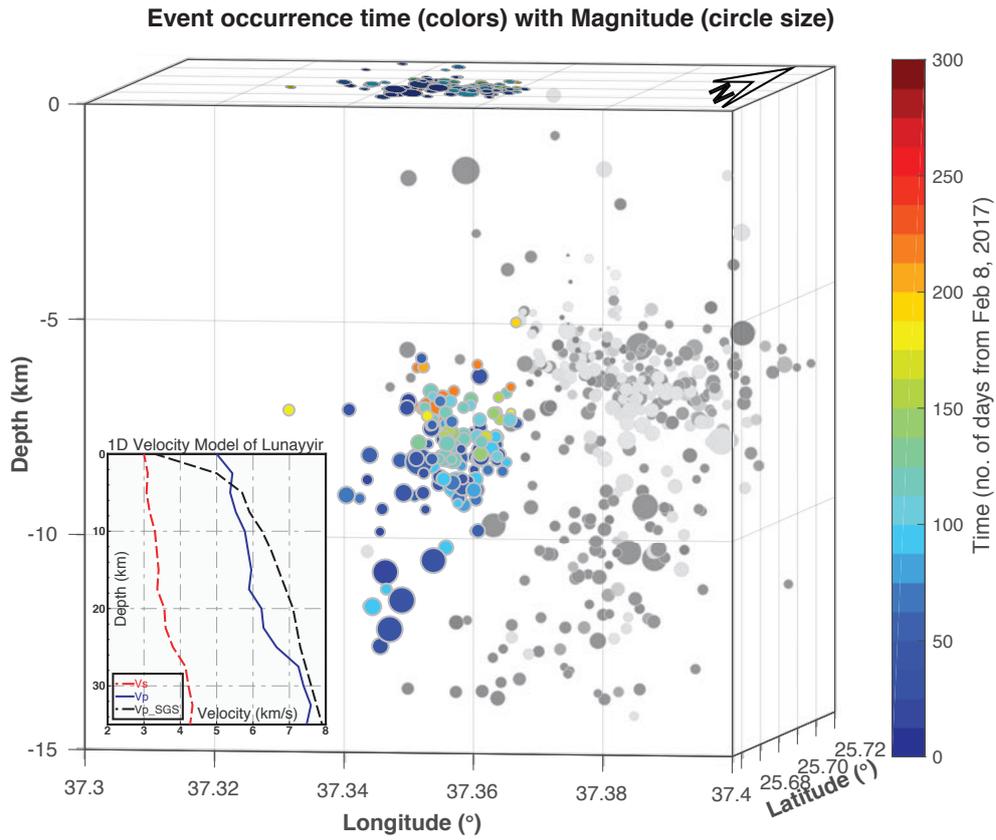


Figure 3: 3D representation showing both the double-difference relative relocation (using differential traveltimes of both P- and S-phases) and the standard network techniques (grey scale circles). The circle size represents the magnitude. The cluster of circles on the zero-level depth shows the surface projections of this swarm events. The lower inset is the velocity models used by SGS (dashed black) and this study (colored [Tang et al. 2016]).

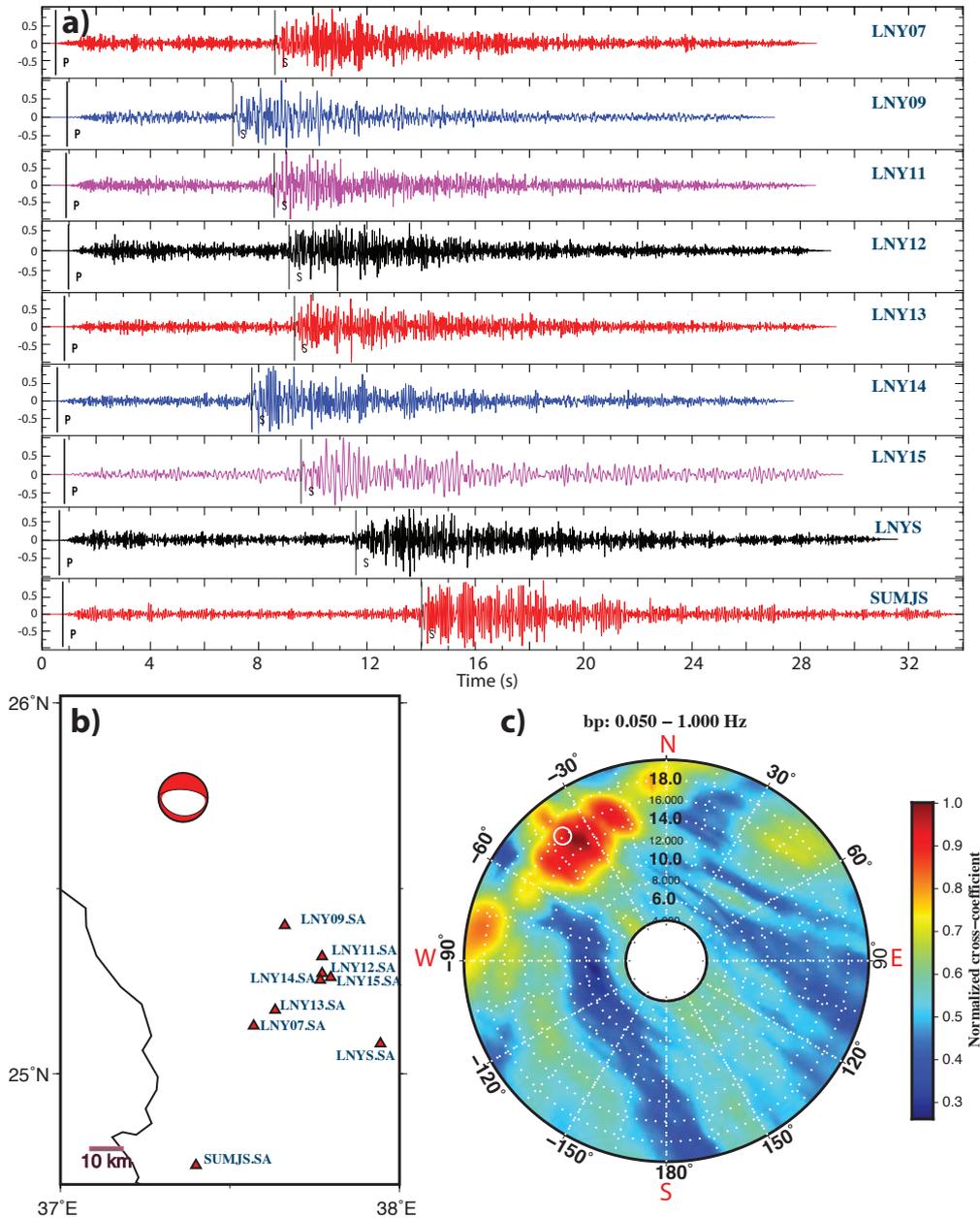


Figure 4: a) Vertical component signals (the largest event of 2017) for selected stations, which covers the southeastern sector (with respect to the 2017 swarm location), with stations average interspace of 6.11 km. b) Location map of the stations with the above (a) records. The beachball marks the location for one single event as a reference point for the fracture zone. The reference point between stations used to evaluate the inter-station spacing is at 0.66° from the selected event (the geographic mid-point between the event and all stations). c) The fk -analysis diagram shows a P-wave arriving with an average slowness of 15.65 s° along a back azimuth of 324° . The slowness from 4 to 20 s° is displayed on the radial axis; the back azimuth is shown clockwise from 0° to 360° . The observed slowness and back azimuth of the maximum power is marked with the darkest red color. Theoretical mean slowness and back azimuth values are marked by a white circle for the P arrival of event on March 10, 2017 (see supplementary Table S2).

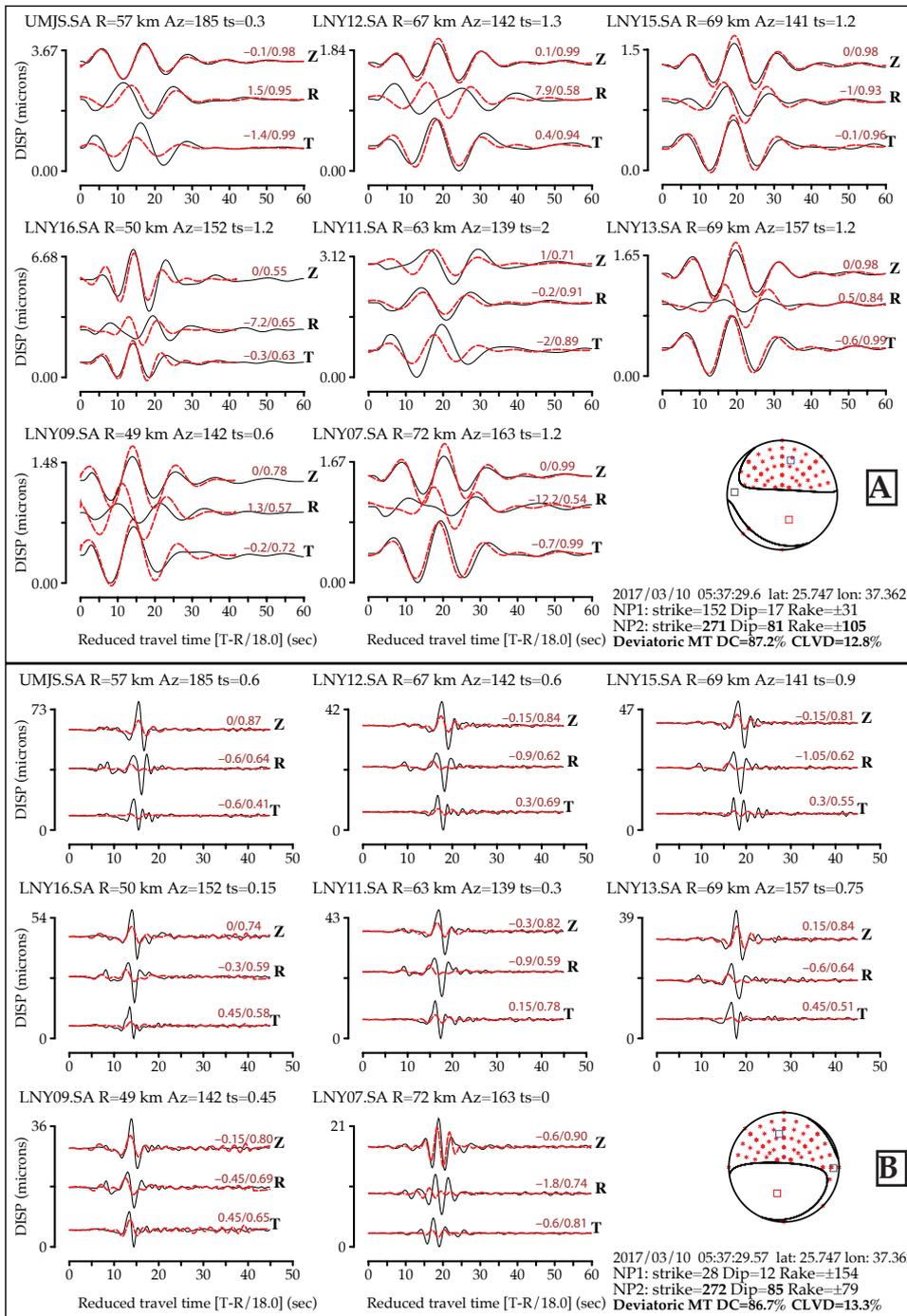


Figure 5: Moment tensor solution for the largest event (March 10th, 2017 at 17:37:10), with waveform fits (red synthetics; black observed displacement). We used a bandpass filter of A) 5-25 s (0.04 Hz – 0.2 Hz) for the low-frequency data, and B) 1-10 s (0.1 Hz – 1 Hz) for the high-frequency data. The values to the right of each seismogram component show the variance reduction (%) and time-shift(sec). The x-axis represents reduced travel-time which includes a reduction velocity.

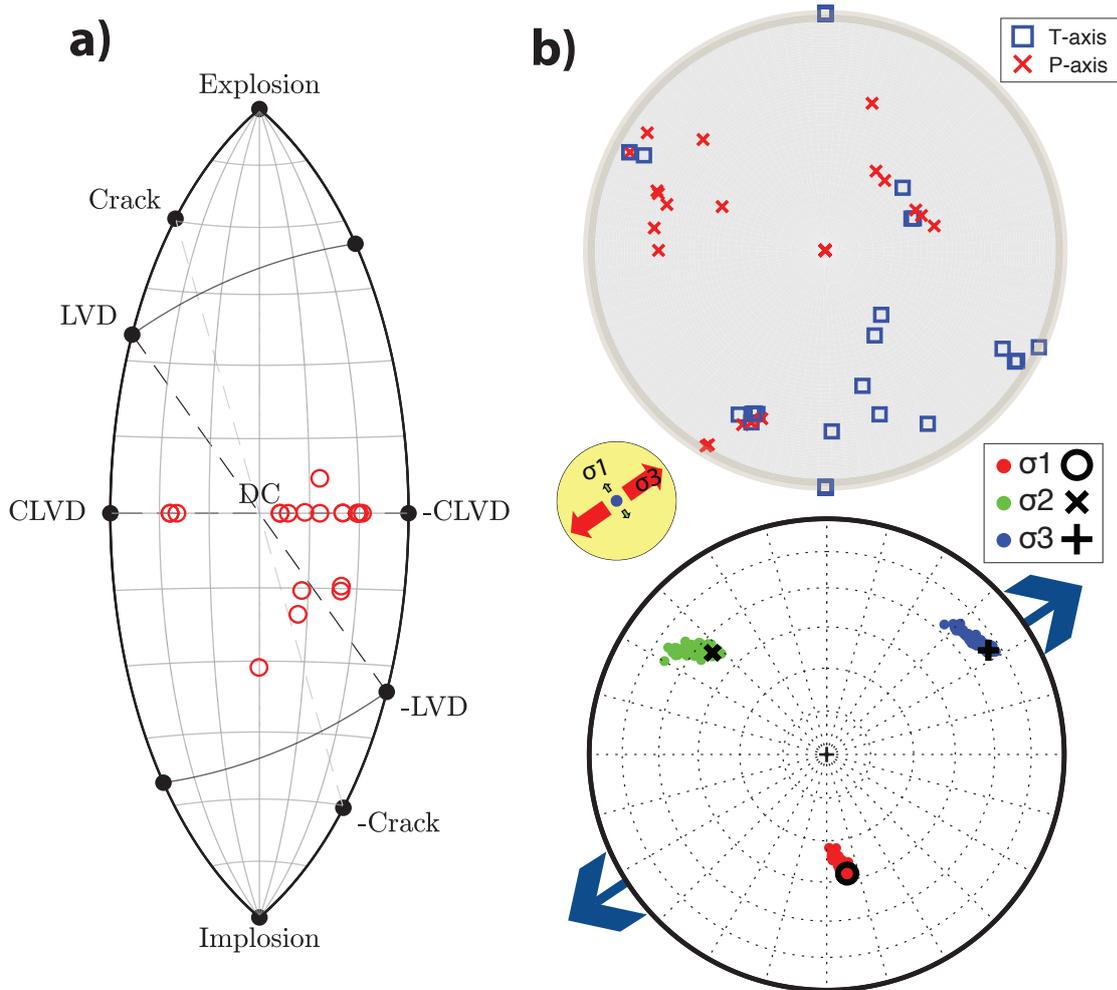


Figure 6: Full moment tensor datasets plotted on the fundamental lune representation of source types. Stress inversion results from focal mechanism solutions of events $\geq M_L$ 2.8. The yellow circular diagram shows the horizontal stress axes (σ_3 Sh_{max} and σ_1 Sh_{min}).

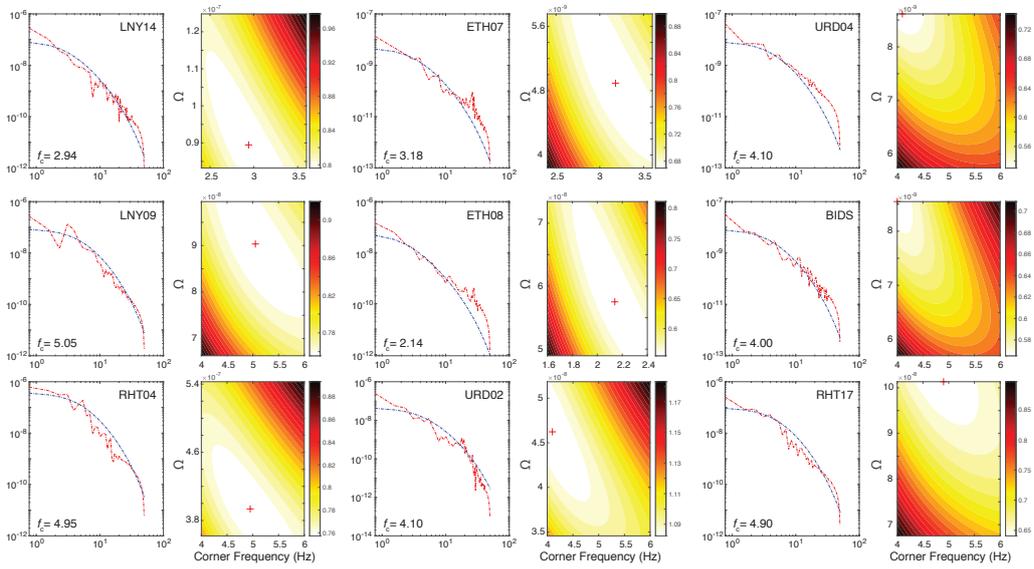


Figure 7: Fourier spectra of the M_L 3.73 event of March 2017 recorded at different stations (SH-component) that were used for calculation of earthquake source parameters. Corner frequencies f_c values are shown in the plots.

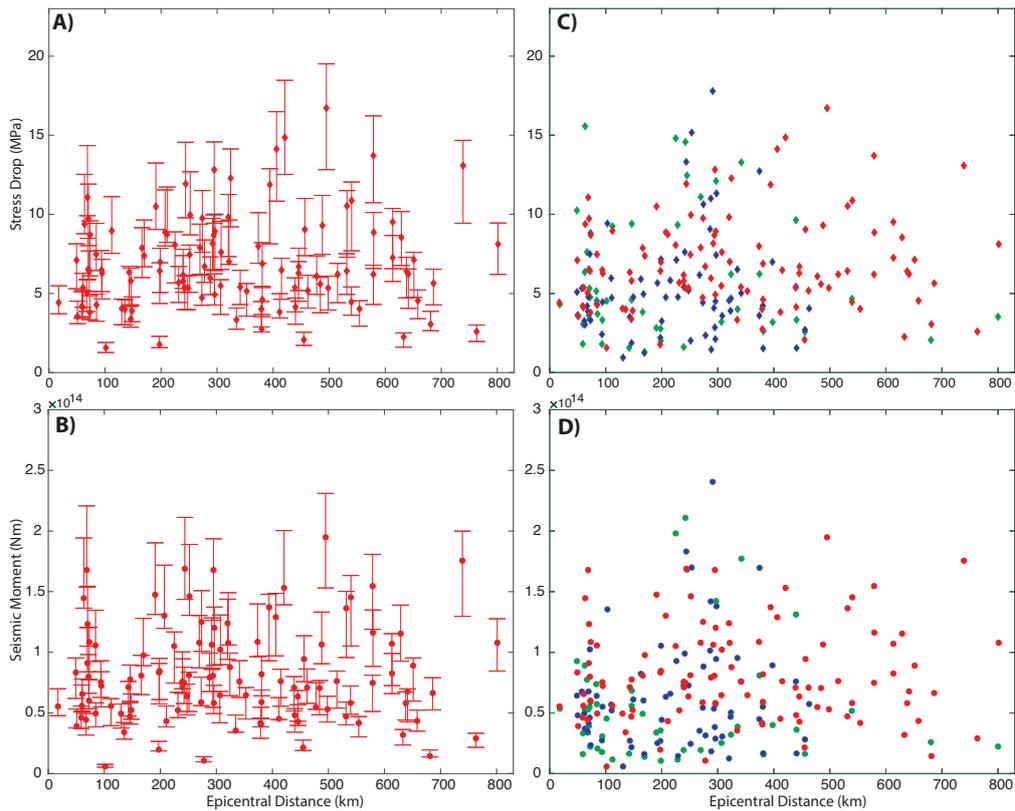


Figure 8: A) Stress drop and B) seismic moment vs epicentral distances calculated for the second largest event of April 3rd, 2017 (M_L 3.59). The uncertainty estimates are data-driven using both P- and S-wave spectra as the limits of the mean values as shown. C) Stress drop and D) seismic moment vs epicentral distances calculated for the three largest events with M_L 3.73 (blue), 3.59 (red), 3.34 (green).

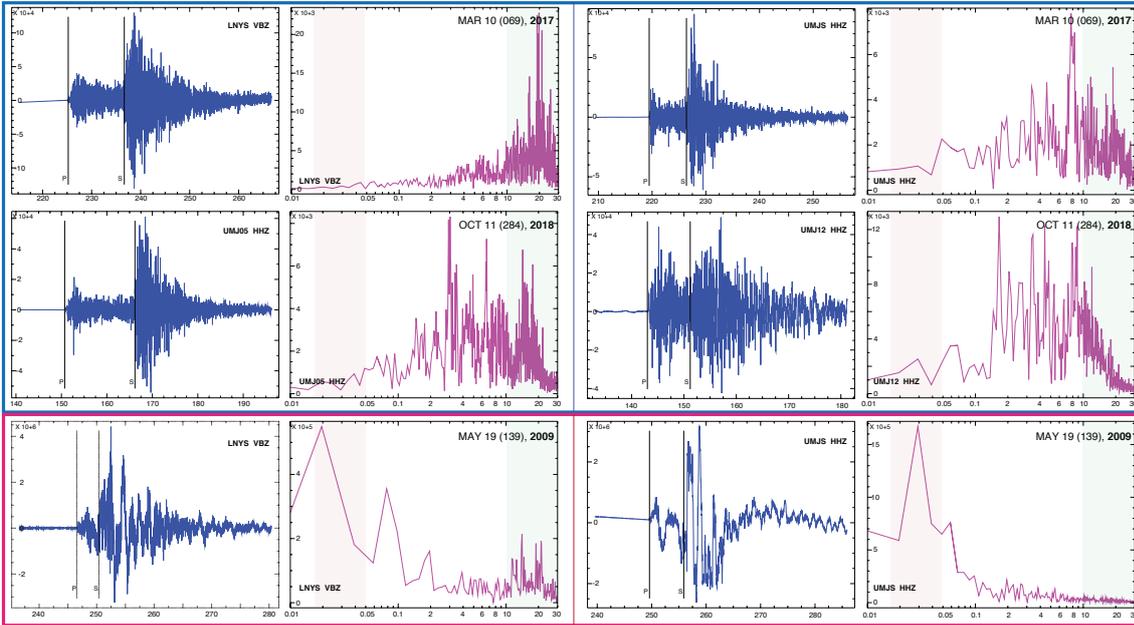


Figure 9: Comparison between frequency spectra for the tectonic event of March 10, 2017 (upper panel) and the magmatic event of May 19, 2009 (lower panel). The event of October 11, 2018 (lower panel) which took place at the southwestern edge of Harrat Lunayyir might indicate a hybrid nature. Waveforms and corresponding frequency spectra show differences between the three cases. The low- and high-frequency bands are defined as 0.015-0.045 Hz and 10-30 Hz, respectively.

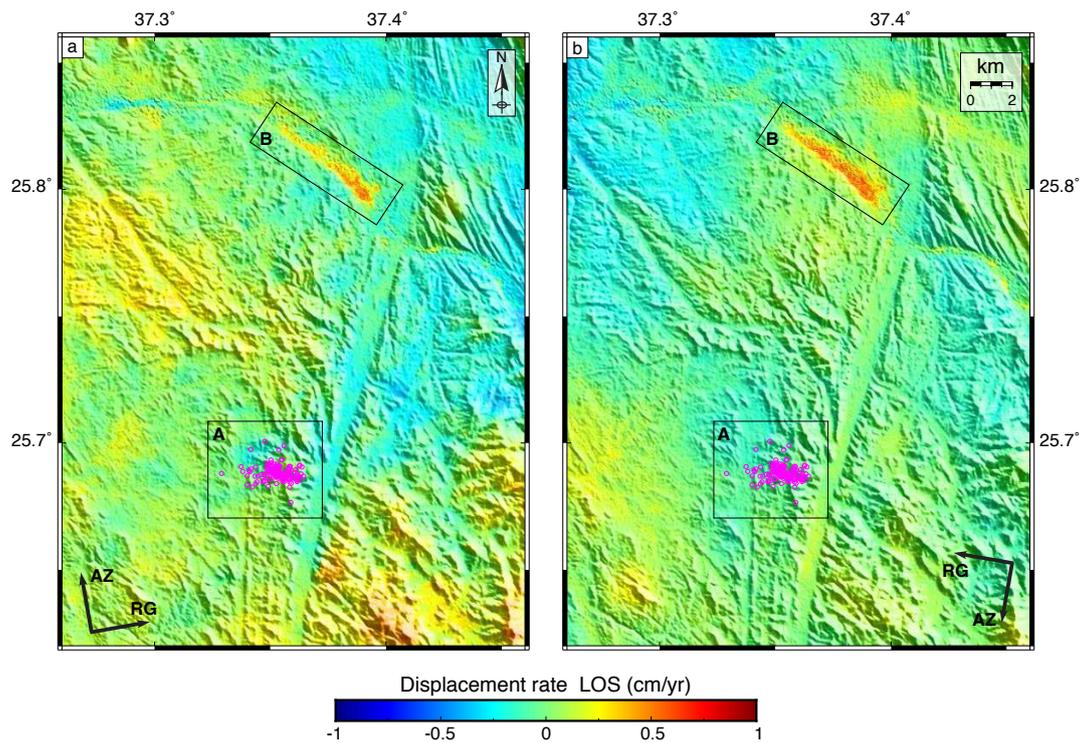


Figure 10: InSAR deformation rate maps in LOS obtained with Sentinel-1 images in ascending (a) and descending (b) orbits. Zone A is a $5 \times 4 \text{ km}^2$ area where the seismic swarm occurred. Here, no evident signal associated to ground deformation is observed. Zone B is a $\sim 6 \text{ km}$ North-West trending valley that corresponds to a dry riverbed. In both maps the area moves toward the satellite experiences uplift of $\sim 1.2 \text{ cm/yr}$, likely related to water level changes in a shallow aquifer.

[Non-Volcanic Earthquake Swarm Near the Harrat Lunayyir Volcanic Field, Saudi Arabia]

[M. Youssef, P. M. Mai, A. Nobile, and S. Jónsson]

[King Abdullah University of Science and Technology (KAUST), Thuwal 23955-6900, Saudi Arabia]

Contents of this file

Text S1 to S4
Figures S1 to S8
Tables S1 to S4

Introduction

The supporting information contains some details for different sections of the main article. This section mainly documents some technical notes about 1) M_L formulas, 2) sensitivity analysis by backprojecting the Incident rays, 3) stress drop estimations, and 4) frequency index analysis.

Text S1. Appendix (A): M_L Formulas

The network-based standard magnitudes produced by the SGS are based on two definitions for M_L formulas (Soliman et al. 2019), depending on the area distances and tectonics. One equation is Eq. (1) of Alsaker et al. (1991), which can only be used for the shield structure and local distances, with epicentral distances ≤ 200 km, while Eq. (2) of Bormann (2012), which is recommended for crustal earthquakes within a regional scale (i.e. $200 \text{ km} \leq D \leq 700 \text{ km}$).

$$M_L \simeq \log A_{Smax} + 1.11 \log R + 0.00189R - 2.09 \quad (1)$$

$$M_L \simeq 0.925 \log A_{Smax} + 0.91 \log R + 0.00087R - 1.31 \quad (2)$$

where A_{Smax} is the maximum of the ground velocity absolute amplitude of highpass corner frequency 2Hz, for the S-wave ground velocity (measured in $\mu\text{m/s}$). R is a distance from the hypocenter to the station in km. Eq. (1) defines the local magnitude scale for the whole region of Lunayyir, thus, we are able to compare the MFDs of any seismic activities within our area of interest as shown in Figure (2).

Text S2. Appendix (B): Sensitivity Analysis by Backprojecting the Incident Rays

The arrival time delays required to bring the signals into phase is a direct estimate of the back azimuth and the slowness of the seismic signals. The total energy (defined by the power spectral density– PSD) recorded at the array can be calculated by the integration of the squared summed amplitudes over time.

The fk diagram (Figure 4c) shows the array response of a selection of our dense-aperture southeastern array (Map of Figure 4b).

For comparison, we calculate the expected phase travel times and ray parameters given the known stations and relocated source coordinates for the same event, using the velocity model we applied in the relocation. As listed in Table S2, the average theoretical values for the ray parameters are 16.1, and 28.7 s/° for P- and S-wave, respectively. These predicted values agree well with the fk-analysis, confirming the observed and relocated hypocenter of the event of interest.

The uncertainty is reasonable between the synthetics and fk results. The difference in slowness values is minimal (0.45 s/°), where the backazimuth values have about 2° difference. A source of such difference is attributed to using one pair of event-station for the synthetics while using several stations in the fk measurements, beside the lack of an accurate 3D velocity model contributes to such difference.

Text S3. Appendix (C): Stress Drop Estimation

We calculate stress drop ($\Delta\sigma$) using the Eshelby (1957) relationship:

$$\Delta\sigma = \left(\frac{7}{16}\right) \times \left(\frac{M_0}{r^3}\right) \quad (3)$$

in which M_0 is the seismic moment (in N.m), and r is the source radius (in meters). Note that M_0 and M_W values depends on the seismic wave types used in the analysis. Despite using models for both P- and S-phases here, we computed the mean values which further used to estimate ($\overline{\Delta\sigma}$).

To measure the source radius, we use the circular source model of Madariaga (1976) and Brune (1970)

$$r = \frac{\kappa\beta}{f_c} \quad (4)$$

in which we assume β , the shear-wave velocity as 3.4 km/s (based on our velocity model and average source depth). While κ depends on the spectra of P and S waves and on the choice of source model.

To estimate the model parameters, we use a standard grid-search method.

Previous studies suggested that estimation of source parameters is robust using this technique (e.g., Tusa et al., 2006a, 2006b; Edwards et al., 2008; de Lorenzo et al., 2010). Here, we implement a grid-search over ω_0 , and f_c , fixing $\kappa=0.3$. To optimize the search process, we adopt a two-level procedure (e.g., Lomax et al. 2000). In the first step, the entire relevant model parameter space is subdivided into a coarse grid. The range of model parameter values in the coarse grid is based on a priori analysis of the dataset (assigning realistic values from the earlier studies on source characterization). At each point of this grid, the observed and the theoretical spectrum are matched using a misfit function of mean absolute error (MAE) performance, which helps to measure accuracy for this continuum of variables. In the second step, a refined grid is built around the initial best estimations, and MAE function is recalculated at each point of this grid to resolve the best-fitting parameters.

Text S4. Appendix (D): Frequency Index Analysis

Buurman and West (2010) developed a measure to discriminate between different types of seismic events, defines the frequency index (FI) based on the ratio of energy in low and high frequency windows:

$$FI = \log \left(\frac{\overline{A(hf)}}{\overline{A(lf)}} \right) \quad (5)$$

where $\overline{A(hf)}$ and $\overline{A(lf)}$ are, respectively, the average spectral amplitudes across selected bands of high and low range of frequencies.

Because the resulting measure spans many orders of magnitude, we use a base-ten logarithm to reduce the index to a simple number.

To calculate the FI in a consistent manner, we first pick the P- and S-onsets, minimizing the time window to approximately the P-S duration, followed by removing the average amplitude from the selected waveforms signals, with a fixed time series duration of 40 seconds: 10 second prior to the earthquake P-onset and 30 seconds after, ensuring that the high frequency signal is fully captured in the Fourier analysis. This is a sufficient time window over which to sample both the shorter duration, smaller magnitude earthquakes recorded, as well as the more emergent, lower frequency events. Linear trends and offsets are removed from the waveforms, and they are transformed to the frequency domain using a tapered Fourier transform.

To avoid problems with noise contamination, only earthquakes with high SNR ≥ 3.0 were included. The average noise amplitude was measured and a noise window was chosen such that it ended 15 s before the P-first arrival pick. After calculating the SNR using the amplitudes measured in the earthquake and noise windows, the amplitude of the noise window was subtracted from the earthquake window amplitude.

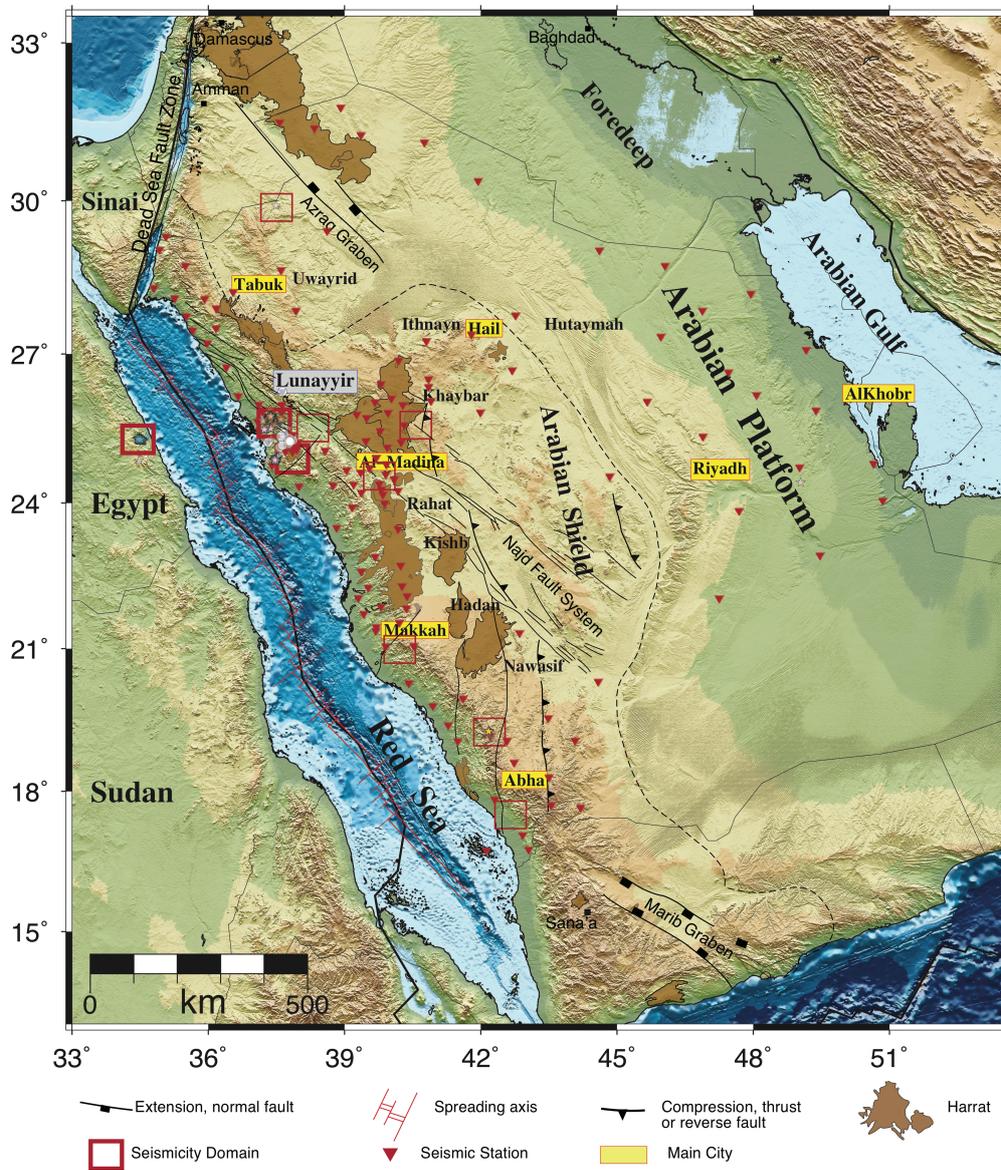


Figure S1: Regional map of Arabia. Volcanic fields (harrats) are shown in brown color. Permanent stations are shown in red color. The places of long- and short-term swarms discussed in the text are highlighted in bold and light red boxes, respectively.

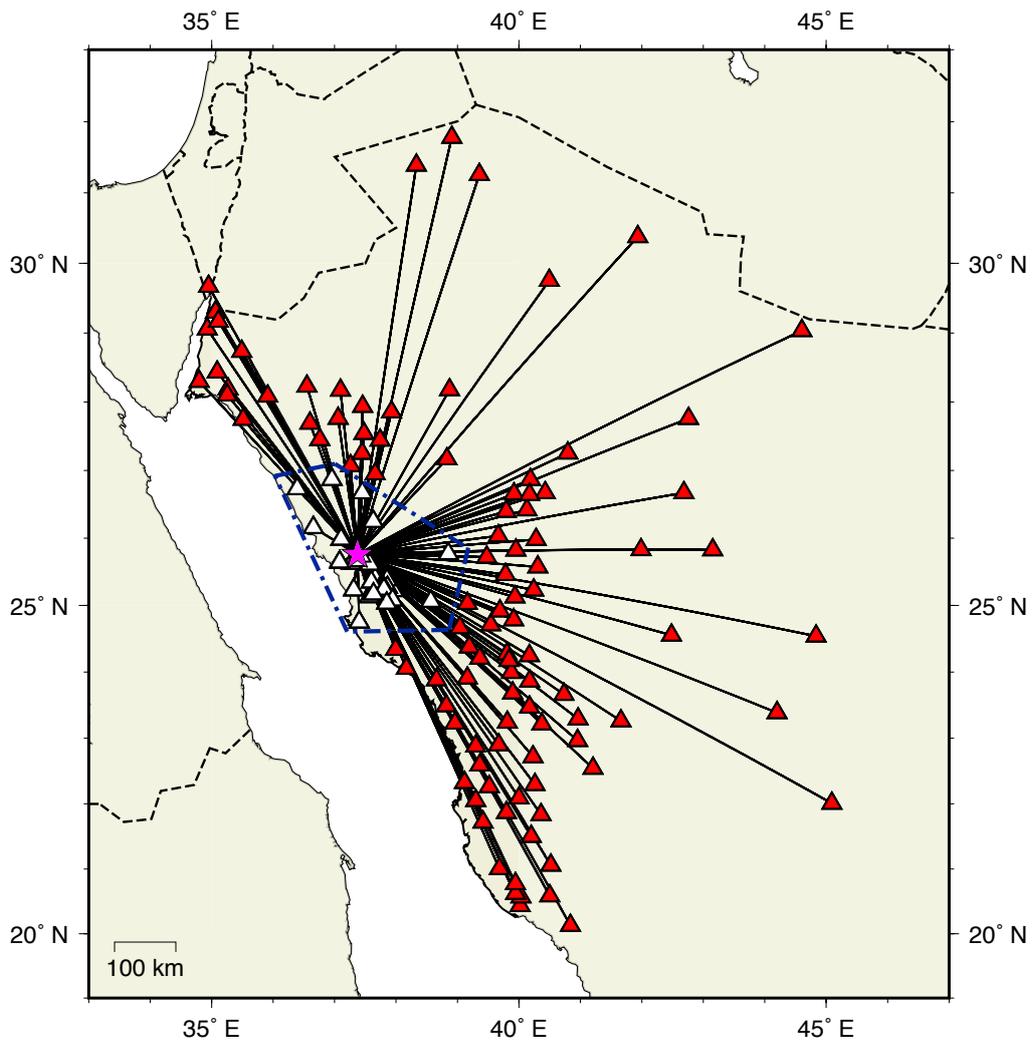


Figure S2: All stations (color-coded with both red and dark yellow colors) used in the first trial to assemble as much records as we can for the first inversion run before selecting the optimum records (from stations in dark yellow) which involved in the final solution.

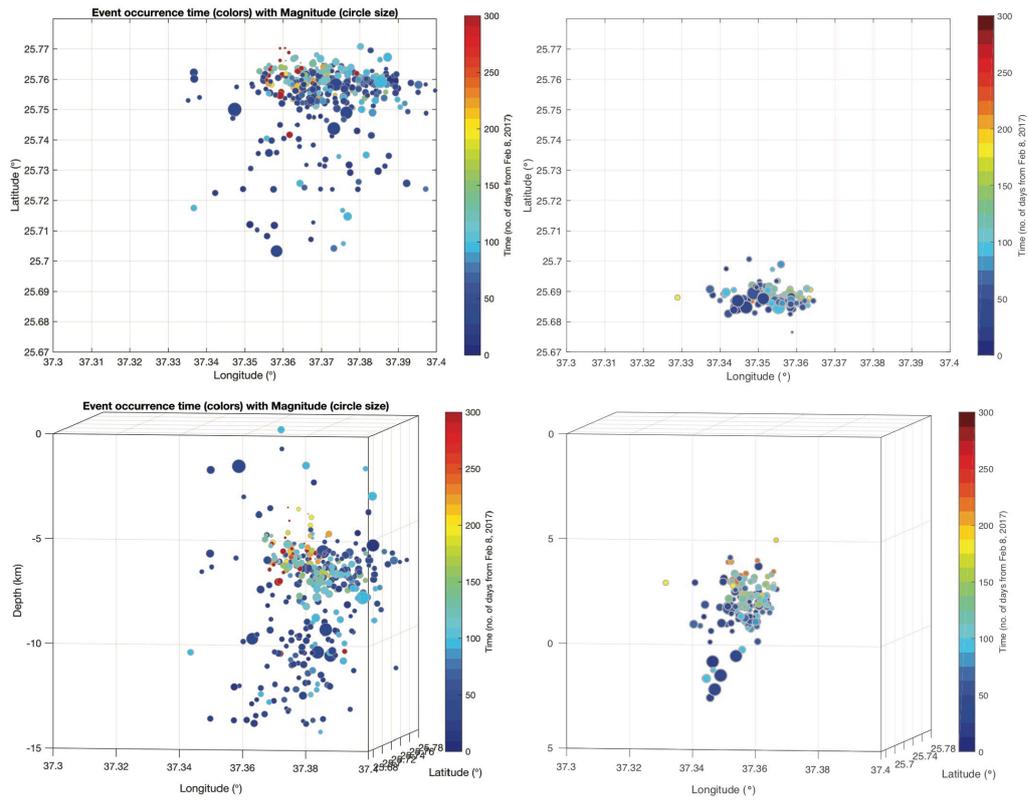


Figure S3: Plan views and vertical cross sections for double-difference relative relocation results using differential traveltimes of both P- and S-phases. The circle size represents the magnitude. An upward migration is clearly observed.

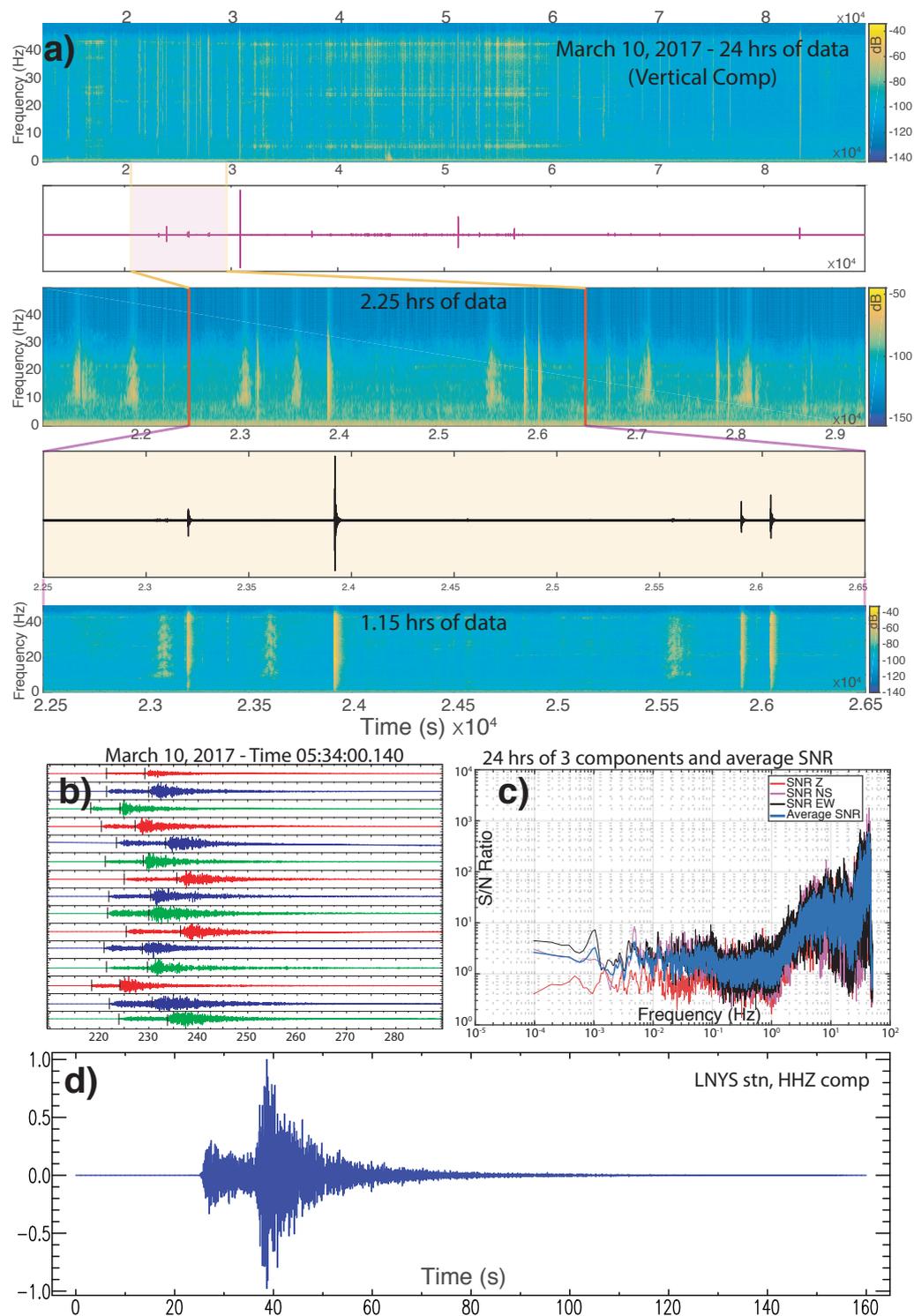


Figure S4: a) Example of data plotted as time-series and frequency domain, helping to identify event-based waveforms, b) Raw seismic data at 15 stations (vertical component), of the largest event with M_W 3.6. The picks of P- and S- phases, used as input for the double-difference analysis. c) SNR example for 24 hours of continuous data at one station. d) One trace out of c) showing a high-quality signal.

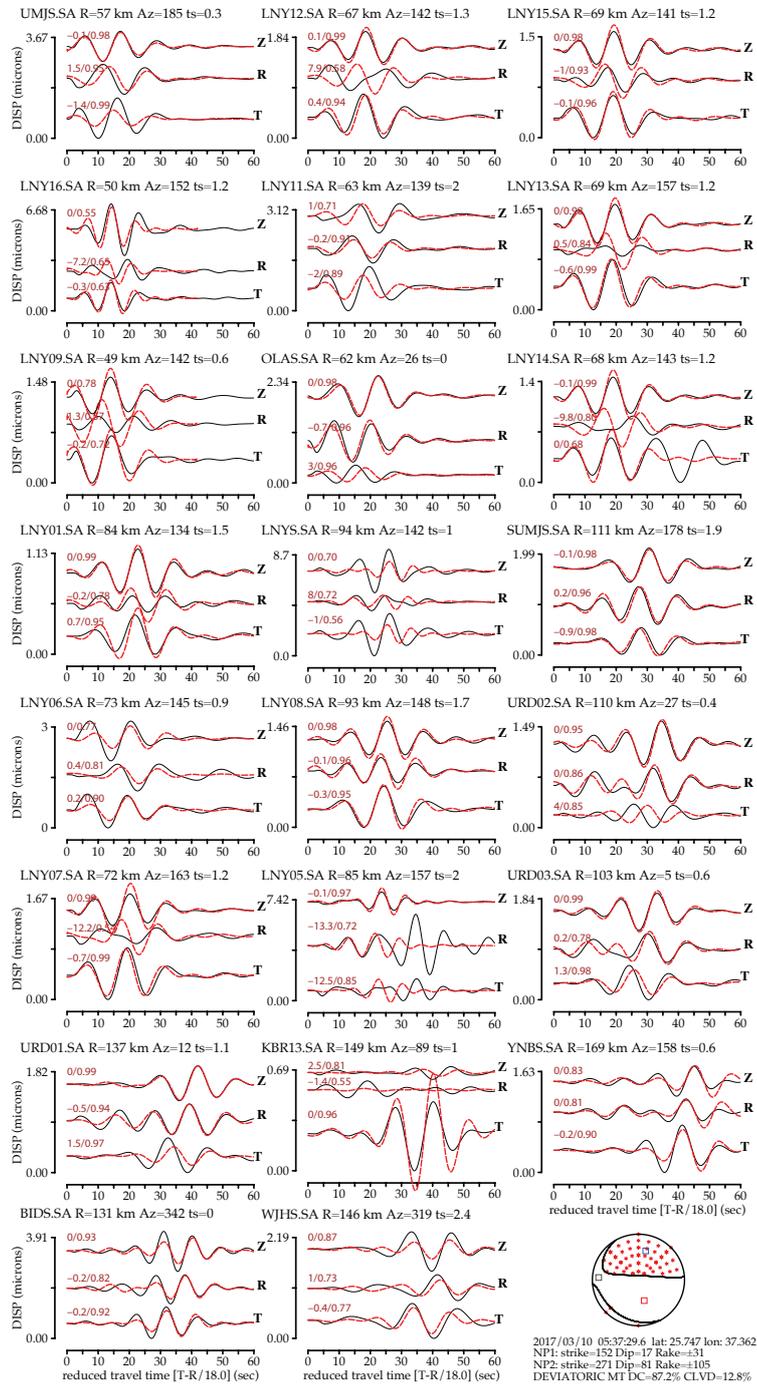


Figure S5: Moment tensor solution for the largest event (March 10th, 2017 at 17:37:10, Umm-Lujj), with waveform fits for the low-frequency data 5-25 s (0.04 Hz – 0.2 Hz), (red synthetic; black observed displacement data).

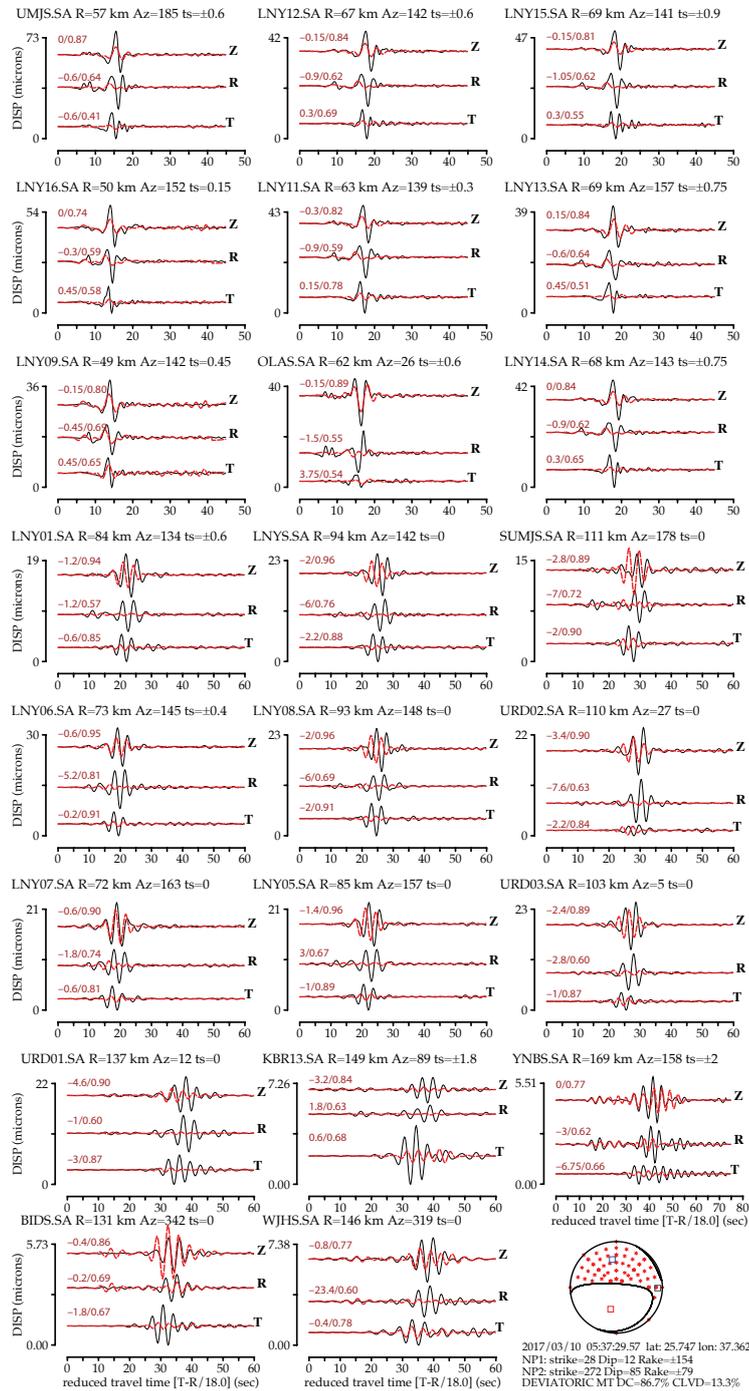


Figure S6: Moment tensor solutions for the largest event (March 10th, 2017 at 17:37:10, Umm-Lujj), with waveform fits for the high-frequency data 1-10 s (0.1 Hz – 1 Hz), (red synthetic; black observed displacement data).

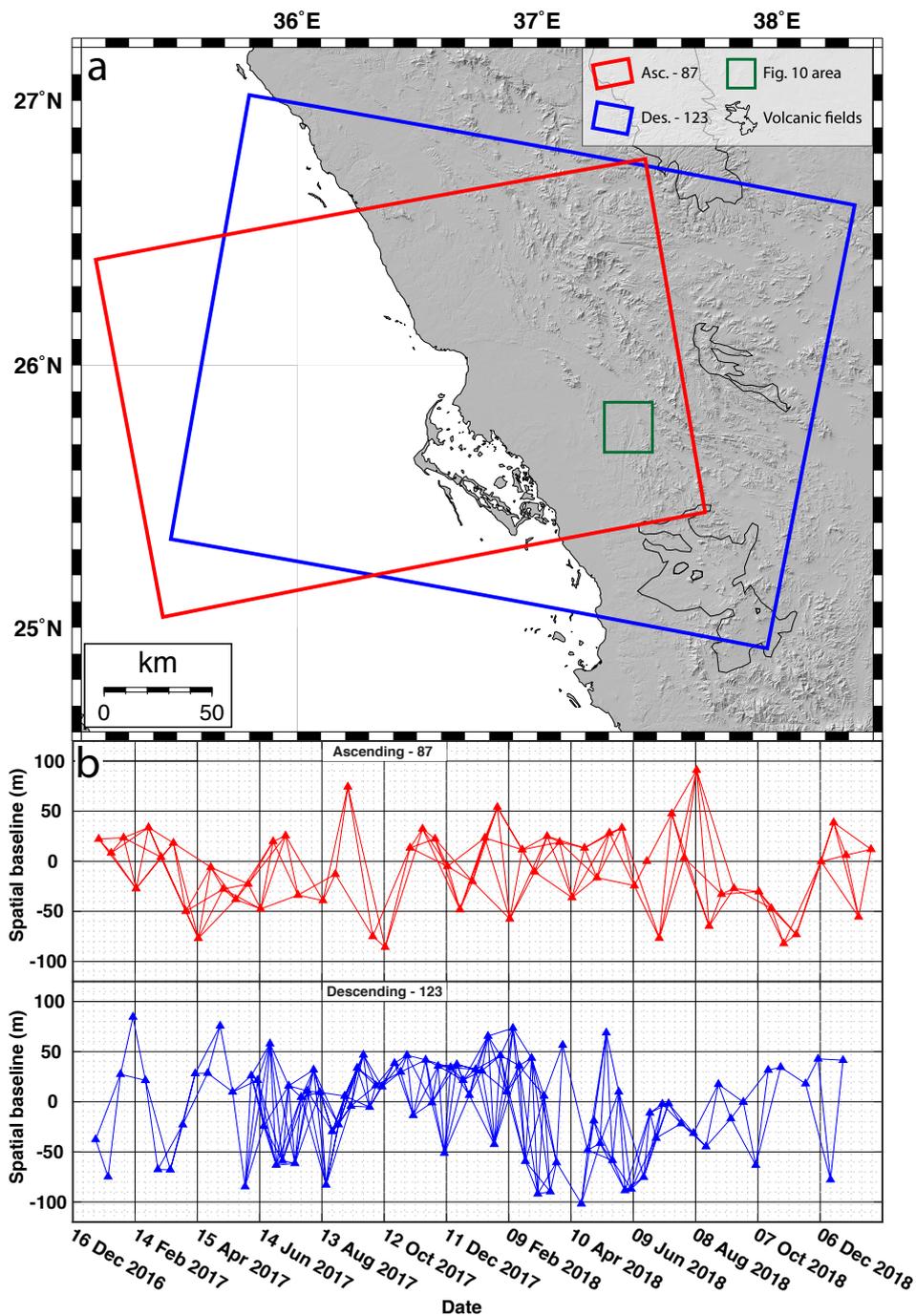


Figure S7: a) The area of interest (green square) is located to the north of Harrat Lunayyir and is covered by two SAR frames in both ascending (red box) and descending (blue box) orbits. b) the interferogram networks used to evaluate the deformation rate maps in figure 10.

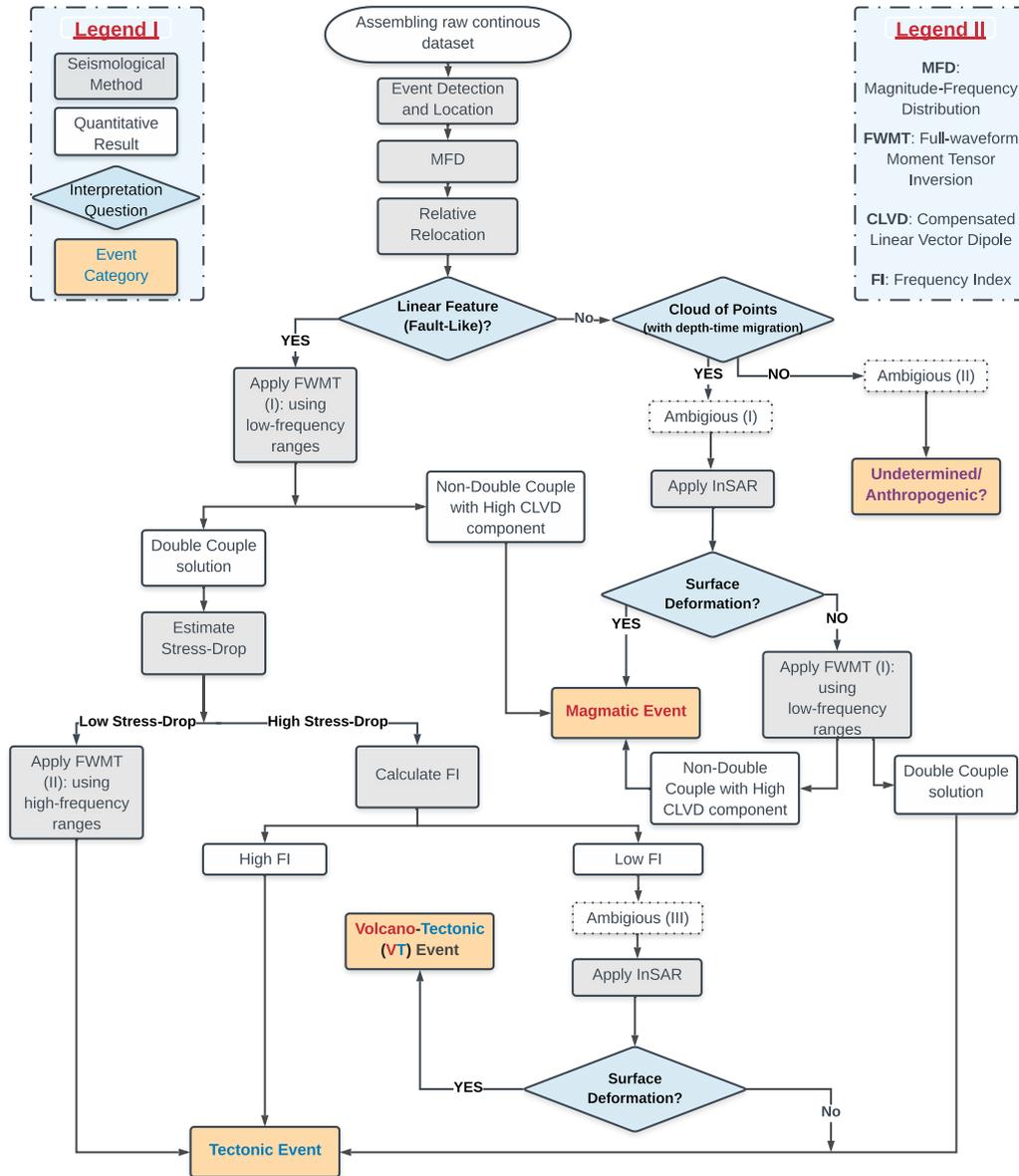


Figure S8: Proposed flow-chart of data processing to implement for distinguishing events nature in such cases of seismic swarm activities.

Table S1: Seismic activity around the Red Sea

Location	Max. Mag. / Time	Duration
¹ Tabuk eq., North of KSA	M_L 5.1 July 27, 2004	3 months (June - August 2004)
² Al-Ays Swarm, North of KSA	M_L 4.2 Nov. 11, 2007	7 months (Oct. 2007 - May 2008)
³ Badr eq., North of KSA	M_b 4.0 August 23, 2009	1 month (August - Sept. 2009)
⁴ Jizan Swarm, South of KSA	M_L 4.5 Jan. 23, 2014	11 months (Jan. - Dec. 2014)
⁵ Al-Namas Swarm, North of KSA	M_L 4.1 Nov. 03, 2017	3 months (Nov. 2017 - Jan. 2018)
⁶ Abu-Dabab Swarm, South of Egypt	M_L 5.1 July 02, 2004	continuous for decades

1011 ¹Al-Damegh et al. (2009). ²Mukhopadhyay et al. (2012), Saibi et al. (2019).
1012 ³Aldamegh et al. (2012). ^{4,5}Abdelfattah et al. (2017, 2020). ⁶Badawy et al.
1013 (2008).

Table S2: Theoretical travel times and ray parameters of corresponding seismic phases computed using 1-D velocity model

Dist. (°)	Phase	TT (s)	Ray Param. (s/°)	Inc. (°)
<i>LNY09</i>				
0.44	P	8.65	18.597	75.94
	S	14.94	32.103	75.95
<i>LNY11</i>				
0.57	P	11.86	17.05	62.81
	S	20.47	29.56	63.28
<i>LNY12</i>				
0.6	P	12.46	17.04	62.79
	S	21.52	29.53	63.23
<i>LNY14</i>				
0.61	P	12.65	17.03	62.77
	S	21.85	29.51	63.21
<i>LNY13</i>				
0.62	P	12.84	17.02	62.75
	S	22.17	29.5	63.2
<i>LNY15</i>				
0.62	P	12.85	17.01	62.75
	S	22.19	29.5	63.2
<i>LNY07</i>				
0.64	P	14.95	13.75	45.84
	S	22.78	29.35	63.17
<i>LNYS</i>				
0.85	P	17.74	13.73	45.78
	S	30.86	24.74	48.38
<i>SUMJS</i>				
0.99	P	19.76	13.68	45.72
	S	34.5	24.5	48.33
<i>Average</i>				
0.66	P	13.75	16.10	58.57
	S	23.48	28.70	61.33

Table S3: Azimuthal dependency of source parameters for the largest event recorded by 68 stations

Azimuth	Mean/Median	Distance	SD	Moment	Mw	fc
All clusters	Median	202.583	5.116	6.308E+13	3.122	2.6
	Mean	192.059	5.491	6.668E+13	3.466	2.8
NE cluster	Median	242.403	5.745	6.073E+13	3.156	2.3
	Mean	232.359	6.995	8.149E+13	3.198	2.6
NW cluster	Median	315.967	4.898	6.544E+13	2.677	3.2
	Mean	305.338	4.920	6.313E+13	4.405	3.1
SE cluster	Median	162.764	4.457	4.803E+13	3.088	2.4
	Mean	173.124	4.714	5.443E+13	3.075	2.7
SW cluster	Single Stn.	57.416	5.334	6.767E+13	3.187	2.8

Table S4: Frequency Index parameters for the stations shown in Figure 9

Event	Station	SNR	<i>FI</i>
2009 event	LNYS	86.52	-0.41
	UMJS	81.47	-0.64
2017 event	LNYS	92.93	0.26
	UMJS	94.32	0.18
2018 event	LNYS	92.46	0.37
	UMJ05	86.41	0.025
	UMJ12	89.89	0.011